

CURIOUS SYMMETRY OF CHIRAL FERMIONS

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‘‘Chiral fermions as classical massless
spinning particles,’’ arXiv:1406.0718
[hep-th].

September 12, 2014

Stephanov & Yin PRL 2012 (semi)classical model derived from Weyl Hamiltonian $H = \boldsymbol{\sigma} \cdot \mathbf{p} \rightsquigarrow$ spin-1/2 system with positive helicity and energy, described by phase-space action $S =$

$$\int \left((\mathbf{p} + e\mathbf{A}) \cdot \frac{d\mathbf{x}}{dt} - (|\mathbf{p}| + e\phi) - \mathbf{a} \cdot \frac{d\mathbf{p}}{dt} \right) dt. \quad (1)$$

$\mathbf{a}(\mathbf{p})$ “momentum-dependent vector potential” for “Berry monopole” in \mathbf{p} -space

$$\nabla_{\mathbf{p}} \times \mathbf{a} = \boldsymbol{\Theta} \equiv \frac{\hat{\mathbf{p}}}{2|\mathbf{p}|^2} \quad (2)$$

where $\hat{\mathbf{p}} = \mathbf{p}/|\mathbf{p}|$, $\mathbf{A}(\mathbf{x}, t)$ & $\phi(\mathbf{x}, t)$ “ordinary” vector & scalar potentials.

Fundamental assumption: Spin “enslaved”

$$\mathbf{s} = s\hat{\mathbf{p}} \quad s = \frac{1}{2}. \quad (3)$$

Lack of manifest Lorentz symmetry .

Chen-Son-Stephanov-Yee-Yin , arXiv:1404.5963

Manuel, Torres-Rincon , arXiv:1404.6409 :

Lorentz algebra acts as

$$\begin{cases} \delta \mathbf{p} = \boldsymbol{\omega} \times \mathbf{p} + |\mathbf{p}| \boldsymbol{\beta}, \\ \delta \mathbf{x} = \boldsymbol{\omega} \times \mathbf{x} + \boldsymbol{\beta} \times \left(s \frac{\mathbf{p}}{|\mathbf{p}|^2} \right) - \boldsymbol{\beta} \cdot \mathbf{x} \frac{\mathbf{p}}{|\mathbf{p}|}. \end{cases} \quad (4)$$

N.B. : **not** a space-time action !

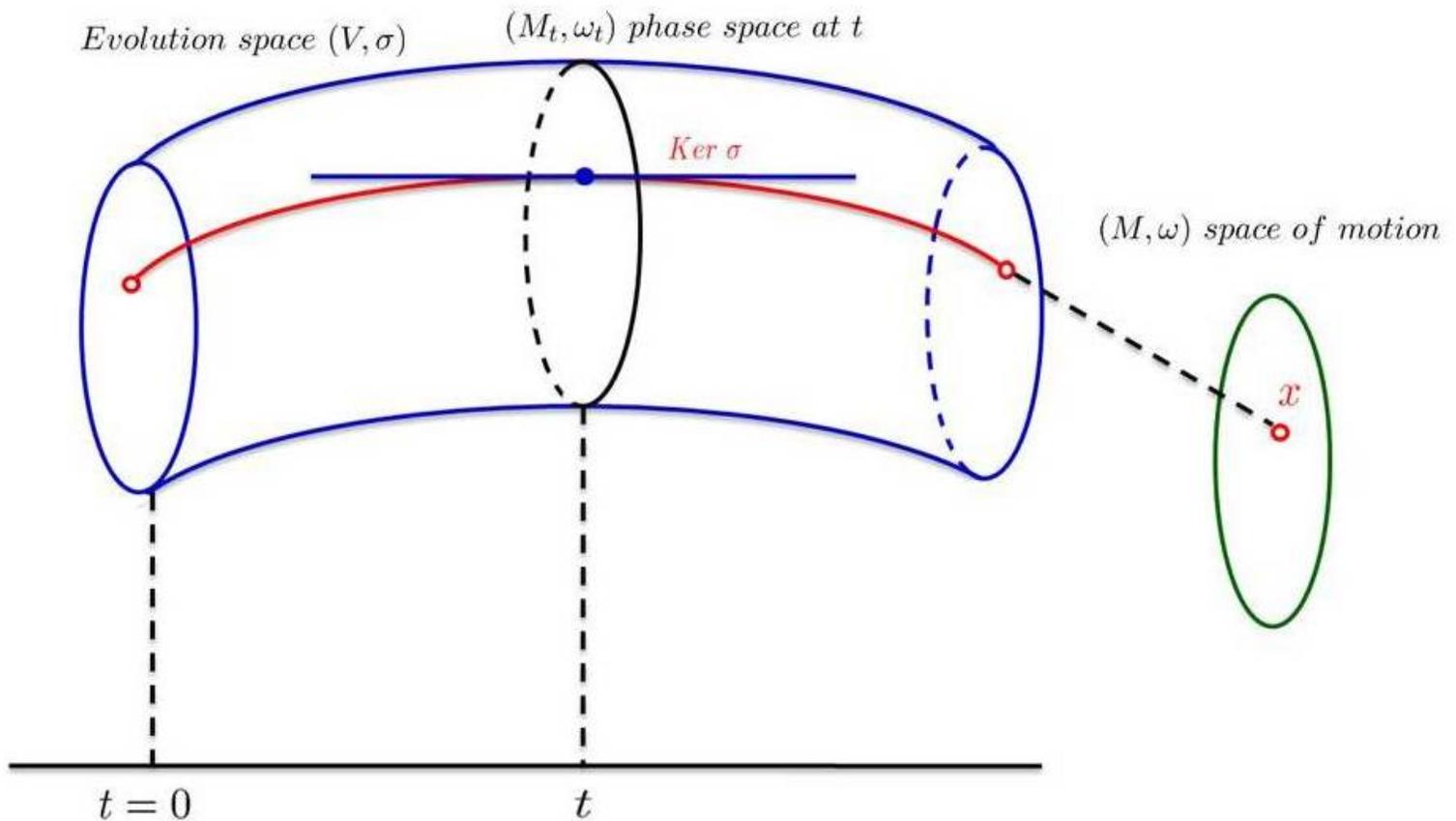
“dynamical” symmetry !

Duval, PAH arXiv: 1406.0718 :

1. *chiral system embedded into Souriau's massless spinning particle with Poincaré symmetry;*
2. Spin enslavement **inconsistent** with natural Lorentz symmetry
3. **twisted** (4) but **no natural** Lorentz symmetry for chiral system (1).

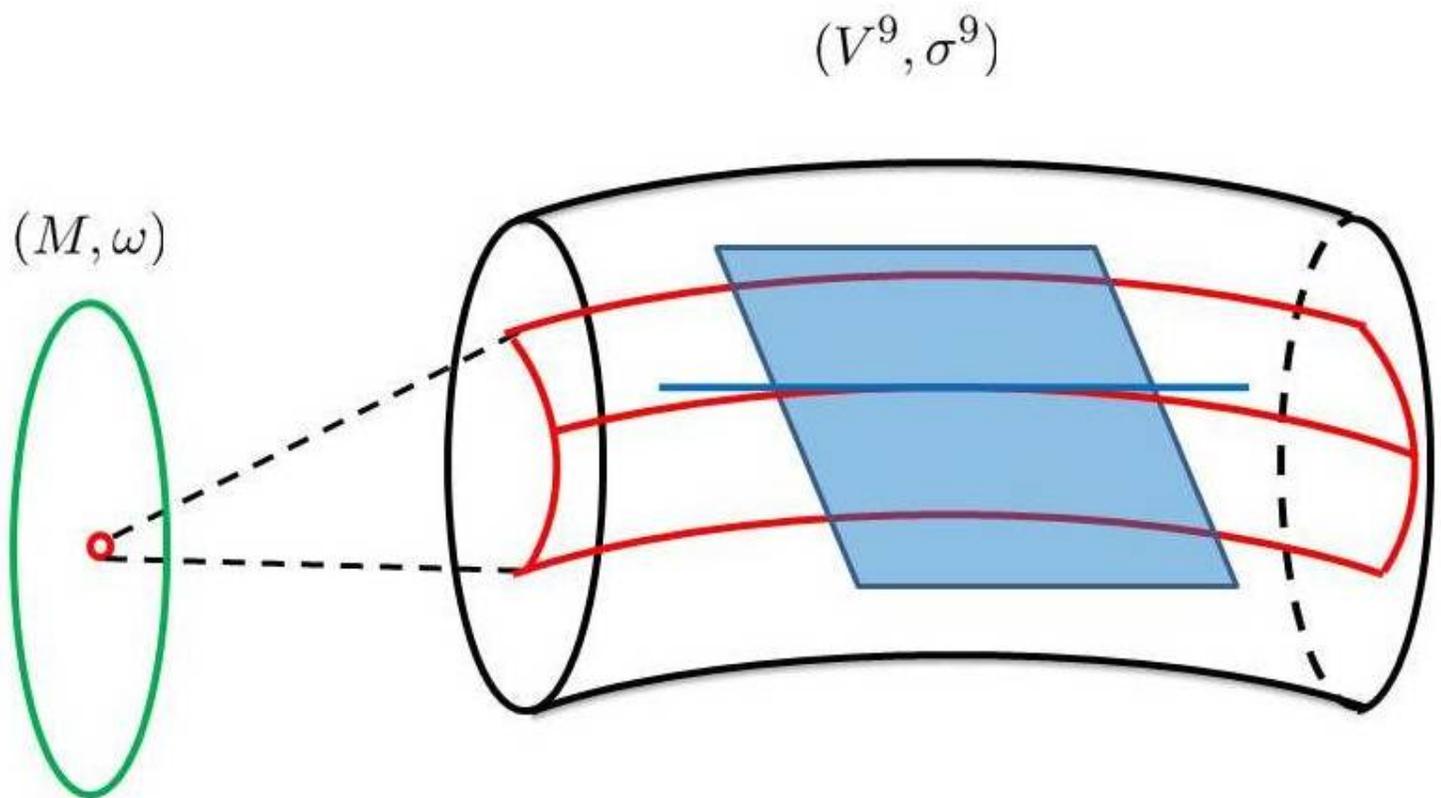
Souriau's mechanics

Dynamics determined by closed two-form σ of constant rank defined on “evolution space” V . Motions [curves or surfaces] \equiv “characteristic leaves”, tangent to kernel of two-form σ .



Can be viewed as generalized variational calculus.

Space of motions (M, ω) symplectic manifold, obtained from (V, σ) by factoring out “characteristic submanifolds” .



phase space [at time t] is section of evolution space for fixed value t .

Symplectic description of chiral model

Variation of chiral action (1) yields eqns of motion for position x and momentum $p \neq 0$ in 3-space,

$$\begin{cases} \left(1 + e\Theta \cdot B\right) \frac{dx}{dt} = \hat{p} + e\mathbf{E} \times \Theta + (\Theta \cdot \hat{p}) e B \\ \left(1 + e\Theta \cdot B\right) \frac{dp}{dt} = e\mathbf{E} + e\hat{p} \times B + e^2(\mathbf{E} \cdot B)\Theta, \end{cases} \quad (5)$$

\mathbf{E} , \mathbf{B} electric/magnetic field. “anomalous velocity”
 \rightsquigarrow “transverse shifts” or “side jumps” in spin-Hall-type effects.

Described within Souriau’s framework. Evolution space

$$V^7 = T(\mathbb{R}^3 \setminus \{0\}) \times \mathbb{R}, \quad (6)$$

endowed with two-form,

$$\omega = \omega_0 + \frac{e}{2} \epsilon_{ijk} B^i dx^j \wedge dx^k, \quad (7)$$

$$\omega_0 = dp_i \wedge dx^i - \frac{s}{2|\mathbf{p}|^3} \epsilon^{ijk} p_i dp_j \wedge dp_k, \quad (8)$$

$$h = |\mathbf{p}| + e\phi, \quad (9)$$

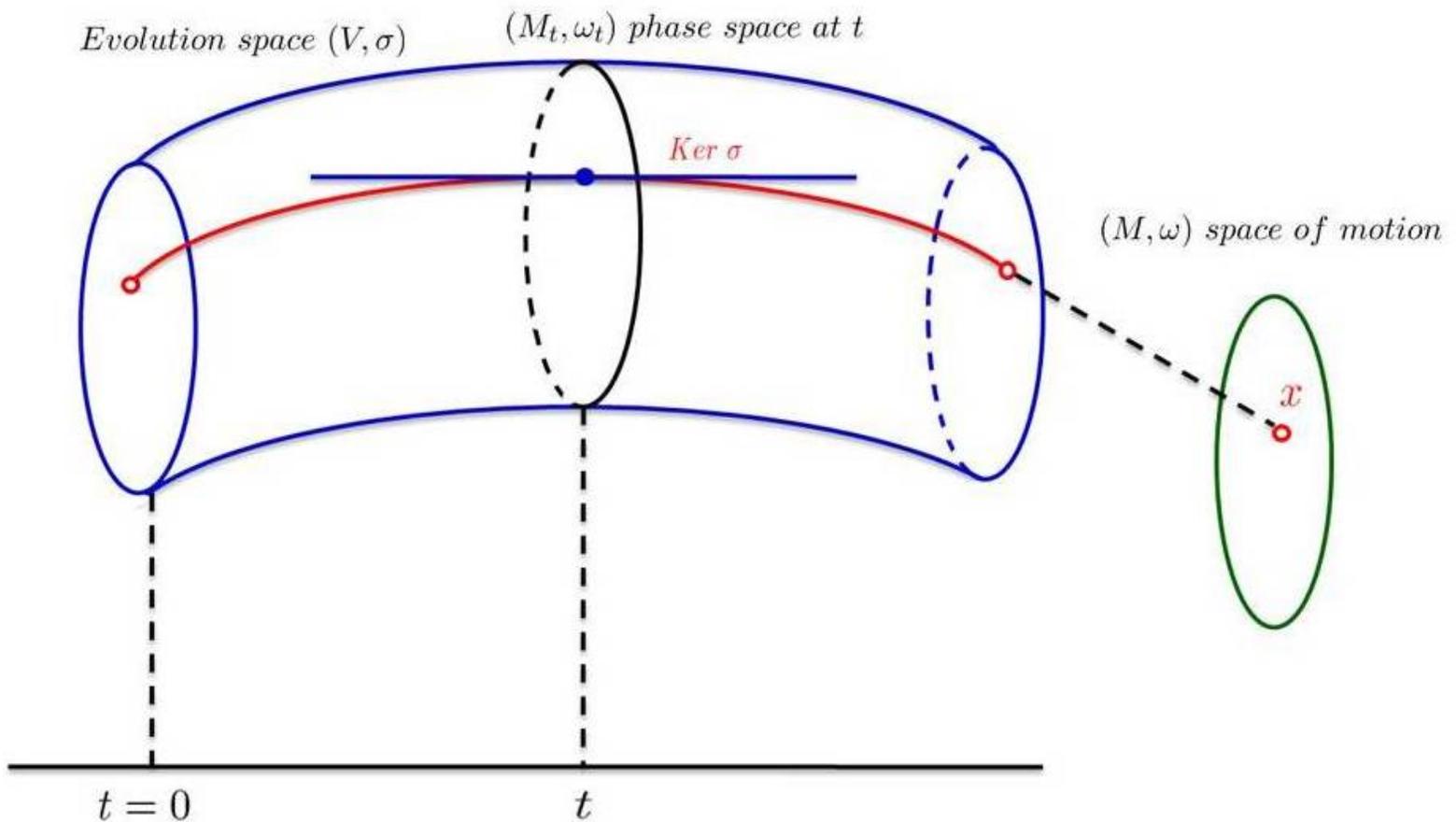
($s = 1/2$). $\omega \Rightarrow \sigma$ closed since

$$\nabla_{\mathbf{x}} \cdot \mathbf{B} = 0, \quad \nabla_{\mathbf{p}} \cdot \Theta = 0 \quad (10)$$

see (2). Where

$$\det(\omega_{\alpha\beta}) = (1 + e \Theta \cdot \mathbf{B})^2 \neq 0, \quad (11)$$

kernel of σ 1-dimensional; curve $(\mathbf{x}(\tau), \mathbf{p}(\tau), t(\tau))$ tangent to it iff eqns of motion (5) satisfied.



Souriau's Massless Spinning Particles

Souriau 1969 Free relativistic massless spinning particle described by 9-dim evolution space V^9 . Start with three four-vectors R, I, J in Minkowski space $\mathbb{R}^{3,1}$ with signature $(-, -, -, +)$.

$$V^9 = \left\{ R, I, J \in \mathbb{R}^{3,1} \mid I_\mu I^\mu = J_\mu J^\mu = 0, I_\mu J^\mu = -1 \right\} \quad (12)$$

I future-directed. Lightlike vectors I, J generate null 2-plane. R represents spacetime event.

Equivalent but more convenient description uses the spin tensor. Renaming $P = I$ (will be interpreted as linear momentum),

$$S_{\mu\nu} = -s \epsilon_{\mu\nu\rho\sigma} P^\rho J^\sigma. \quad (13)$$

Satisfies $\frac{1}{2} S_{\mu\nu} S^{\mu\nu} = s^2$, where $s \neq 0$ scalar spin (also called helicity). Pauli-Lubanski condition $S_{\mu\nu} P^\nu = 0$ satisfied.

Identifying $S = (S_{\mu\nu})$ with element of Lorentz algebra $\mathfrak{o}(3, 1)$, evolutions space also presented as

$$\left\{ R, P, S \mid P_\mu P^\mu = 0, S_{\mu\nu} P^\nu = 0, \frac{1}{2} S_{\mu\nu} S^{\mu\nu} = s^2 \right\} \quad (14)$$

V^9 endowed with closed two-form

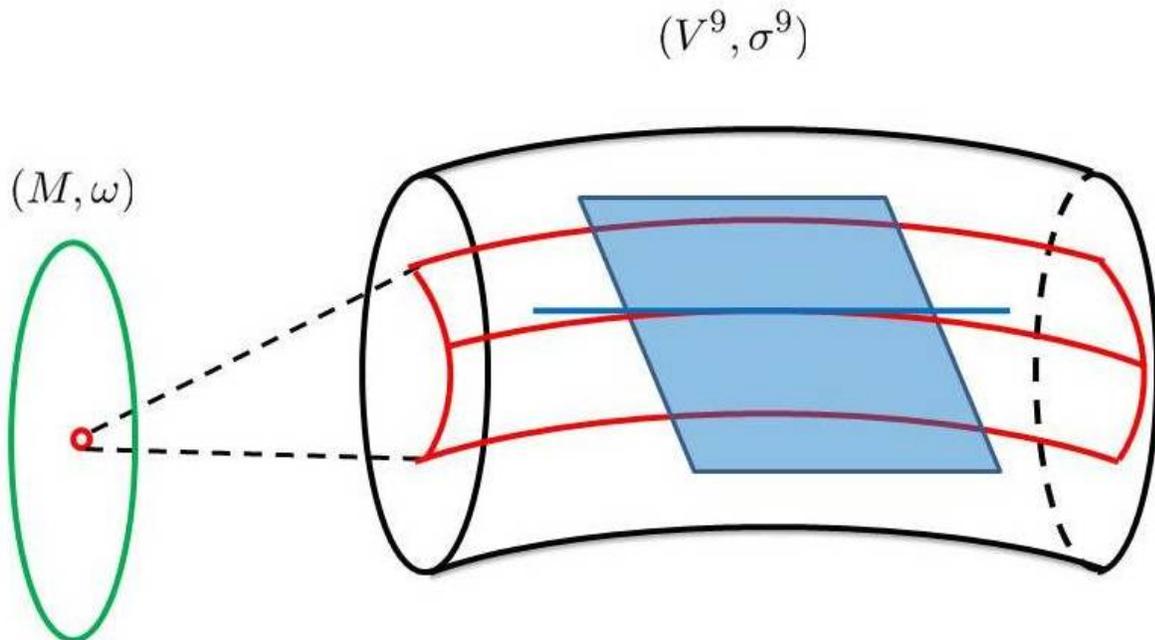
$$\sigma = -dP_\mu \wedge dR^\mu - \frac{1}{2s^2} dS^\mu_\lambda \wedge S^\lambda_\rho dS^\rho_\mu. \quad (15)$$

Dynamics given by foliation whose leaves are tangent to $\ker \sigma$; “world-sheet” [or world-line] obtained by projecting leaf to Minkowski space, yielding corresponding spacetime track.

Calculating kernel using constraints which define evolution space \Rightarrow curve $(R(\tau), P(\tau), S(\tau))$ tangent to $\ker \sigma$ iff

$$\begin{cases} P_\mu \dot{R}^\mu = 0, \\ \dot{P}^\mu = 0, \\ \dot{S}^{\mu\nu} = P^\mu \dot{R}^\nu - P^\nu \dot{R}^\mu, \end{cases} \quad (16)$$

where “dot” $\equiv d/d\tau$. Spacetime velocity \dot{R} orthogonal to momentum P .



Wigner translations

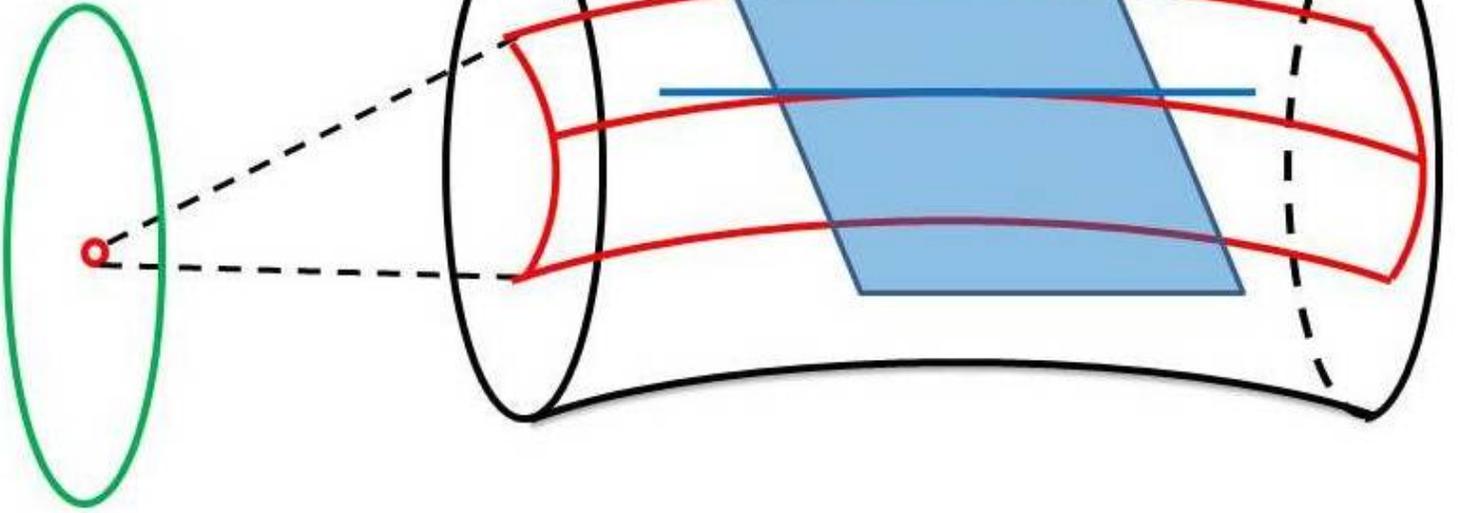
(16) integrated using spacetime vectors Z orthogonal to P , $P_\mu Z^\mu = 0$,

$$\begin{cases} R^\mu & \rightarrow R^\mu + Z^\mu, \\ P^\mu & \rightarrow P^\mu, \\ S^{\mu\nu} & \rightarrow S^{\mu\nu} + (P^\mu Z^\nu - P^\nu Z^\mu). \end{cases} \quad (17)$$

Any point in leaf reached by choosing suitable $Z \Rightarrow$ at each point kernel is 3-dim, projects to spacetime as affine subspace of $\mathbb{R}^{3,1}$, spanned by all vectors orthogonal to linear momentum $P \Rightarrow$ “motions” of free massless spinning particle take place on 3-dimensional “plane-wave” tangent to light-cone at each spacetime event R : particle is *not localized* in spacetime.

Wigner 39, Penrose 67, Souriau 69.

All curves which lie in a leaf (left invariant by a “ Z -shift” in (17)) should be considered as same motion. Each (3-dimensional) leaf defines therefore “motion” of particle. Space of motions is collection $M^6 = V^9 / \ker \sigma$ of leaves. Explained below: spin is responsible for space-time delocalization of massless particles.

(V^9, σ^9) (M, ω) 

Free massless spinning particle has 9-dim evolution space V^9 . Dynamics defined by two-form σ . $\ker \sigma$ 3-dimensional, whose points reached by a “Z-shift” (“Wigner translation”). Characteristic leaves project into Minkowski as 3-planes orthogonal to momentum : massless spinning particle can not be localized.

Put $R = (\mathbf{r}, t)$ where \mathbf{r} and t are position & time in chosen Lorentz frame. Null-vectors $P = (\mathbf{p}, |\mathbf{p}|)$ and $J = (\mathbf{q}, -|\mathbf{q}|)$, where \mathbf{p} and \mathbf{q} two 3-vectors which satisfy $\mathbf{p} \cdot \mathbf{q} + |\mathbf{p}| |\mathbf{q}| = 1$.

$$S_{ij} = \epsilon_{ijk} s^k, \quad \mathbf{s} = s(\mathbf{p}|\mathbf{q}| + \mathbf{q}|\mathbf{p}|), \quad (18)$$

$$S_{j4} = s(\mathbf{p} \times \mathbf{q})_j = (\hat{\mathbf{p}} \times \mathbf{s})_j. \quad (19)$$

Important observation

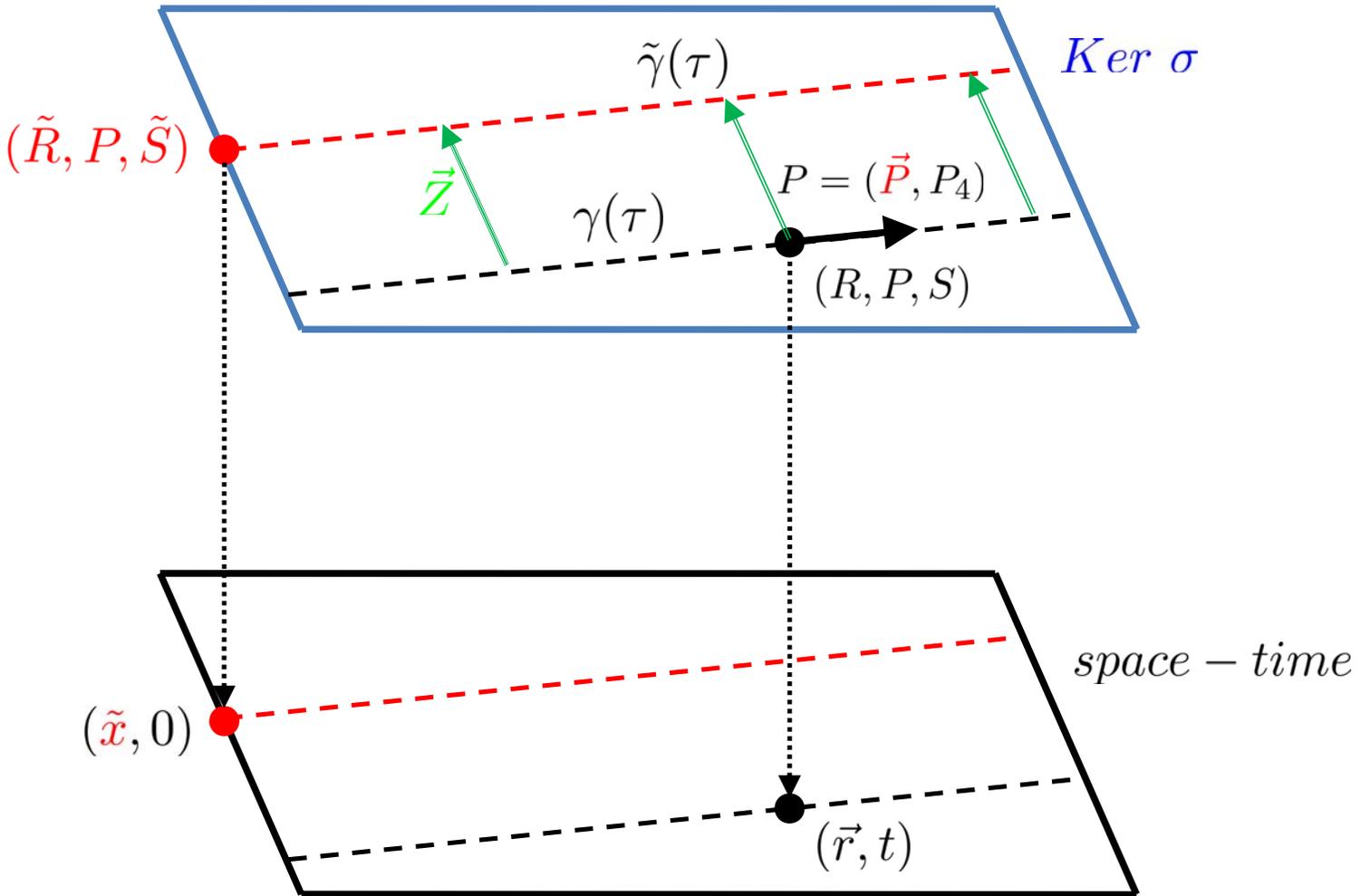
$$\hat{p} \cdot s = s \quad (20)$$

in general. s **not** length of s but *projection onto \hat{p}* , which is a constant.

Wigner transl. used to **“enslave” spin**

$$s = s\hat{p} \quad (21)$$

SPIN eliminated as indept degree of freedom.



The spin of a motion tangent to (3d) *kernel of σ* can be “enslaved” by suitable “Z-shift”. Choosing point on shifted curve with vanishing time coordinate provides us with coordinate \tilde{x} of “motion”. The displaced motions sweep a 3-plane under all Z-shifts.

Two-form σ descends to space of motions M^6 as *symplectic two-form*

$$\omega = d\tilde{p}_i \wedge d\tilde{x}^i - \frac{s}{2|\tilde{\mathbf{p}}|^3} \epsilon^{ijk} \tilde{p}_i d\tilde{p}_j \wedge d\tilde{p}_k. \quad (22)$$

Poincaré symmetry

Poincaré algebra spanned by (Λ, Γ) where $\Lambda = (\Lambda_{\mu\nu})$ belongs to Lorentz algebra and $\Gamma = (\Gamma^\mu)$ is translation in Minkowski space. Acts on V^9 :

$$\begin{cases} \delta R^\mu = \Lambda^\mu_\nu R^\nu + \Gamma^\mu, \\ \delta P^\mu = \Lambda^\mu_\nu P^\nu, \\ \delta S_{\mu\nu} = \Lambda^\mu_\rho S^{\rho\nu} - \Lambda^\nu_\rho S^{\rho\mu}. \end{cases} \quad (23)$$

Leaves two-form (15) invariant \rightsquigarrow symmetry. Action descends to space of motions (M^6, ω) . Noetherian conserved quantities

$$P^\mu = I^\mu, \quad M^{\mu\nu} = R^\mu P^\nu - R^\nu P^\mu + S^{\mu\nu} \quad (24)$$

conserved linear and angular momentum.

In (3+1)-decomposition: parametrize Poincaré algebra by $\Lambda_{ij} = \epsilon_{ijk} \omega^k$, $\Lambda_{i4} = \beta^i$, $\Gamma = (\gamma, \varepsilon)$, $\omega, \beta, \gamma \in \mathbb{R}^3$, $\varepsilon \in \mathbb{R}$ (rotations, boosts and space- resp. time-translations). Inf. Poincaré-action on V^9 given by

$$\begin{cases} \delta \mathbf{r} = \boldsymbol{\omega} \times \mathbf{r} + \boldsymbol{\beta} t + \boldsymbol{\gamma}, \\ \delta t = \boldsymbol{\beta} \cdot \mathbf{r} + \varepsilon, \\ \delta \mathbf{p} = \boldsymbol{\omega} \times \mathbf{p} + \boldsymbol{\beta} |\mathbf{p}|, \\ \delta \mathbf{s} = \boldsymbol{\omega} \times \mathbf{s} - \boldsymbol{\beta} \times (\hat{\mathbf{p}} \times \mathbf{s}), \end{cases} \quad (25)$$

projects as natural action on Minkowski space.

To write down explicit form of Poincaré momenta, present matrix $M = (M_{\mu\nu})$ as $M_{ij} = \epsilon_{ijk} \ell^k$, $M_{j4} = g^j$ with ℓ , g 3-vectors.

$$\begin{cases} \ell = \mathbf{r} \times \mathbf{p} + \mathbf{s}, \\ \mathbf{g} = |\mathbf{p}| \mathbf{r} - \mathbf{p}t + \hat{\mathbf{p}} \times \mathbf{s}. \end{cases} \quad (26)$$

Then

$$\boxed{\tilde{\mathbf{x}} = \frac{\mathbf{g}}{|\mathbf{p}|} = \mathbf{r} - \hat{\mathbf{p}}t + \frac{\hat{\mathbf{p}}}{|\mathbf{p}|} \times \mathbf{s}} \quad (27)$$

itself conserved.

Full Poincaré algebra acts on space of motions as

$$\begin{cases} \delta \tilde{\mathbf{p}} = \boldsymbol{\omega} \times \tilde{\mathbf{p}} + |\tilde{\mathbf{p}}| \boldsymbol{\beta}, \\ \delta \tilde{\mathbf{x}} = \boldsymbol{\omega} \times \tilde{\mathbf{x}} + \boldsymbol{\beta} \times \left(s \frac{\tilde{\mathbf{p}}}{|\tilde{\mathbf{p}}|^2} \right) - \boldsymbol{\beta} \cdot \tilde{\mathbf{x}} \frac{\tilde{\mathbf{p}}}{|\tilde{\mathbf{p}}|} \\ \quad + \gamma - \varepsilon \frac{\tilde{\mathbf{p}}}{|\tilde{\mathbf{p}}|}. \end{cases} \quad (28)$$

10-parameter vectorfield (28) leaves free symplectic structure invariant, i.e., generates family of symmetries \rightsquigarrow Noether thm \rightsquigarrow 10 constants of the motion,

$$\left\{ \begin{array}{ll} \boldsymbol{\ell} = \tilde{\boldsymbol{x}} \times \tilde{\boldsymbol{p}} + s\hat{\boldsymbol{p}} & \text{angular momentum} \\ \boldsymbol{g} = |\tilde{\boldsymbol{p}}| \tilde{\boldsymbol{x}} & \text{boost momentum} \\ \boldsymbol{p} = \tilde{\boldsymbol{p}} & \text{linear momentum} \\ \mathcal{E} = |\tilde{\boldsymbol{p}}| & \text{energy} \end{array} \right. \quad (29)$$

Conservation follows also directly from free eqns of motions. NB: two terms in angular momentum $\boldsymbol{\ell}$ separately conserved.

Poisson brackets of calculated using (22),

$$\begin{aligned} \{\ell_i, \ell_j\} &= -\epsilon_{ij}^k \ell_k, & \{\ell_i, g_j\} &= -\epsilon_{ij}^k g_k, & \{\ell_i, p_j\} &= -\epsilon_{ij}^k p_k, \\ \{\ell_i, \mathcal{E}\} &= 0, & \{g_i, g_j\} &= \epsilon_{ij}^k \ell_k, & \{g_i, p_j\} &= -\mathcal{E} \delta_{ij}, \\ \{g_i, \mathcal{E}\} &= -p_i, & \{p_i, p_j\} &= 0, & \{p_i, \mathcal{E}\} &= 0, \end{aligned} \quad (30)$$

Poincaré algebra. Casimir invariants

$$m^2 = -\boldsymbol{p}^2 + \mathcal{E}^2 = 0, \quad \boldsymbol{\ell} \cdot \hat{\boldsymbol{p}} = s, \quad (31)$$

zero-mass and spin- s representation .

Twisted Poincaré symmetry of free chiral model

Poincaré symmetry for free chiral system ?

In free case $\mathbf{E} = \mathbf{B} = 0$ motions determined explicitly: Θ -term drops out from (5),

$$\begin{cases} \left(1 + e\Theta \cdot \mathbf{B}\right) \frac{d\mathbf{x}}{dt} = \hat{\mathbf{p}} + e\mathbf{E} \times \Theta + (\Theta \cdot \hat{\mathbf{p}}) e\mathbf{B}, \\ \left(1 + e\Theta \cdot \mathbf{B}\right) \frac{d\mathbf{p}}{dt} = e\mathbf{E} + e\hat{\mathbf{p}} \times \mathbf{B} + e^2(\mathbf{E} \cdot \mathbf{B}) \Theta, \end{cases}$$

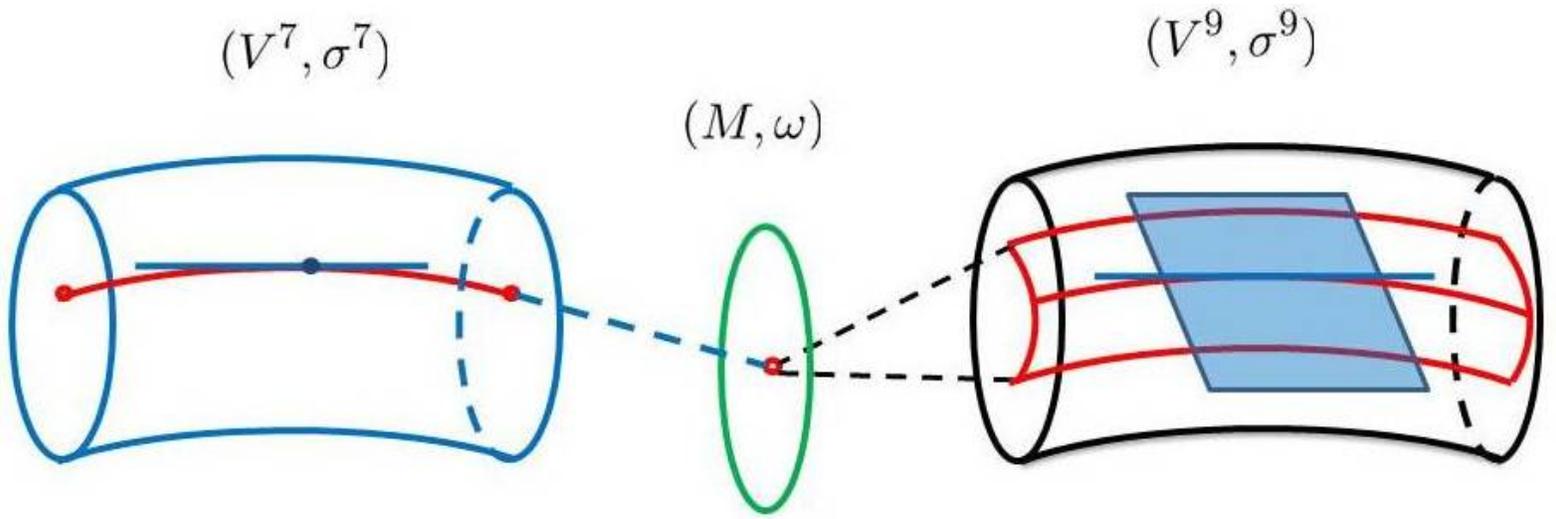
integrated as:

$$\mathbf{x}(t) = \tilde{\mathbf{x}} + \hat{\mathbf{p}}t, \quad \mathbf{p}(t) = \tilde{\mathbf{p}}, \quad (32)$$

with $\tilde{\mathbf{x}}, \tilde{\mathbf{p}}$ const. Space of motions $M^6 = V^7 / \ker \sigma$ described by $\tilde{\mathbf{x}} = \mathbf{x}(t) - \hat{\mathbf{p}}t$ and $\tilde{\mathbf{p}}$. Two-form ω is (22) :

$$\omega = \omega_0 = dp_i \wedge dx^i - \frac{s}{2|\mathbf{p}|^3} \epsilon^{ijk} p_i dp_j \wedge dp_k.$$

Free chiral model has same space of motions as massless spinning particle.



Chiral and massless spinning systems admit same space of motions.

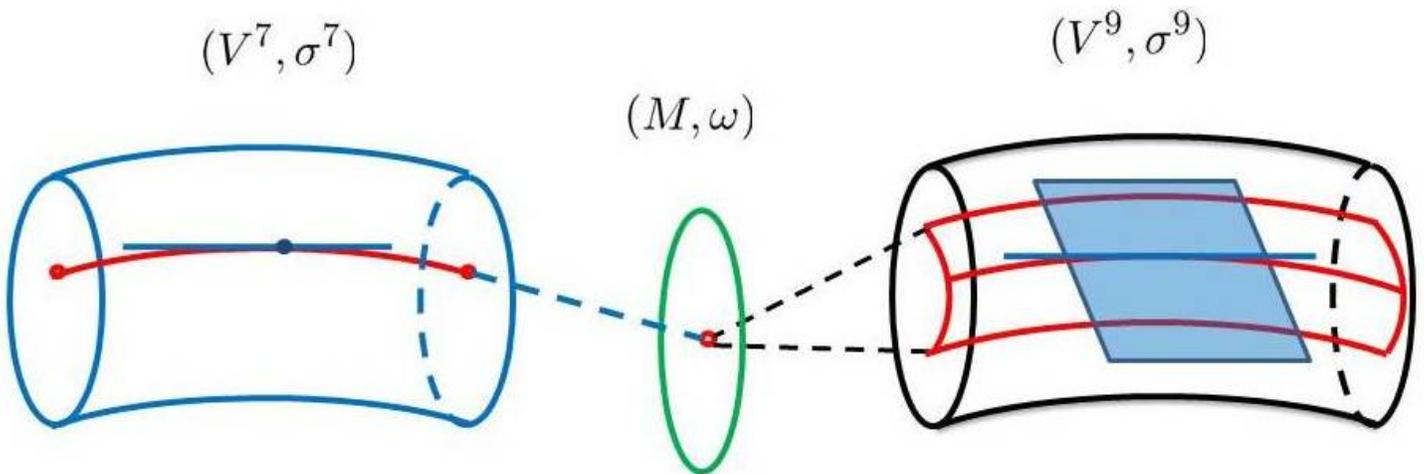
Strategy: import" natural Poincaré symmetry of Souriau model to chiral system through their common space of motions (M, ω) .

From identity of space of motions coordinates conclude that position of chiral particle and of massless Poincaré model are same, $\mathbf{x} = \mathbf{r}$. In terms of coordinates $(\mathbf{x}, \mathbf{p}, t)$ on chiral evolution space V^7 , Poincaré action (28) becomes "twisted",

$$\begin{cases} \delta \mathbf{x} = \boldsymbol{\omega} \times \mathbf{x} + \boldsymbol{\beta} \times \frac{\hat{\mathbf{p}}}{2|\mathbf{p}|} + \boldsymbol{\beta} t + \boldsymbol{\gamma}, \\ \delta \mathbf{p} = \boldsymbol{\omega} \times \mathbf{p} + |\mathbf{p}| \boldsymbol{\beta}, \\ \delta t = \boldsymbol{\beta} \cdot \mathbf{x} + \varepsilon. \end{cases} \quad (33)$$

Vector fields generate same (Poincaré) algebra as in (28). Eqn (33) confirms and extends recently proposed action of Lorentz subalgebra **Chen, Son, Stephanov ...** Conserved Lorentz quantities,

$$\begin{cases} \ell = \mathbf{x} \times \mathbf{p} + \frac{1}{2} \hat{\mathbf{p}}, \\ \mathbf{g} = |\mathbf{p}| \mathbf{x} - \mathbf{p}t. \end{cases} \quad (34)$$



Identity of spaces of motion allows to “export” natural Poincaré symmetry of latter to become “twisted” (“dynamical”) symmetry of former.

N.B. action (33) is **not** usual, natural one on ordinary spacetime. In fact, it is **not** action on spacetime at all, since it also involves momentum variable \mathbf{p} ; is sort of “dynamical symmetry”. \mathbf{x} should **not** be viewed as position variable, because *does not transform* under a boost as positions should.

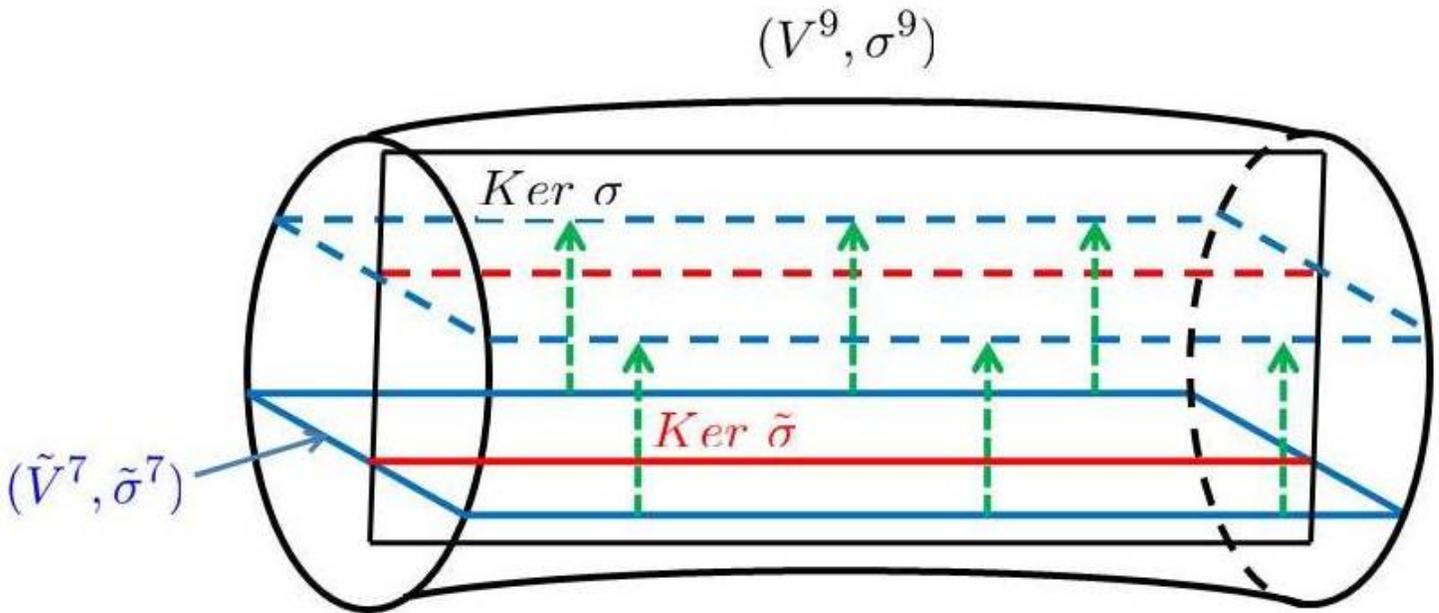
Embedding chiral system into massless spinning model

Further insight gained by *embedding evolution space of chiral system, V^7* , into that, V^9 , of *massless spinning particle* by “enslaving spin”.
Condition

$$S_{j4} = (\hat{\mathbf{p}} \times \mathbf{s})_j = 0 \quad (35)$$

defines submanifold \tilde{V}^7 of V^9 identified with V^7 ; restriction of free two-form of V^9 to \tilde{V}^7 is (7)–(8).

Dynamics consistent with embedding. Motions of chiral system lie within leaves of 3-dimensional foliation of V^9 and remain therefore motions also for extended system.



Chiral evolution space can be embedded, $V^7 \rightarrow \tilde{V}^7$, into that, V^9 , of massless spinning particle by enslaving spin by requiring $\mathbf{p} \times \mathbf{s} = 0$. Constraint is **not** Lorentz-invariant : Lorentz boost β does not leave $\mathbf{p} \times \mathbf{s}$ invariant.

Lorentz symmetry inherited from embedding ?

no : consider chiral motion $\mathbf{r} - \hat{\mathbf{p}}t$ embedded into V^9 s.t. $\mathbf{s} = s\hat{\mathbf{p}}$ and boost. Although spin is boost-invariant,

$$\delta(\mathbf{p} \times \mathbf{s}) = s\beta \times \mathbf{p} \neq 0 \quad (36)$$

in general, \Rightarrow spin and momentum do not remain parallel : embedding chiral system through spin enslavement **is not** boost invariant.

While full Poincaré group acts on V^9 , it is only its “Aristotle subgroup” generated by space rotations and by space- and time translations which leaves \tilde{V}^7 invariant : *natural boosts are broken* by spin enslavement.

Lorentz symmetry restored by “unchaining spin” .
 Boosting a trajectory in V^9 according to (25),

$$\begin{aligned}\delta(\mathbf{r} - \hat{\mathbf{p}}t) &= -\hat{\mathbf{p}}(\boldsymbol{\beta} \cdot (\mathbf{r} - \hat{\mathbf{p}}t)), \\ \delta\left(\frac{\hat{\mathbf{p}}}{|\mathbf{p}|} \times \mathbf{s}\right) &= \boldsymbol{\beta} \times \left(s \frac{\mathbf{p}}{|\mathbf{p}|^2}\right) - \frac{\mathbf{p}}{|\mathbf{p}|^2}(\boldsymbol{\beta} \cdot (\hat{\mathbf{p}} \times \mathbf{s})).\end{aligned}$$

Terms combine to yield action (28) on space of motions,

$$\delta\tilde{\mathbf{x}} = \delta\left(\mathbf{r} - \hat{\mathbf{p}}t + \frac{\hat{\mathbf{p}}}{|\mathbf{p}|} \times \mathbf{s}\right) = \boldsymbol{\beta} \times \left(s \frac{\hat{\mathbf{p}}}{|\mathbf{p}|}\right) - \hat{\mathbf{p}}(\boldsymbol{\beta} \cdot \tilde{\mathbf{x}}). \quad (37)$$

Coupling to electromagnetic field

Conventional “minimal coupling” rule : 4-momentum shifted by 4-potential,

$$p_\mu \rightarrow p_\mu - eA_\mu. \quad (38)$$

not what is proposed in (1): rule (38) used for 4-momentum (p, h) , momentum in “Berry term” $\Theta(p)$ **not** shifted.

“half-way-rule” in (1) is instead consistent with Souriau’s prescription, who works with same evolution space as for free particle, but adds electromagnetic field strength F to free two-form (15) ,

$$\sigma \rightarrow \sigma + eF, \quad (39)$$

where e is electric charge. σ closed, $d\sigma = 0$, because F is closed 2-form of Minkowski space.

Rules (38) and (39) equivalent in spinless case only. Why should (39) be chosen ? Argument in its favor comes from experience in plane, where yielded insight into Hall-type phenomena cf. Duval et al Phys. Lett. **B 479**, 284 (2000).

Minimal coupling of massless spinning model

Souriau's prescription (39) applied to massless spinning model yields, on evolution space V^9 , closed two-form

$$\sigma = -dP_\mu \wedge dR^\mu - \frac{1}{2s^2} dS^\mu_\lambda \wedge S^\lambda_\rho dS^\rho_\mu + \frac{e}{2} F_{\mu\nu} dR^\mu \wedge dR^\nu. \quad (40)$$

Lengthy calculation \Rightarrow eqns of motions (16) change into

$$\left\{ \begin{array}{l} \dot{R}^\mu = P^\mu + \frac{S^{\mu\nu} F_{\nu\rho} P^\rho}{\frac{1}{2} S \cdot F}, \\ \dot{P}^\mu = -e F^\mu_\nu \dot{R}^\nu, \\ \dot{S}^{\mu\nu} = P^\mu \dot{R}^\nu - P^\nu \dot{R}^\mu \end{array} \right. \quad (41)$$

assuming

$$S \cdot F \equiv S_{\alpha\beta} F^{\alpha\beta} \neq 0. \quad (42)$$

Dim of $\ker \sigma$ drops $3 \rightarrow 1$: spin-field coupling breaks "Z-shift" invariance \Rightarrow spin degree can not be eliminated, we are left with $9 - 1 = 8$ -dim. *space of motions* (phase space locally).

In terms of (3 + 1)-decomposition : assuming

$$(a) \quad \frac{1}{2}S \cdot F \equiv \frac{1}{2}S_{\alpha\beta}F^{\alpha\beta} = s \cdot (\mathbf{B} - \hat{\mathbf{p}} \times \mathbf{E}) \neq 0,$$

$$(b) \quad \hat{\mathbf{p}} \cdot \mathbf{B} \neq 0,$$

$$\left\{ \begin{array}{l} \frac{d\mathbf{r}}{dt} = \frac{\mathbf{B} - \hat{\mathbf{p}} \times \mathbf{E}}{\hat{\mathbf{p}} \cdot \mathbf{B}}, \\ \frac{d\mathbf{p}}{dt} = e \left(\mathbf{E} + \frac{d\mathbf{r}}{dt} \times \mathbf{B} \right) = e \frac{\mathbf{E} \cdot \mathbf{B}}{\hat{\mathbf{p}} \cdot \mathbf{B}} \hat{\mathbf{p}}, \\ \frac{ds}{dt} = \mathbf{p} \times \frac{d\mathbf{r}}{dt} = \frac{\mathbf{p} \times \mathbf{B}}{\hat{\mathbf{p}} \cdot \mathbf{B}} - \frac{\mathbf{p} \times (\hat{\mathbf{p}} \times \mathbf{E})}{\hat{\mathbf{p}} \cdot \mathbf{B}}. \end{array} \right. \quad (43)$$

Unusual features :

- $\hat{\mathbf{p}}$ one would have expected on r.h.s. of velocity relation cancels out;
- electric charge drops out;
- Dynamics of the momentum decouples from spin as long as latter does not vanish; scalar spin $s \neq 0$ disappears also from all equations.

Eqs (43) $\Rightarrow d\hat{\mathbf{p}}/dt = 0 \Rightarrow$ direction of \mathbf{p} unchanged during motion.

Put, for example, into crossed constant electric & magnetic fields (like in Hall effect),

$$\mathbf{E} = E \hat{x}, \quad \mathbf{B} = B \hat{z}$$

$\mathbf{E} \cdot \mathbf{B} = 0 \Rightarrow \mathbf{p}$ itself const. of motion, \Rightarrow so is angle θ between \mathbf{B} and \mathbf{p} ($\theta \neq \pi/2$ for $\mathbf{p} \cdot \mathbf{B} \neq 0$).

Assume for simplicity that initial momentum lies in x - z plane. Then eqns of motion solved by

$$\left\{ \begin{array}{l} \mathbf{r}(t) = \left((\cos \theta)^{-1} \hat{z} - \frac{E}{B} \hat{y} \right) t + \mathbf{r}_0, \\ \mathbf{p}(t) = \mathbf{p}_0, \\ \mathbf{s}(t) = |\mathbf{p}| \left(-\tan \theta \hat{y} + \frac{E}{B} (\cos \theta \hat{x} - \sin \theta \hat{z}) \right) t + \mathbf{s}_0. \end{array} \right. \quad (44)$$

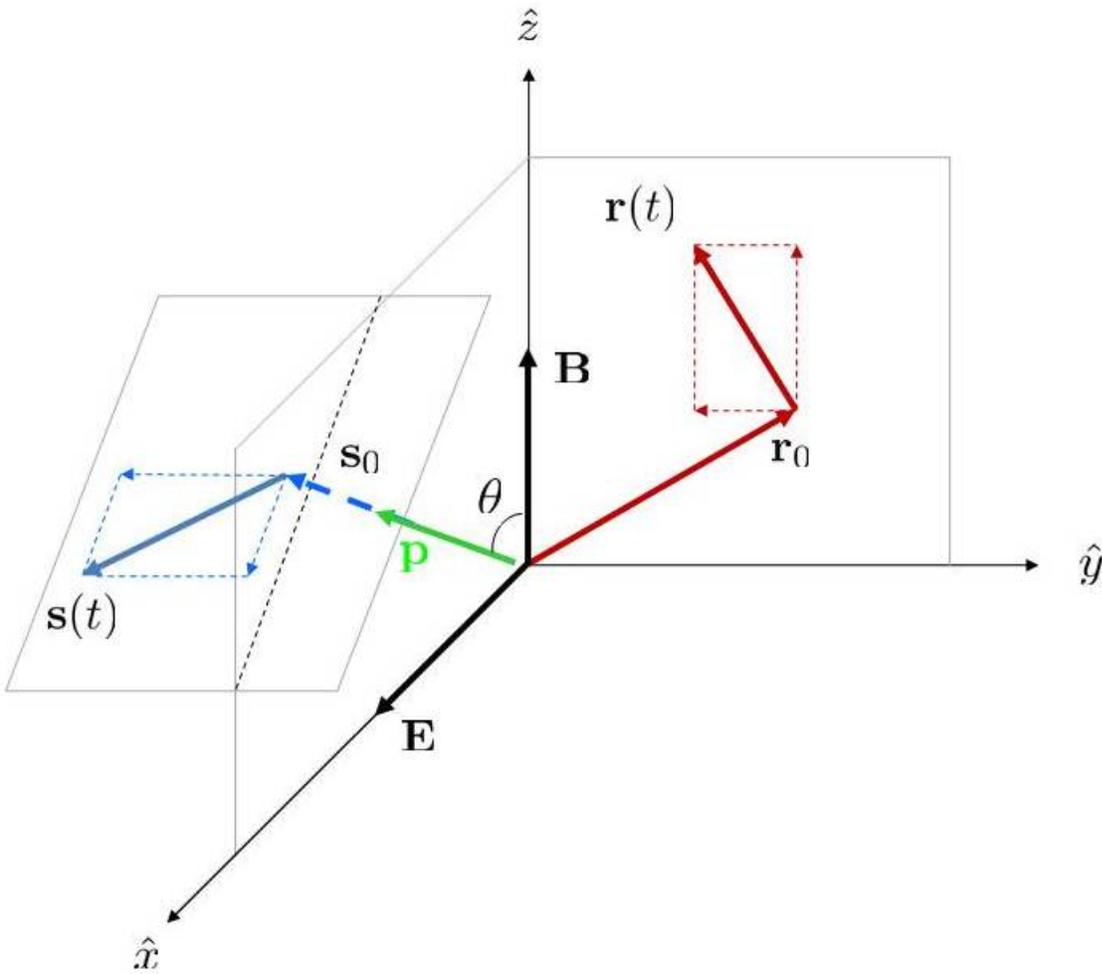
Thus, in addition to const.-speed vertical motion, particle also drifts perpendicularly to electric field with *Hall velocity* E/B .

Spin follows curious motion perpendicularly to

$\hat{\mathbf{p}} \Rightarrow$ projection on $\hat{\mathbf{p}}$ remains const, $\mathbf{s}(t) \cdot \hat{\mathbf{p}} = \mathbf{s}_0 \cdot \hat{\mathbf{p}}$.

Spin decoupled, but can **not** consistently “enslaved”

\mathbf{s} and $\hat{\mathbf{p}}$ do **not** remain parallel.



Motion in constant Hall-type electric and magnetic fields described by eqn (44). Initial position, \mathbf{r}_0 , chosen in y - z plane, initial momentum and spin chosen parallel and in x - z plane. Spatial motion, $\mathbf{r}(t)$, is combination of constant-velocity Hall drift perp. to \mathbf{E} and \mathbf{B} , combined with const-velocity vertical drift. Momentum, \mathbf{p} , conserved, but spin, $\mathbf{s}(t)$, has curious, Hall-type motion in plane perp to momentum.

Velocity **superluminal** ($|\mathbf{dr}/\mathbf{dt}| > 1$), diverges as $\theta \rightarrow \pi/2$; for $\hat{\mathbf{p}} \cdot \mathbf{B} = 0$ instantaneous motions with infinite velocity, parallel to z axis.

Anomalous coupling

Minimal model curious, but not completely satisfactory. Generalize. Clue : allow “mass-square” to depend on coupling of spin to e.m. field,

$$P_\mu P^\mu = -\frac{eg}{2} S \cdot F, \quad (45)$$

where used shorthand $S \cdot F \equiv S_{\alpha\beta} F^{\alpha\beta}$. Real constant g : gyromagnetic ratio.

$$P^\mu = I^\mu + \frac{eg}{4} (S \cdot F) J^\mu, \quad (46)$$

where I, J still as in (12). Allows to implement (45). **Pauli-Lubanski** condition

$$S_{\mu\nu} P^\nu = 0 \quad (47)$$

automatically satisfied.

Novel evolution space

$$\left\{ P, R, S \mid P_\mu P^\mu = -\frac{eg}{2} S \cdot F, S_{\mu\nu} P^\nu = 0, \frac{1}{2} S_{\mu\nu} S^{\mu\nu} = s^2 \right\}, \quad (48)$$

endowed with closed two-form,

$$\sigma = -dP_\mu \wedge dR^\mu - \frac{1}{2s^2} dS_\lambda^\mu \wedge S_\rho^\lambda dS_\mu^\rho + \frac{1}{2} e F_{\mu\nu} dR^\mu \wedge dR^\nu. \quad (49)$$

formally same as (40) up to different mass-shell constraint.

New eqns of motion from kernel using constraints which define \tilde{V}^9 .

$$\begin{aligned} \dot{R}^\mu &= P^\mu - \frac{1}{(g+1)} \frac{1}{S_{\alpha\beta} F^{\alpha\beta}} \left[(g-2) S^{\mu\nu} F_{\nu\rho} P^\rho - g S^{\mu\nu} \partial_\nu F_{\rho\sigma} S^{\rho\sigma} \right], \\ \dot{P}^\mu &= -e F_{\nu}^\mu \dot{R}^\nu - \frac{eg}{4} \partial^\mu F_{\rho\sigma} S^{\rho\sigma}, \\ \dot{S}^{\mu\nu} &= P^\mu \dot{R}^\nu - P^\nu \dot{R}^\mu + \frac{eg}{2} \left[S_\rho^\mu F^{\rho\nu} - S_\rho^\nu F^{\rho\mu} \right]. \end{aligned} \quad (50)$$

Zero-rest-mass counterparts of Bargmann-Michel-Telegdi eqns for massless particle. Reduce to (40) for $g = 0$. In “normal” case $g = 2 \sim$ Dirac eqn anomalous velocity canceled, but new contribution involving derivative of em field.

(3 + 1)-decomposition:

$$R = (\mathbf{r}, t), \quad P = (\mathbf{p}, \mathcal{E}), \quad S_{j4} = \left(\frac{\mathbf{p}}{\mathcal{E}} \times \mathbf{s} \right)_j \quad (51)$$

where spin tensor still defined as in (13), but new dispersion relation,

$$\mathcal{E} = \sqrt{|\mathbf{p}|^2 - \frac{eg}{2} S \cdot F}. \quad (52)$$

Decomposing em field into electric and magnetic components,

$$\frac{1}{2} S \cdot F = \mathbf{s} \cdot \left(\mathbf{B} - \frac{\mathbf{p}}{\mathcal{E}} \times \mathbf{E} \right). \quad (53)$$

Tedious calculation yields (3 + 1)-form of eqns of motion (50).

N.B. For weak pure magnetic field, $\mathbf{s} = \frac{1}{2} \hat{\mathbf{p}}$

$$\mathcal{E} \approx |\mathbf{p}| - \frac{eg}{4} S \cdot F = |\mathbf{p}| - e \frac{\hat{\mathbf{p}} \cdot \mathbf{B}}{2|\mathbf{p}|}, \quad (54)$$

also proposed recently by **Chen et al**, **Manuel et al**.

Consider $g = 2 +$ fields constant; field-derivative terms & anomalous velocity drop out. Simplifies to one reminiscent of a massive relativistic particle,

$$(g=2) \left\{ \begin{array}{l} \mathcal{E} \frac{d\mathbf{r}}{dt} = \mathbf{p}, \\ \frac{d\mathbf{p}}{dt} = e \left(\mathbf{E} + \frac{d\mathbf{r}}{dt} \times \mathbf{B} \right), \\ \frac{d\mathbf{s}}{dt} = e \left(\left(\frac{\mathbf{p}}{\mathcal{E}^2} \times \mathbf{s} \right) \times \mathbf{E} + \frac{\mathbf{s}}{\mathcal{E}} \times \mathbf{B} \right) \end{array} \right. \quad (55)$$

assuming that $\mathcal{E} \neq 0$ [sort of “effective mass”] is real. [NB: $\mathbf{p} \neq 0$ implies that \mathcal{E} can not vanish].

In pure magnetic field momentum and spin satisfy eqns of identical form,

$$\frac{d\mathbf{p}}{dt} = \frac{e}{\mathcal{E}} \mathbf{p} \times \mathbf{B}, \quad \frac{d\mathbf{s}}{dt} = \frac{e}{\mathcal{E}} \mathbf{s} \times \mathbf{B}. \quad (56)$$

Multiplying by \mathbf{p} , \mathbf{s} and by \mathbf{B} , resp. \Rightarrow

$$\begin{aligned} |\mathbf{p}| &= \text{const} \neq 0, & \mathbf{p} \cdot \mathbf{B} &= \text{const}, \\ |\mathbf{s}| &= \text{const} \neq 0, & \mathbf{s} \cdot \mathbf{B} &= \text{const}, \\ p_z &= \text{const}, & s_z &= \text{const}, \\ \mathcal{E} &= \sqrt{|\mathbf{p}|^2 - e\mathbf{s} \cdot \mathbf{B}} = \text{const}. \end{aligned}$$

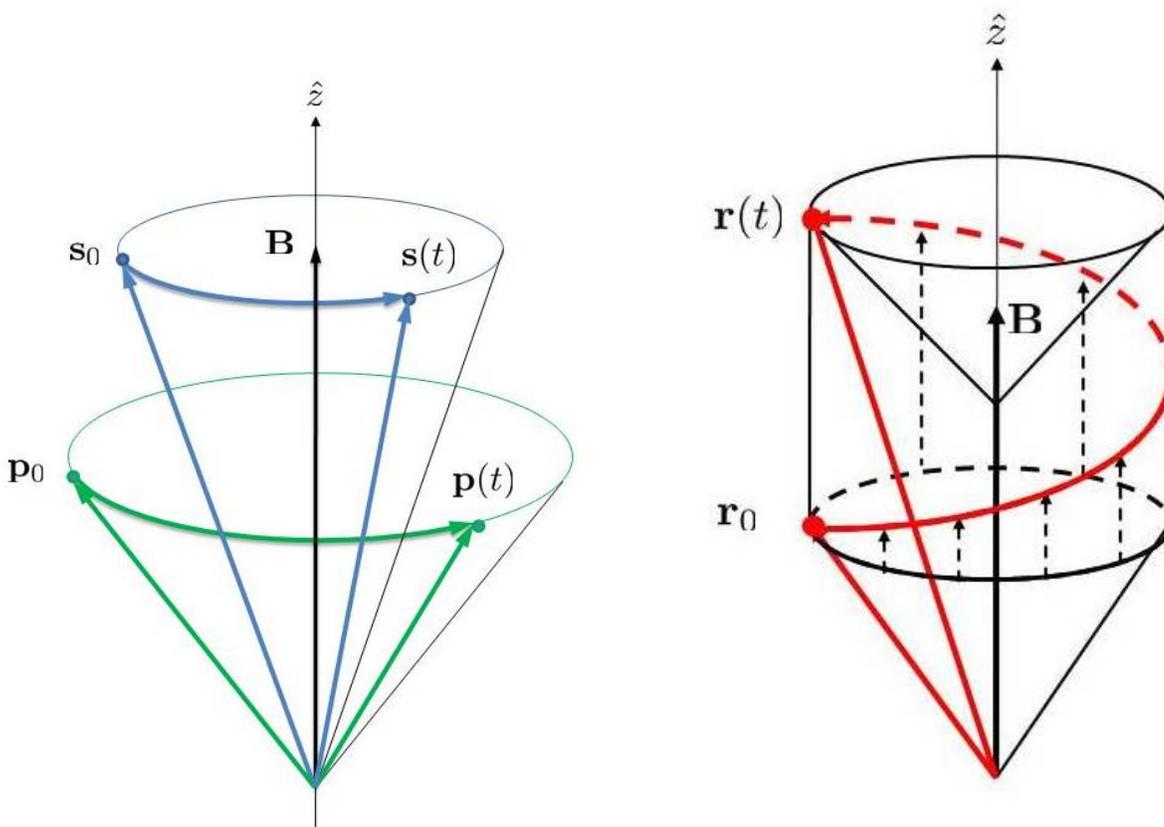
Both momentum and spin vectors precess around direction of magnetic field, $\mathbf{B} = B\hat{z}$ with common angular velocity $\omega = -eB/\mathcal{E}$,

$$\mathbf{p}(t) = (p_0 e^{-i(eB/\mathcal{E})t}, p_z), \quad (57)$$

$$\mathbf{s}(t) = (s_0 e^{-i(eB/\mathcal{E})t}, s_z), \quad (58)$$

where $p_0 = p_x + ip_y$, $s_0 = s_x + is_y$. Therefore

$$\mathbf{r}(t) = \left(\frac{ip_0}{eB} e^{-i(eB/\mathcal{E})t}, \frac{p_z}{\mathcal{E}} t \right) + \mathbf{r}_0. \quad (59)$$



Mo-

tion in pure const magnetic field \mathbf{B} . Both **momentum** $\mathbf{p}(t)$ and **spin** $\mathbf{s}(t)$ precess around \mathbf{B} direction. **Position** $\mathbf{r}(t)$ spirals on cylinder around \mathbf{B} -axis.