

Production of hypernuclei from excited nuclear residues in relativistic ion collisions: New opportunities for BM@N and MPD@NICA

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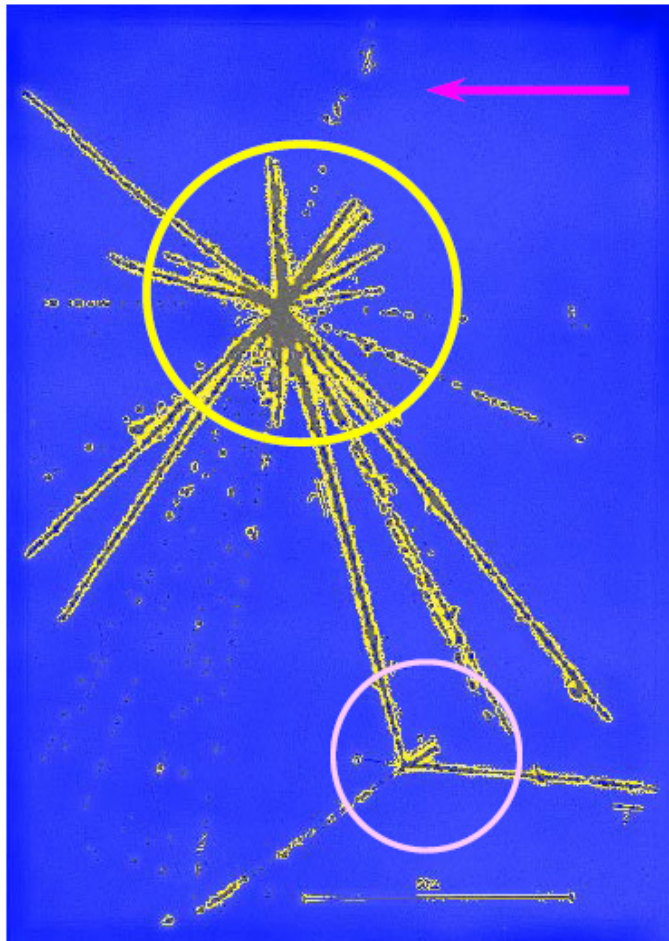
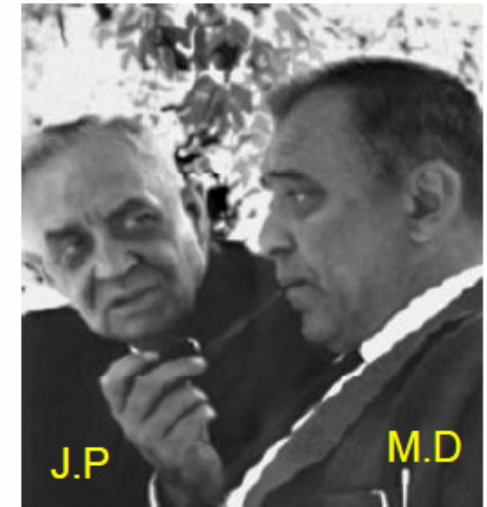
(Collaboration with M.Bleicher, E.Bratkovskaya, K.Gudima,
J.Pochodzalla, J.Steinheimer)

**Seminar of BLTP (JINR),
Dubna, Russia
July 7, 2014**

Discovery of a Strange nucleus: Hypernucleus

M. Danysz and J. Pniewski, *Philos. Mag.* 44 (1953) 348

First-hypernucleus was observed in a stack of photographic emulsions exposed to cosmic rays at about 26 km above the ground.



Incoming high energy proton from cosmic ray

colliding with a nucleus of the emulsion, breaks it in several fragments forming a star. **Multifragmentation !**

All nuclear fragments stop in the emulsion after a short path

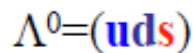
From the first star, 21 Tracks => $9\alpha + 11\text{H} + 1\Lambda\text{X}$

The fragment ΛX disintegrates later, makes the bottom star. Time taken $\sim 10^{-12}$ sec (typical for weak decay)

This particular nuclear fragment, and the others obtained afterwards in similar conditions, were called **hyperfragments or hypernuclei**.

Hyperons: Baryons with Strangeness

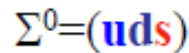
Lambda



$$m(\Lambda^0) = 1115.683 \pm 0.006 \text{ MeV}$$

$$S = -1$$

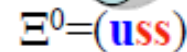
Sigma



$$m(\Sigma^0) = 1192.642 \pm 0.024 \text{ MeV}$$

$$S = -1$$

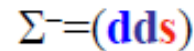
Cascade or Xi



$$m(\Xi^0) = 1314.86 \pm 0.2 \text{ MeV}$$

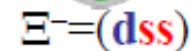
$$S = -2$$

Quark	Symbol	charge (e)	Strangeness (S)
Up	(u)	2/3	0
Down	(d)	-1/3	0
Strange	(s)	-1/3	-1
Charm	(c)	2/3	0
Bottom	(b)	-1/3	0
Top	(t)	2/3	0



$$m(\Sigma^-) = 1197.449 \pm 0.030 \text{ MeV}$$

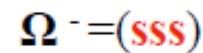
$$S = -1$$



$$m(\Xi^-) = 1321.71 \pm 0.07 \text{ MeV}$$

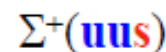
$$S = -2$$

Omega



$$m(\Omega^-) = 1672.45 \pm 0.29 \text{ MeV}$$

$$S = -3$$



$$m(\Sigma^+) = 1189.37 \pm 0.07 \text{ MeV}$$

$$S = -1$$

$$\text{lifetime of } \sim 8.2 \times 10^{-11} \text{ s}$$

lifetimes of $\sim 1 \times 10^{-10} \text{ s}$

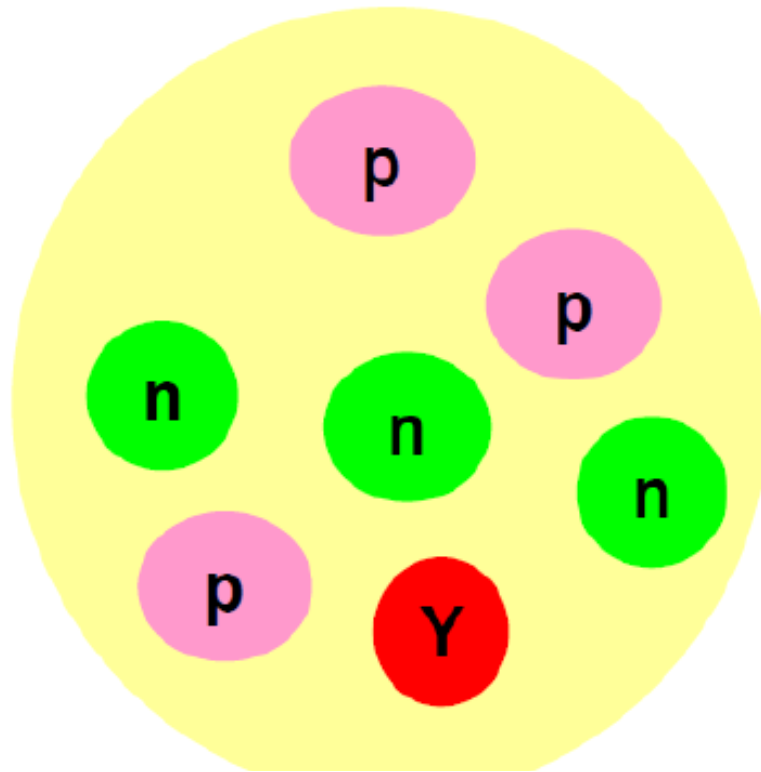
with the exception of Σ^0

whose lifetime is

shorter than $1 \times 10^{-10} \text{ s}$

Hypernucleus: Hyperons Bound in Nuclei

Hypernucleus: consists of nucleons (n, p) + hyperon (Y)



Notation:

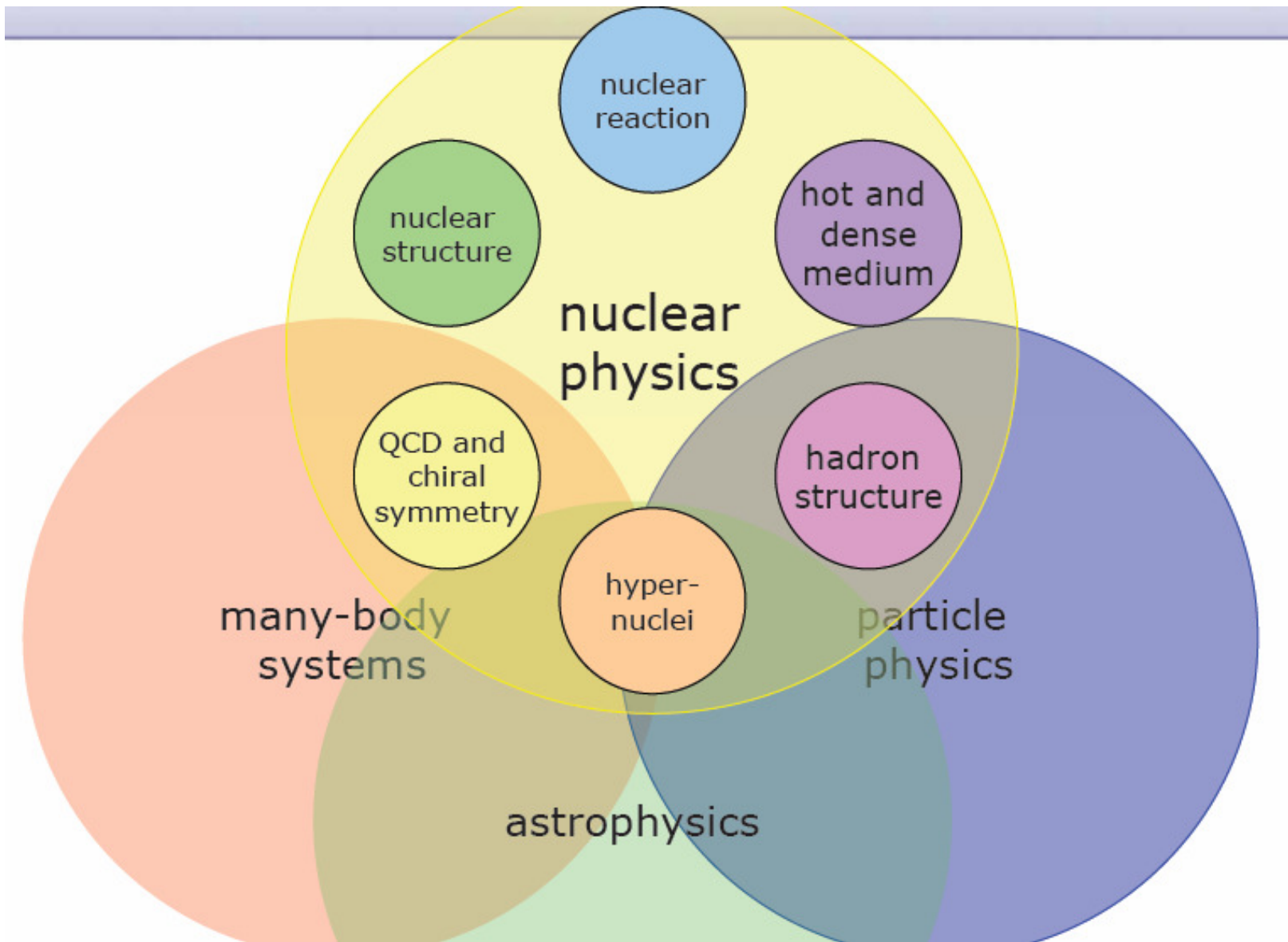


\mathbf{Y} = Hyperon

$$Z = Z_p + (N_Y \cdot q_Y)$$

$$A = N_n + N_p + N_Y$$

Hypernuclei within the research fields



Why hypernuclei ?

QCD theory development

Micro-laboratory with protons, neutrons, and hyperons;

YN & YY interaction can be investigated (strangeness sector of hadronic EoS); ...

Astrophysics

Hyperons are important for cosmology, physics of neutron stars , "strange stars", black holes, ...

Nuclear physics

Phenomenology: extension of nuclear charts into strangeness, exotic nuclei, limits of nuclear stability

Structure theory -- new degree of freedom for investigating interaction of baryons in nuclei (hyperons - without Pauli blocking)

Reaction theory - new probe for fragmentation of nuclei, phase transitions and EoS in hypermatter and finite hypernuclei

5 decades of hyperons in neutron stars

NEUTRON STAR MODELS

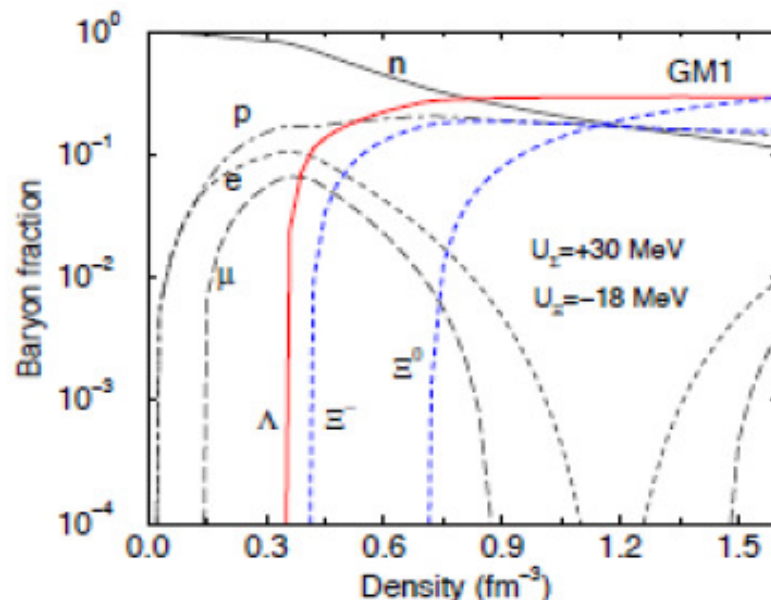
A. G. W. CAMERON

Atomic Energy of Canada Limited, Chalk River, Ontario, Canada

Received June 17, 1959

Another reason why the writer has not taken into account complications inherent in using a relativistic equation of state is that no such things as pure neutron stars can be expected to exist. The neutrons must always be contaminated with some protons and sometimes with other kinds of nucleons (hyperons or heavy mesons).

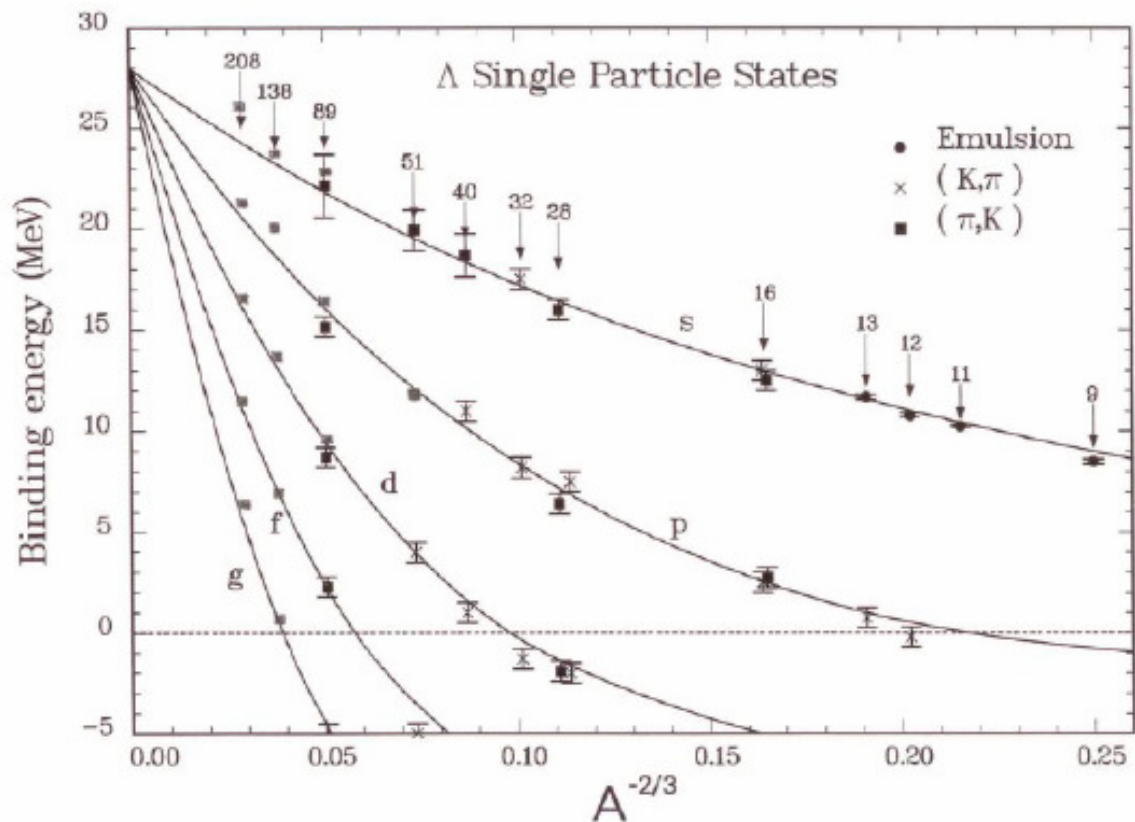
Alastair G.W. Cameron, Astrophysical Journal, vol. 130, p.884 (1959)



n-rich hypernucleus is a doorway to n-star



insight into halo nuclear structure through hypernuclei



Hyperon can be put deep inside - no Pauli blocking

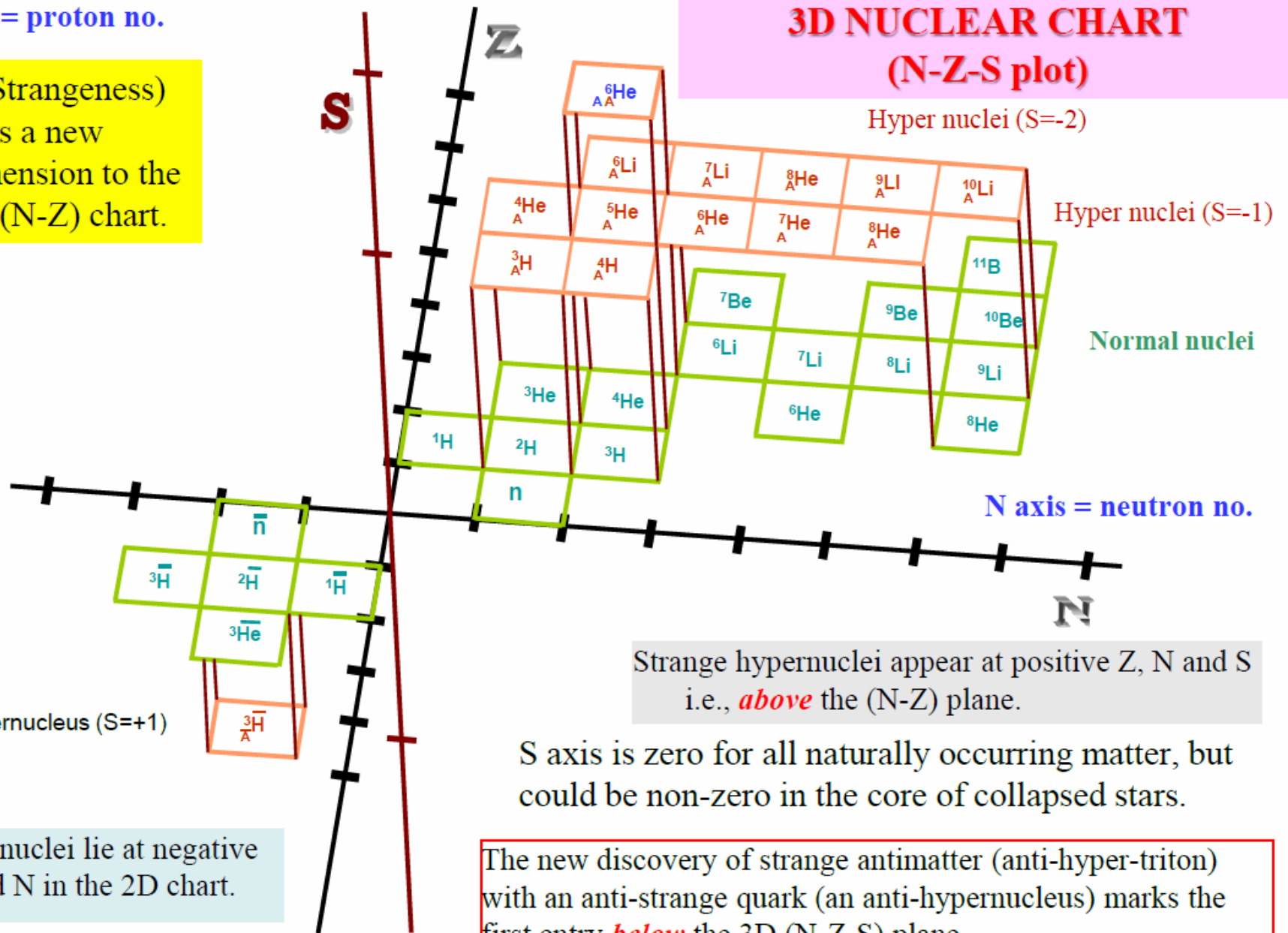
In heavy nuclei effects of the additional hyperon binding will be larger:

Production of nuclei beyond the drip lines

Z axis = proton no.

S (Strangeness) adds a new dimension to the 2D (N-Z) chart.

3D NUCLEAR CHART (N-Z-S plot)



Nuclear reactions: production mechanisms for hypernuclei

Traditional way for production of hypernuclei:

Conversion of Nucleons into Hyperons

by using hadron and electron beams

(CERN, BNL, KEK, CEBAF, DAΦNE, JPARC, MAMI, ...)

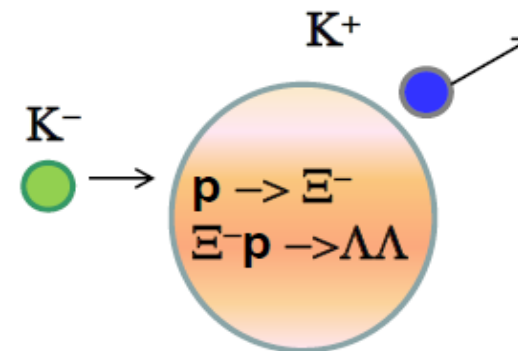
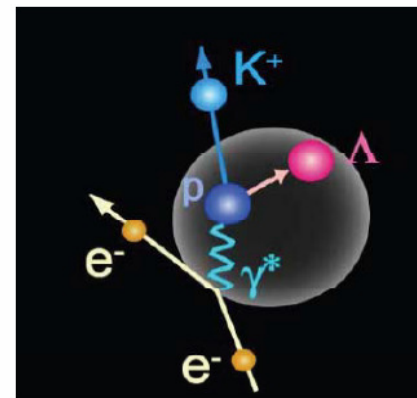
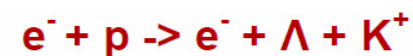
Advantages: rather precise determination of masses

(e.g., via the missing mass spectroscopy) :

good for nuclear structure studies !

Disadvantages: very limited range of nuclei in A and Z can be investigated; the phase space of the reaction is narrow (since hypernuclei are produced in ground and slightly excited states), so production probability is low; it is difficult to produce multi-strange nuclei.

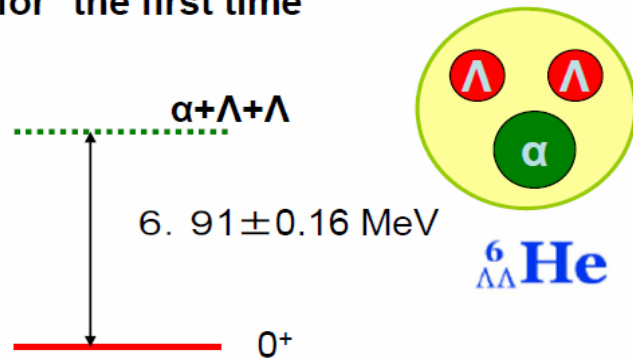
What reactions can be used to produce exotic strange nuclei and nuclei with many hyperons ?



(K^-, K^+) reactions

Ξ^- hyperons at the emulsion

Uniquely identified without ambiguity for the first time



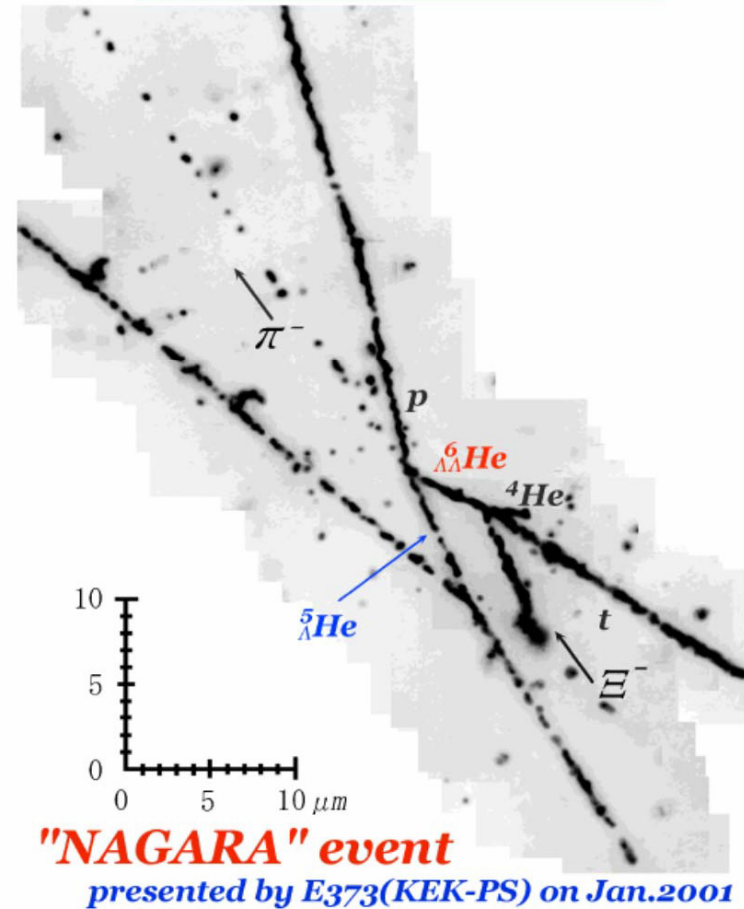
Possible mechanism of this reaction:



Break-up of excited hyper-system (~28MeV)
[Fermi-Break-up calculated probability~0.01]

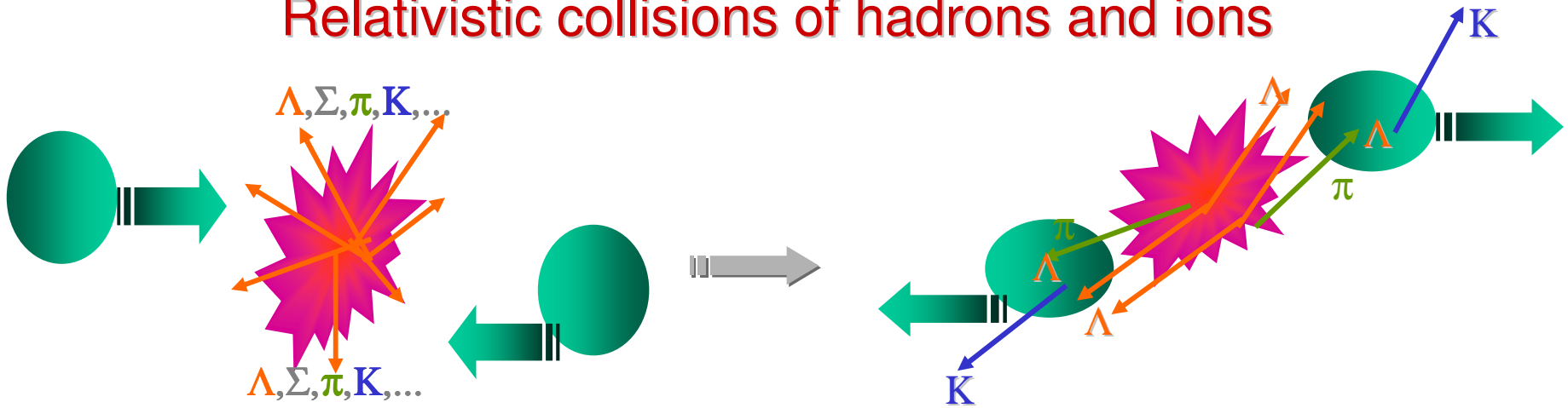
A.Sanchez Lorente et al., Phys. Lett. B697 (2011)222

${}^6_{\Lambda\Lambda}\text{He}$ double-hypernucleus U.S. et al. Unique interpretation!!



H. Takahashi et al., PRL 87, 212502-1 (2001)

Relativistic collisions of hadrons and ions

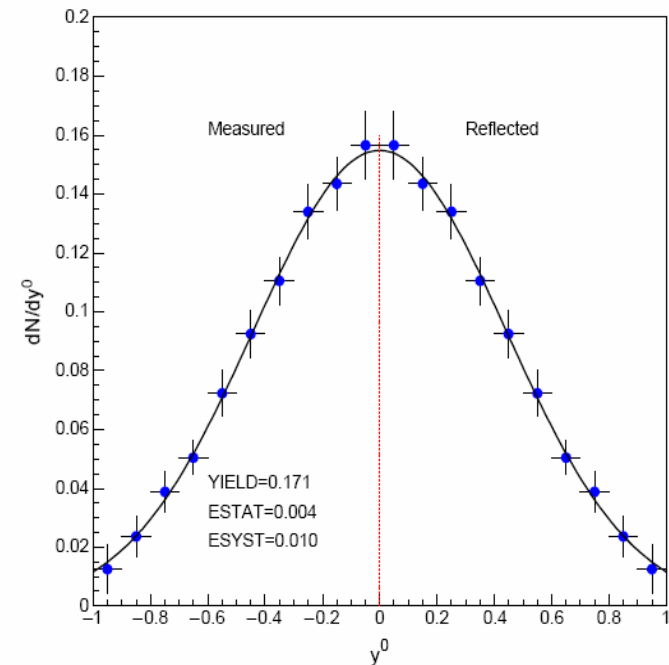


Production of hypermatter in relativistic HI and hadron collisions

- Production of strange particles and hyperons by "participants",
- Rescattering and absorption of hyperons by excited "spectators"

X. Lopez / Progress in Particle and Nuclear Physics 53 (2004) 149–151

Reconstructed Λ rapidity distribution for central
 ($\sigma_{\text{geo}} = 350 \text{ mb}$) Ni + Ni reactions at 1.93 A GeV.



Central collisions of relativistic ions

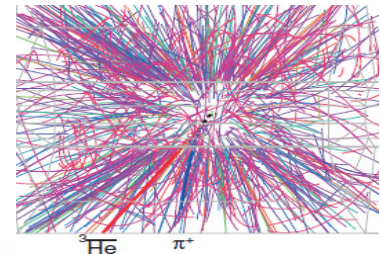
Production of ${}^3_{\Lambda}\text{H}$ and ${}^4_{\Lambda}\text{H}$ in central 11.5 GeV/c Au+Pt heavy ion collisions

PHYSICAL REVIEW C 70, 024902 (2004) (AGS) $N_{event} 13.5 \times 10^9$ ${}^3_{\Lambda}\text{H}$
 Rapidity 1.6–2.6 coalescence mechanism $N_{count} 1220 \pm 854$
 p_t (GeV/c) 0–1.5

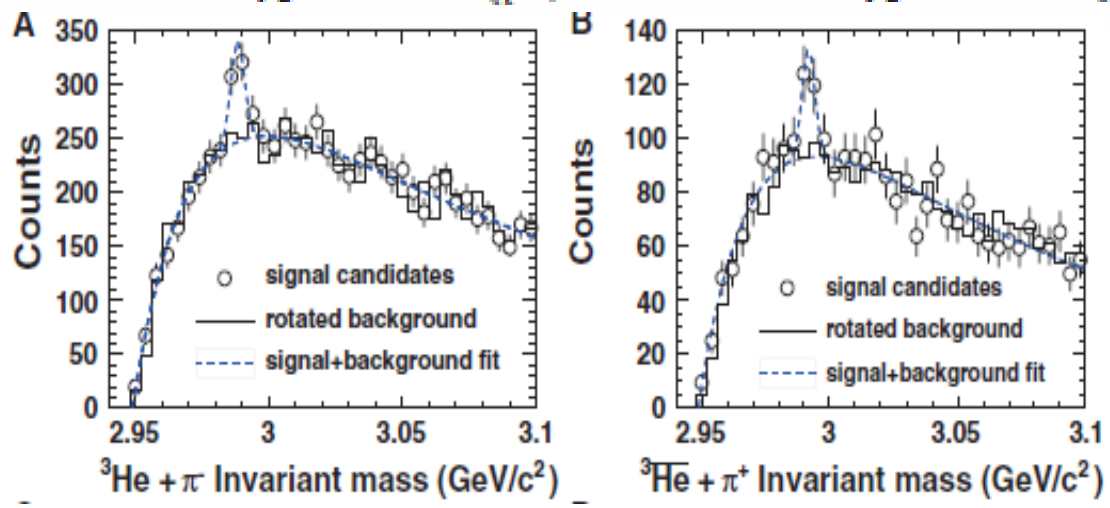
STAR collaboration (RHIC): Science, 238 (2010) 58

Au + Au collisions at 200 A GeV

gas-filled cylindrical Time Projection Chamber

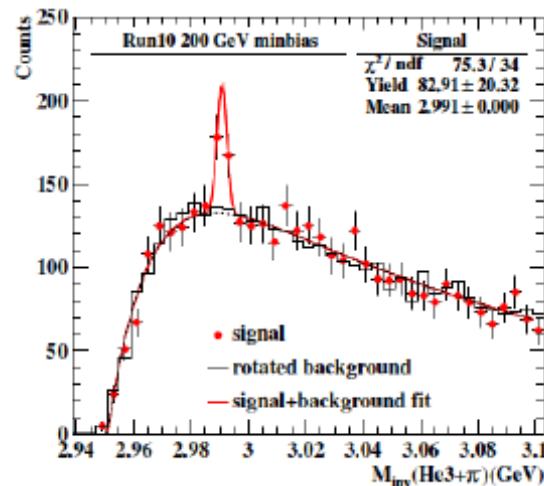
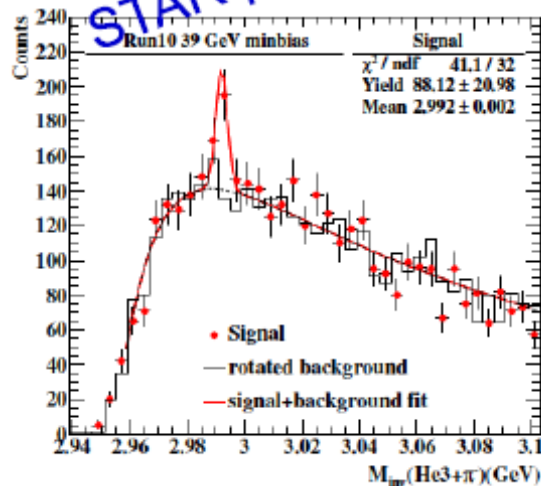
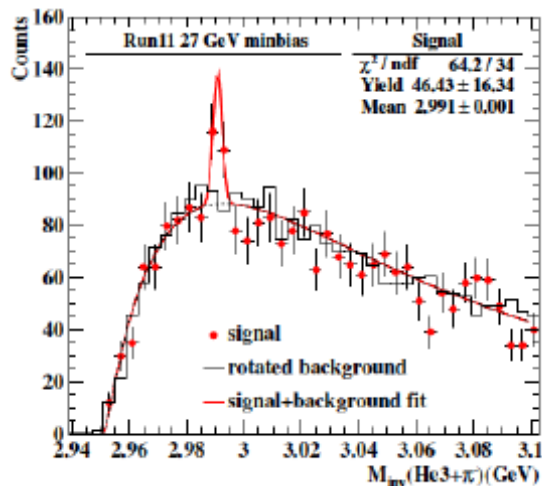
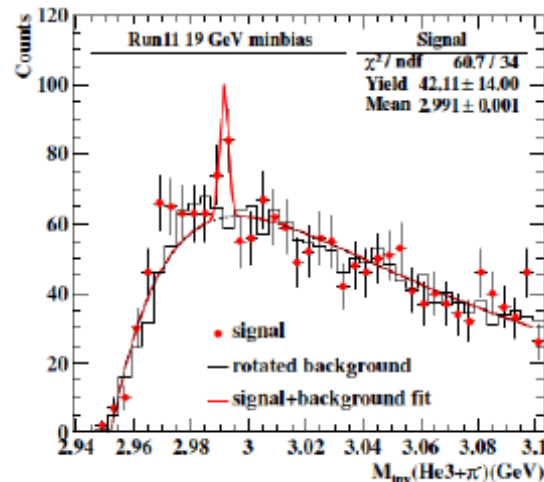
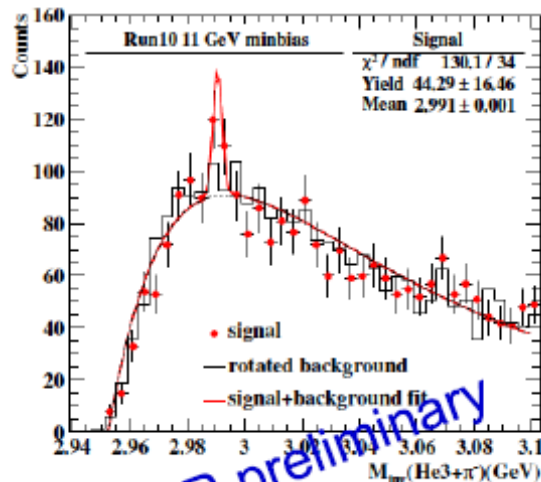
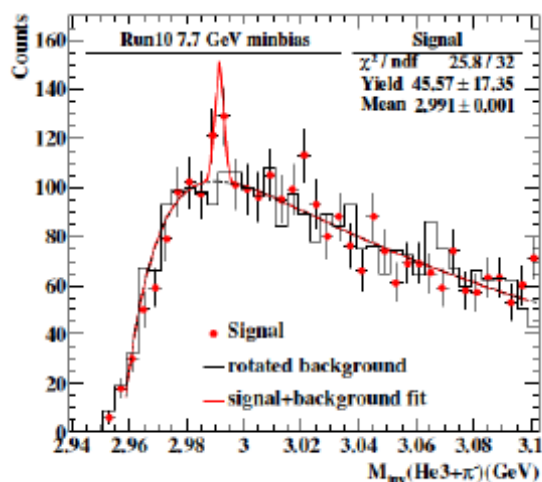


70 ± 17 antihypertritons (${}^3_{\bar{\Lambda}}\text{H}$) and 157 ± 30 hypertritons (${}^3_{\Lambda}\text{H}$).



Y.-G. Ma , talk at NUFRA2013 (Kemer, Turkey)

Hypertriton signal from STAR BES at $\sqrt{s_{NN}}=7.7, 11.5, 19.6, 27, 39, 200\text{GeV}$



STAR preliminary

ALICE's observation for (anti-)hypertriton

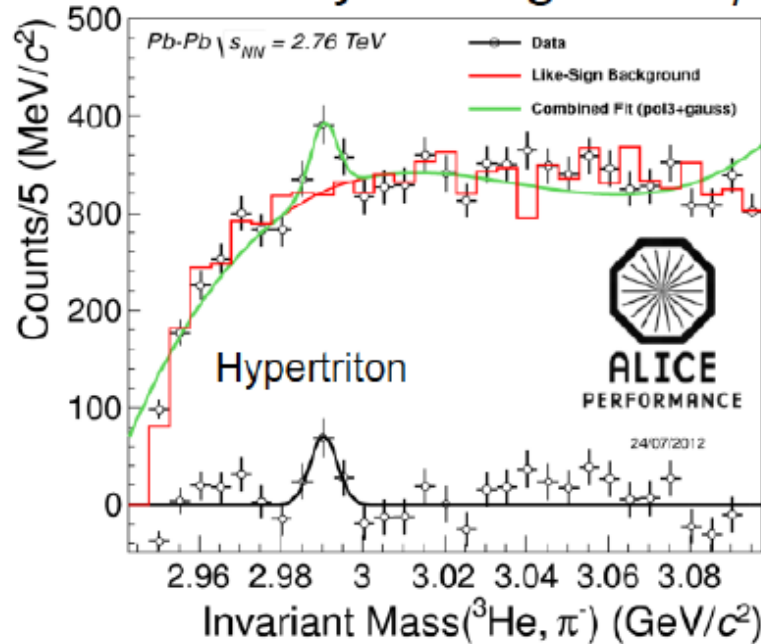


Hypertriton

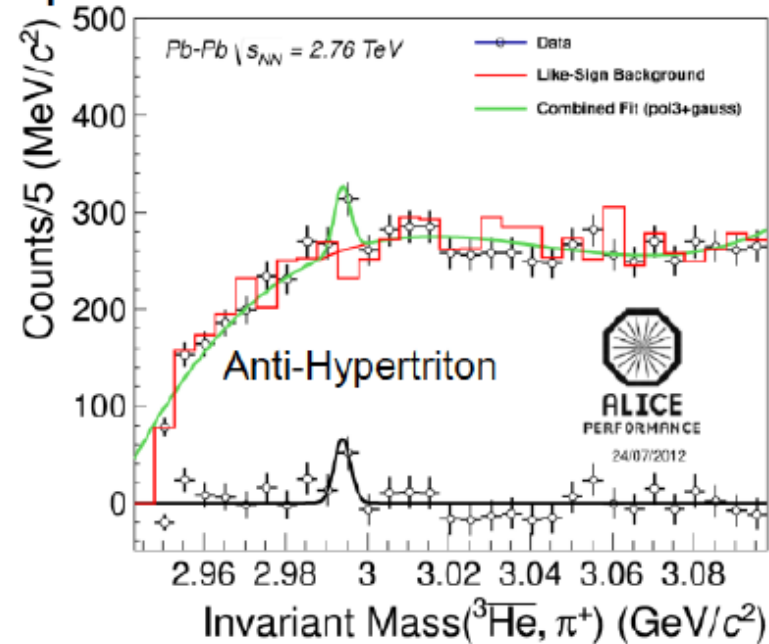


Signal of the hypertriton from the 2011 run

→ currently working on the p_T spectra extraction



$$\mu = (2.990 \pm 0.001) \text{ GeV}/c^2$$
$$\sigma = (3.35 \pm 0.7) \times 10^{-3} \text{ GeV}/c^2$$



$$\mu = (2.993 \pm 0.001) \text{ GeV}/c^2$$
$$\sigma = (2.00 \pm 1.2) \times 10^{-3} \text{ GeV}/c^2$$

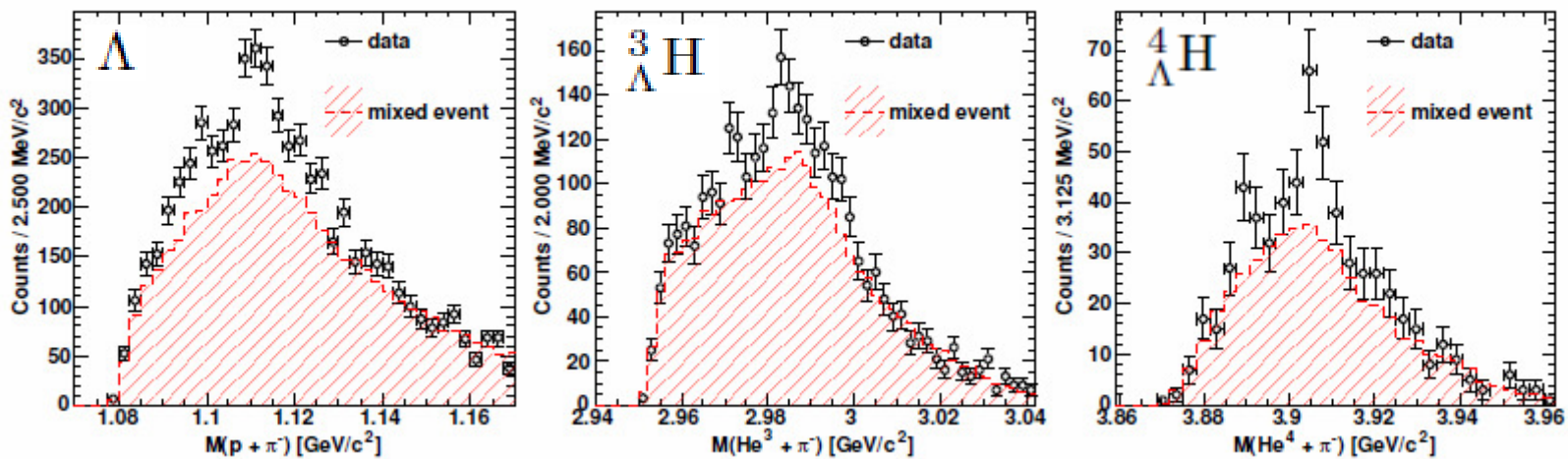
Production of hypernuclei in peripheral HI collisions: The HypHI project at GSI

T.Saito, (for HypHI),
NUFRA2011 conference, and
Nucl. Phys. A881 (2012) 218;
Nucl. Phys. A913 (2013) 170.

C. Rappold et al.,
Phys. Rev. C88 (2013) 041001:
Ann bound state ?

T.R. Saito^{a,b,c}, D. Nakajima^{a,d}, C. Rappold^{a,c,e}, S. Bianchin^a, O.
Borodina^{a,b}, V. Bozkurt^{a,f}, B. Göküzüm^{a,f}, M. Kavatsyuk^g, E. Kim^{a,h}, Y.
Ma^{a,b}, F. Maas^{a,b,c}, S. Minami^a, B. Özel-Tashenov^a, P. Achenbach^b, S.
Ajimuraⁱ, T. Aumann^a, C. Ayerbe Gayoso^b, H.C. Bhang^f, C. Caesar^a, S.
Erturk^f, T. Fukuda^j, E. Guliev^h, Y. Hayashi^k, T. Hiraiwa^k, J. Hoffmann^a,
G. Ickert^a, Z.S. Ketenci^f, D. Khanefte^{a,b}, M. Kim^h, S. Kim^h, K. Koch^a, N.
Kurz^a, A. Le Fevre^{a,1}, Y. Mizoi^j, M. Moritsu^k, T. Nagae^k, L. Nungesser^b, A.
Okamura^k, W. Ott^a, J. Pochodzalla^b, A. Sakaguchi^m, M. Sako^k, C.J.
Schmidt^a, M. Sekimotoⁿ, H. Simon^a, H. Sugimura^k, T. Takahashiⁿ, G.J.
Tambave^g, H. Tamura^o, W. Trautmann^a, S. Voltz^a, N. Yokota^k, C.J. Yoon^h,
K. Yoshida^m,

Projectile fragmentation: ⁶Li beam at 2 A GeV on ¹²C target



For the first, they have also observed a large correlation of ${}^2\text{H} + \pi^-$
i.e., considerable production of Λn bound states

Theoretical descriptions of strangeness production within transport codes

- old models :* e.g., Z.Rudy, W.Cassing et al., *Z. Phys.A351(1995)217*
INC, QMD, BUU
- GiBUU model:* Th.Gaitanos, H.Lenske, U.Mosel , *Phys.Lett. B663(2008)197,*
(+SMM) Phys.Lett. B675(2009)297
- PHSD model:* E.Bratkovskaya, W.Cassing, ... *Phys. Rev. C78(2008)034919*
- DCM (INC) :* JINR version: K.K.Gudima et al., *Nucl. Phys. A400(1983)173, ...*
(+QGSM+SMM) Phys. Rev. C84 (2011) 064904
- UrQMD approach:* S.A. Bass et al., *Prog. Part. Nucl. Phys. 41 (1998)255.*
M.Bleicher et al. *J. Phys. G25(1999)1859, ...* , J.Steinheimer ...

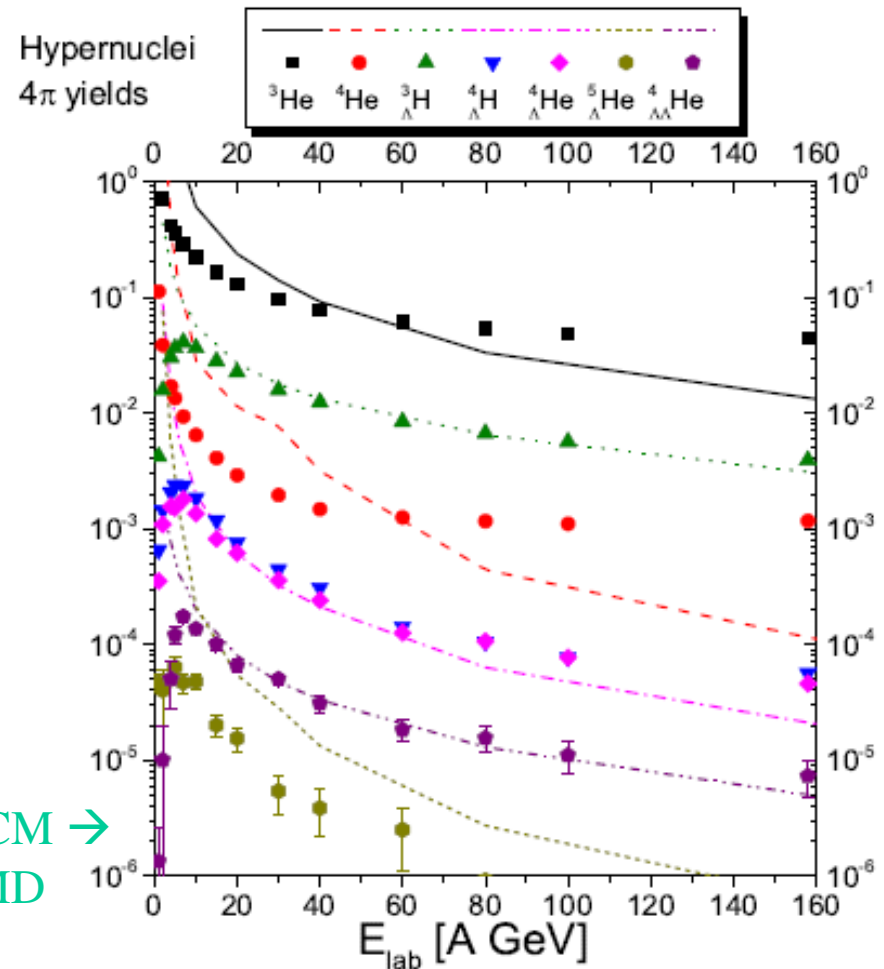
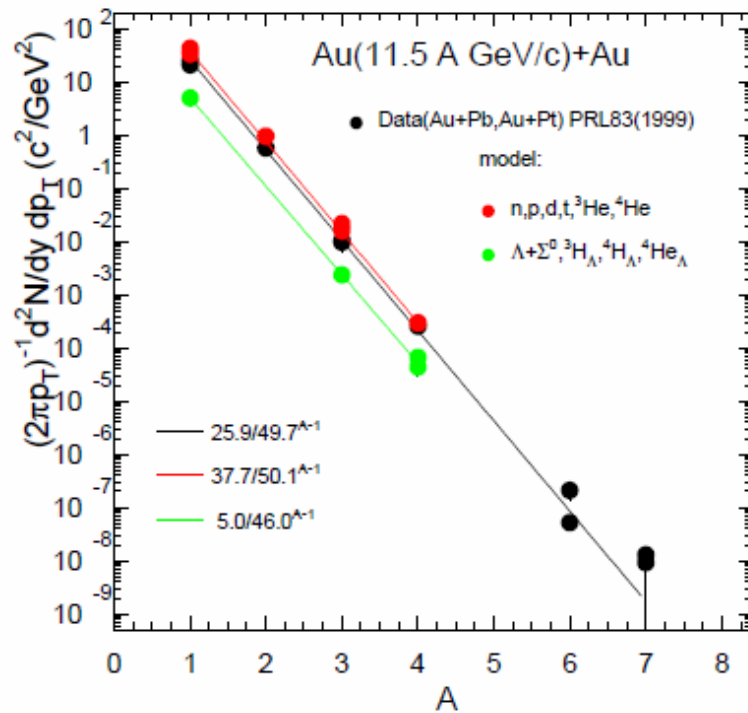
Main channels for production of strangeness in individual hadron- nucleon collisions: $BB \rightarrow BYK$, $B\pi \rightarrow YK$, ... (like $p+n \rightarrow n+\Lambda+K^+$, and secondary meson interactions, like $\pi+p \rightarrow \Lambda+K^+$). Rescattering of hyperons is important for their capture by spectators. Expected decay of produced hyperons and hypernuclei: 1) mesonic $\Lambda \rightarrow \pi+N$; 2) in nuclear medium nonmesonic $\Lambda+N \rightarrow N+N$.

Production of light nuclei in central collisions : Au+Au

DCM and UrQMD calculations - J.Steinheimer et al., Phys. Lett. B714, 85 (2012)

DCM versus experiment :
coalescence mechanism

Also predictions for hybrid approach :
UrQMD + thermal hydrodynamics



It is not possible to
produce big nuclei !

Symbols - DCM →
Lines - UrQMD

Physical picture of peripheral relativistic HI collisions:

nucleons of projectile interact with nucleons of target, however, in peripheral collisions many nucleons (spectators) are not involved. All products of the interactions can also interact with nucleons and between themselves. The time-space evolution of all nucleons and produced particles can be calculated with transport models.

All strange particles: Kaons, Lambda, Sigma, Xi, Omega are included in the transport models

ABSORPTION of LAMBDA :

The residual spectator nuclei produced during the non-equilibrium stage may capture the produced Lambda hyperons if these hyperons are (a) inside the nuclei and (b) their energy is lower than the hyperon potential in nuclear matter (~ 30 MeV). In the model a depletion of the potential with reduction of number of nucleons in nucleus is taken into account by calculating the local density of spectator nucleons.

Λ -Hypernucleus formation in proton-nucleus reactions

Z.Rudy, W.Casing et al., *Z. Phys. A351(1995)217*
BUU approach (Λ -potential in matter ~ 30 MeV)

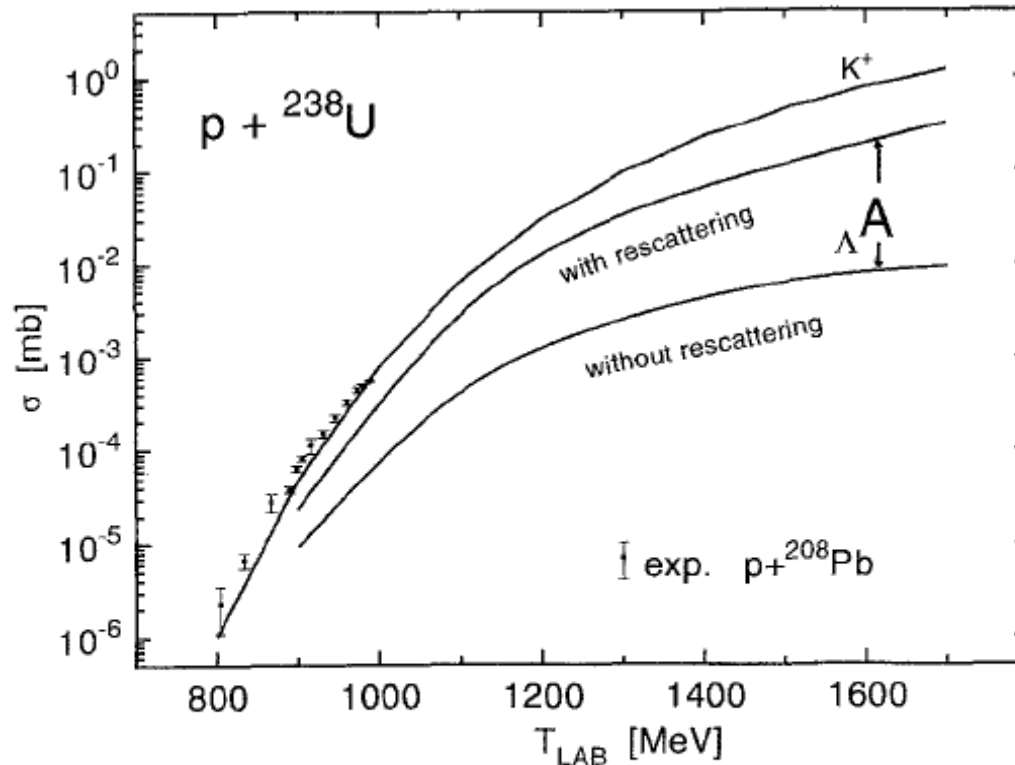


Fig. 5. The energy dependence of the ΛA production cross section for the cases with and without Λ hyperon- N rescattering. For relative orientation we also compare the calculated total K^+ cross section with the inclusive K^+ data from [30] for $p + {}^{208}\text{Pb}$

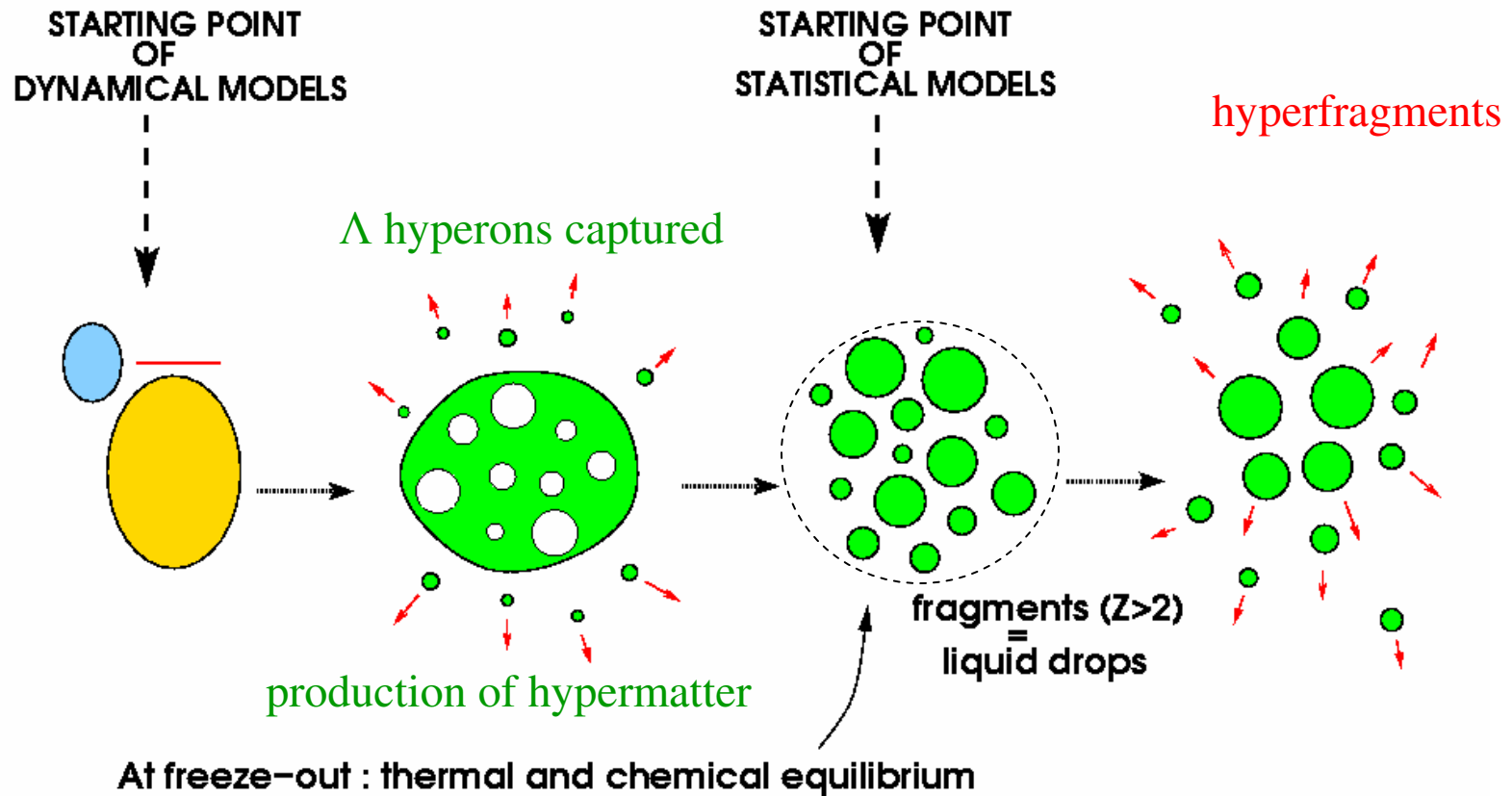
However, lifetime of these hypernuclei is too short ~ 200 ps, experimental identification using products of their decay is very difficult, because of background of other particles.

A.S.Botvina and J.Pochodzalla, Phys. Rev.C76 (2007) 024909

Generalization of the statistical de-excitation model for nuclei with Lambda hyperons

In these reactions we expect analogy with multifragmentation in intermediate and high energy nuclear reactions

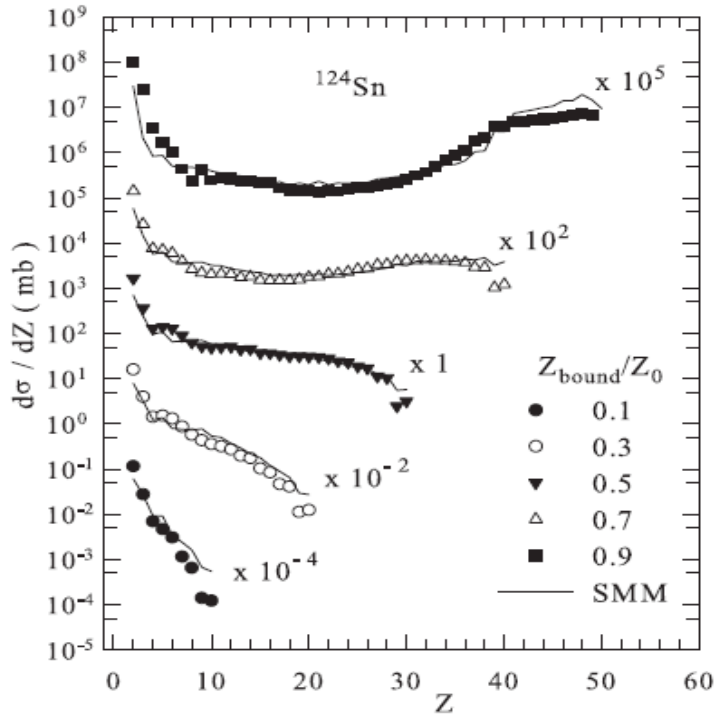
+ nuclear matter with strangeness



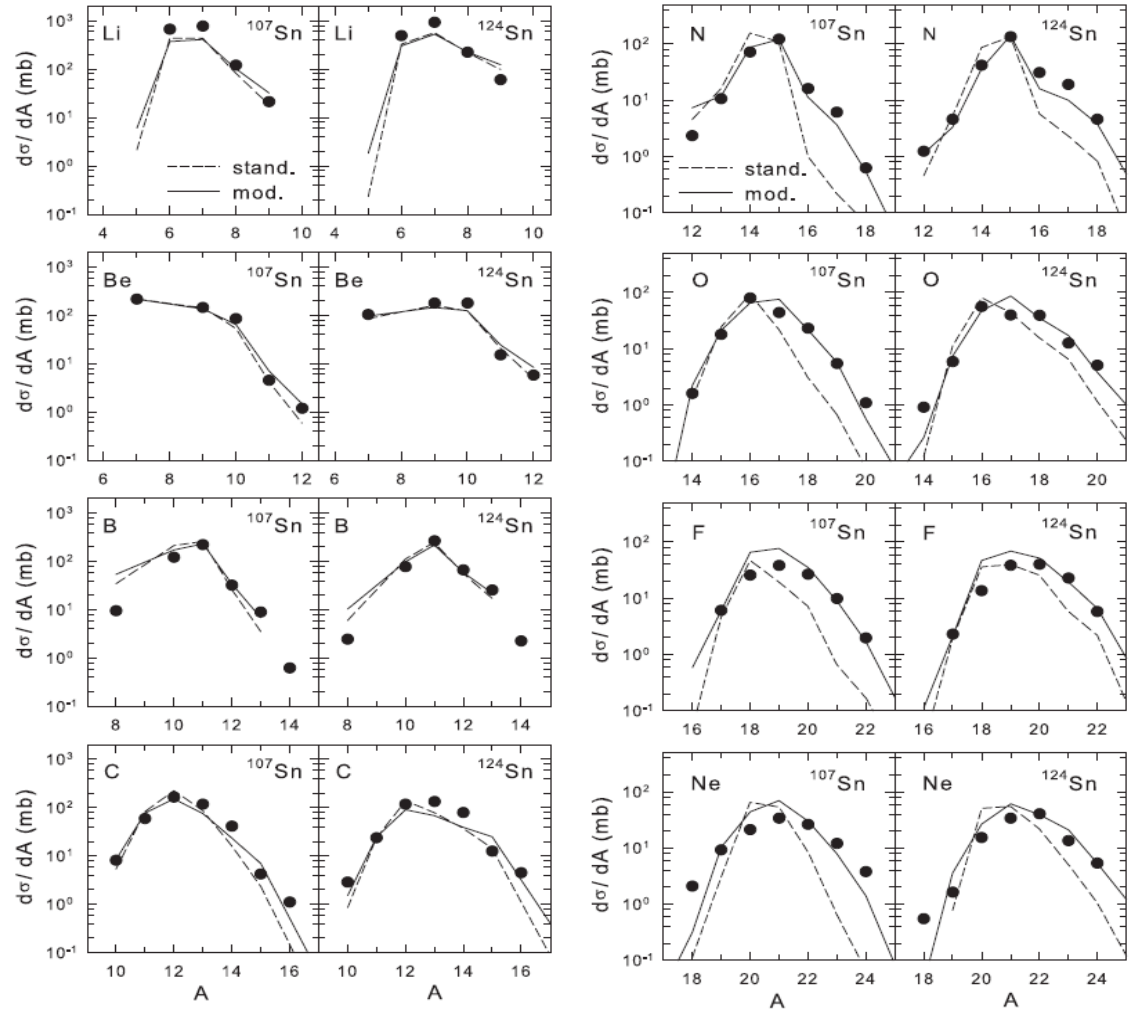
Isospin-dependent multifragmentation of relativistic projectiles

$^{124,107}\text{-Sn}, ^{124}\text{-La}$ (600 A MeV) + Sn \rightarrow projectile (multi-)fragmentation

Very good description is obtained within Statistical Multifragmentation Model, including fragment charge yields, isotope yields, various fragment correlations.



Statistical (chemical) equilibrium is established at break-up of hot projectile residues ! In the case of strangeness admixture we expect it too !



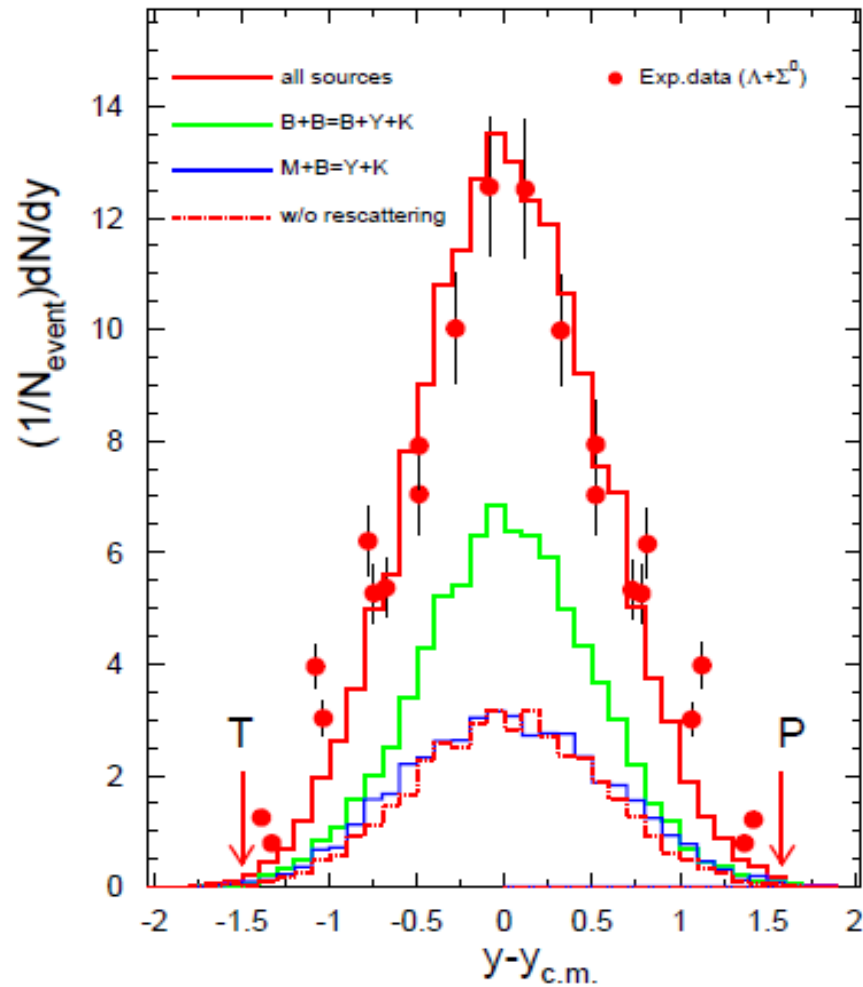
Verification of the models DCM

Hyperon production
in central collisions
Au(11 A GeV/c)+Au

experiment: S.Albergo et al.,
E896: PRL88(2002)062301

A.S.Botvina, K.K.Gudima,
J.Steinheimer, M.Bleicher,
I.N.Mishustin.

PRC **84** (2011) 064904

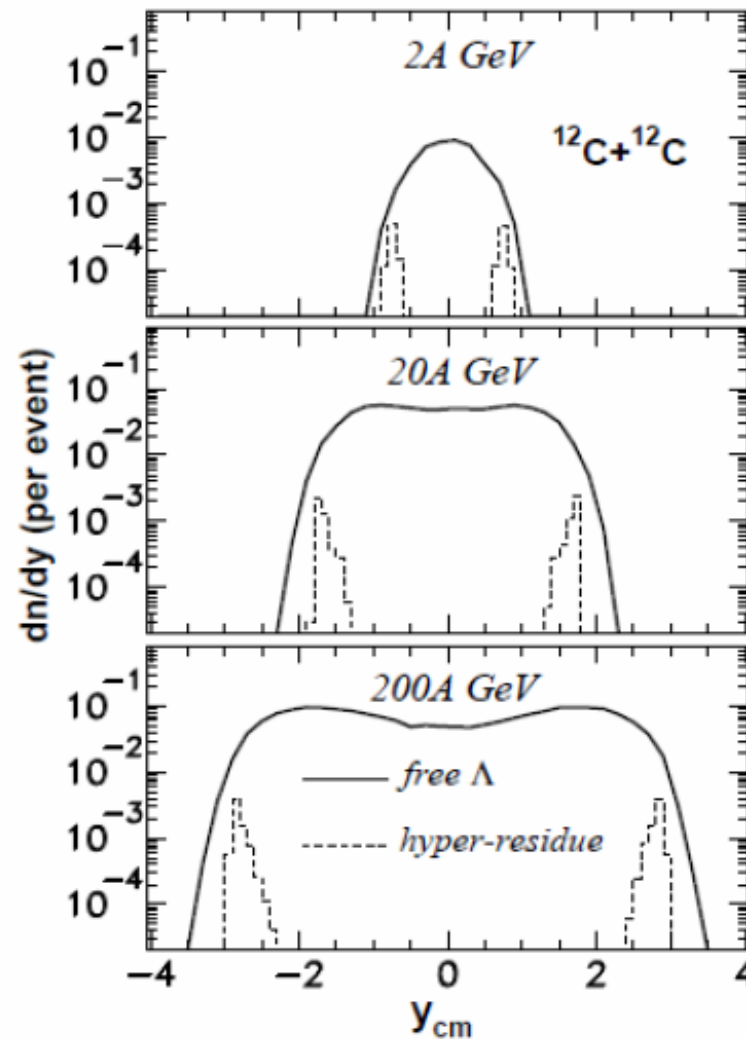


Rapidity distribution of free hyperons and hyper-residues in relativistic ion collisions (DCM calculations)

A.S.Botvina, K.K.Gudima, J.Pochodzalla PRC 88 (2013) 054605

Wide distributions of produced Lambda-hyperons up to spectator rapidities at all incident energies. A stochastic process related to secondary interactions leads to the hyperon capture by residues.

The evolution to a double peak distribution with increasing energy tell us that the Lambda production is mainly caused by secondary processes too.

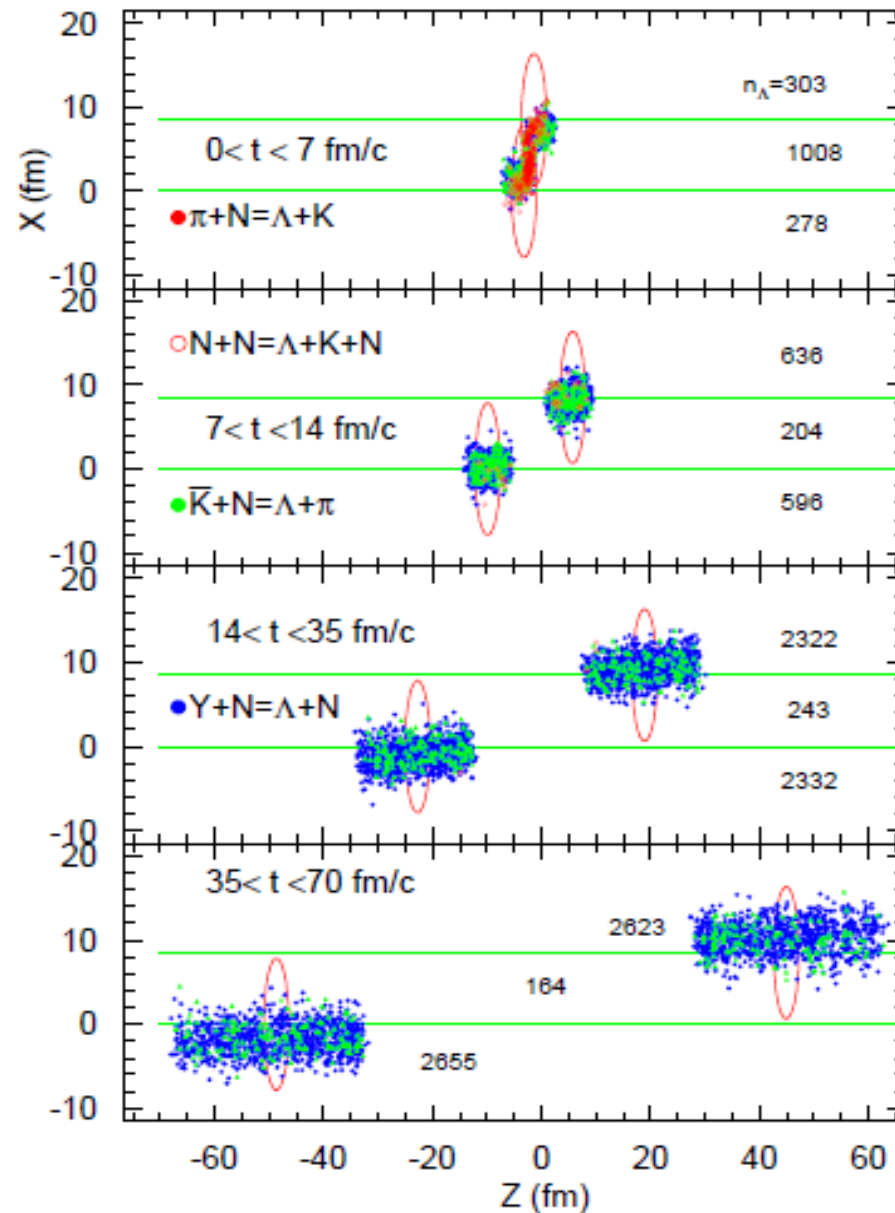


Peripheral relativistic ion collisions: Au (20 A GeV) + Au impact parameter= 8.5fm

Absorption of Lambda hyperons by residual nuclei within DCM and UrQMD model description (times/coordinates of the absorption are given on the panels)

Secondary interactions of the particles dominates in the process.

A.S.Botvina, K.K.Gudima,
J.Steinheimer, M.Bleicher,
I.N.Mishustin.
PRC 84 (2011) 064904



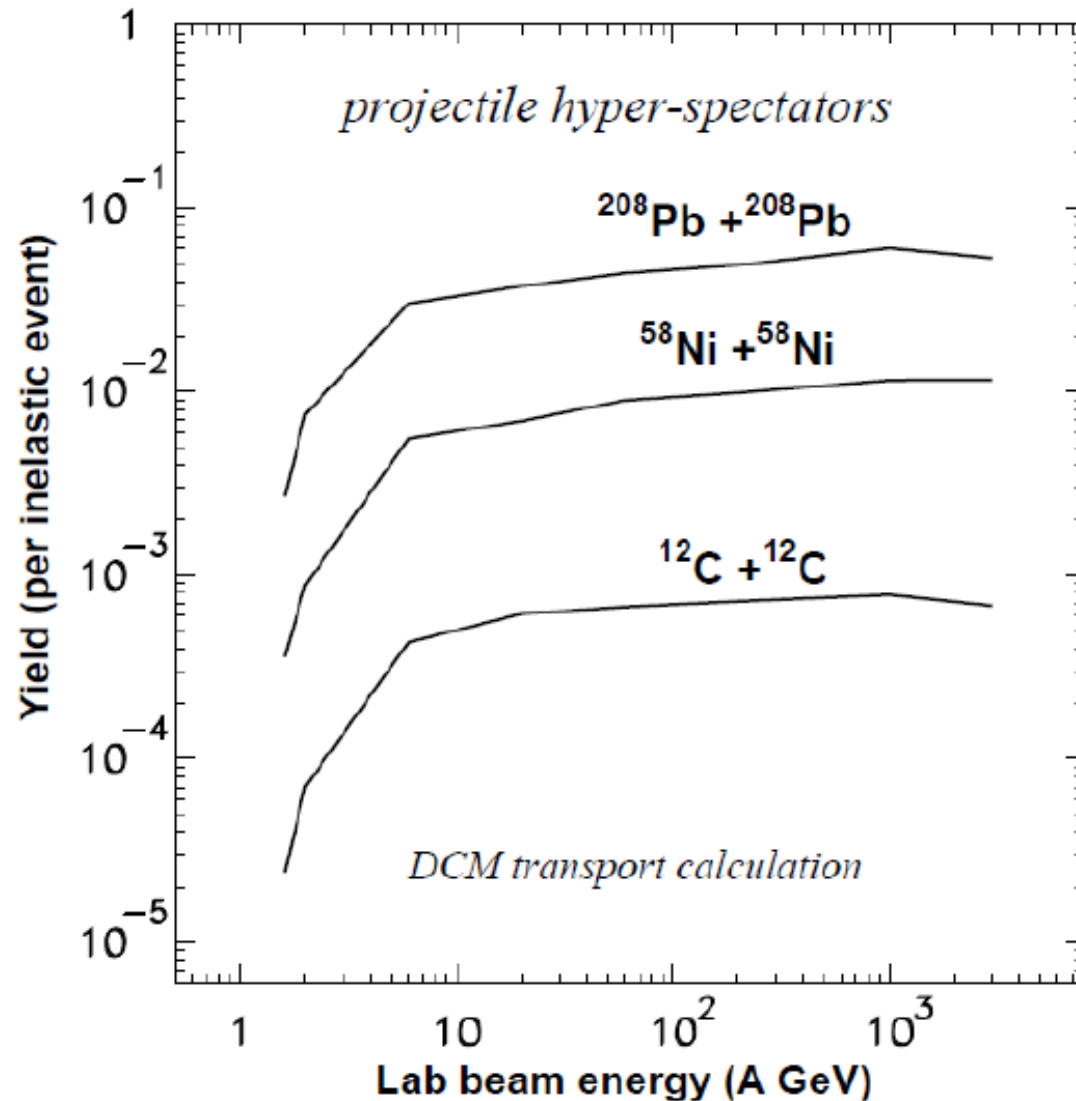
Yield of hypernuclei in peripheral collisions

A.S.Botvina, K.K.Gudima, J.Pochodzalla PRC 88 (2013) 054605

Threshold behavior with saturation at high energies (for single hypernuclei)

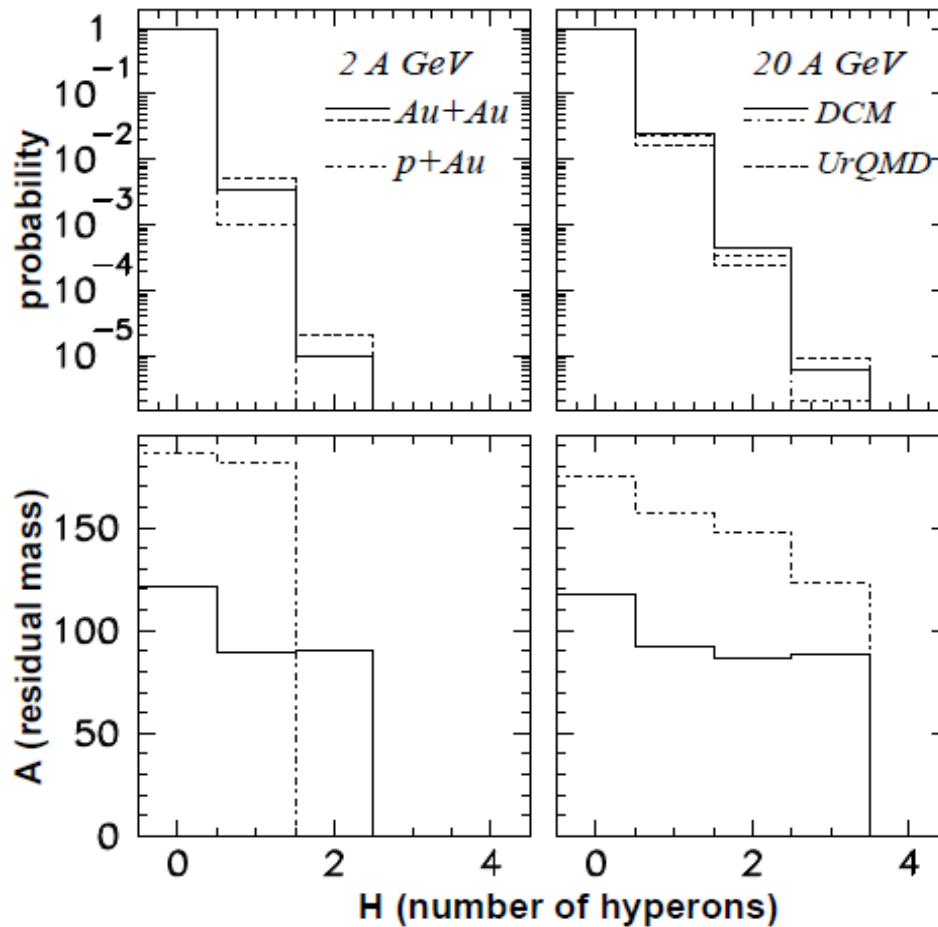
Yield is integrated over all impact parameters.

Reactions can be studied at GSI/FAIR and JINR/NICA facilities as well on operating RHIC and LHC.



projectile residuals produced after non-equilibrium stage

total yield of residuals with single hyperons $\sim 1\%$, with double ones $\sim 0.01\%$,
at 2 GeV per nucleon, and considerably more at 20 GeV per nucleon



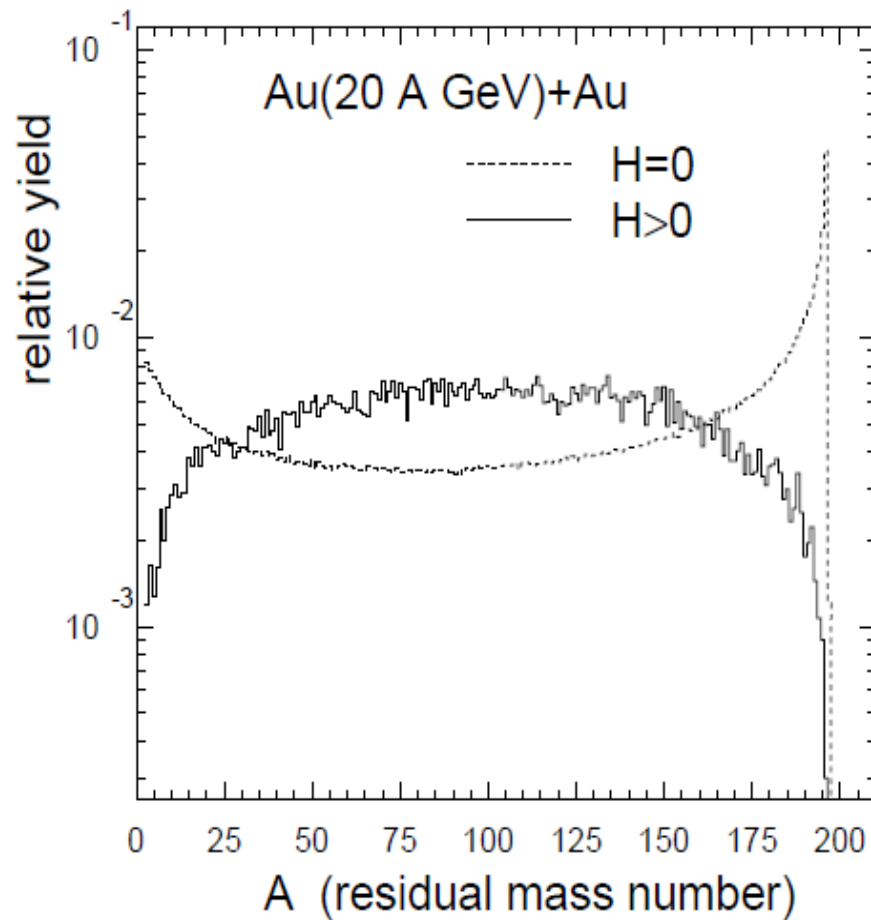
Integrated over all impact parameters

Formation of multi-strange nuclear systems ($H > 2$) is possible!

The disintegration of such systems can lead to production of exotic hypernuclei.

Masses of projectile residuals produced after DCM

different hyper-residuals (with large cross-section) can be formed
(from studies of conventional matter: expected temperatures - up to 5-8 MeV)

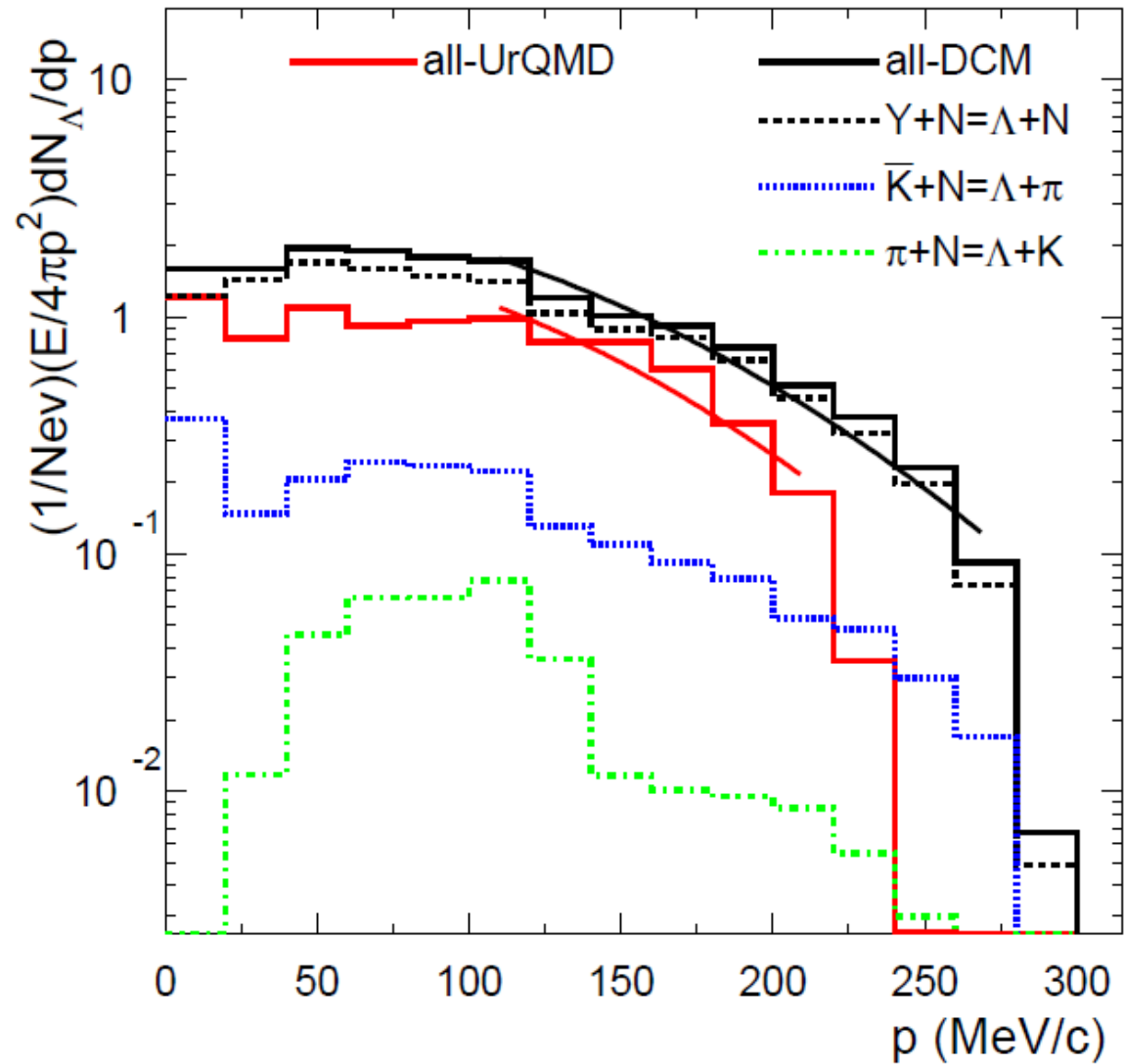


6b : H=0

200mb: H>0

Momentum distribution of Lambda captured in the spectators

(Connection of the potential capture and the coalescence)



W.Neubert and A.S.Botvina, EPJ A7 , 101 (2000)

Coalescence of Baryons (CB) Model :

Development of the coalescence for formation of clusters of all sizes

- 1) Relative velocities between baryons and clusters are considered,
if $(|\mathbf{V}_b - \mathbf{V}_A|)^2 < C$ the particle b is included in the A-cluster.
- 2) Step by step numerical approximation.

A.Botvina, E.Bratkovskaya, J.Steinheimer, ...

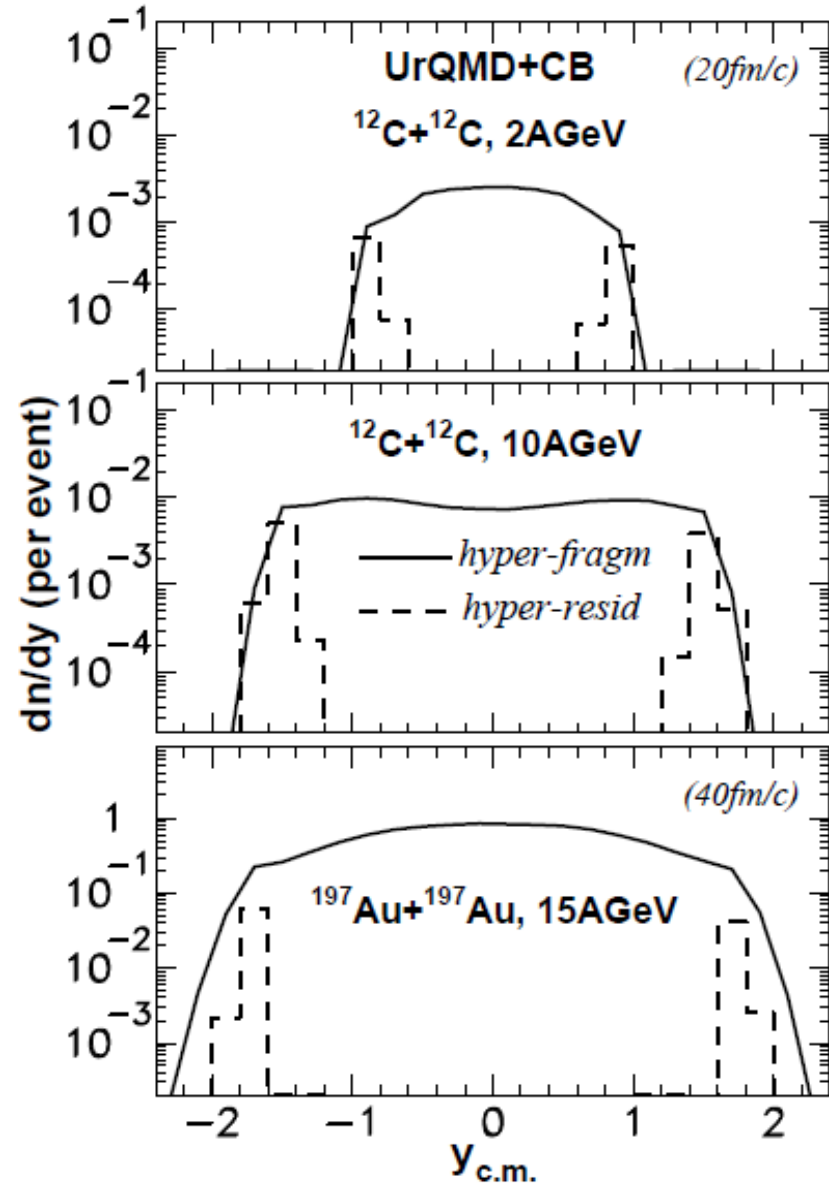
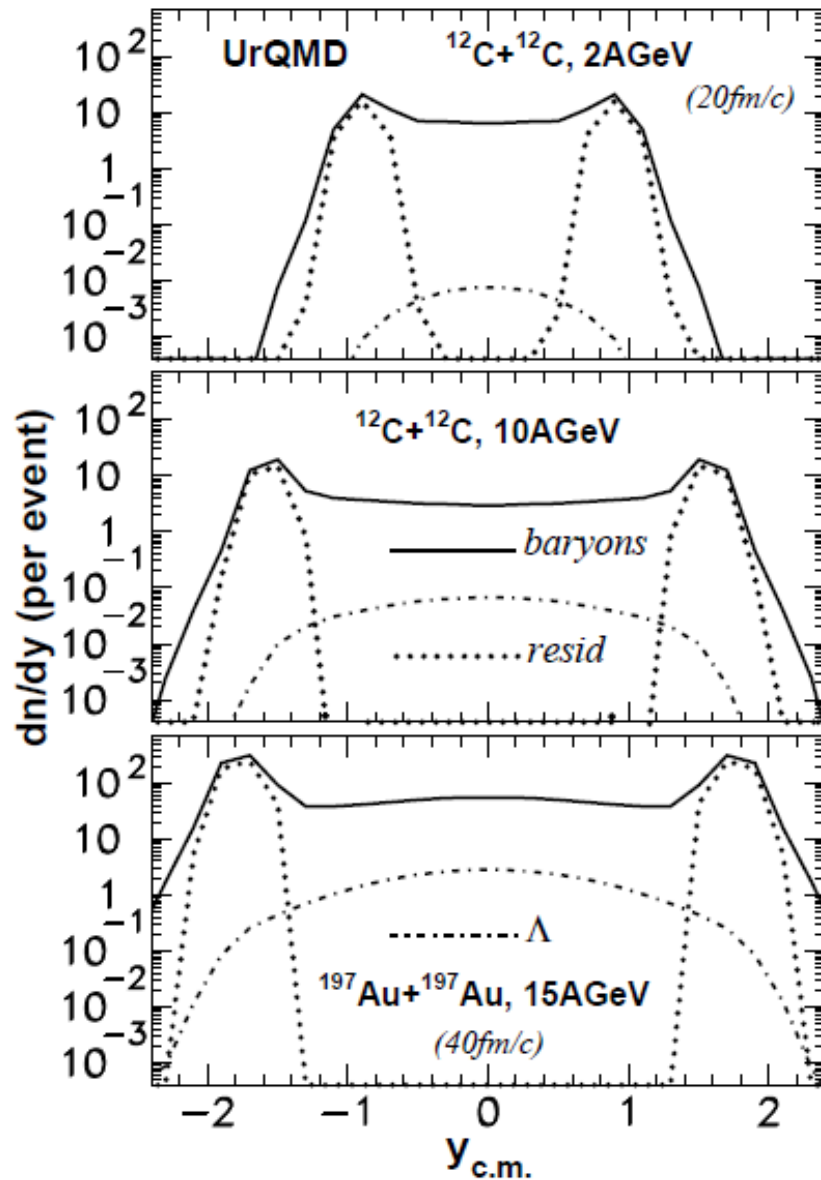
Combination of transport UrQMD and HSD models with CB:

- 1) In addition, coordinates of baryons and clusters are considered,
if $|\mathbf{X}_b - \mathbf{X}_A| < R * A^{1/3}$ the particle b may be included in A-cluster.
- 2) Spectators' nucleons are always included in the residues.

Investigation of fragments/hyperfragments at all rapidities !

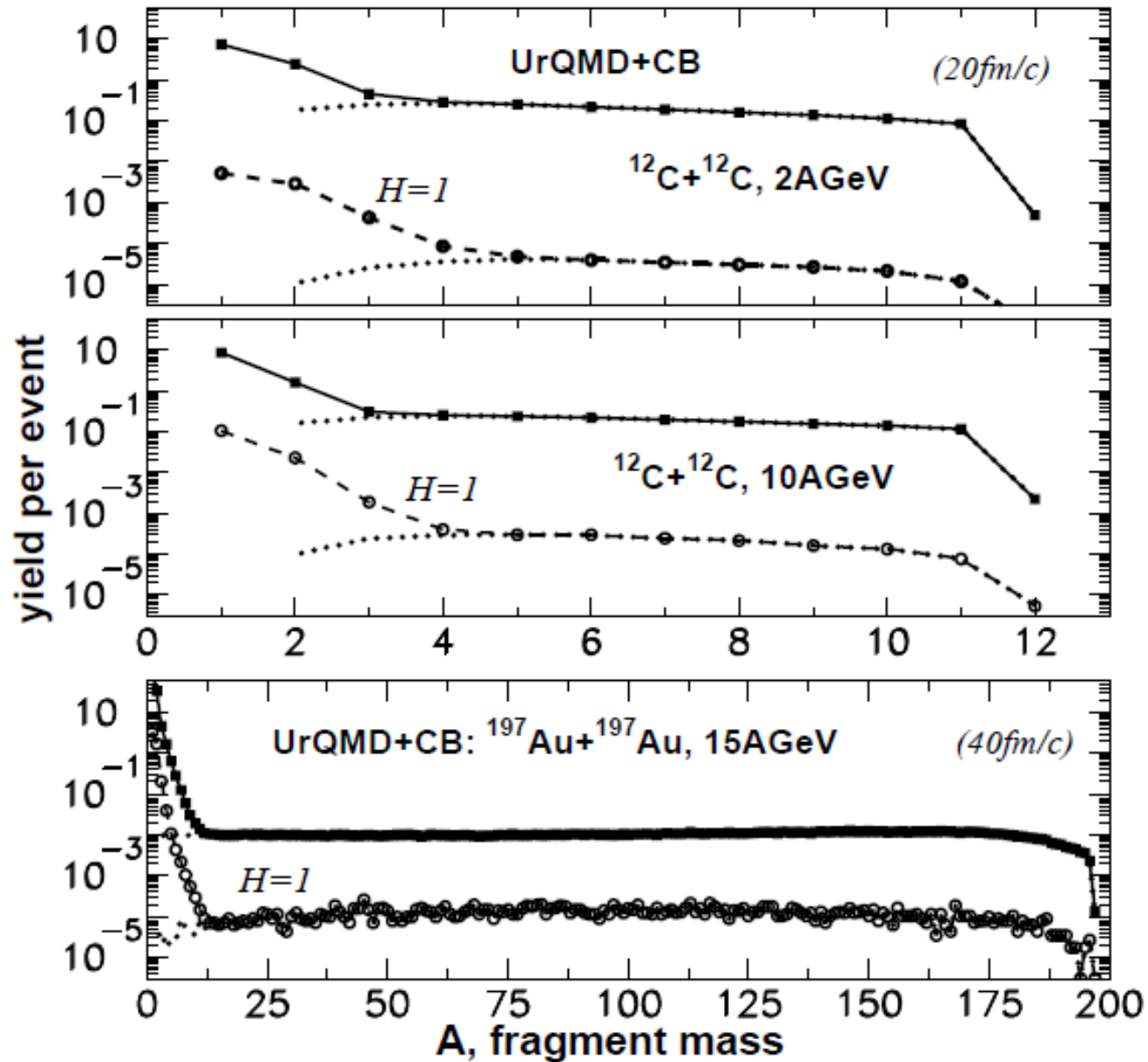
preliminary

baryons, Lambdas, hyper-fragments, hyper-residues



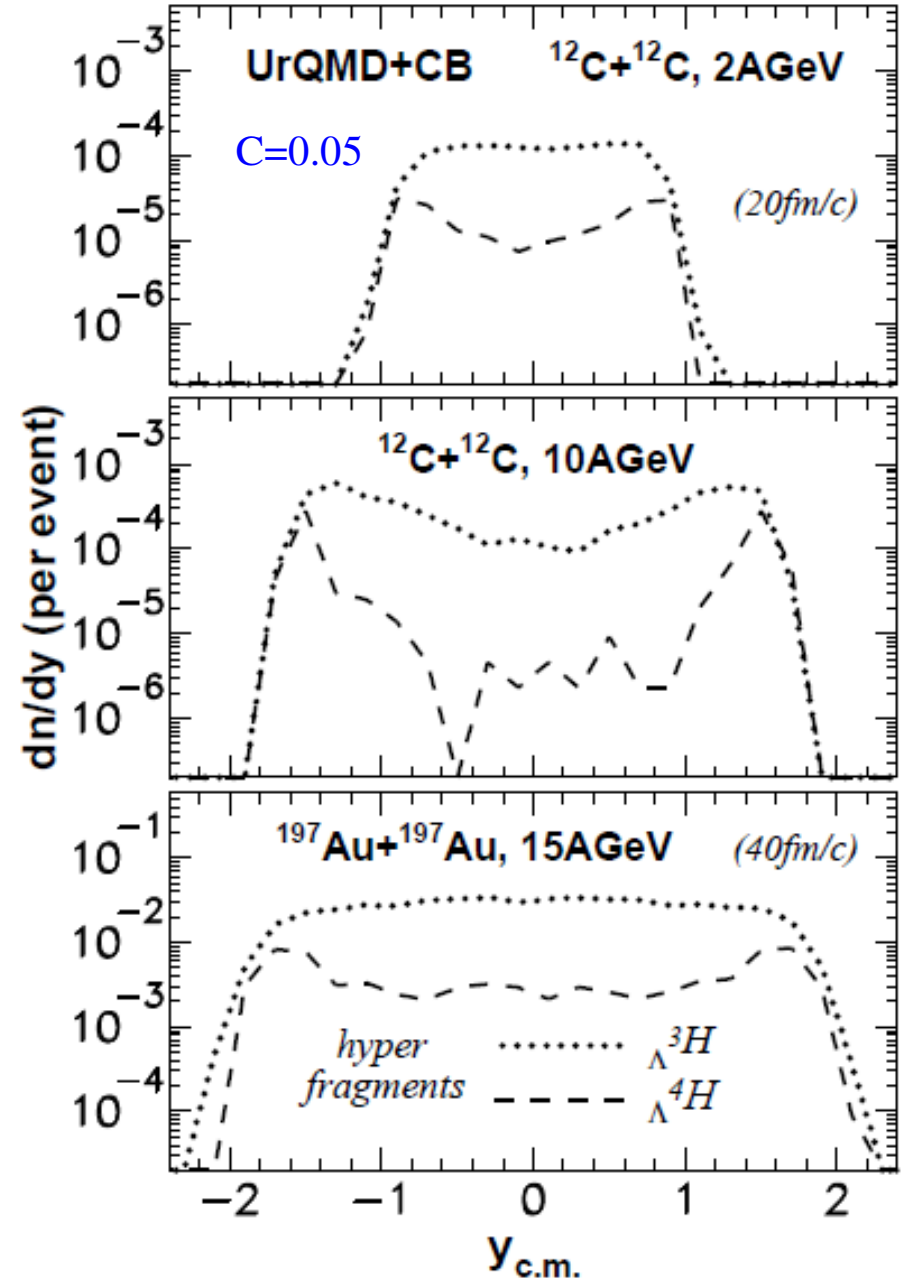
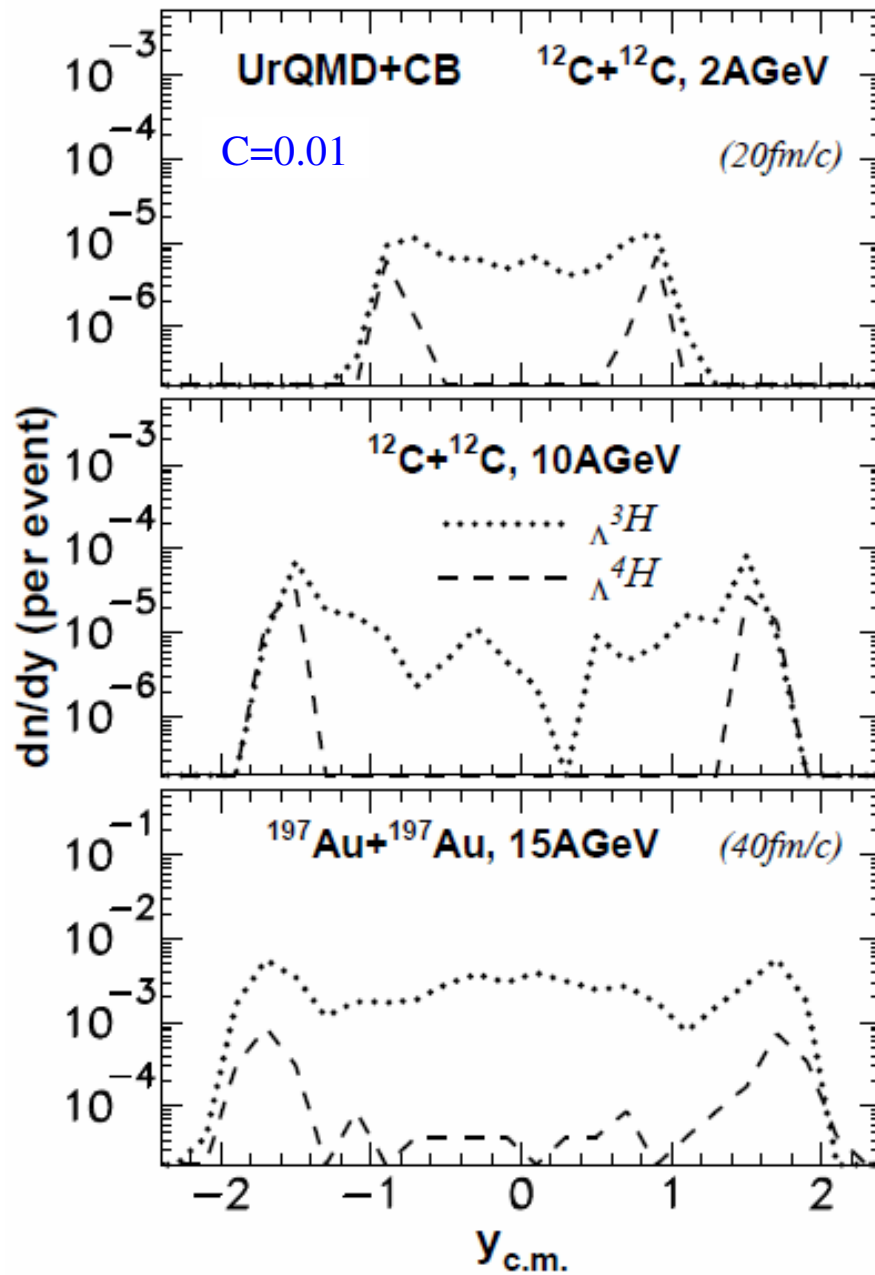
preliminary

normal fragments and hyper-fragments (with residue contribution)



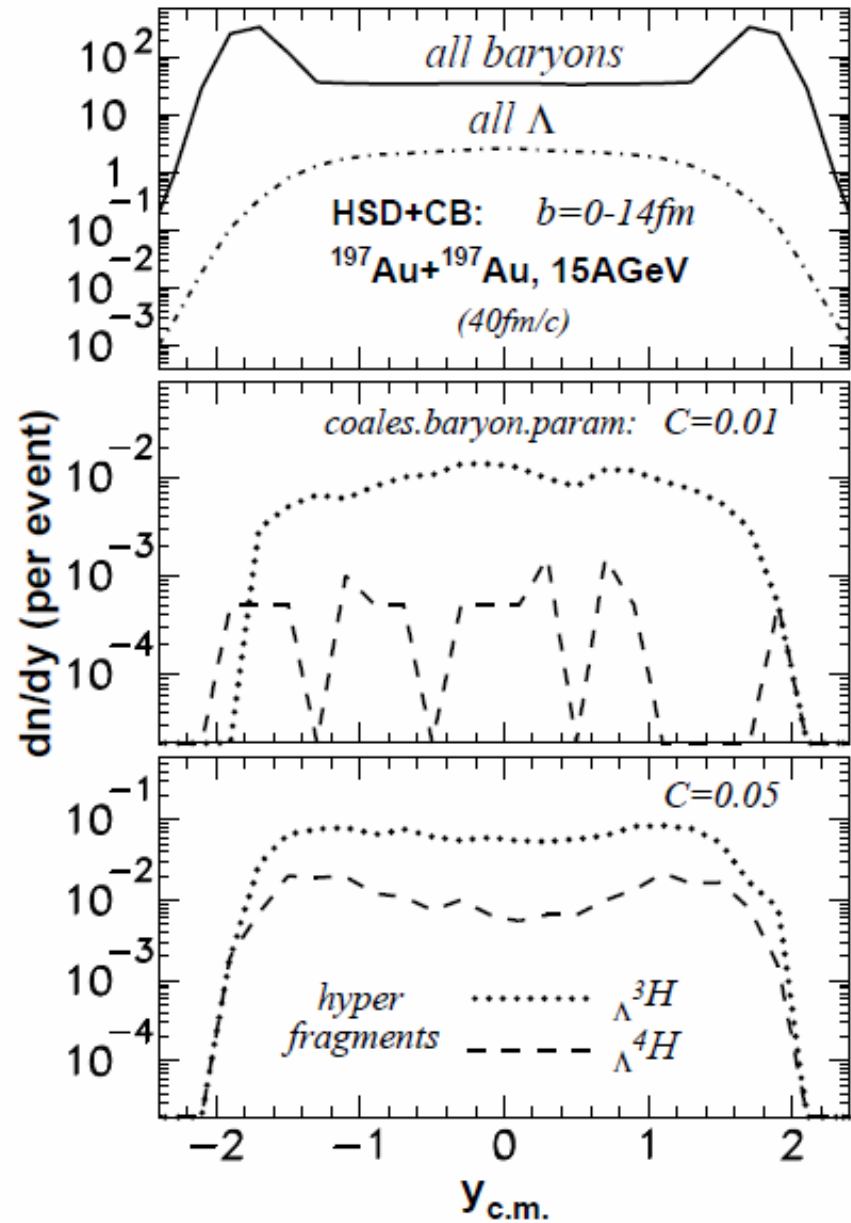
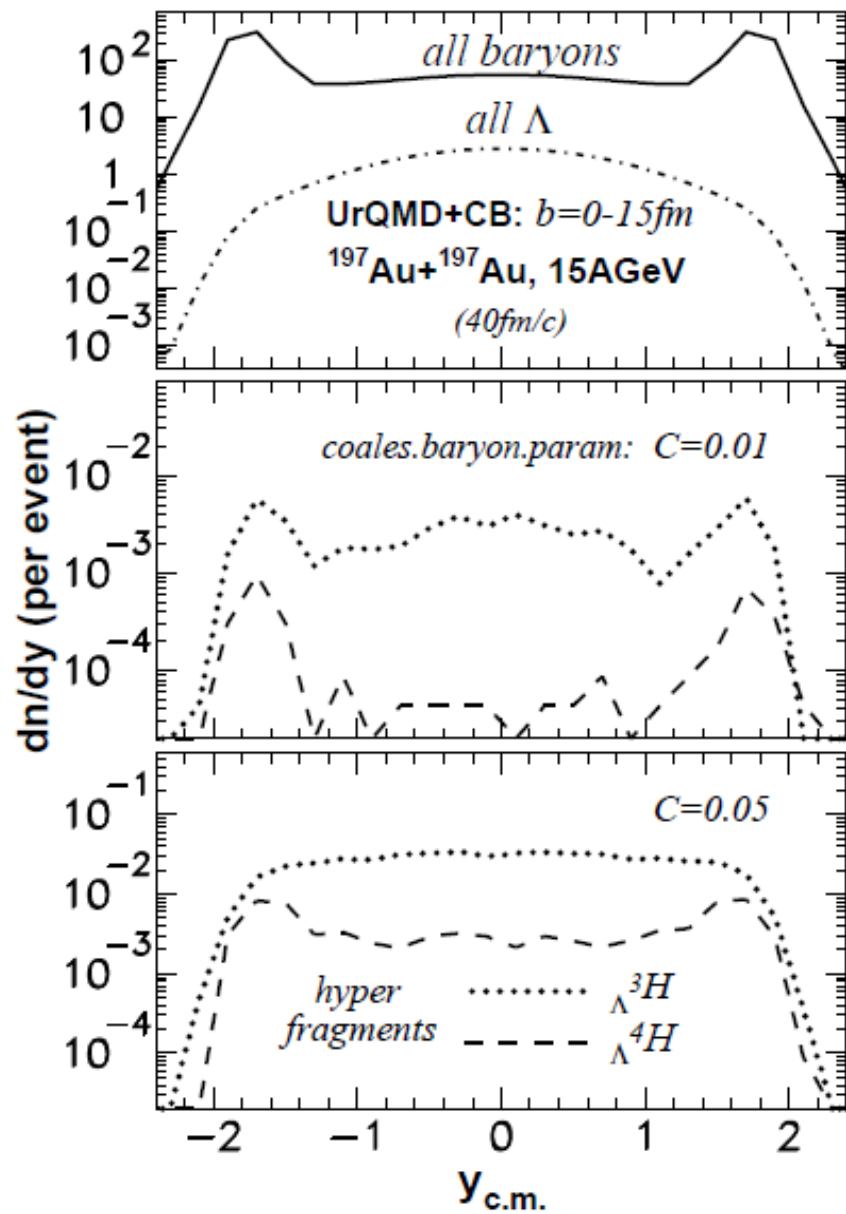
preliminary

light hyper-fragments



preliminary

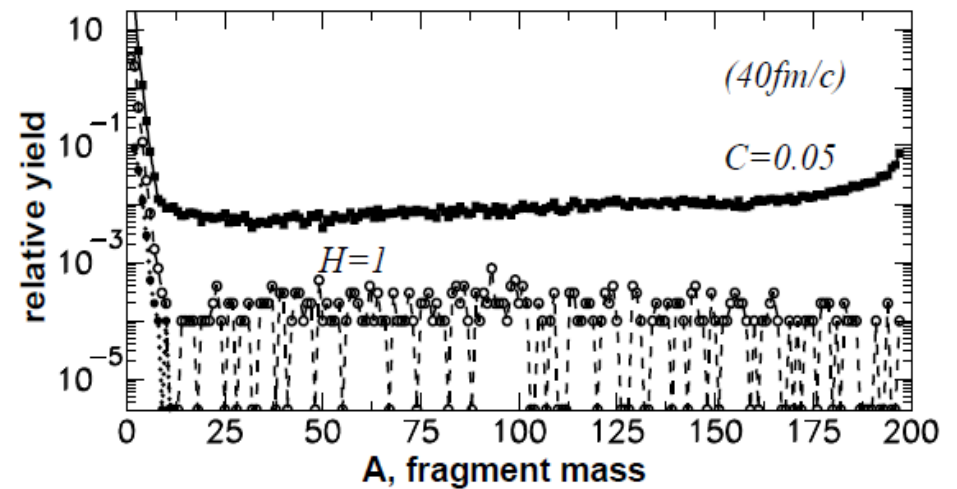
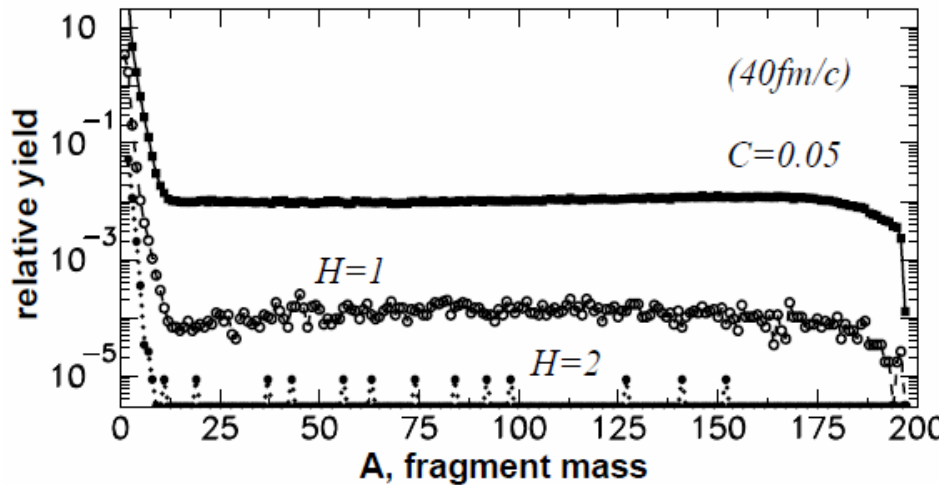
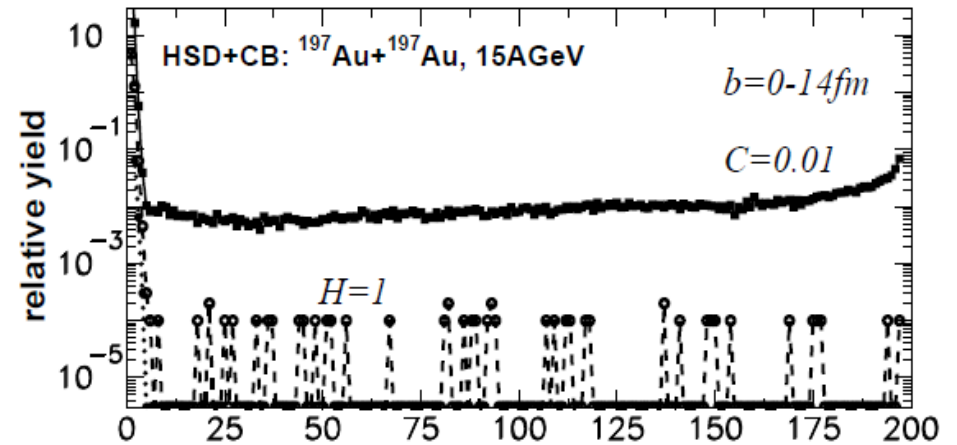
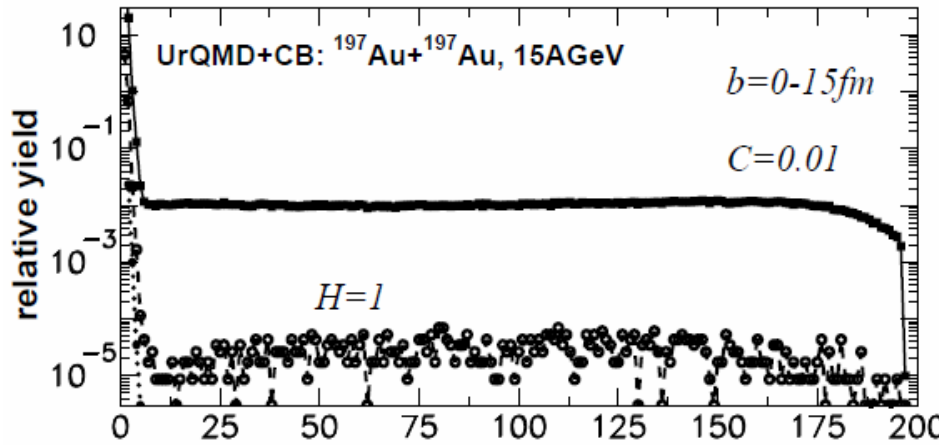
Transport models are consistent (UrQMD, HSD)



preliminary

Transport models are consistent (UrQMD, HSD)

Large fragments and hyperfragments:



Statistical approach for fragmentation of hyper-matter

$$Y_{AZH} = g_{AZH} V_f \frac{A^{3/2}}{\lambda_T^3} \exp \left[-\frac{1}{T} (F_{AZH} - \mu_{AZH}) \right]$$

$$\mu_{AZH} = A\mu + Z\nu + H\xi$$

mean yield of fragments with mass number A , charge Z , and Λ -hyperon number H

$$F_{AZH}(T, V) = F_A^B + F_A^S + F_{AZH}^{sym} + F_{AZ}^C + F_{AH}^{hyp}$$

liquid-drop description of fragments: bulk, surface, symmetry, Coulomb (as in Wigner-Seitz approximation), and hyper energy contributions

J.Bondorf et al., Phys. Rep. **257** (1995) 133

$$F_A^B(T) = \left(-w_0 - \frac{T^2}{\varepsilon_0} \right) A$$

$$F_A^S(T) = \beta_0 \left(\frac{T_c^2 - T^2}{T_c^2 + T^2} \right)^{5/4} A^{2/3}$$

parameters \approx Bethe-Weizsäcker formula:

$$w_0 = 16 \text{ MeV}, \beta_0 = 18 \text{ MeV}, T_c = 18 \text{ MeV}$$

$$F_{AZH}^{sym} = \gamma \frac{(A - H - 2Z)^2}{A - H}$$

$$\gamma = 25 \text{ MeV} \quad \varepsilon_0 \approx 16 \text{ MeV}$$

$$\sum_{AZH} AY_{AZH} = A_0, \quad \sum_{AZH} ZY_{AZH} = Z_0, \quad \sum_{AZH} HY_{AZH} = H_0.$$

chemical potentials are from mass, charge and Hyperon number conservations

$$F_{AH}^{hyp} = E_{sam}^{hyp} = H \cdot (-10.68 + 48.7/(A^{2/3})).$$

-- C.Samanta et al. J. Phys. G: 32 (2006) 363 (motivated: single Λ in potential well)

$$F_{AH}^{hyp} = (H/A) \cdot (-10.68A + 21.27A^{2/3}).$$

-- liquid-drop description of hyper-matter

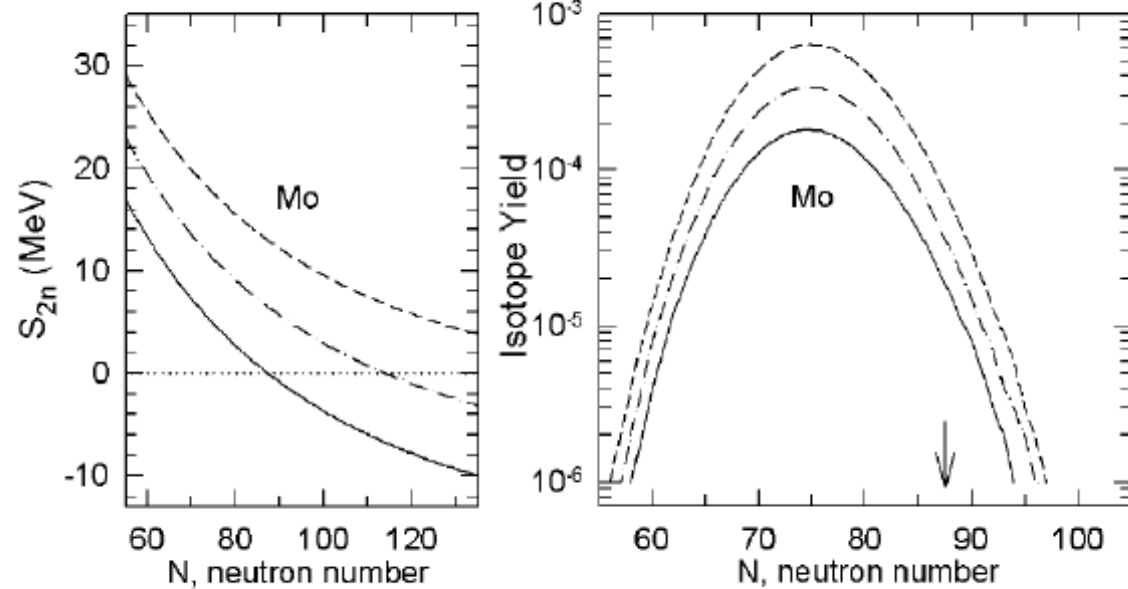
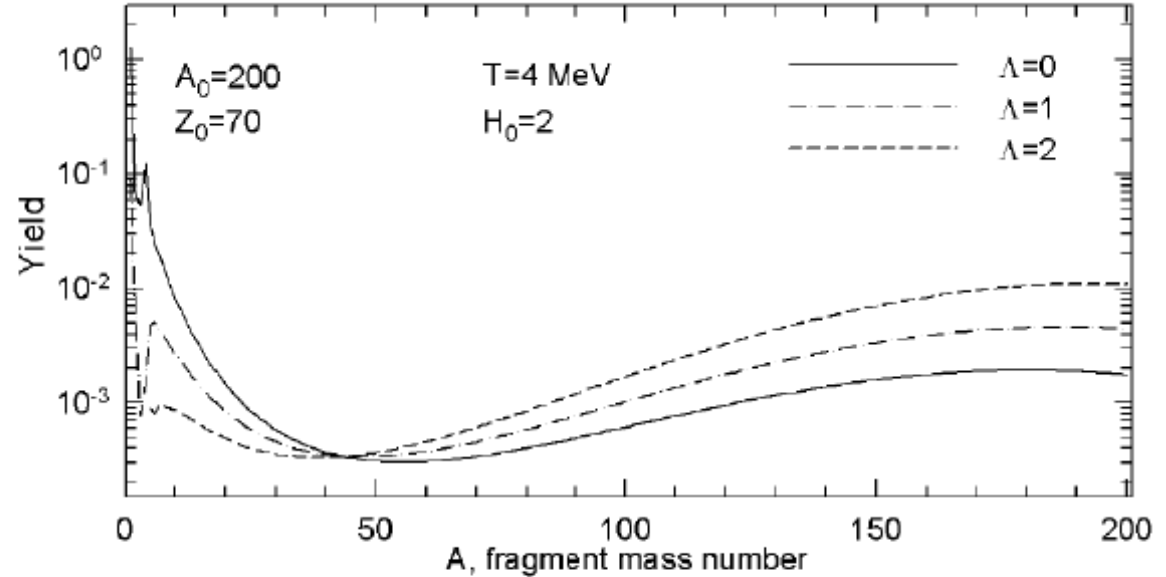
Break-up of excited hyper-residues

Normal nuclei + hypernuclei can be formed via evaporation, fission and multifragmentation processes.

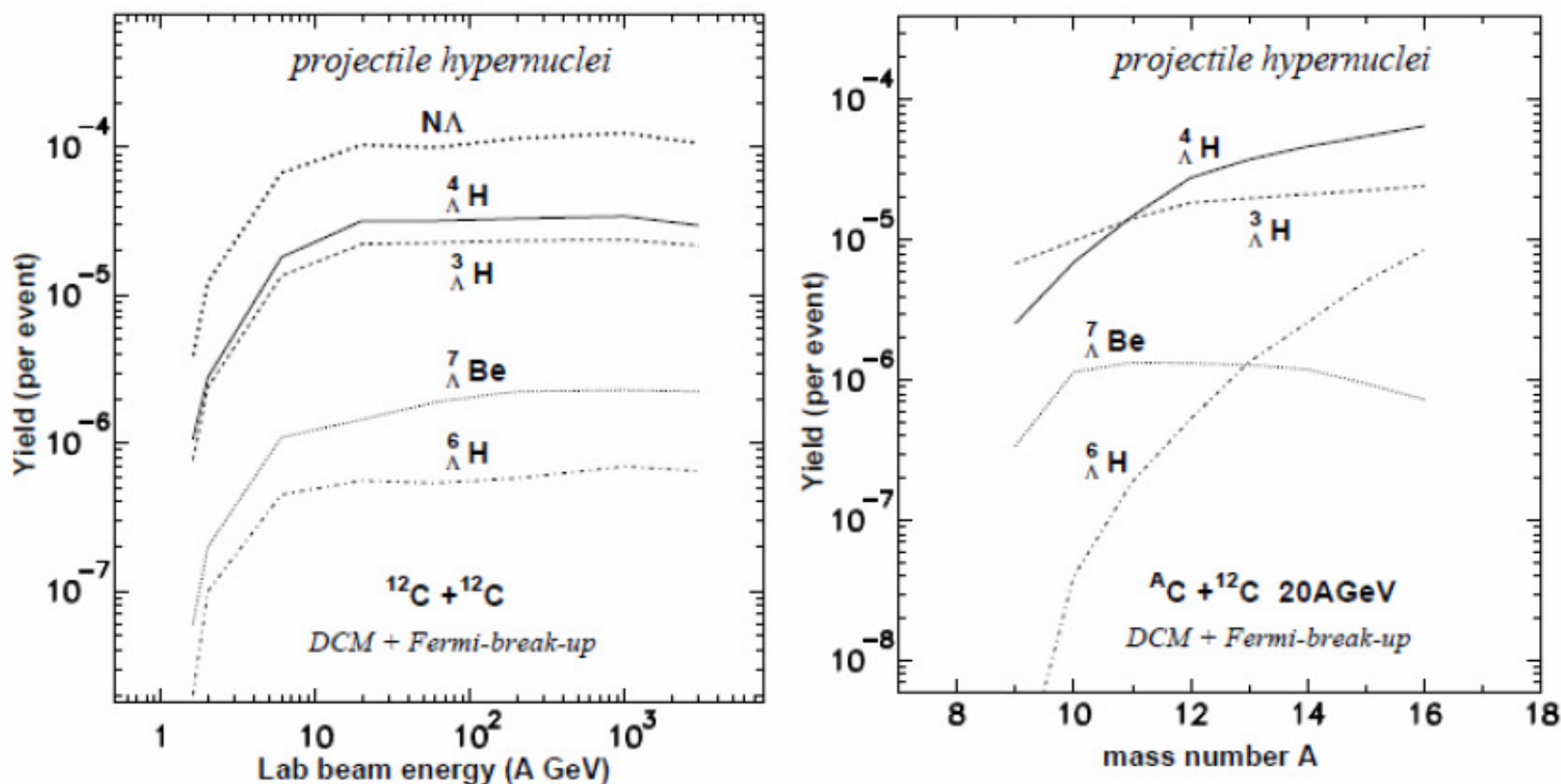
Liquid-gas type phase transition in hyper-matter is expected at subnuclear densities.

Very broad distributions of nuclei similar to ones in normal nuclear matter. At moderate temperatures hyperons concentrate in large species

Important: formed hypernuclei can reach beyond traditional neutron and proton drip-lines



Production of light hypernuclei in relativistic ion collisions



One can use exotic neutron-rich and neutron-poor projectiles, which are not possible to use as targets in traditional hyper-nuclear experiments, because of their short lifetime. Comparing yields of hypernuclei from various sources we can get info about their binding energies and properties of hyper-matter.

A.S.Botvina, K.K.Gudima, J.Pochodzalla, PRC 88 (2013) 054605

Conclusions

Peripheral collisions of relativistic ions and hadrons with nuclei are promising reactions for novel nuclear physics research at BM@N.

A mechanism of formation of hypernuclei:

Relativistic hadron and peripheral ion collisions : Strange baryons (Λ , Σ , Ξ , ...) are transported to the spectator residues and are captured in nuclear matter. These strange systems are excited and after decay of such systems hypernuclei of all sizes (and isospin), including exotic weakly-bound states, multi-strange nuclei, and those beyond the drip-lines can be produced.

Advantages over other reactions producing hypernuclei: there is no limit on sizes and isotope content of produced nuclei; probability of their formation is very high; a large strangeness can be deposited in nuclei. After decay of such hypernuclei exotic normal nuclei can be obtained. EOS of hypermatter at subnuclear density can be investigated.

Description of elementary interactions in DCM transport code

