


Quark-Gluon Plasma Formation in Heavy Ion Collisions in Holographic Description



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Outlook

- Quark-Gluon Plasma(QGP) in heavy-ions collisions(HIC)
- Holography description of QGP in equilibrium
- Holography description of formation of
QGP in HIC \Leftrightarrow Black Holes formation in AdS
- Thermalization time/Dethermalization time
- Non-central collisions in holography description

Quark-Gluon Plasma (QGP): a new state of matter

QGP is a state of matter formed from deconfined quarks, antiquarks, and gluons at high temperature

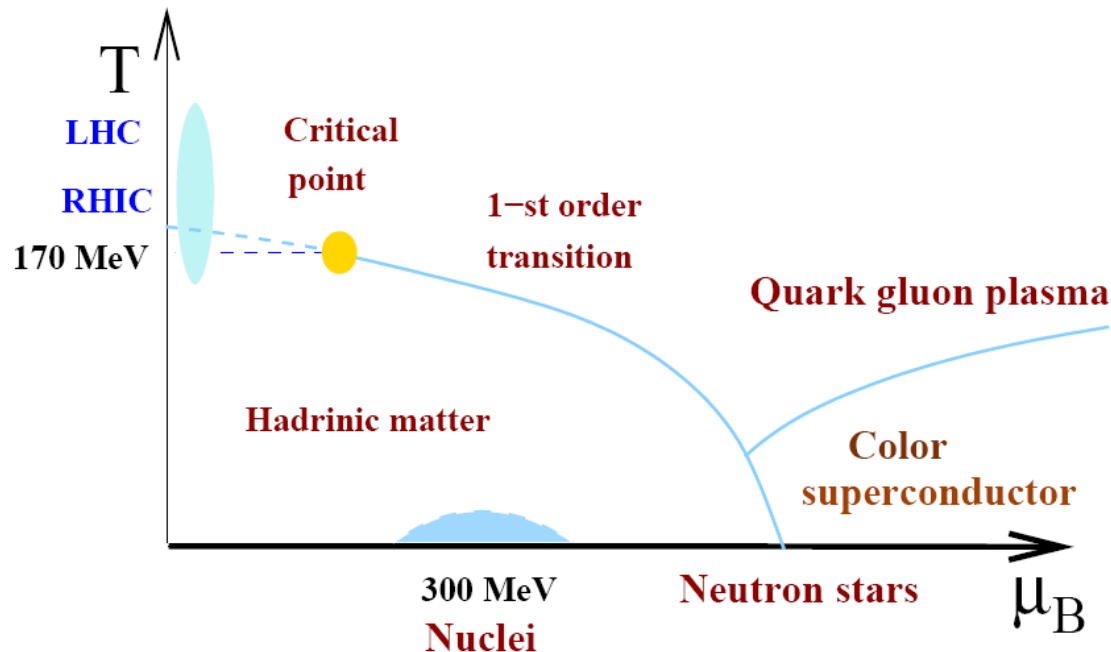
QCD: asymptotic freedom, quark confinement

T increases, or
density increases

nuclear
matter



Deconfined
phase



Experiments: Heavy Ions collisions produced a medium

HIC are studied in several **experiments:**

- started in the 1990's at the Brookhaven Alternating Gradient Synchrotron (AGS),
- the CERN Super Proton Synchrotron (SPS)
- the Brookhaven Relativistic Heavy-Ion Collider (RHIC)
- the LHC collider at CERN.

$$\sqrt{s_{NN}} = 4.75 GeV$$

$$\sqrt{s_{NN}} = 17.2 GeV$$

$$\sqrt{s_{NN}} = 200 GeV$$

$$\sqrt{s_{NN}} = 2.76 TeV$$

Fireball at the LHC is denser, larger and longer lived than at RHIC.

$$\epsilon \sim 10 GeV/fm^3, V \sim 4800 fm^3, \tau_{life} \sim 10 fm/c$$

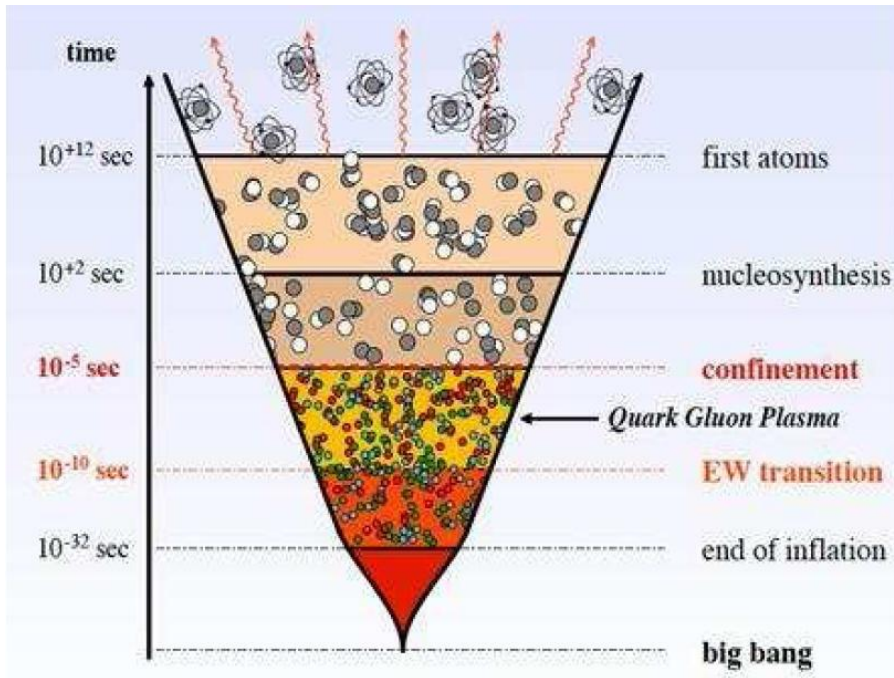
There are strong experimental evidences that RHIC or LHC have created some medium which behaves collectively:

- modification of particle spectra (compared to p+p)
- jet quenching
- high p_T -suppression of hadrons
- elliptic flow
- suppression of quarkonium production

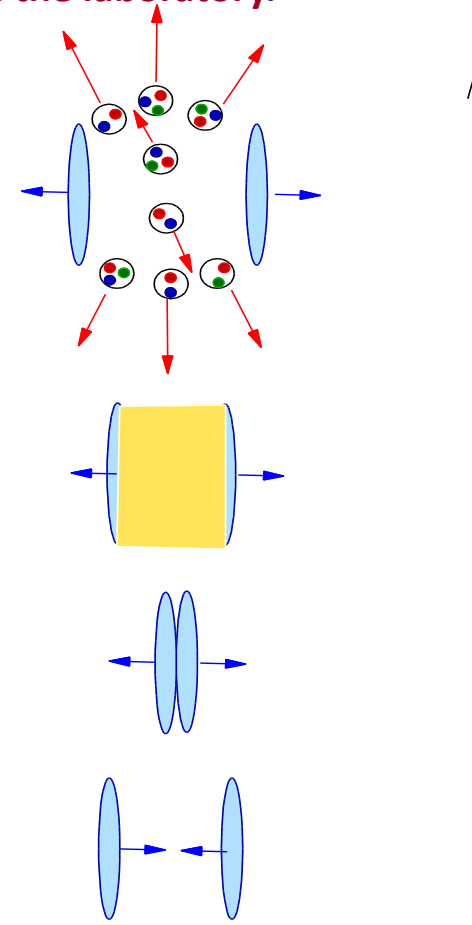
Study of this medium is also related with study of Early Universe

QGP in Heavy Ion Collision and Early Universe

- One of the fundamental questions in physics is: what happens to matter at extreme densities and temperatures as may have existed in the first microseconds after the Big Bang
- The aim of heavy-ion physics is to create such a state of matter **in the laboratory**.

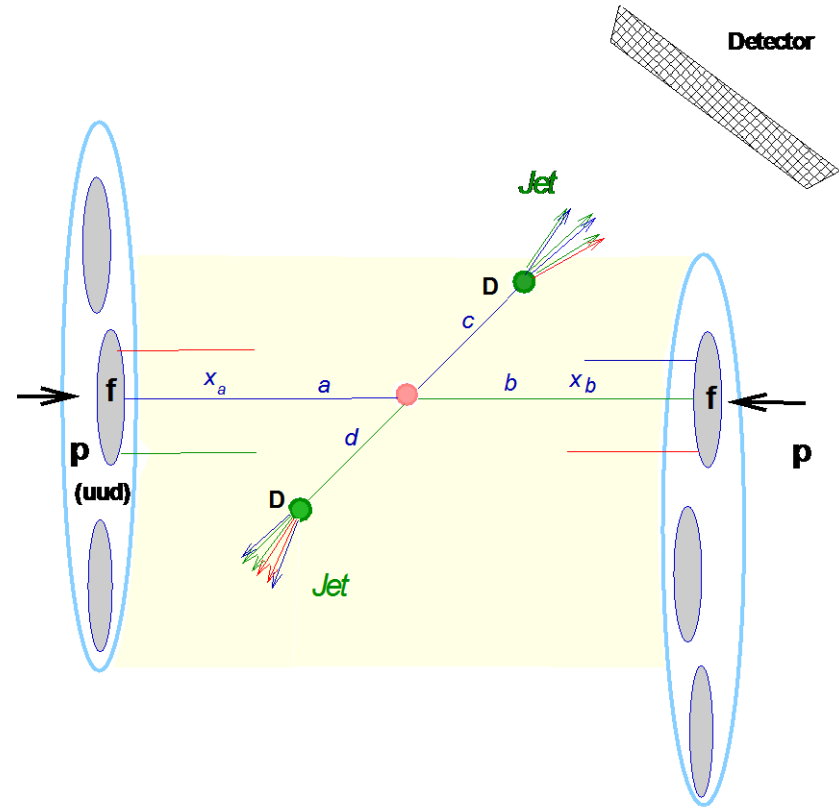
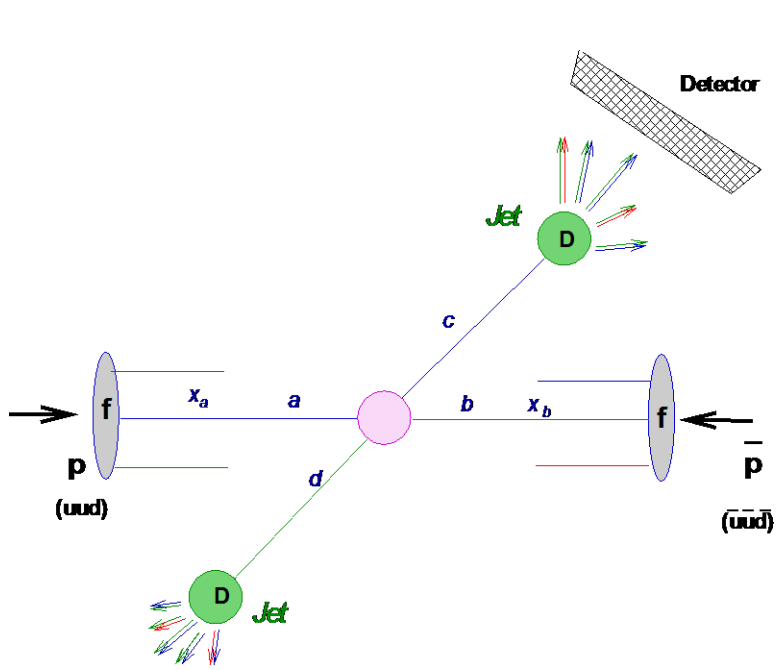


Evolution of the Early Universe

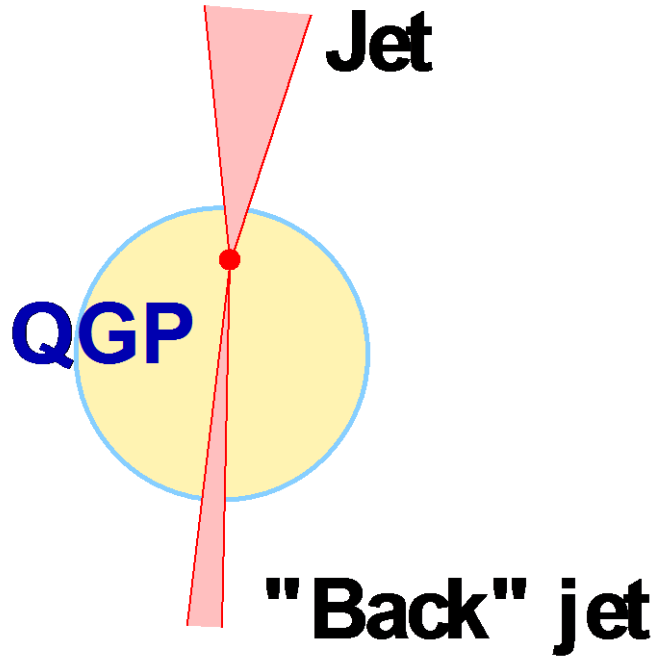


Evolution of a Heavy Ion Collision

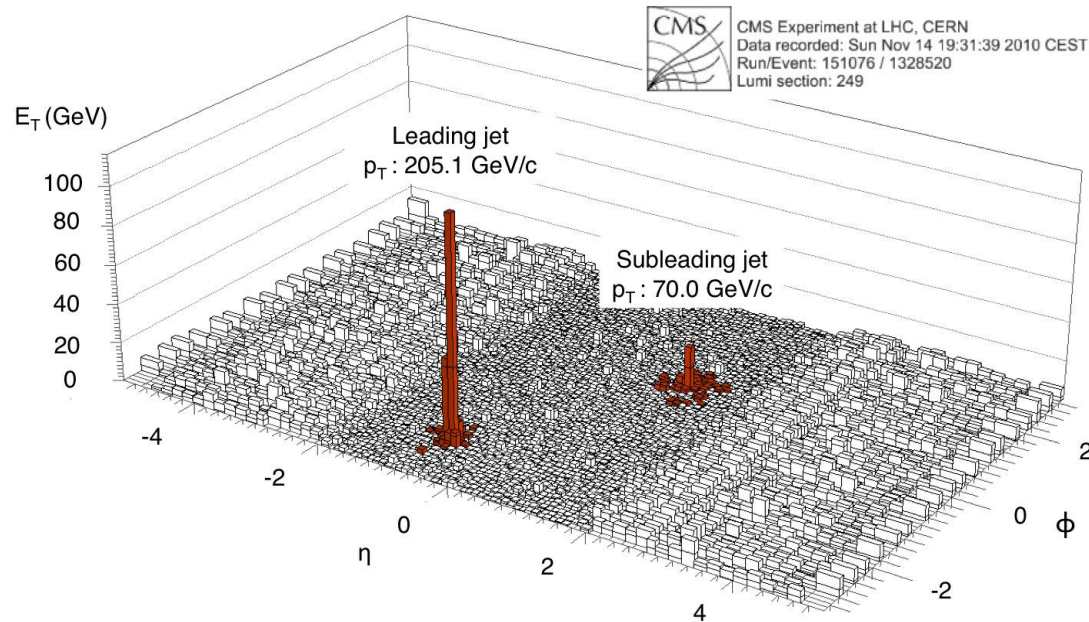
pp collisions vs heavy ions collisions



Jet quenching

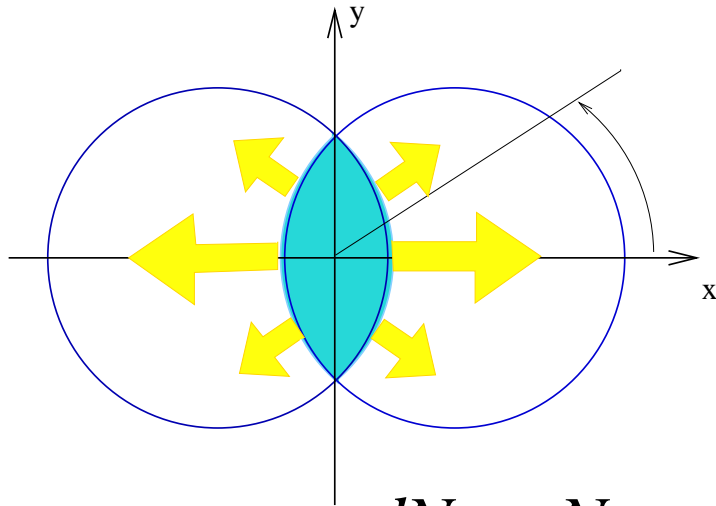


Central collision



P. Sorensen, Highlights from Heavy Ion Collisions at RHIC....., 1201.0784[nucl-ex]

Elliptic flow



Non-central collision

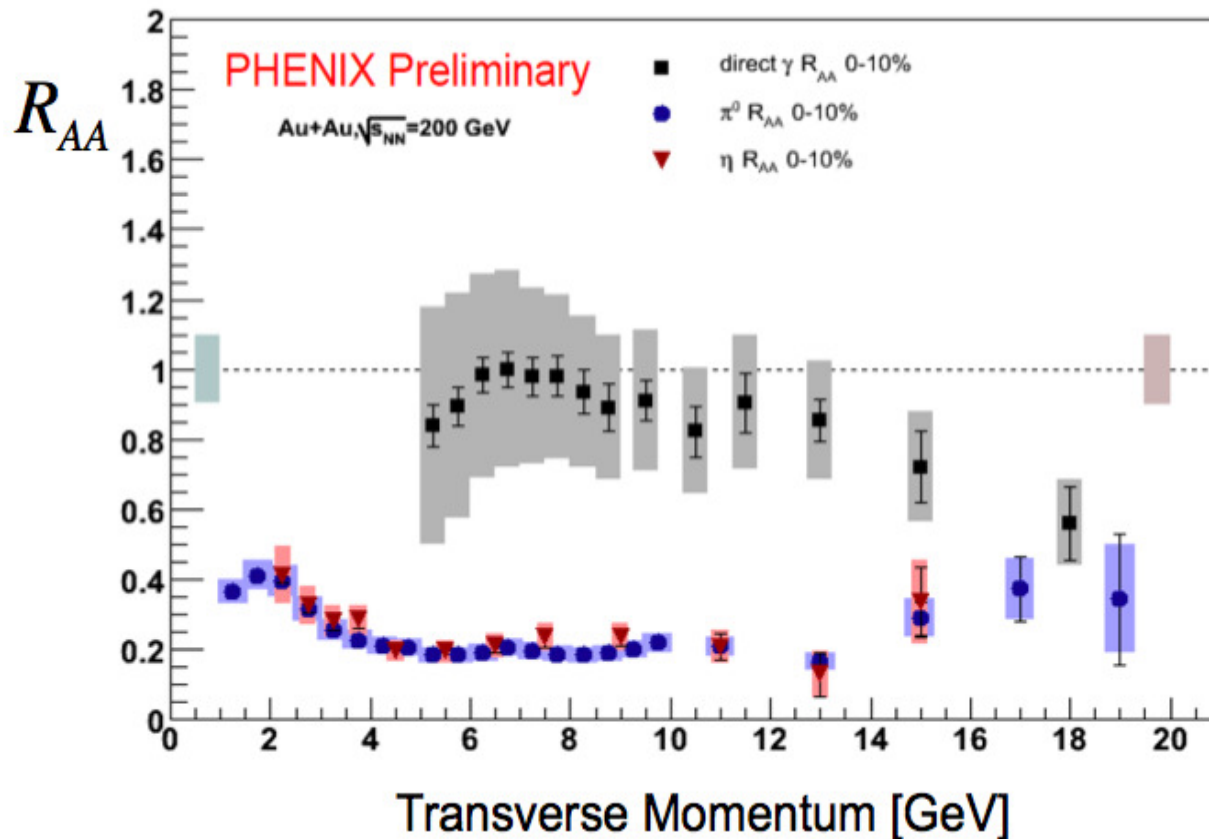
$$\frac{dN}{d\varphi} = \frac{N}{2\pi} (1 + v_2(p_{\perp}, b) \cos(2\varphi) + \dots)$$

Imprints of anisotropies are more essential for small shear viscosity, since usually large viscosity erases stronger irregularity

$$\eta/s \approx 0.03 - 0.15$$

The nuclear modification factor

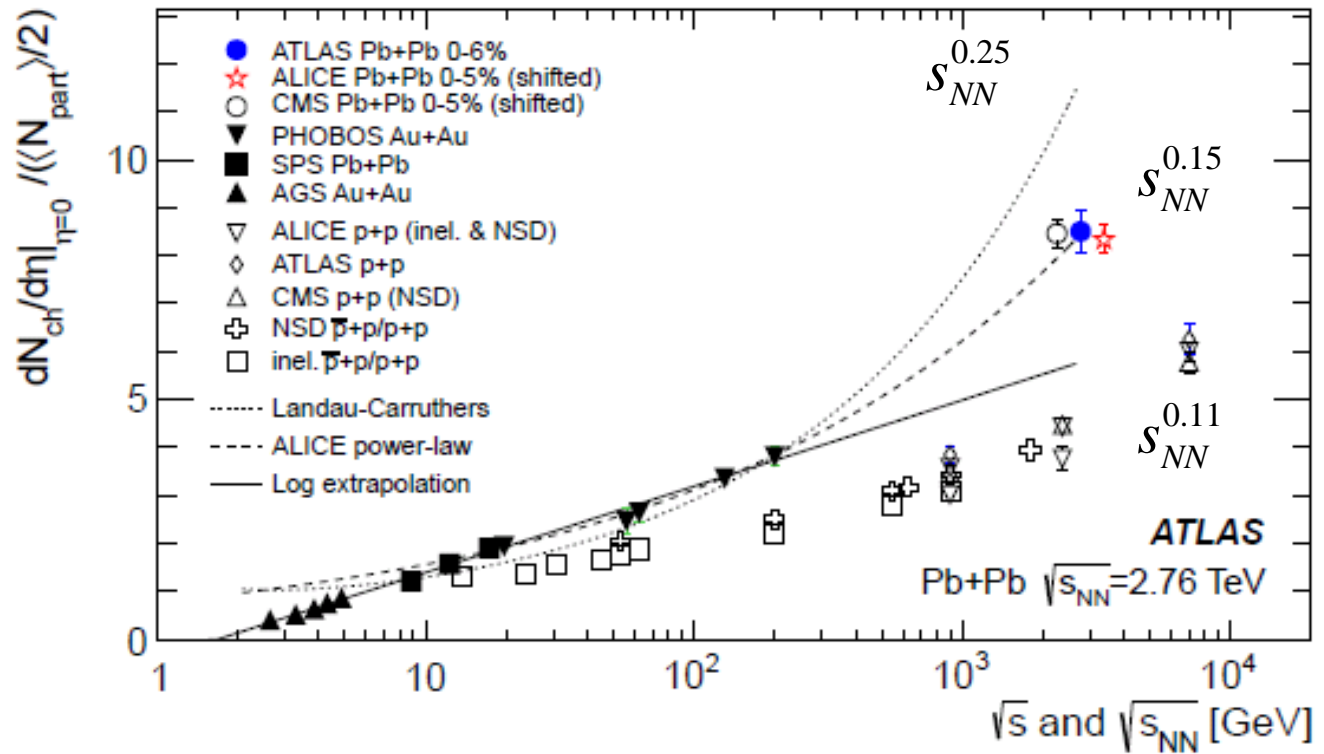
$$R_{AB}^h(p_T, \eta, \text{centrality}) = \frac{\frac{dN_{\text{medium}}^{AB \rightarrow h}}{dp_T d\eta}}{\langle N_{\text{coll}}^{AB} \rangle \frac{dN_{\text{vacuum}}^{pp \rightarrow h}}{dp_T d\eta}}$$



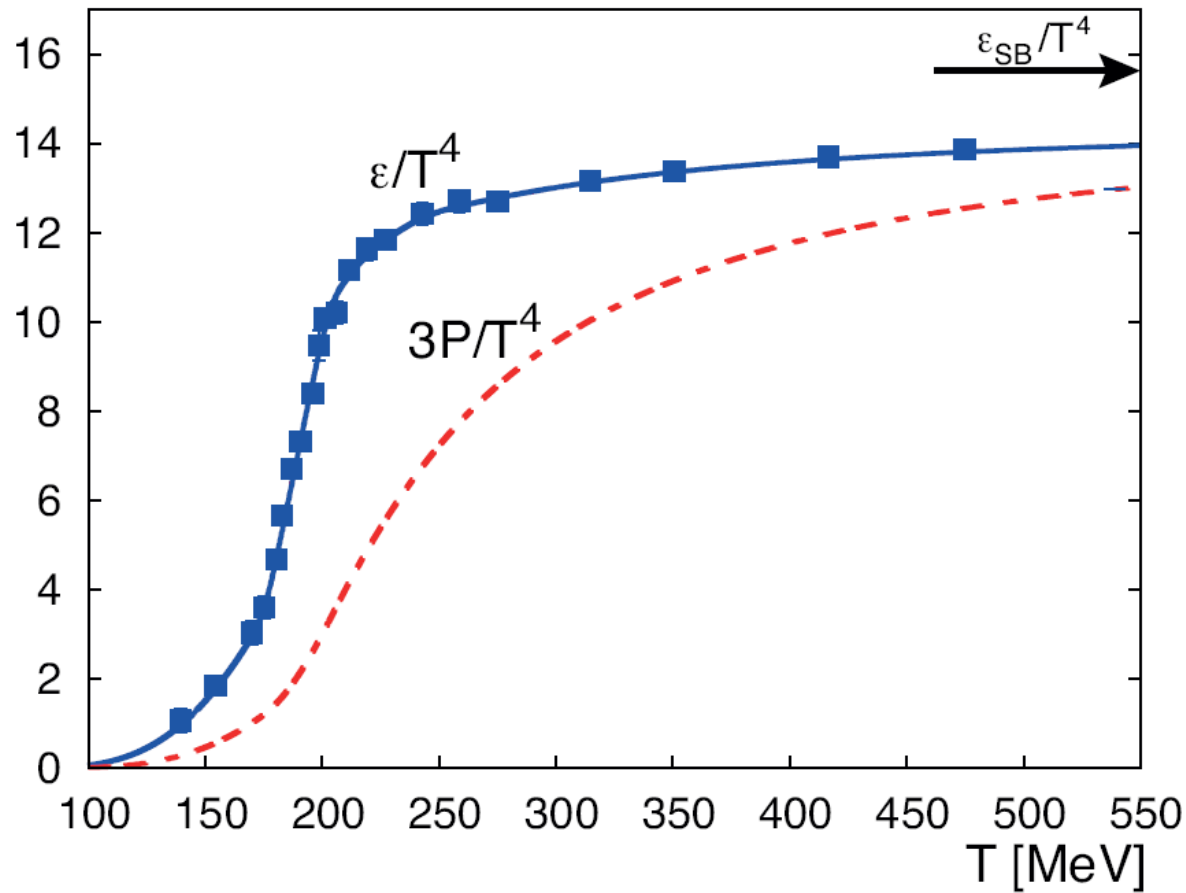
Multiplicity: Landau's/Holographic formula vs experimental data

Landau formula

$$\mathcal{M} \sim S_{NN}^{1/4}$$



Plot from: ATLAS
Collaboration
1108.6027



lattice calculation of QCD thermodynamics $N_f = 3$

S. Borsanyi et al., "The QCD equation of state with dynamical quarks," arXiv:1007.2580

QGP as a strongly coupled fluid

- Conclusion from the RHIC and LHC experiments: appearance of QGP (not a weakly coupled gas of quarks and gluons, but **a strongly coupled fluid**).
- This makes perturbative methods inapplicable
- The lattice formulation of QCD does not work, since we have to study real-time phenomena.
- This has provided a motivation to try to understand the dynamics of QGP through the **gauge/string duality**

Dual description of QGP as a part of Gauge/string duality

- There is not yet exist a gravity dual construction for QCD.
- Differences between $N = 4$ SYM and QCD are less significant, when quarks and gluons are in the deconfined phase (because of the conformal symmetry at the quantum level $N = 4$ SYM theory does not exhibit confinement.)
- Lattice calculations show that QCD exhibits a quasi-conformal behavior at temperatures $T > 300$ MeV and the equation of state can be approximated by $E = 3 P$ (a traceless conformal energy-momentum tensor).
- The above observations, have motivated to use the AdS/CFT correspondence as a tool to get non-perturbative dynamics of QGP.
- There is the considerable success in description of the static QGP.

“Holographic description of quark-gluon plasma”



- Holographic description of quark-gluon plasma **in equilibrium**
- Holography description of quark-gluon plasma **formation** in heavy-ions collisions

Holographic description of QGP

(QGP in equilibrium)

Holography for thermal states

TQFT in
 M_D -spacetime

=

Black hole
in AdS_{D+1} -space-time

TQFT = QFT with temperature

AdS/CFT correspondence in Euclidean space. $T \neq 0$

To compute the Matsubara correlator at finite temperature

$$G^E(k_E) = \int d^4 x_E e^{-ik_E \cdot x_E} \langle T_E \hat{O}(x_E) \hat{O}(0) \rangle$$

Here T_E denotes Euclidean time ordering

Euclidean time coordinate τ is periodic, $\tau \sim \tau + T^{-1}$

$$\begin{aligned} & \langle e^{\int_{\partial M} \phi_0 \mathcal{O}} \rangle \\ &= e^{S_g[\phi_c(\phi_0)]} \end{aligned}$$

$$\underbrace{\phi(\tau, \vec{x}, z)}_{\mathcal{X}_E}, \quad S_g[\phi], \quad \delta S_g[\phi_c] = 0 \quad \phi_c|_{\partial M} = \phi_0 \quad \phi_c = \phi_c(\phi_0)$$

boundary condition $\phi|_{z=0} = \phi_0$

+ requirement of regularity at horizon

$$\mathbf{g}: \quad ds^2 = \frac{R^2}{z^2} \left(f(z) d\tau^2 + d\mathbf{x}^2 + \frac{dz^2}{f(z)} \right) + R^2 d\Omega_5^2$$

$$f(z) = 1 - z^4/z_H^4 \quad z_H = (\pi T)^{-1}$$

T is the Hawking temperature

$$0 < z < z_H$$

Correlators with $T \neq 0$ AdS/CFT

Example. $D=2$

$$\langle \mathcal{O}(t, \mathbf{x}) \mathcal{O}(t, \mathbf{x}') \rangle_{ren} \sim e^{-\Delta \delta \mathcal{L}} \quad \mathbf{x} - \mathbf{x}' = \ell$$

$$\delta \mathcal{L} \equiv \mathcal{L} + 2 \ln(z_0/2)$$

Vacuum correlators $M=\text{AdS}$

$$\delta \mathcal{L}_{\text{vacuum}}(\ell) = 2 \ln \frac{\ell}{2}$$

Temperature $M=\text{BHAdS}$ with

r_H

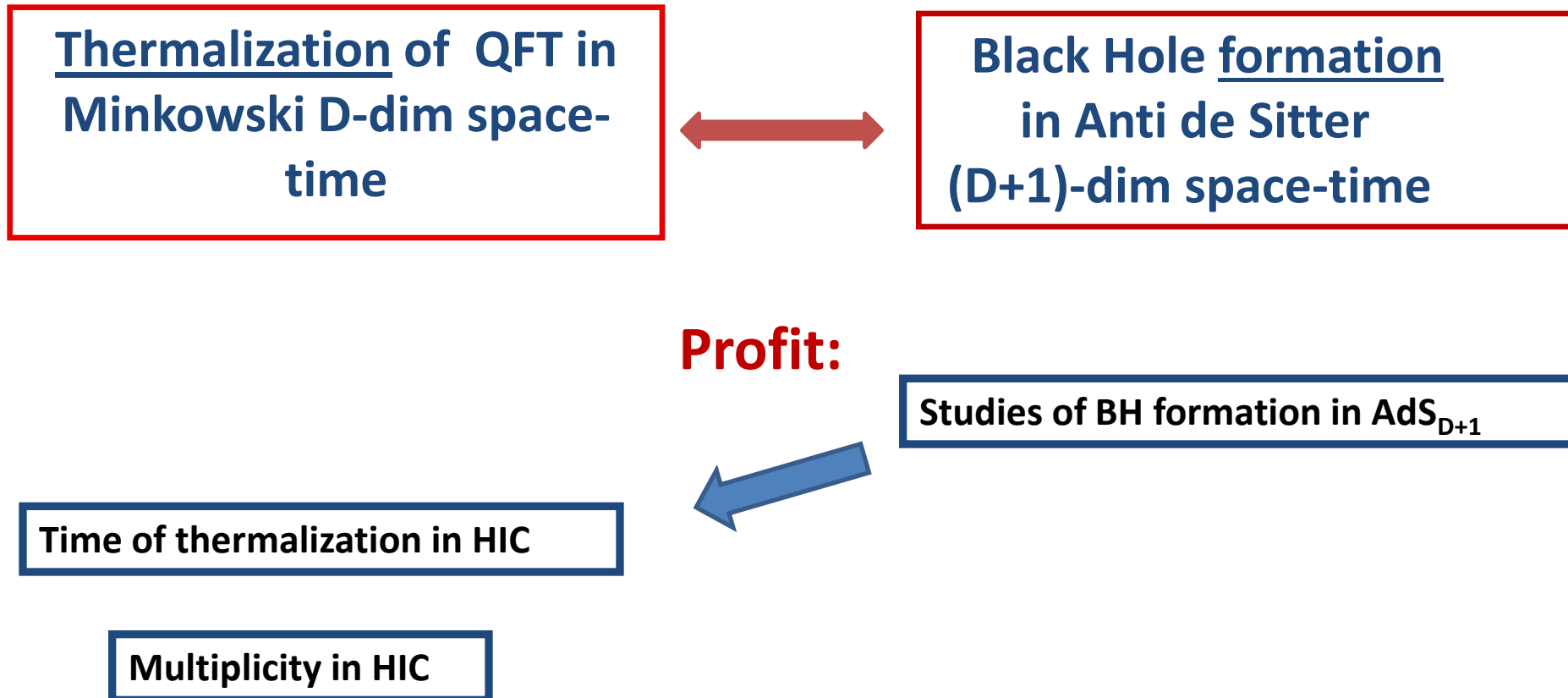
$$t \geq \ell/2$$

$$\delta \mathcal{L}_{\text{thermal}}(\ell) = 2 \ln \frac{\sinh \frac{r_H \ell}{2}}{r_H}$$

Bose gas

Holographic Description of Formation of QGP

Holographic thermalization



Formation of BH in AdS. Deformations of AdS metric leading to BH formation

- colliding gravitational shock waves

Gubser, Pufu, Yarom, Phys.Rev. , 2008 (I)
Gubser, Pufu, Yarom, JHEP , 2009 (II)
Alvarez-Gaume, C. Gomez, Vera,
Tavanfar, Vazquez-Mozo, PRL, 2009
IA, Bagrov, Guseva, Joukowskaya, E.Pozdeeva
2009, 2010, 2012 JHEP
Kiritsis, Taliotis, 2011 JHEP

- drop of a shell of matter with vanishing rest mass

("null dust"),

infalling shell geometry = Vaidya metric

Danielsson, Keski-Vakkuri , Kruczenski, 1999

.....

Balasubramanian +9. PRL, 2011, Phys.Rev.2011

- sudden perturbations of the metric near the boundary that propagate into the bulk

Chesler, Yaffe, PRL, 2011

Deformations of AdS metric by infalling shell

d+1-dimensional infalling shell geometry is described in Poincar'e coordinates by the Vaidya metric

Danielsson, Keski-Vakkuri and Kruczenski

$$ds^2 = \frac{1}{z^2} \left[- \left(1 - m(v) z^d \right) dv^2 - 2dz dv + d\mathbf{x}^2 \right] \quad \star$$

v labels ingoing null trajectories

- 1) For constant $m(v) = M$, the coordinate transformation $dv = dt - \frac{dz}{1 - M z^d}$ brings \star in the form

$$ds^2 = \frac{1}{z^2} \left[- \left(1 - M z^d \right) dt^2 + \frac{dz^2}{1 - M z^d} + d\mathbf{x}^2 \right]$$

- 2) $m(v) = \frac{M}{2} \left(1 + \tanh \frac{v}{v_0} \right)$

\star interpolates between vacuum AdS inside the shell and an AdS black brane

Correlators via Geodesics in AdS/CFT

$$\langle \mathcal{O}_\Delta(\tau_1, \vec{x}_1) \mathcal{O}_\Delta(\tau_2, \vec{x}_2) \rangle = \int_{\mathcal{P} \in M} D\mathcal{P} e^{i\Delta L(\mathcal{P})}$$

$$\mathcal{P} \in M$$

$$(\tau_1, \vec{x}_1) \in \partial M$$

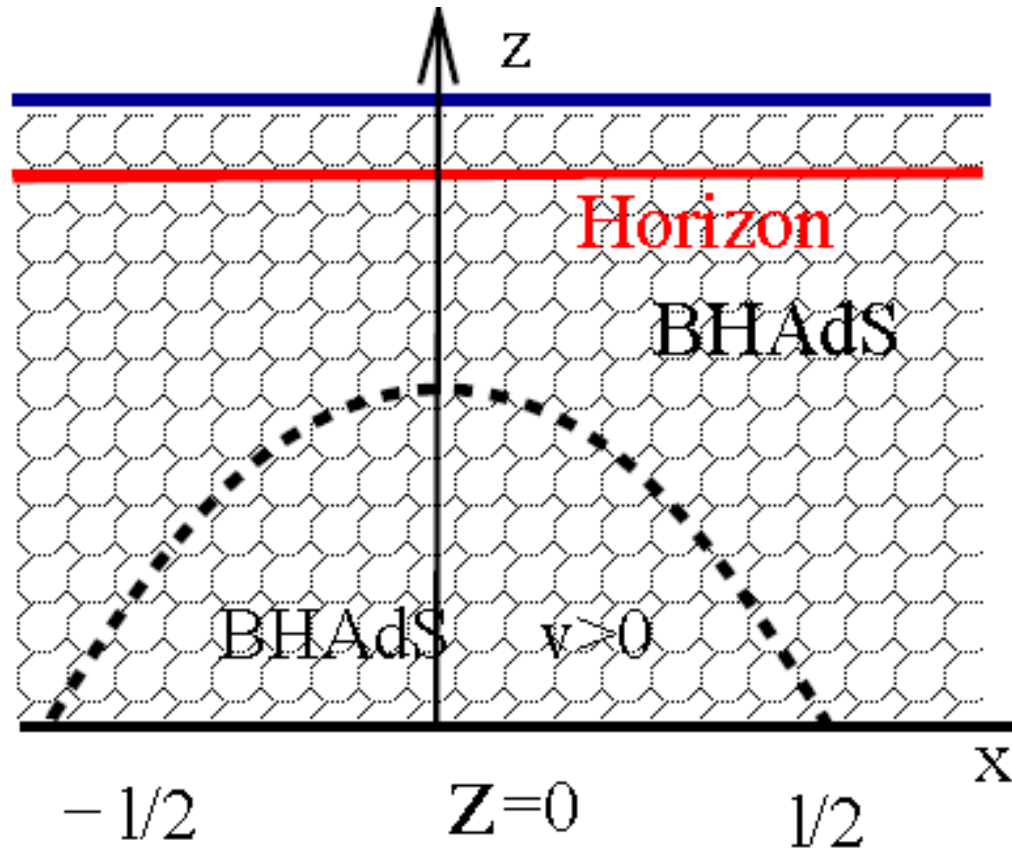
$$L(\mathcal{P}) = \int_{(\tau_2, \vec{x}_2) \in \partial M} (-g_{\mu\nu} \dot{X}^\mu \dot{X}^\nu)^{1/2}$$

$$(\tau_2, \vec{x}_2) \in \partial M$$

Vacuum correlators: $M = \text{AdS}$

Temperature: $M = \text{BHAdS}$

Thermalization with Vadya AdS



Equal-time correlators

$t_{\text{dethermalization}}$ / $t_{\text{thermalization}}$

$$\tau_{\text{therm}} = \int_J^\infty \frac{dr}{r^2 \left(1 - \frac{M}{r^d}\right)}$$

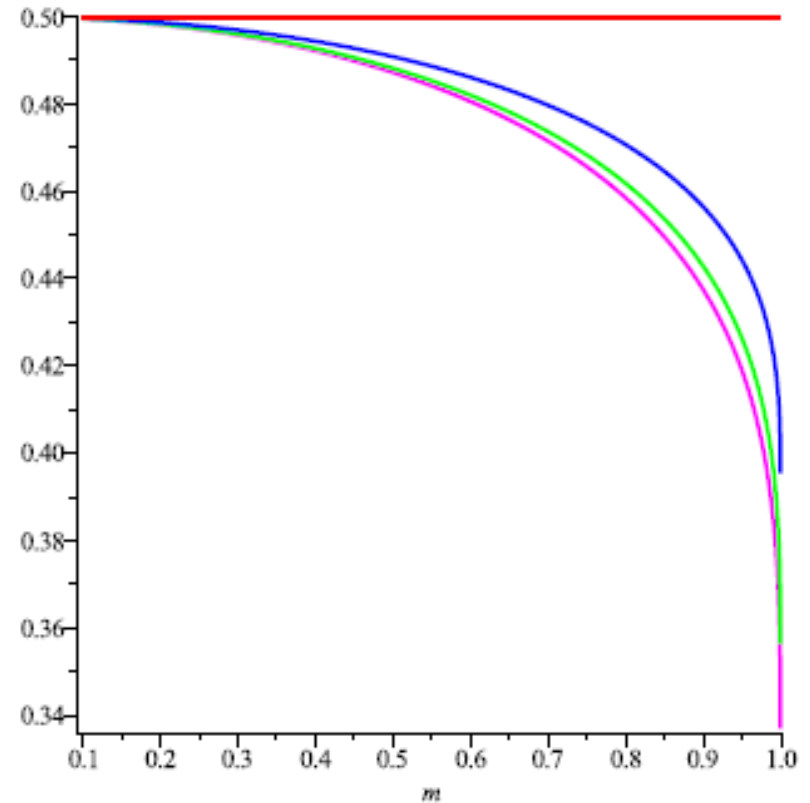
$$\ell = 2J \int_J^\infty \frac{dr}{r^2 \sqrt{(r^2 - J^2) \left(1 - \frac{M}{r^d}\right)}}.$$

$$\frac{\tau_{\text{ther}}}{\tau_{\text{dether}}} = F(m^2, d) \quad F(m^2, d) = \frac{\int_1^\infty \frac{d\rho}{\rho^2 \left(1 - \frac{m^2}{\rho^d}\right)}}{2 \int_1^\infty \frac{d\rho}{\rho^2 \sqrt{(\rho^2 - 1) \left(1 - \frac{m^2}{\rho^d}\right)}}}.$$

$\tau_{dethermalization}$ / $\tau_{thermalization}$

$$\frac{\tau_{ther}}{\tau_{dether}} = F(m^2, d)$$

$$0.78 < \frac{\tau_{ther}}{\tau_{det}} < 1$$



Data: $\tau_{ther} / \tau_{det} \sim 0.1 - 0.05$

thermalization

Data: $\tau_{ther} \sim 1 \text{ fm}/c$

Balasubramanian +9,PRL, 2011,Phys.Rev.2011

thermal scale $l \sim \hbar/T$ $T \sim 300 - 400 \text{ MeV}$

$\tau_{therm} \sim 0.3 \text{ fm}/c,$

$l \sim 2 \text{ fm}$

I.A., I.Volovich,1211.6041

$r_{Pb} \approx 7 \text{ fm}$ can pack 208 (A=208 for Pb) balls with radius

$$r_n = \sqrt[3]{\frac{\eta_K}{208}} r_{Pb} \approx 1.07 \text{ fm} \quad l \sim r_n$$

η_K is the Kepler number $\eta_K = \pi/\sqrt{18} \approx 0.74$

$t_{\text{dethermalization}} / t_{\text{thermalization}}$

Data: $\tau_{ther} / \tau_{det} \sim 0.1 - 0.05$

$$l_{det} \sim 2r_{Pb} \sim 14 \text{ fm.}$$

$$l_{therm} \sim 2 \text{ fm}$$

$$\tau_{det} \sim 7 \text{ fm}/c$$

$$\frac{\tau_{ther}}{\tau_{det}} = \frac{\tau_{ther}}{0.5 \cdot l_{ther}} \cdot \frac{l_{ther}}{l_{det}} = 0.39 \cdot \frac{2}{14} \approx 0.056$$

Thermalization Time and Centricity

In progress with A.Koshelev, A.Bagrov

Non-centricity

Kerr-ADS-BH

$$ds^2 = -(N^\perp(r))^2 dt^2 + \frac{1}{(N^\perp(r))^2} dr^2 + r^2 (N^\phi(r) dt + d\phi)^2$$

$$N^\perp = \left(-M + \left(\frac{r}{l} \right)^2 + \frac{a^2}{r^2} \right)^{1/2}, \quad N^\phi(r) = -\frac{a}{r^2}$$

$$ds^2 = - \left(-M + \frac{r^2}{l^2} \right) dv^2 + 2dvdr - 2advd\hat{\phi} + r^2 d\hat{\phi}^2$$

$$dv = dt + \frac{dr}{(N^\perp)^2}$$

$$d\hat{\phi} = d\phi - \frac{N^\phi}{(N^\perp)^2} dr$$

$$M(v) = M\theta(v),$$

$$j(v) = 2a\theta(v)$$

Kerr-ADS-BH Geometry

$$\beta_{1,2} = \frac{l^2 M}{2} \left(1 \pm \sqrt{1 - \frac{4a^2}{l^2 M^2}} \right)$$

Geodesics

$$t(r) = t_0 - \frac{\mathcal{E}l^3}{2} I_{\pm} \Big|_{\alpha = -\frac{\mathcal{J}a}{\mathcal{E}}}, \quad \phi(r) = \phi_0 + \frac{\mathcal{J}l}{2} I_{\pm} \Big|_{\alpha = \frac{\mathcal{E}a + \mathcal{J}M}{\mathcal{J}/l^2}}$$

$$-I_+ = \frac{1}{(\beta_1 - \beta_2)} \left[\frac{\alpha - \beta_1}{\sqrt{B_1}} \ln(X_1 - \text{sign}(x - \beta_1) \sqrt{X_1^2 - (\gamma_1 - \gamma_2)^2}) - \frac{\alpha - \beta_2}{\sqrt{B_2}} \ln(X_2 - \text{sign}(x - \beta_2) \sqrt{X_2^2 - (\gamma_1 - \gamma_2)^2}) \right] + C = I_- + C$$

$$X_i = (2\beta_i - \gamma_1 - \gamma_2) + \frac{2B_2}{x - \beta_i}$$

$$x \equiv r^2$$

$$\gamma_1 + \gamma_2 = Ml^2 - l^2 \mathcal{E}^2 - J^2$$

$$\gamma_1 \gamma_2 = l^2 a^2 - l^2 J(MJ + 2a\mathcal{E})$$

$$C = -\frac{2}{(\beta_1 - \beta_2)} \left[\frac{\alpha - \beta_1}{\sqrt{B_1}} - \frac{\alpha - \beta_2}{\sqrt{B_2}} \right] \ln(\gamma_1 - \gamma_2)$$

Geodesics which start and finish at

$$r = \infty$$

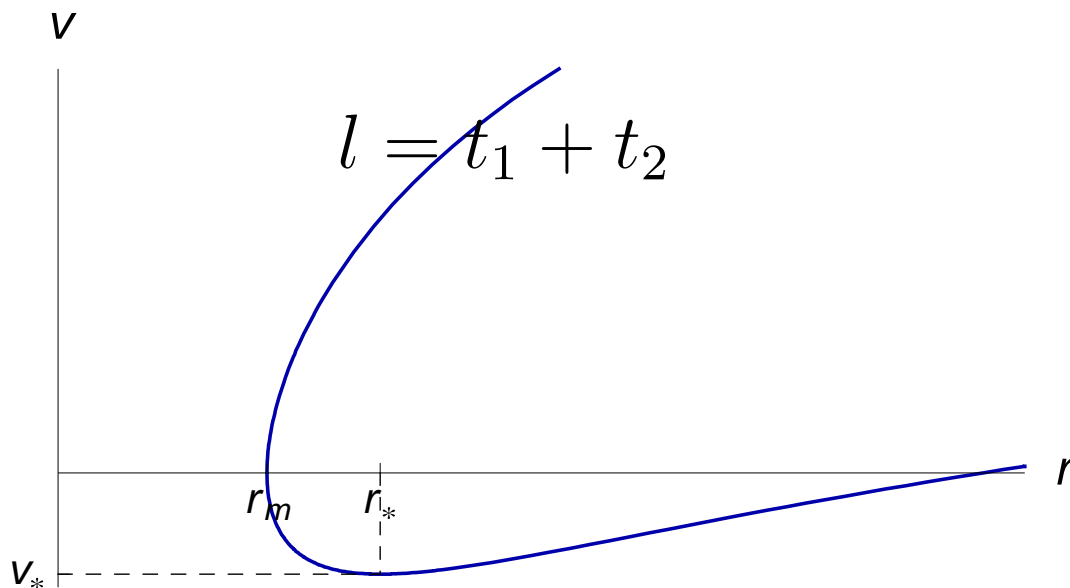
$$t_1 \equiv t_-(\infty) \quad t_2 \equiv t_+(\infty)$$

$$t_1 = t_0 - \frac{c}{2} I_-(\infty) \Big|_{\alpha = -\frac{\mathcal{J}a}{\mathcal{E}}}$$

$$\phi_1 = \phi_-(\infty) = \phi_0 + \frac{\mathcal{J}}{2} I_-(\infty) \Big|_{\alpha = \frac{\mathcal{E}a + \mathcal{J}M}{\mathcal{J}}}$$

$$t_2 = t_0 - \frac{\mathcal{E}}{2} I_+(\infty) \Big|_{\alpha = -\frac{\mathcal{J}a}{\mathcal{E}}}$$

$$\phi_2 = \phi_+(\infty) = \phi_0 + \frac{\mathcal{J}}{2} I_+(\infty) \Big|_{\alpha = \frac{\mathcal{E}a + \mathcal{J}M}{\mathcal{J}}}$$



Thermal point

$$\dot{v}_* = 0$$

$$\mathbf{a} = 0$$

$$l = t_1 + t_2 - 2v_*$$

Conclusion

Formation of QGP of 4-dim QCD \Leftrightarrow Black Hole formation in AdS_5

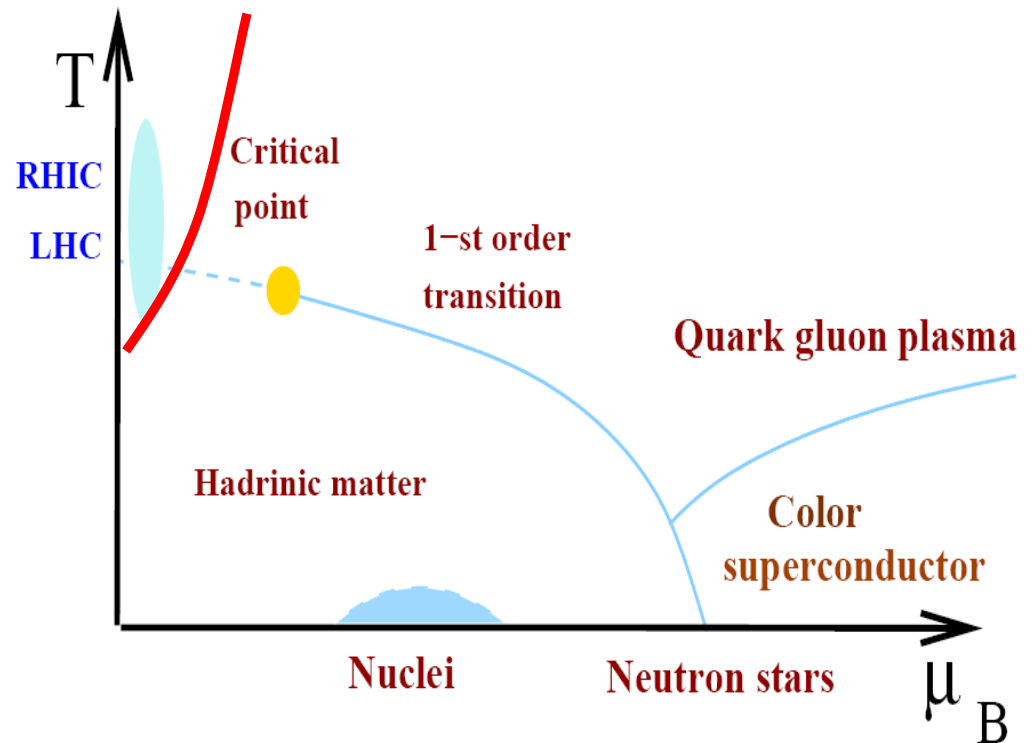
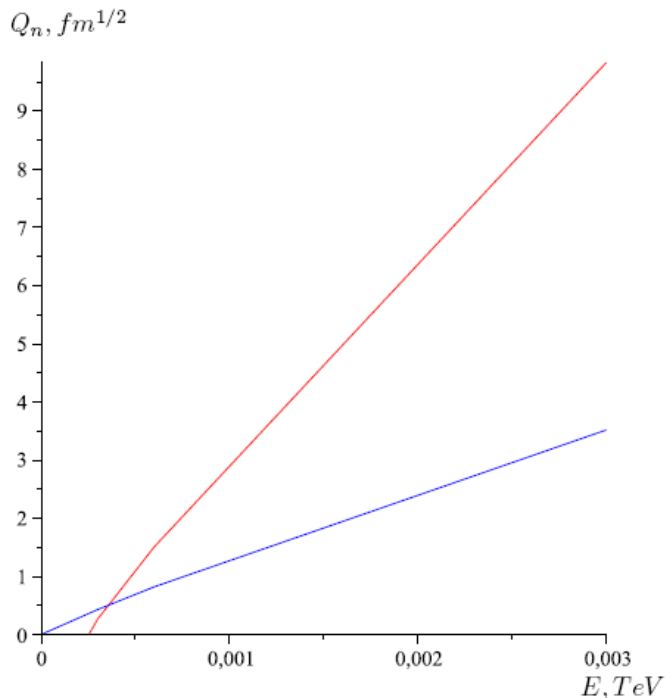
- **Multiplicity: AdS-estimations fit experimental data**

$$S_{data} \propto s_{NN}^{0.15}$$

- $\tau_{ther}/\tau_{det} \sim 0.1 - 0.05$
- **Non-centrality decreases thermalization time.**
- **New phase transition (T vs μ_B)**

BACKUP: Phase diagram from dual approach

Formation of trapped surfaces is only possible when $Q < Q_{cr}$



Red for a smeared matter
Blue for a point-like source

I.A., A.Bagrov, Joukovskaya, 0909.1294
I.A., A.Bagrov, E.Pozdeeva, 1201.6542