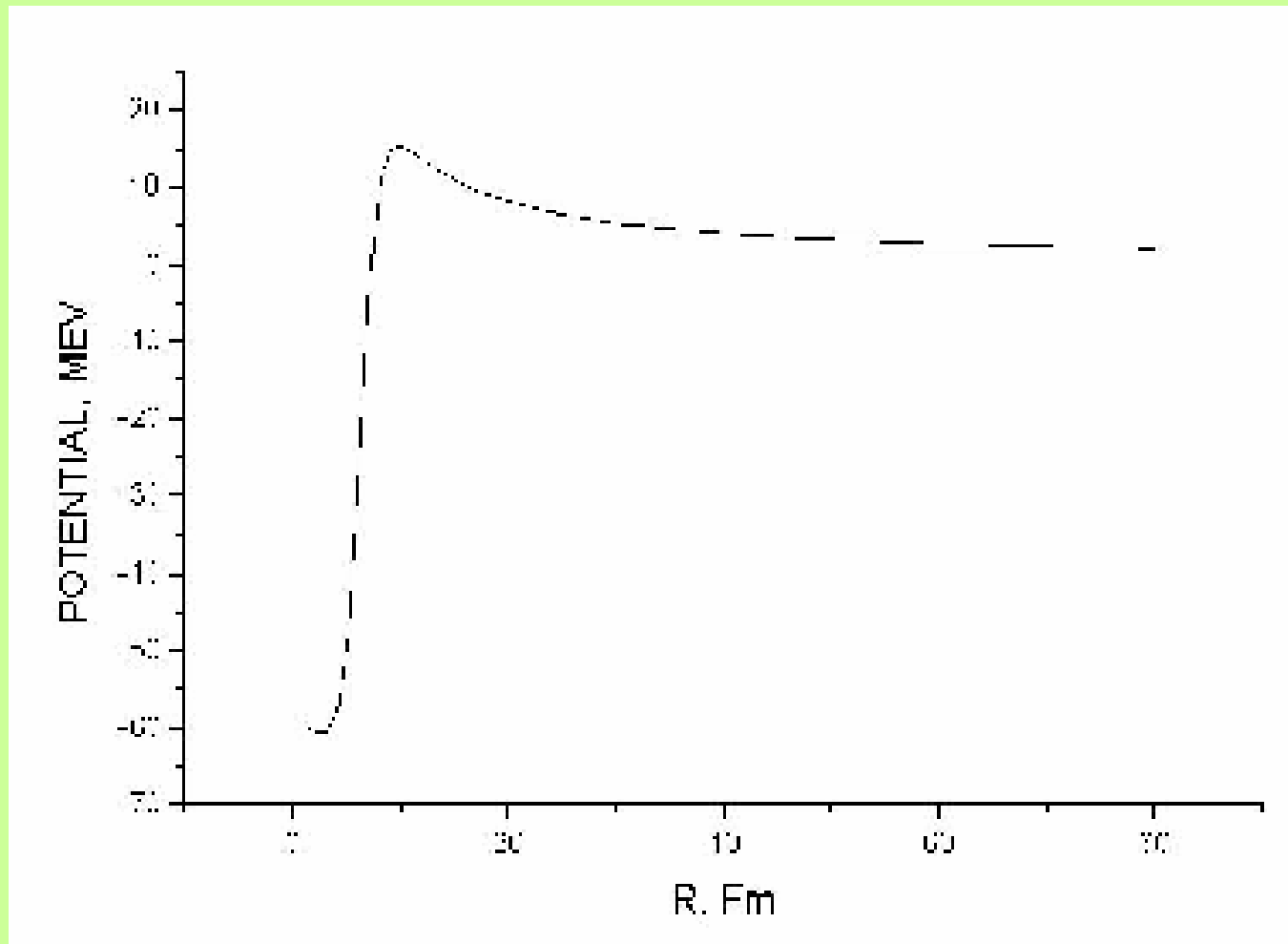


GIANT RESONANCES and ANGULAR DISTRIBUTION IN ALPHA DECAY

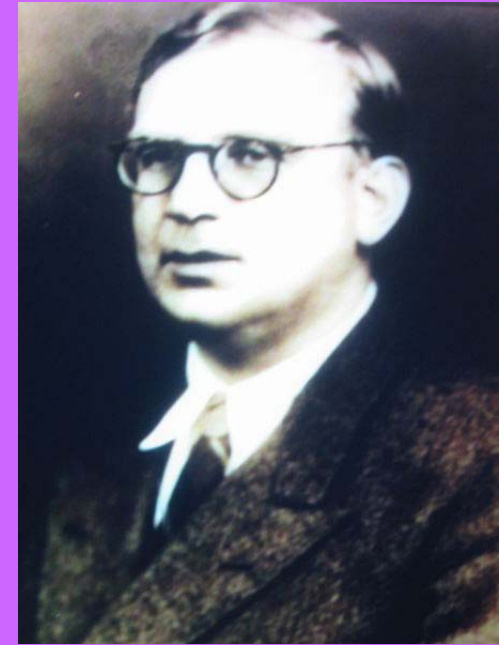
F.F.Karpeshin

Tunneling of α particles through the Coulomb barrier is consecutively treated. The effect of sharp peaks arising in the case of coincidence of the α energy with that of a quasistationary state within the barrier is elucidated. Peaks' energy depend on the α -nucleus potential. The peaks can also be observed in the incoming α -nucleus channel. The method is also applied for calculation of the angular distribution of the emitted α particles from the α decay of a compound nucleus of ^{135}Pr .

α -nucleus potential for the system of $^{131}\text{La} + \alpha$



- The theory was proposed by **Gamov** as an extension of the theory of *quasistationary states* by introducing the complex energy as an eigenvalue. In this theory, the real part of the energy has the physical sense of the α eigenvalue, and the imaginary part gives the probability of tunneling through the barrier.

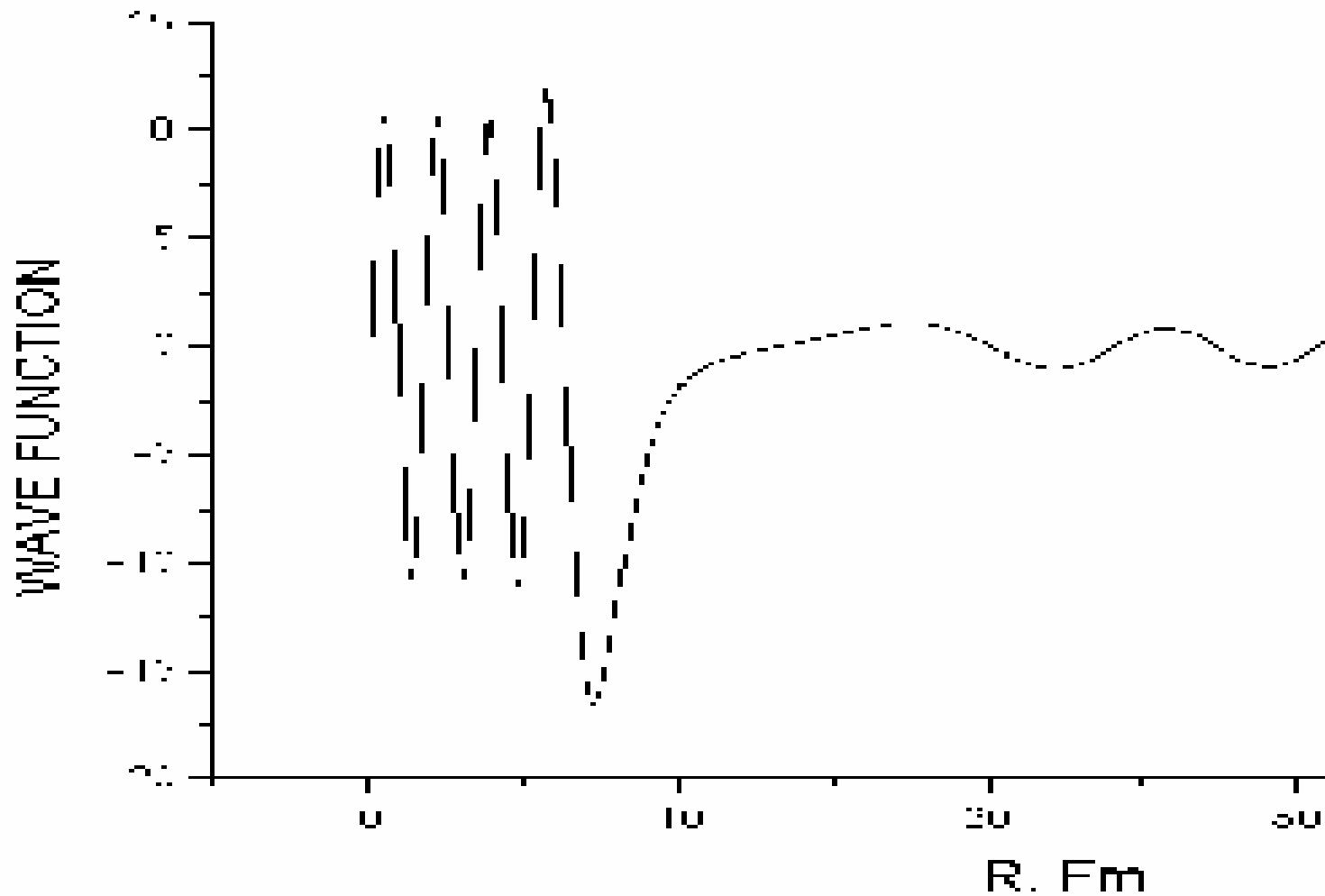


BUT:

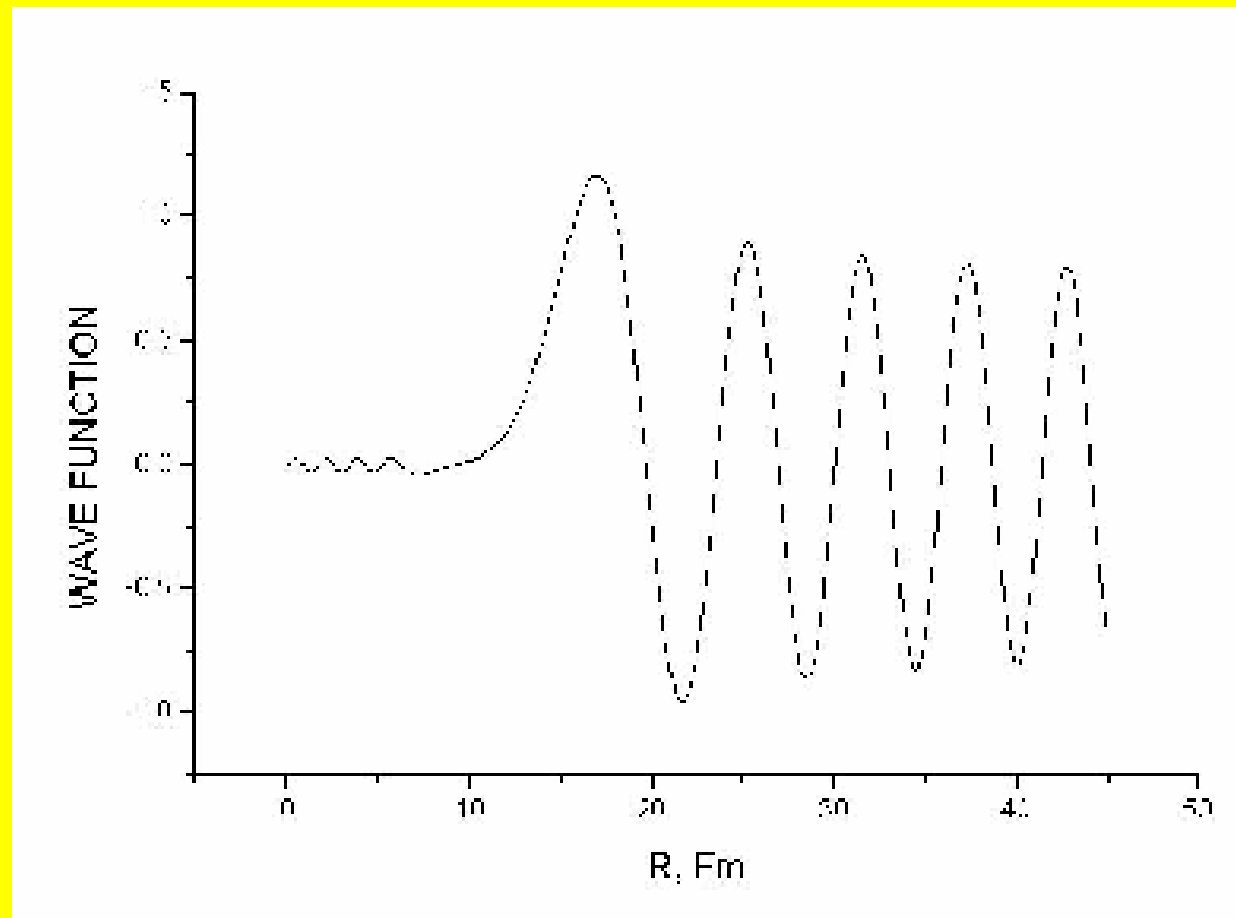
*there are **NO quasistationary α orbits** in the nuclei. They are only formed **virtually**, mainly on the nuclear surface. Then they have a chance to escape through the barrier, otherwise they are immediately destroyed and absorbed by the bulk of the nuclear matter.*

*This means that α 's may be with an **arbitrary energy**.*

α wavefunction for the system of $^{131}\text{La} + \alpha$ with the *resonance* α energy of 10.79 MeV



α wavefunction for the system of $^{131}\text{La} + \alpha$,
with α *nonresonance* energy of 14 MeV



HOW to DEFINE the tunneling probability?

Natural idea which can be expressed from
the **first principles**:

to *project* the shell-model wavefunction
on to the **CHANNEL w.f.**

Transparency acquires a new
sense: the channel w.f. may be
small or great in the nuclear
region.

CONCLUSION

- There is no need to fit the α -nucleus potential by adjusting the α energy to an eigen state inside the barrier.
- GR are predicted in the α -nucleus interaction. Discovering them will help to fix the parameters of the potential.

SUMMARY

- Finding the predicted GR looks a challenging but fairly feasible problem at the present stage of investigations.

nucleons constituting the α particle:

$$\langle \varphi_f(\mathbf{r}_1, \dots, \mathbf{r}_{A-4}) | \Psi(\mathbf{r}_1, \dots, \mathbf{r}_A) \rangle = \eta(\mathbf{r}_{A-3}, \dots, \mathbf{r}_A). \quad (6)$$

The resulting product can be called α cluster in the adiabatic representation. Within the framework of the shell model, its wavefunction vanishes outside the nuclear surface.

Therefore, we have to relate the shell-model wavefunction of the source nucleus to the channel wavefunction (1). Then the re-expansion coefficients give the transition amplitude under consideration. This way is similar to that found by Migdal when solving his classical problem of shake of an atom in β decay.¹³ To find the transition amplitude, one has to project the right part of eq. (1) onto the primary wavefunction Ψ :

$$M_{f\mathbf{p}} = \langle g_{\mathbf{p}}(\mathbf{r}_{A-3} - \mathbf{R}, \dots, \mathbf{r}_A - \mathbf{R}) | \eta(\mathbf{r}_{A-3}, \dots, \mathbf{r}_A) \rangle. \quad (7)$$

According to,⁹ the α decay amplitude is thus determined by the formation-penetration factor (7), which includes both the formation and penetration amplitudes. One can further break this factor into these conventional amplitudes by inserting the set of the α wavefunctions, reduced to the single ground state ξ of an α particle which is in the point R , and then integrate over R . As a result, we get the following expression:

$$\begin{aligned} M_{f\mathbf{p}} &= \langle g_{\mathbf{p}}(\mathbf{r}_{A-3} - \mathbf{R}, \dots, \mathbf{r}_A - \mathbf{R}) | \xi(\mathbf{r}_{A-3} - \mathbf{R}, \dots, \mathbf{r}_A - \mathbf{R}) \rangle \times \\ &\quad \times \langle \xi(\mathbf{r}_{A-3} - \mathbf{R}, \dots, \mathbf{r}_A - \mathbf{R}) | \eta(\mathbf{r}_{A-3}, \dots, \mathbf{r}_A) \rangle \equiv \\ &\quad \equiv \int_0^\infty \mathcal{F}_p(R) f(R) dR. \quad (8) \end{aligned}$$

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In (8), $f(R)$ can be considered as a wavefunction of an α particle inside the nucleus. The preformation probability P is therefore included in this factor:

$$P = \left| \int_0^\infty f(R) dR \right|^2. \quad (9)$$

The conventional spectroscopic factor s for α decay is given by the overlapping integral

$$s = \int_0^\infty \langle \xi(\mathbf{r}_{A-3} - \mathbf{R}, \dots, \mathbf{r}_A - \mathbf{R}) | \eta(\mathbf{r}_{A-3}, \dots, \mathbf{r}_A) \rangle dR = \int_0^\infty f(R) dR. \quad (10)$$

On the other hand, the other overlapping integral in (8) gives the tunneling probability t :

$$t = \left| \int_0^\infty \mathcal{F}_{\mathbf{p}}(R) f(R) dR \right|^2. \quad (11)$$

Furthermore, taking into account the asymptotics (3), the wavefunction $\mathcal{F}_{\mathbf{p}}(\mathbf{R})$ can be expressed in terms of the spherical harmonics in a usual way:

$$\mathcal{F}_{\mathbf{p}}(\mathbf{R}) = \sum_{\ell=0}^{\infty} i^\ell (2\ell + 1) e^{-i\delta_\ell} R_{p\ell}(R) Y_{\ell m}(\theta, \varphi). \quad (12)$$