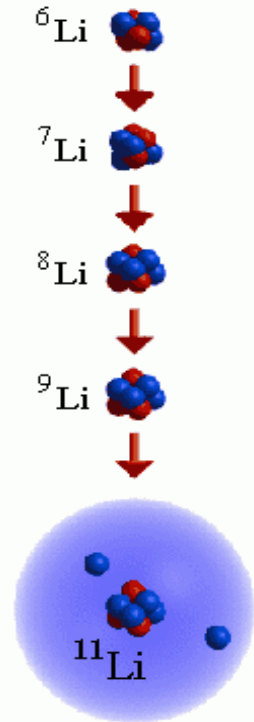


Helmholtz International Summer School
“Nuclear Theory and Astrophysical Applications”
Dubna, Russia, July 10 – 22, 2017

S. N. Ershov

Joint Institute for Nuclear Research



HALO NUCLEI

Halo is a widespread name for specific phenomena in different fields of physics



in optics:

Halo is the name for a family of optical phenomena produced by light interacting with ice crystals suspended in the atmosphere. Halos can have many forms, ranging from colored or white rings to arcs and spots in the sky.

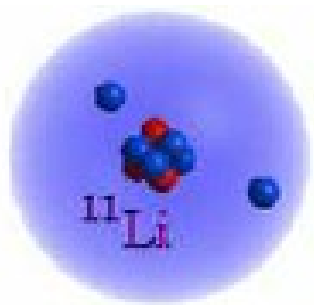
in astrophysics:



Galactic halo is a region of scattered stars that surrounds spiral galaxies.

A dark matter halo is a hypothetical component of a galaxy that envelops the galactic disc and extends well beyond the edge of the visible galaxy.

in nuclear structure:



The new structural dripline phenomenon with clusterization into an ordinary core nucleus and a veil of halo nucleons forming very dilute neutron matter

STRUCTURE

ATOMIC

electromagnetic forces between
negatively charged **electrons**
and positively charged **nucleus**

NUCLEAR

Interactions between
neutrons and **protons** via:

- **strong** forces
- **electromagnetic** forces
- **weak** forces

SPACE SCALES ($1 \text{ fm} = 10^{-13} \text{ cm}$) :

$\sim (0.3 - 2.3) \cdot 10^5 \text{ fm}$

$\sim (2 - 6) \text{ fm}$

ENERGY SCALES :

\sim electronvolts

(**Ionization energy** for $^1\text{H} = 13.6 \text{ eV}$)

$\sim \text{MeV} = 10^6 \text{ electronvolts}$

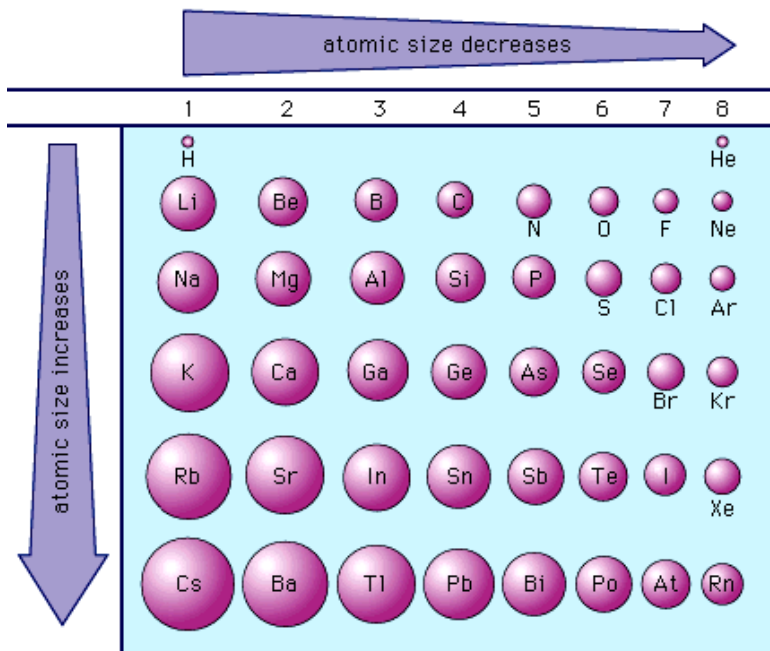
(**Binding energy** for deuteron = **2.2 MeV**)

Noble gases correspond to the magic numbers : 2, 10, 18, 36, 54, 86

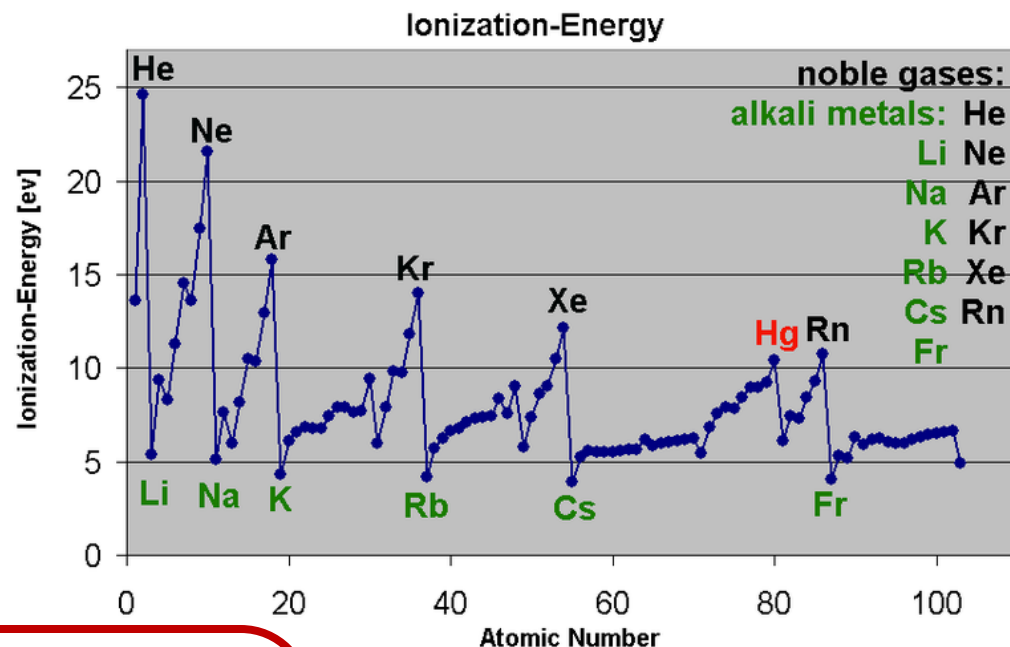
$$\begin{array}{cccccccccccccccccccc} \color{red}{2} & \color{blue}{8} & \color{red}{8} & & \color{blue}{18} & & \color{red}{18} & & \color{blue}{32} & & \color{red}{32} \\ \color{red}{1s^2} & \color{blue}{2s^2} & \color{blue}{2p^6} & \color{red}{3s^2} & \color{red}{3p^6} & \color{blue}{4s^2} & \color{blue}{3d^{10}} & \color{blue}{4p^6} & \color{red}{5s^2} & \color{red}{4d^{10}} & \color{red}{5p^6} & \color{blue}{6s^2} & \color{blue}{4f^{14}} & \color{blue}{5d^{10}} & \color{blue}{6p^6} & \color{red}{7s^2} & \color{red}{5f^{14}} & \color{red}{6d^{10}} & \color{red}{7p^6} \end{array}$$

<div><div>H1 1s</div><div>He2 1s</div></div>																	
<div>Electron Configuration Table</div>																	
<div><div><div>Li1 2s</div><div>Na1 3s</div><div>K1 4s</div><div>Rb1 5s</div><div>Cs1 6s</div><div>Fr1 7s</div></div><div><div>Be2</div><div>Mg2</div><div>Ca2</div><div>Sr2</div><div>Ba2</div><div>Ra2</div></div><div><div>Sc1</div><div>Ti2</div><div>V3</div><div>Cr4</div><div>Mn5</div><div>Fe6</div><div>Co7</div><div>Ni8</div><div>Cu9</div><div>Zn10</div></div><div><div>Y1</div><div>Zr2</div><div>Nb3</div><div>Mo4</div><div>Tc5</div><div>Ru6</div><div>Rh7</div><div>Pd8</div><div>Ag9</div><div>Cd10</div></div><div><div>La*1</div><div>Hf2</div><div>Ta3</div><div>W4</div><div>Re5</div><div>Os6</div><div>Ir7</div><div>Pt8</div><div>Au9</div><div>Hg10</div></div><div><div>+Ac1</div><div>Rf2</div><div>Db3</div><div>Sg4</div><div>Bh5</div><div>Hs6</div><div>Mt7</div><div>Ds8</div><div>Rg9</div><div></div><div></div></div><div><div>←</div><div>←</div><div>←</div><div>←</div><div>←</div><div>←</div></div><div><div>3d</div><div>4d</div><div>5d</div><div>6d</div></div></div> <div><div><div>B1</div><div>C2</div><div>N3</div><div>O4</div><div>F5</div><div>Ne6</div></div><div><div>Al1</div><div>Si2</div><div>P3</div><div>S4</div><div>Cl5</div><div>Ar6</div></div><div><div>Ga1</div><div>Ge2</div><div>As3</div><div>Se4</div><div>Br5</div><div>Kr6</div></div><div><div>In1</div><div>Sn2</div><div>Sb3</div><div>Te4</div><div>I5</div><div>Xe6</div></div><div><div>Tl1</div><div>Pb2</div><div>Bi3</div><div>Po4</div><div>At5</div><div>Rn6</div></div><div><div>←</div><div>←</div><div>←</div><div>←</div><div>←</div><div>←</div></div><div><div>2p</div><div>3p</div><div>4p</div><div>5p</div><div>6p</div><div>7p</div></div></div>																	
<div><div><div>Ce1</div><div>Pr2</div><div>Nd3</div><div>Pm4</div><div>Sm5</div><div>Eu6</div><div>Gd7</div><div>Tb8</div><div>Dy9</div><div>Ho10</div><div>Er11</div><div>Tm12</div><div>Yb13</div><div>Lu14</div></div><div><div>←</div><div>←</div><div>←</div><div>←</div><div>←</div><div>←</div><div>←</div><div>←</div><div>←</div><div>←</div><div>←</div><div>←</div><div>←</div><div>←</div></div><div><div>4f</div><div>5f</div></div></div> <div><div><div>Th1</div><div>Pa2</div><div>U3</div><div>Np4</div><div>Pu5</div><div>Am6</div><div>Cm7</div><div>Bk8</div><div>Cf9</div><div>Es10</div><div>Fm11</div><div>Md12</div><div>No13</div><div>Lr14</div></div><div><div>←</div><div>←</div><div>←</div><div>←</div><div>←</div><div>←</div><div>←</div><div>←</div><div>←</div><div>←</div><div>←</div><div>←</div><div>←</div><div>←</div></div><div><div>5f</div></div></div>																	

ATOMIC SIZES



IONIZATION ENERGIES



Atoms lack a **well-defined** outer boundary.

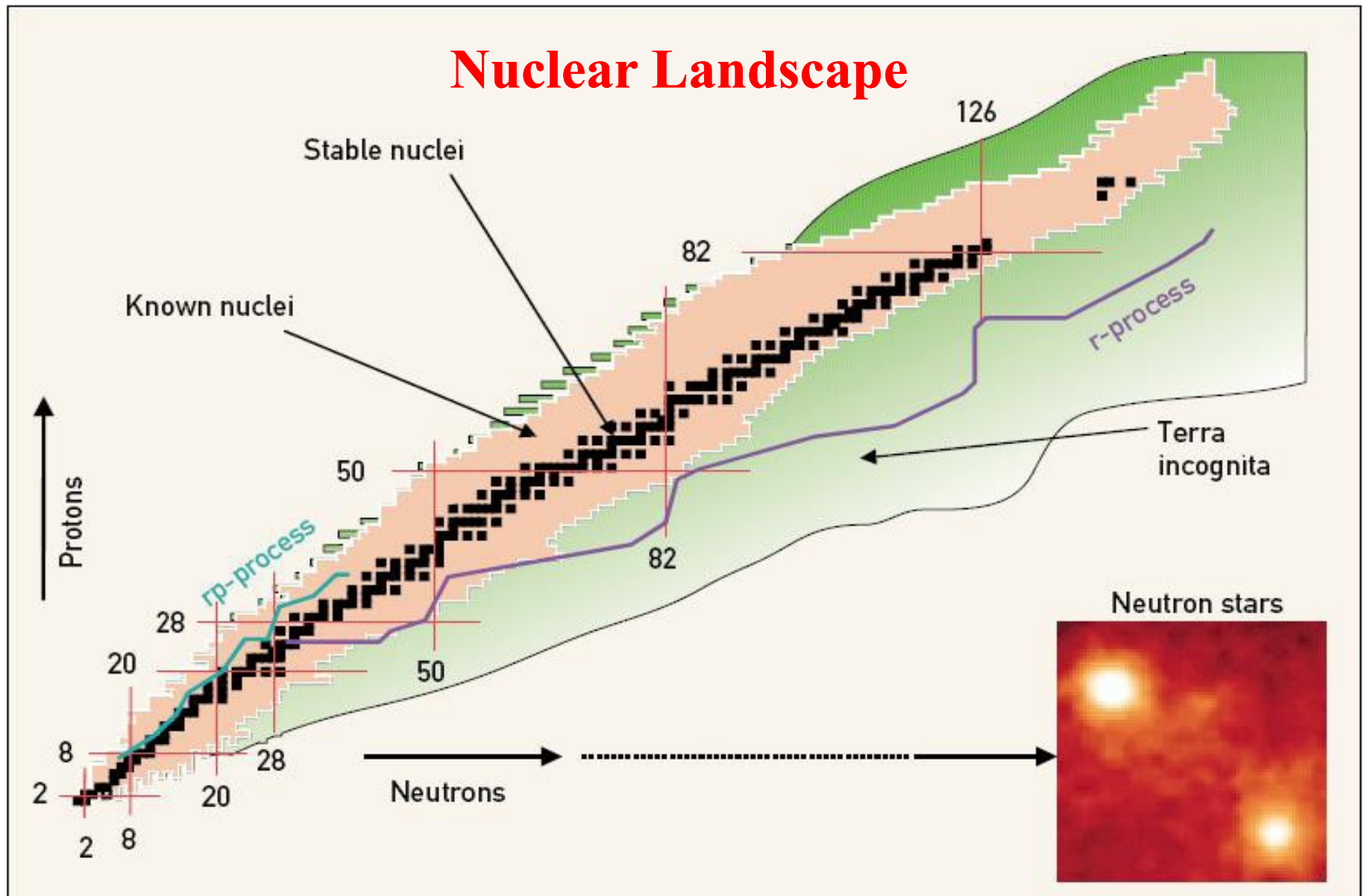
The radius varies with:

- the **location** of an atom on the atomic chart
- the **type of chemical bond**
- the number of **neighboring** atoms

Despite conceptual difficulties **atomic radii** vary in a **predictable** and **explicable** manner across the periodic table

$$H_0 = \sum_{i=1}^Z \left(T_i - \frac{Ze^2}{r_i} \right) + \sum_{i<j}^Z \frac{e^2}{r_{ij}}$$

Nuclear Landscape



Presently ~ 3600 nuclei have been observed,
less than 300 nuclei are **stable** (with a lifetime greater than 10^9 years)

NUCLEAR STRUCTURE NEAR THE VALLEY OF STABILITY

exhibit similar binding for neutrons and protons
density and diffuseness of the surface are nearly constant

the resulting **shell structure** is well established



M. Goeppert-Mayer, J.H.D. Jensen, nobel prize in 1963
"for their discoveries concerning nuclear shell structure"

Nobel Prize Medal

magic numbers

(2, 8, 20, 28, 50, 82, 126) are the same for neutrons and protons

stable **double-magic** nuclei: ${}^4_2\text{He}_2$, ${}^{16}_8\text{O}_8$, ${}^{40}_{20}\text{Ca}_{20}$, ${}^{48}_{20}\text{Ca}_{28}$, ${}^{208}_{82}\text{Pb}_{126}$

nuclear potential is well parametrized

pronounced shell closures define the effective degrees of freedom
needed for a quantitative understanding of atomic nuclei

A.N. Bohr, B.R. Mottelson, L.J. Raynwater , nobel prize in 1975

"for the discovery of the connection between collective motion and
particle motion in atomic nuclei and the development of the theory
of the structure of the atomic nucleus based on this connection"

NUCLEUS: proton/neutron shells

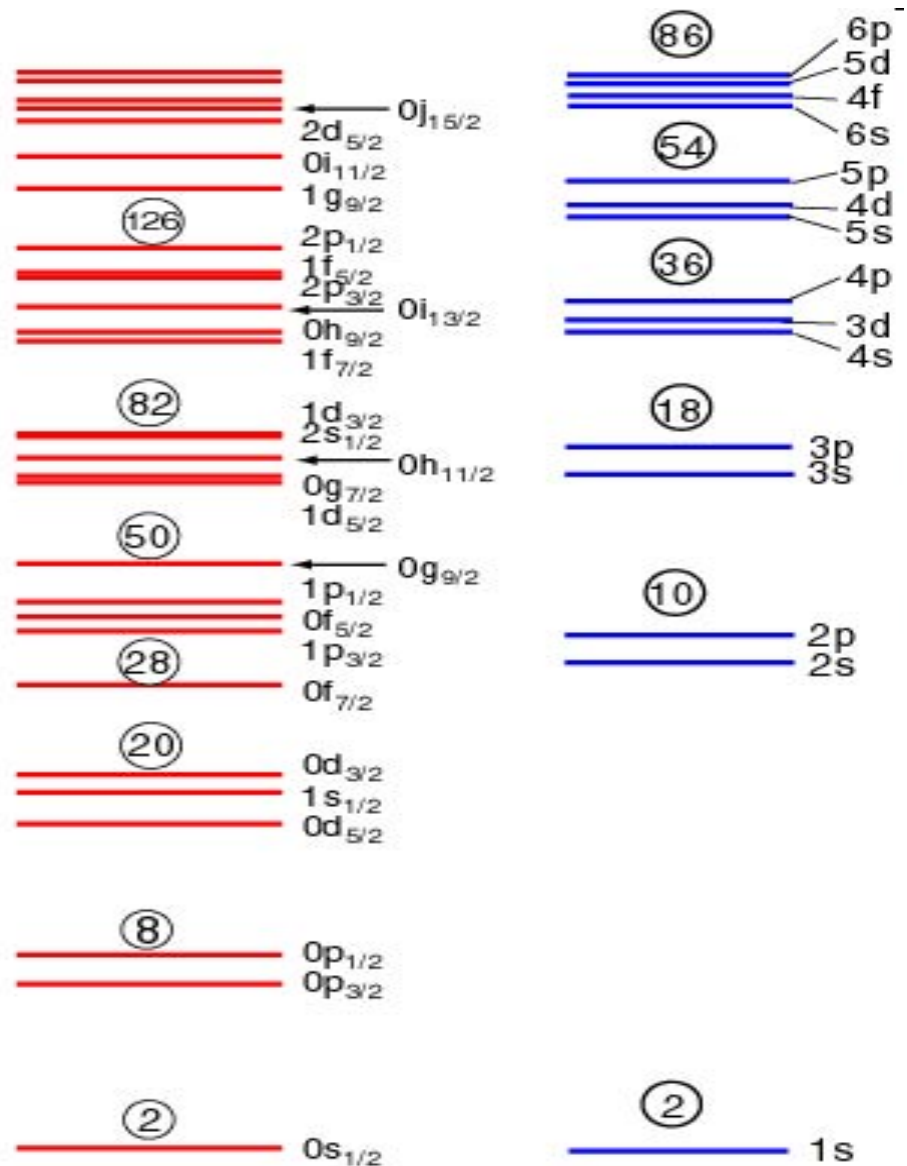
ATOM: electron shells

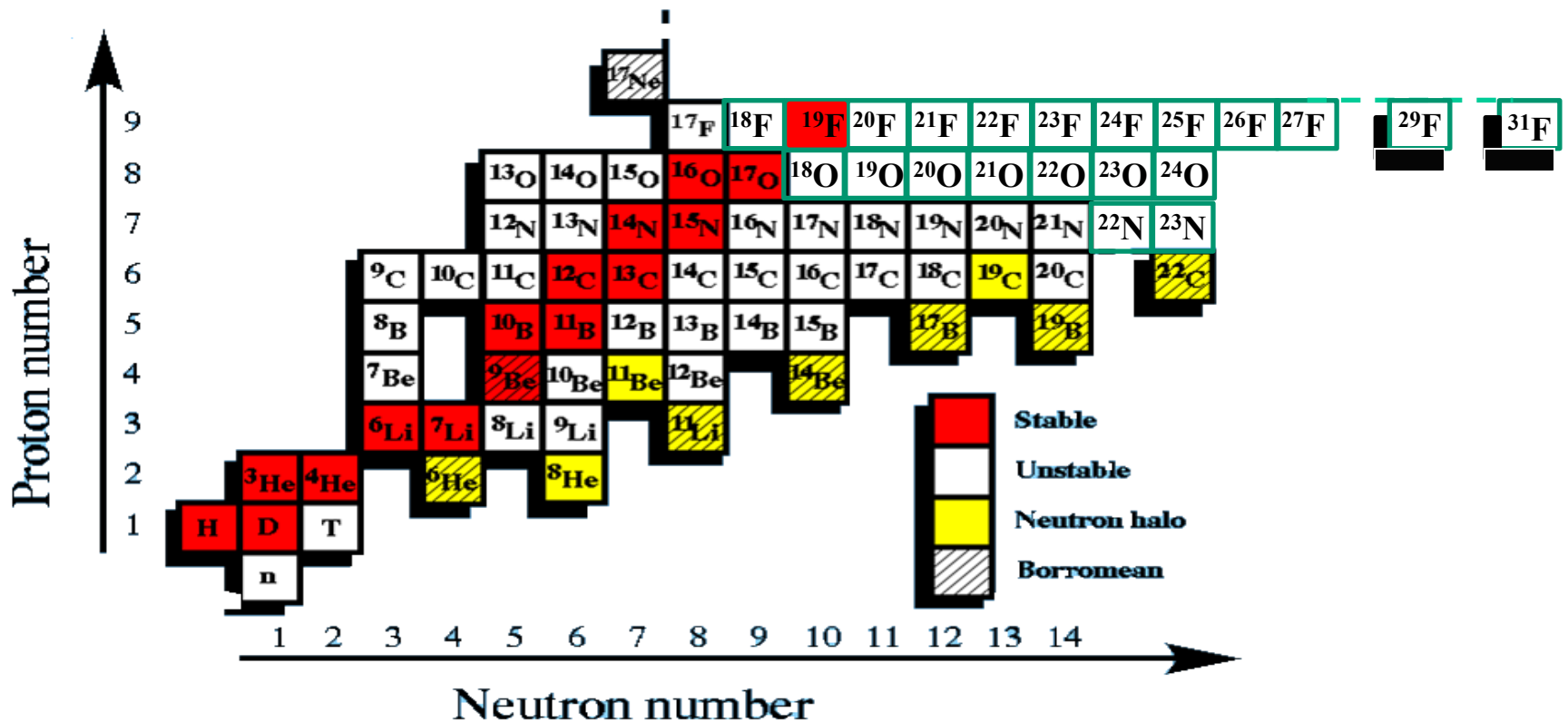
Nuclei with **magic numbers** of **neutrons** and **protons** have

- large binding energies
- high natural abundances
- high energy of first excited state
- ...

Shell-model potential

$$V(r) = V_c(r) + V_{ls}(\mathbf{L} \cdot \mathbf{S})$$





Chains of the lightest isotopes (He , Li , Be , B , ...)
end up with **two neutron halo nuclei**

Two neutron halo nuclei (${}^6\text{He}$, ${}^{11}\text{Li}$, ${}^{14}\text{Be}$, ...) are **Borromean**
systems and break into **three fragments**

One neutron halo nuclei (${}^{11}\text{Be}$, ${}^{19}\text{C}$, ...) break into **two fragments**

Borromean system: **three fragments** are bound together
but **two-body subsystems** are all unbound

DEFINITION OF A NUCLEUS

many combinations of neutrons and protons
can make up a nucleus of a given mass

What system of nucleons
can we call a nucleus ?



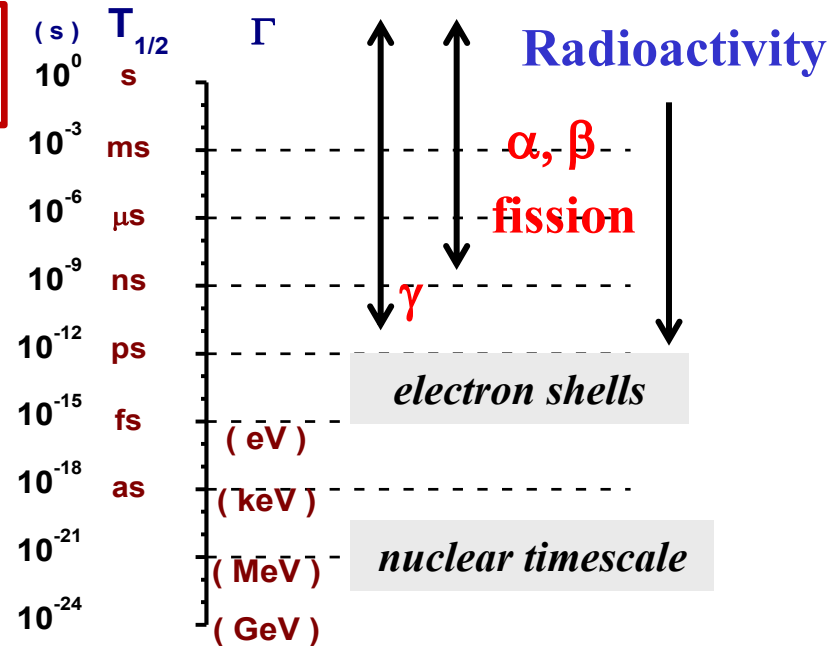
A possible limit could be set by :

- the definition of radioactivity

lifetimes longer than 10^{-12} s 

- definition of the International Union of
Pure and Applied Chemistry (IUPAC)
(Pure Appl. Chem. 63 (1991) 879)

existing for at least 10^{-14} s 



a lower limit for the process
to be called radioactivity

time for a nucleus to
acquire its outer electrons

DEFINITION OF THE DRIPLINES

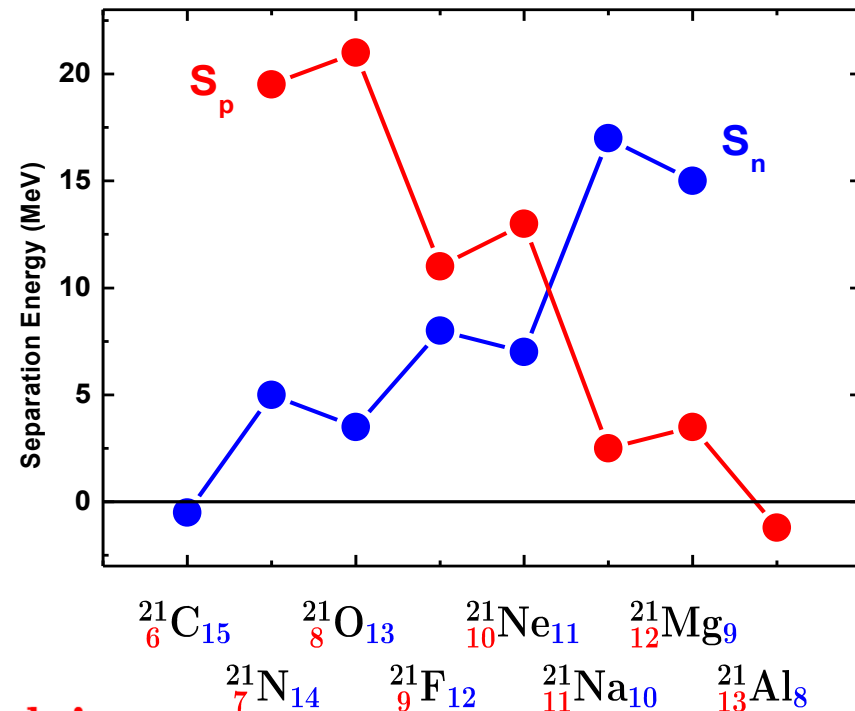
The **nuclear binding** in nuclei with an extreme excess of neutrons or protons may not be sufficient to bind the last neutron or proton

$$BE(N, Z) = ZM_Hc^2 + NM_nc^2 - M(N, Z)c^2$$

$$S_n = BE(N, Z) - BE(N - 1, Z)$$

$$S_p = BE(N, Z) - BE(N, Z - 1)$$

the **dripline** is the boundary where **neutron** S_n or **proton** S_p separation energies cross zero



The unbound nuclei :

- can have fairly **long lifetimes** along the proton dripline (because of the Coulomb barrier)
- extremely **short lived** beyond the neutron dripline (absence of the Coulomb barrier)

the **proton dripline** has been reached for elements up to $Z = 83$
the **neutron dripline** has been reached for elements up to $Z < 10$

Conceptual framework of nuclear structure is the **nuclear shell model**

Qualitative picture:

nucleons are moving almost independently in a **mean-field (self-consistent) potential** obtained by averaging out the interactions between a single nucleon and all remaining protons and neutrons.

Quantitative results:

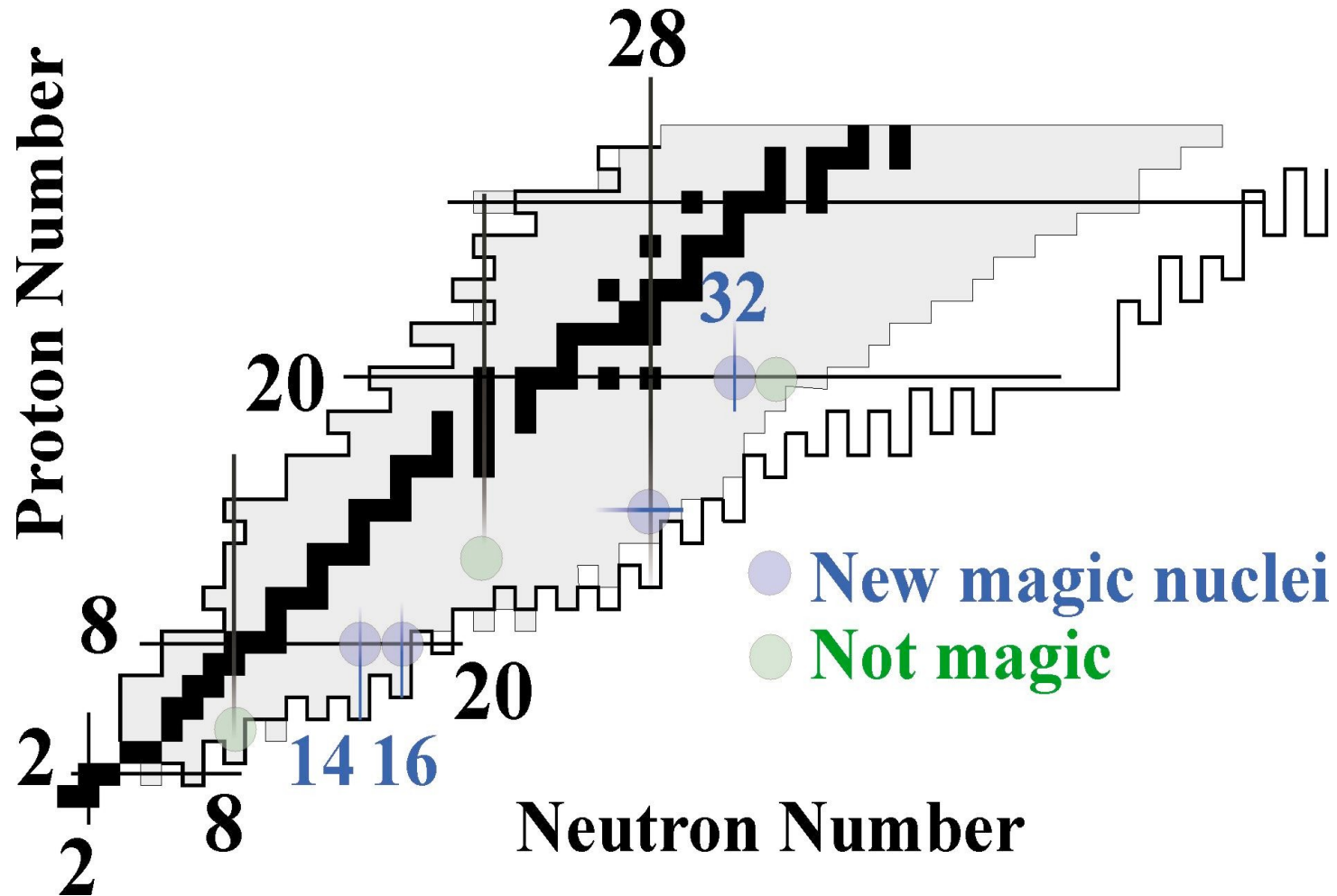
accounting for the residual interaction between the nucleons

Nuclear structure of **exotic nuclei is different** from that around the stability line and represent a formidable challenge for the nuclear many-body theories and their **power to predict** nuclear properties

Unique factors for exotic nuclei

- **the weak binding**, closeness of the particle continuum (a large diffuseness of the nuclear surface, extreme spatial dimensions for the outermost nucleons)
- exotic combinations of proton and neutron numbers (prospects for completely **new structural phenomena**)

EVOLUTION OF THE SHELL STRUCTURE TOWARDS TO THE DRIPLINE



No shell closure for $N = 8$ and 20 for drip-line nuclei
New shell closure at $N = 14, 16, 32, \dots$

Neutron halo nuclei

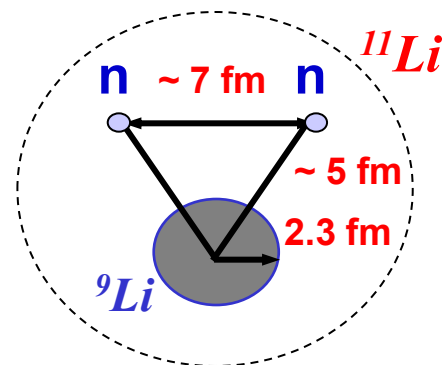
HALO NUCLEI

(^6He , ^{11}Li , ^{11}Be , ^{14}Be , ^{17}B , ...) 

weakly bound systems
with large extension
few-body clusterization

”Residence in **forbidden** regions”

Appreciable probability for dilute nuclear matter
extending far out into **classically forbidden** region



Separation energies of last neutron(s)

<u>halo</u>	<u>stable</u>
< 1	6 - 8 MeV

$$\epsilon(^{11}\text{Li}) = 0.4 \text{ MeV}$$

$$\epsilon(^{11}\text{Be}) = 0.5 \text{ MeV}$$

$$\epsilon(^6\text{He}) = 0.97 \text{ MeV}$$

Large size of halo nuclei

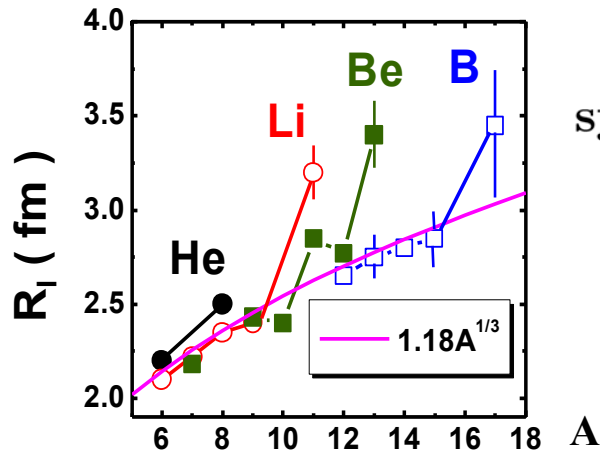
$$\begin{cases} \langle r^2(^{11}\text{Li}) \rangle^{1/2} \sim 3.5 \text{ fm} \\ (r.m.s. \text{ for } A \sim 48) \end{cases}$$

Two-neutron halo nuclei
(^{11}Li , ^6He , ^{14}Be , ^{17}B , ...) 

Borromean systems

Peculiarities of halo nuclei: the example of ^{11}Li

- (i) weakly bound: the two-neutron separation energy (~ 400 KeV) ~ 10 times less than the energy of the first excited state in ^9Li
- (ii) large size: ^{11}Li interaction cross section $\sim 30\%$ larger than for ^9Li



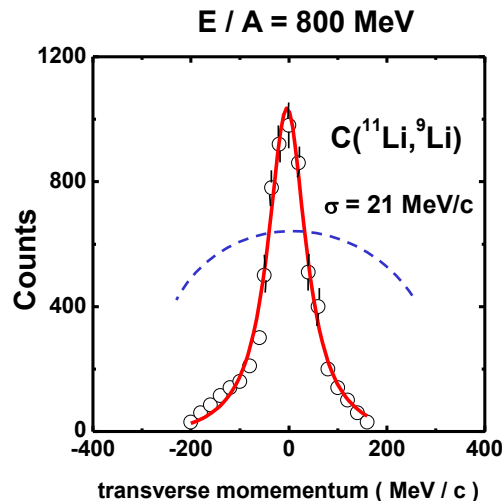
This is very unusual for **strongly interacting** systems held together by **short-range interactions**

Interaction radii: $\sigma_I = \pi(R_{proj} + R_{tar})^2$

light targets: $E/A = 790$ MeV

I. Tanihata et al., PRL 55 (1985) 2676

- (iii) very narrow (in comparison to stable nuclei) **fragment momentum distributions of both neutrons and ^9Li measured in high energy breakup reactions of ^{11}Li**



No narrow fragment distributions in breakup on **other fragments**, say ^8Li or ^8He

(naive picture)

narrow momentum distributions



large spatial extensions

(IV) Relations between interaction and neutron removal cross sections

$A + {}^{12}\text{C}$	σ_I (mb)	σ_{-2n} (mb)	σ_{-4n} (mb)
${}^9\text{Li}$	796 ± 6		
${}^{11}\text{Li}$	1060 ± 10	220 ± 40	
${}^4\text{He}$	503 ± 5		
${}^6\text{He}$	722 ± 5	189 ± 4	
${}^8\text{He}$	817 ± 6	202 ± 17	95 ± 5

$$E/A = 790 \text{ MeV}$$

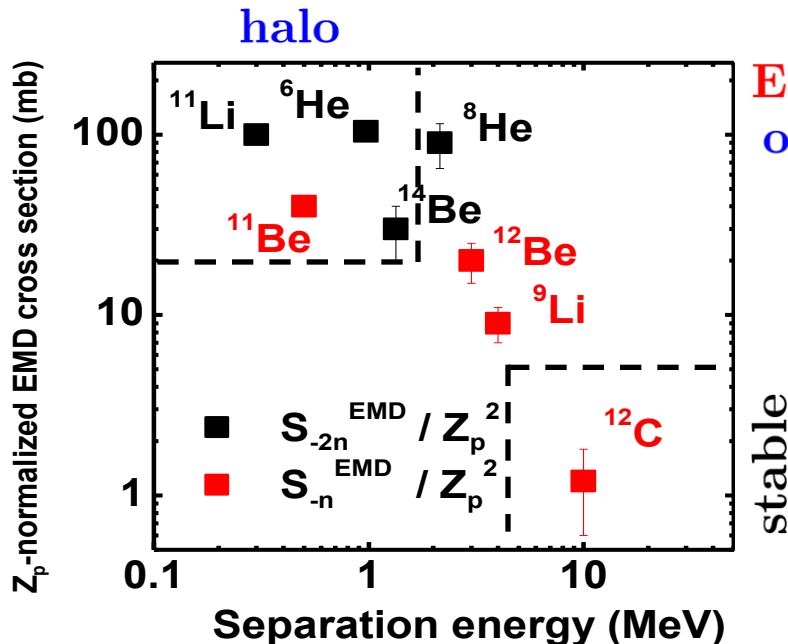
$$\sigma_I(A = C + xn) = \sigma_I(C) + \sigma_{-xn}$$

Strong evidence for the well defined **clusterization** into the **core** and **two neutrons**

Tanihata I. et al.

PRL, 55 (1987) 2670; PL, B289 (1992) 263

(V) Electromagnetic dissociation cross sections (EMDC)



EMDC per unit charge for **halos** are **orders of magnitude larger** than for stable nuclei

Evidence for a rather **large difference** between **charge** and **mass** centers in a body fixed frame

concentration of the dipole strength at **low excitation** energies

T. Kobayashi, Proc. 1st Int. Conf. On Radiative Nuclear Beams, 1990.

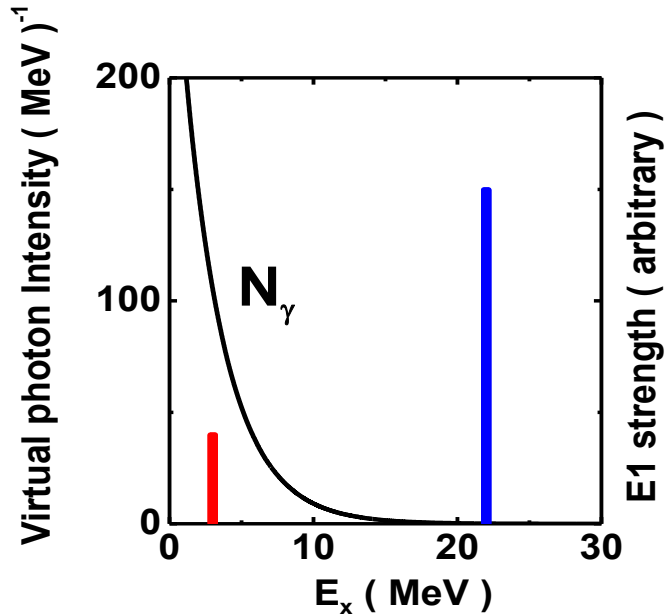
SOFT EXCITATION MODES

(peculiarities of **low energy** halo continuum)

Large EMD cross sections

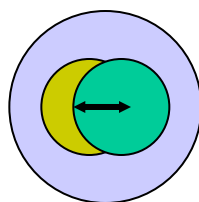
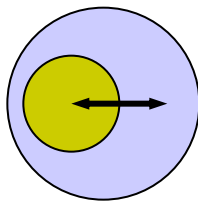


specific nuclear property of
extremely neutron-rich nuclei



soft DR

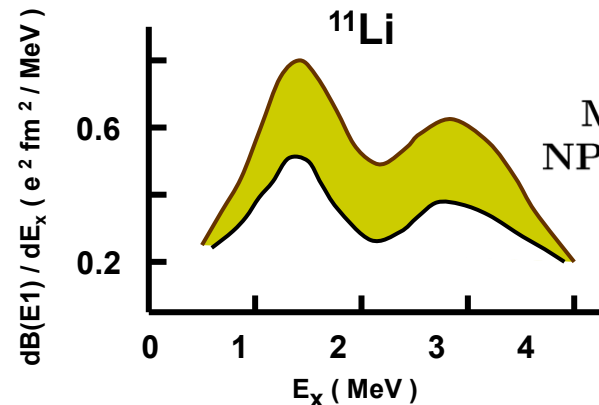
normal GDR



$$E_x \sim 1 \text{ MeV} \sim 20 \text{ MeV}$$

$$\sigma_{EMD} = \int N(E_x) \sigma(E_x) dE_x$$

$$\sigma(E_x) = \frac{16\pi^3}{9\hbar c} E_x \frac{dB(E1)}{dE_x}$$



M. Zinser et al.,
NPA619 (1997) 151

excitations of soft modes with

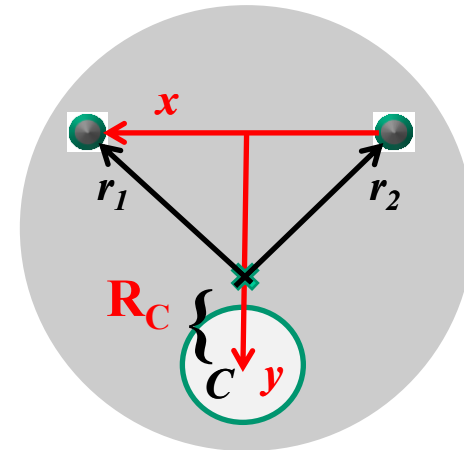
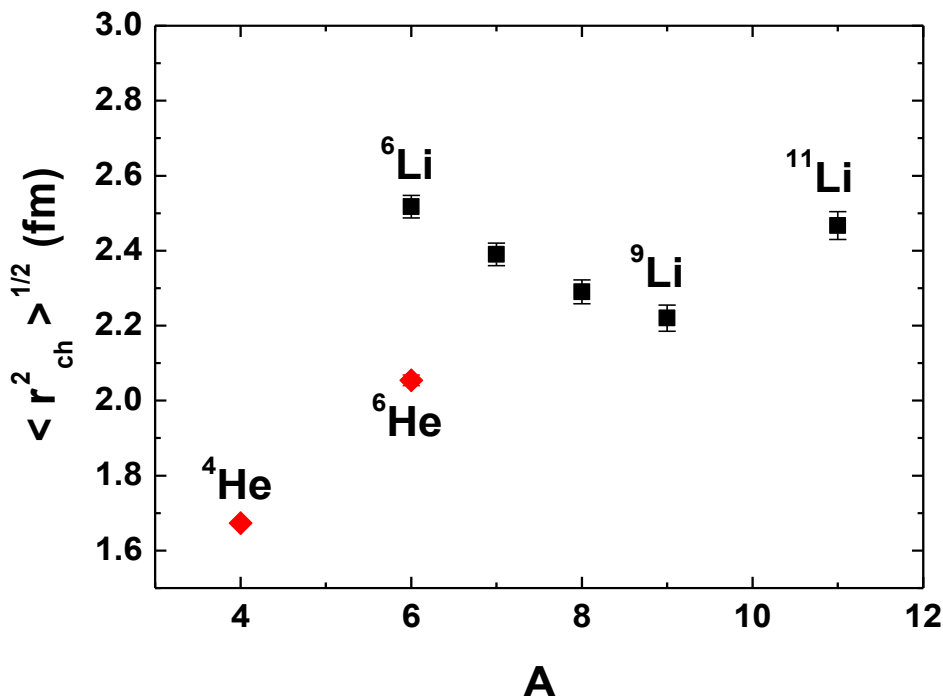
- **different multipolarity**
- **collective** excitations versus **direct** transition from weakly bound to continuum states

(VI) Ground state properties of ^{11}Li and ^9Li

	^9Li	^{11}Li
Spin J^π :	$3/2^-$	$3/2^-$
quadrupole moments :	-27.4 ± 1.0 mb	-31.2 ± 4.5 mb
magnetic moments :	3.4391 ± 0.0006 n.m.	3.6678 ± 0.0025 n.m.

Previous peculiarities **cannot arise** from large **deformations**.
Core is **not significantly perturbed** by the two valence neutrons

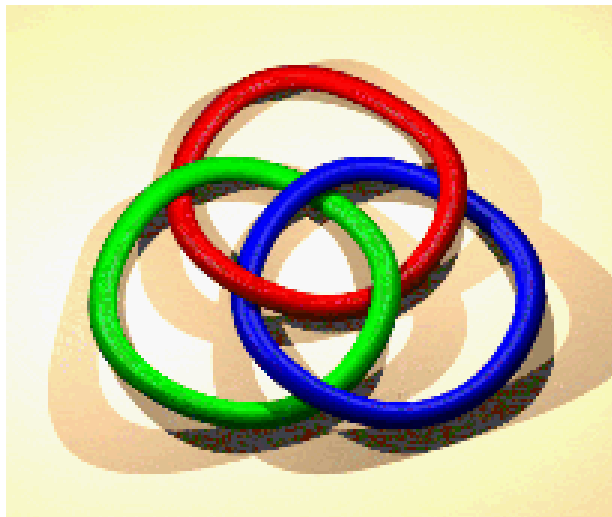
Nuclear charge radii by laser spectroscopy



R. Sanches et al., PRL 96 (2006) 033002
 L.B. Wang et al., PRL 93 (2004) 142501

(VII) The three-body system ^{11}Li ($^9\text{Li} + n + n$) is Borromean (neither the two neutron nor the core-neutron subsystems are bound)

Three-body correlations are the most important,
due to them the system becomes bound.



”The Borromean rings, the heraldic symbol of the Princes of Borromeo, are carved in the stone of their castle in Lake Maggiore in northern Italy. The three rings are interlocked in such a way that any of them were removed, the other two would also fall apart. In nuclear physics ^{11}Li and ^6He have been found to have this property (although for quite different physical reasons) when described in a three-body model. ”

M.V. Zhukov et al., Phys. Rep. 231 (1993) 151

STABLE NUCLEI

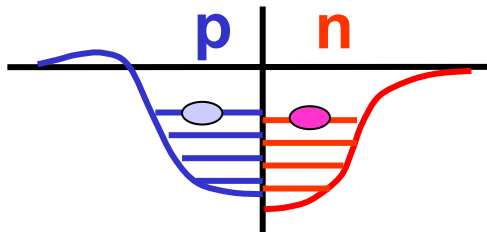
$$N / Z \sim 1 - 1.5$$

$$\epsilon_S \sim 6 - 8 \text{ MeV}$$

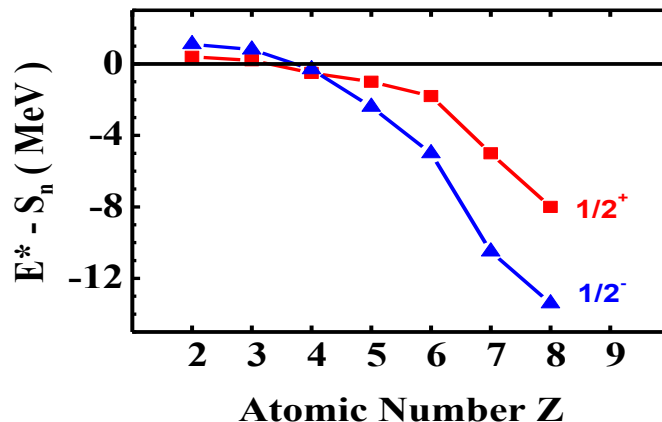


$$\rho_0 \sim 0.16 \text{ fm}^{-3}$$

proton and neutrons
homogeneously mixed,
no decoupling of proton
and neutron distributions



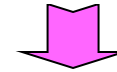
⁹He ¹⁰Li ¹¹Be ¹²B ¹³C ¹⁴N ¹⁵O



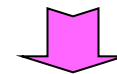
UNSTABLE NUCLEI

$$N / Z \sim 0.6 - 4$$

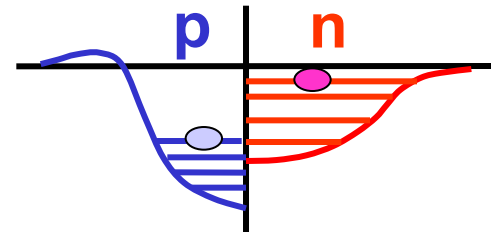
$$\epsilon_S \sim 0 - 40 \text{ MeV}$$



decoupling of proton and
neutron distributions



neutron **halos** and
neutron skins



Prerequisite of the halo formation:

low angular momentum motion for **halo**
particles and **few-body** dynamics

1s – intruder level

¹¹Be **parity inversion** of g.s.

¹⁰Li g.s. : $[\pi 0 p_{3/2} \otimes \nu 1 s_{1/2}]_{2-}$

PECULIARITIES OF HALO

in ground state

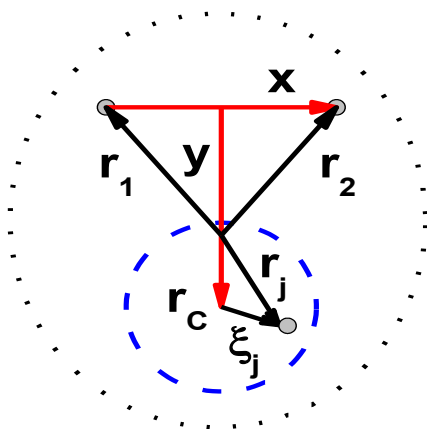
weakly bound,
with large extension
and space granularity

elastic scattering
some inclusive observables
(reaction cross sections, ...)

in low-energy continuum

concentration of the transition
strength near break up threshold
– soft modes

nuclear reactions
(transition properties)



BASIC dynamics
of halo nuclei

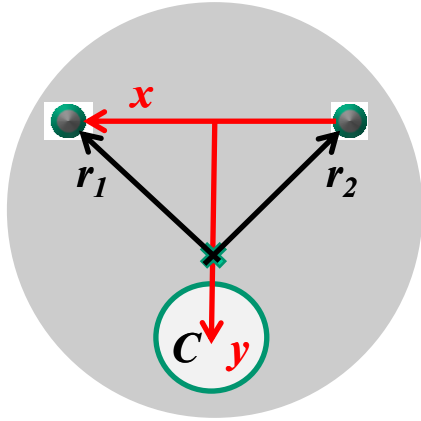


Decoupling of halo and
nuclear core degrees of
freedom

Dominance of few-body dynamics

$$\Phi(\bar{r}_1, \dots, \bar{r}_A) = \varphi(\bar{\xi}_1, \dots, \bar{\xi}_C) \psi(\bar{x}, \bar{y})$$

FEW-BODY CLUSTER MODELS



The Schrodinger equation

$$H_{\mathbf{A}} \Psi(\mathbf{r}_1, \dots, \mathbf{r}_A) = E \Psi(\mathbf{r}_1, \dots, \mathbf{r}_A)$$

Total hamiltonian of the **three-body** cluster models ($A = A_C + 2$)

$$H_{\mathbf{A}} = H_{\mathbf{A}_C} + T_{\mathbf{x}, \mathbf{y}} + V(r_1, r_2) + \sum_{i=1}^{A_C} V(r_1, r_i) + \sum_{i=1}^{A_C} V(r_2, r_i)$$

wave function is **factorized** into a sum of products from two parts

$$\Psi(\mathbf{r}_1, \dots, \mathbf{r}_A) = \sum_i \phi_i(\mathbf{r}_1, \dots, \mathbf{r}_{A_C}) \psi_i(\mathbf{x}, \mathbf{y})$$

The sum may include core excitations

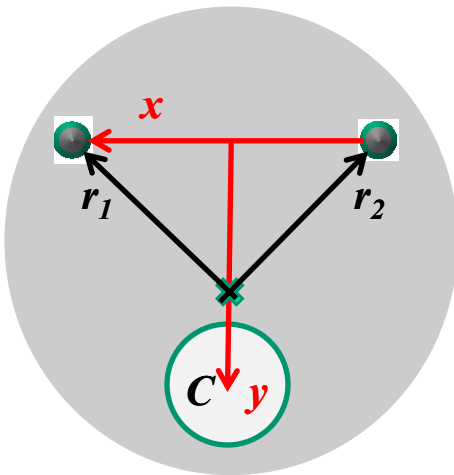
$$H_{\mathbf{A}_C} \phi_i(\mathbf{r}_1, \dots, \mathbf{r}_{A_C}) = \epsilon_i \phi_i(\mathbf{r}_1, \dots, \mathbf{r}_{A_C})$$

Calculations of the **bound** states and **continuum** wave functions

Borromean nature of halo nuclei
(no bound states between pairs of clusters)



one type of the wave function asymptotic behaviour



$$\rho^2 = \mu_x \bar{x}^2 + \mu_y \bar{y}^2$$

$$\alpha_\rho = \arctan\left(\frac{\sqrt{\mu_x} \bar{x}}{\sqrt{\mu_y} \bar{y}}\right)$$

$$\Omega_5^\rho = \{\alpha_\rho, \hat{x}, \hat{y}\}$$

$$\{\bar{x}, \bar{y}\} \Rightarrow \{\rho, \Omega_5^\rho\}$$

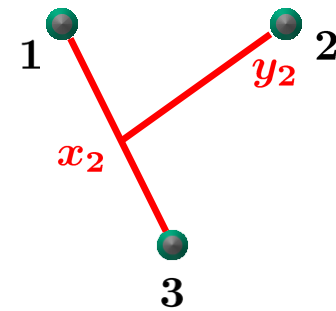
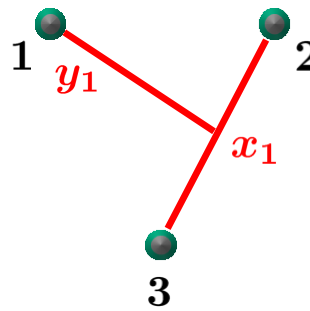
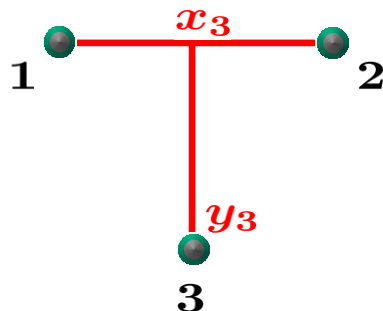
$$\frac{\kappa^2}{2m} = \frac{\bar{k}_x^2}{2\mu_x} + \frac{\bar{k}_y^2}{2\mu_y}$$

$$\alpha_\kappa = \arctan\left(\frac{\sqrt{\mu_y} \bar{k}_x}{\sqrt{\mu_x} \bar{k}_y}\right)$$

$$\Omega_5^\kappa = \{\alpha_\kappa, \hat{k}_x, \hat{k}_y\}$$

$$\{\bar{k}_x, \bar{k}_y\} \Rightarrow \{\kappa, \Omega_5^\kappa\}$$

$\{\rho, \kappa\}$ are independent of the Jacobi system



The **bound state** wave function ($\gamma = \{K, l_x, l_y, L, S, I, j\}$)

$$\Psi_{JM}(\mathbf{r}_1, \dots, \mathbf{r}_A) = \frac{1}{\rho^{5/2}} \sum_{\gamma} \chi_{\gamma}^J(\rho) \left[\Upsilon_{KL}^{l_x l_y}(\Omega_5^{\rho}) \otimes [\chi_s \otimes \phi_{nI}]_j \right]_{JM}$$

The **continuum** wave function at the positive energy

$$\begin{aligned} \Psi_{s\nu IM_I}^{(\pm)}(\mathbf{k}_x, \mathbf{k}_y; \mathbf{r}_i) = & \sum_{\gamma} i^K (s\nu IM_I \mid jm_j) (LM_L jm_j \mid JM_J) \Upsilon_{KL}^{l_x l_y*}(\Omega_5^{\kappa}) \times \\ & \times \frac{1}{\rho^{5/2}} \sum_{\gamma'} \chi_{\gamma', \gamma}^J(\kappa, \rho) \left[\Upsilon_{K'L'}^{l'_x l'_y}(\Omega_5^{\rho}) \otimes [\chi_{s'} \otimes \phi_{n'I'}]_{j'} \right]_{JM} \end{aligned}$$

Set of coupled Schrödinger equations for **radial** wave functions

$$\left(-\frac{\hbar^2}{2\mu} \left[\frac{d^2}{d\rho^2} - \frac{(K+3/2)(K+5/2)}{\rho^2} \right] + \epsilon_{\gamma} - E \right) \chi_{\gamma, \gamma'}^J(\rho) = - \sum_{\gamma''} V_{\gamma, \gamma''}^J(\rho) \chi_{\gamma'', \gamma'}^J(\rho)$$

Hyperspherical harmonics $\Upsilon_{KLM}^{l_x, l_y}(\Omega_5)$ ($K = 2n + l_x + l_y$)

$$\Upsilon_{KLM}^{l_x, l_y}(\Omega_5^{\rho}) = \psi_K^{l_x, l_y}(\alpha_{\rho}) [Y_{l_x}(\hat{x}) \otimes Y_{l_y}(\hat{y})]_{LM}$$

$$\psi_K^{l_x, l_y}(\alpha) = N_K^{l_x, l_y} (\sin \alpha)^{l_x} (\cos \alpha)^{l_y} P_n^{(l_x+1/2, l_y+1/2)}(\cos 2\alpha)$$

$$\Upsilon_{KLM}^{l'_x, l'_y}(\Omega'_5) = \sum_{l_x, l_y} \left\langle l_x, l_y \mid l'_x, l'_y \right\rangle_{KL} \Upsilon_{KLM}^{l_x, l_y}(\Omega_5)$$

boundary condition of the radial wave function at the origin

$$\chi_{\gamma}^J(\rho \rightarrow 0) \rightarrow 0$$

asymptotic behaviour of the bound state radial wave function

$$\chi_{\gamma}^J(\rho \rightarrow 0) \rightarrow \exp(-\kappa_n \rho), \quad \kappa_n = \sqrt{2m | E_n - \epsilon_{\gamma} | / \hbar^2}$$

asymptotic behaviour of the continuum radial wave function

$$\chi_{\gamma',\gamma}^J(\kappa, \rho \rightarrow \infty) \rightarrow \frac{i}{\sqrt{2\pi}} \frac{1}{\sqrt{k_{\gamma} k_{\gamma'}}} \left(H_{K+2}^{(-)}(k_{\gamma} \rho) \delta_{\gamma,\gamma'} - H_{K'+2}^{(+)}(k_{\gamma'} \rho) S_{\gamma',\gamma} \right)$$

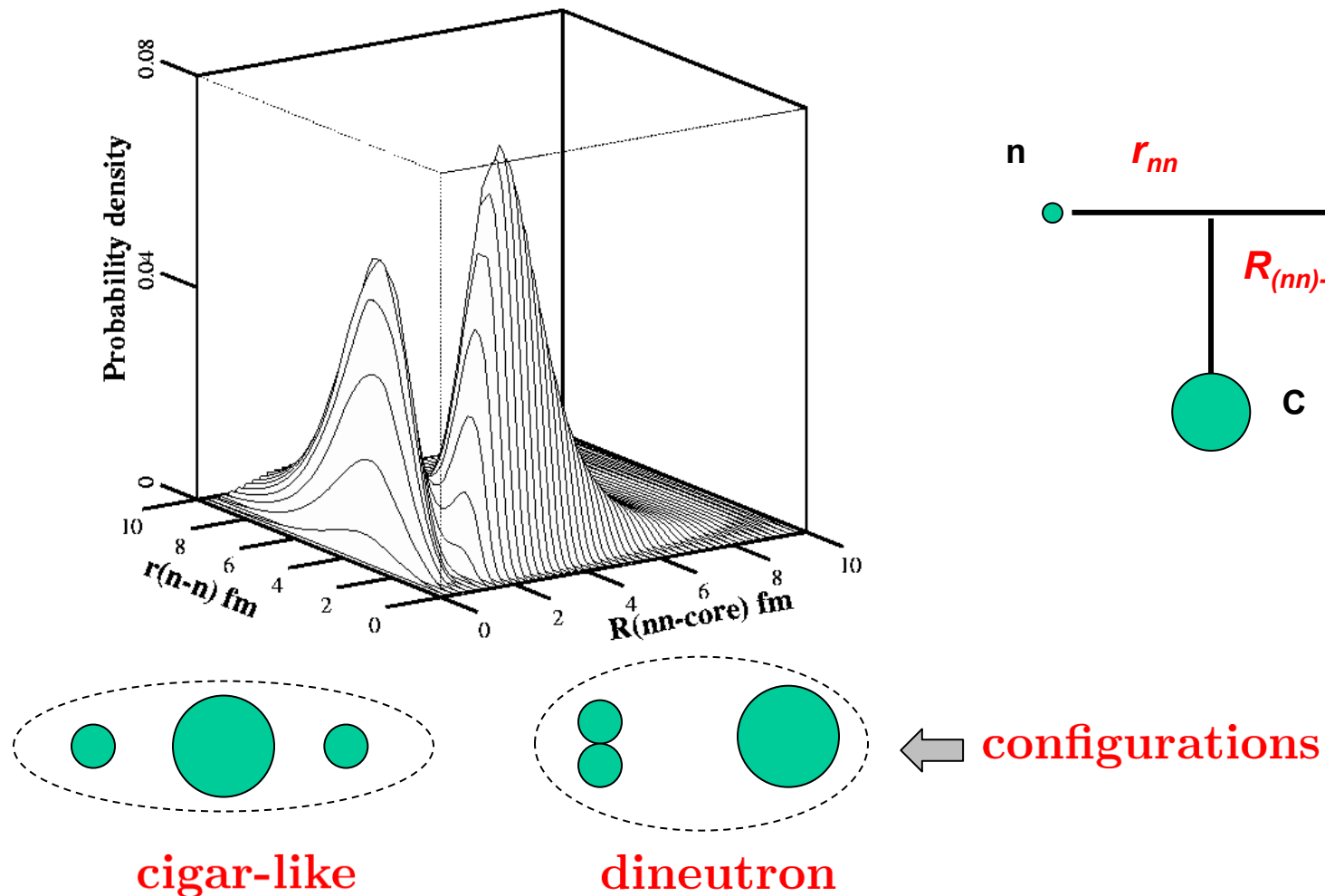
$$k_{\gamma} = \sqrt{2m | E - \epsilon_{\gamma} | / \hbar^2}$$

In collisions we explore the transition properties of nuclei
from ground state to continuum states

$$\langle \Psi^{(-)}(\bar{k}_x, \bar{k}_y) || \sum_p \frac{\delta(r - r_p)}{r r_p} [Y_L(\hat{r}_p) \times \sigma_p]_J || \Psi_{gr.st.} \rangle$$

Correlation density for the ground state of ${}^6\text{He}$

$$P(r_{nn}, R_{nn-C}) = r_{nn}^2 R_{nn-C}^2 \frac{1}{2J+1} \sum_M \int d\Omega_{nn} d\Omega_{nn-C} |\Phi_{JM}(\bar{r}_{nn}^2, \bar{R}_{nn-C}^2)|^2$$



CONCLUSIONS

- Near the dripline nuclear structure may be dramatically different. Reaching the limits of nuclear stability offers unique opportunities to understand basic nuclear properties
- Halo – new structural dripline phenomenon with clusterization into an ordinary core nucleus and a veil of halo nucleons forming very dilute neutron matter, have already been discovered in reaction with radioactive ion beams
- Development of new experimental techniques for production and/or detection of radioactive beams is the way to unexplored

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