

# 20th Annual RDMS CMS Collaboration Conference

Tashkent-Samarkand

Uzbekistan, 12-15 September 2018

Conference dedicated to new physics results,  
Future Physics beyond the LHC era, and the CMS  
Detector Upgrade, including Endcap Calorimetry

Registration: <http://rdms2018.jinr.ru/>

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# WHAT MAKES US THINK THAT THERE IS PHYSICS BEYOND THE STANDARD MODEL?

Dmitry Kazakov

BLTP JINR

RDMS CMS Annual Meeting, Tashkent, Sept 2018

# The Standard Model of Fundamental Interactions

Higgs Sector

Neutrino Sector

Flavour Sector

Dark Matter

New particles and Interactions



# THE LAGRANGIAN

$$\mathcal{L} = \mathcal{L}_{gauge} + \mathcal{L}_{fermions} + \mathcal{L}_{Higgs}$$

$$\mathcal{L}_{gauge} = -\frac{1}{4} G_{\mu\nu}^c G^{\mu\nu c} + i\bar{L}_\alpha \gamma^\mu D_\mu L_\alpha + i\bar{U}_\alpha \gamma^\mu D_\mu U_\alpha + i\bar{N}_\alpha \gamma^\mu \partial_\mu N_\alpha$$

$$+ i\bar{L}_\alpha \gamma^\mu D_\mu L_\alpha + i\bar{U}_\alpha \gamma^\mu D_\mu U_\alpha + i\bar{N}_\alpha \gamma^\mu \partial_\mu N_\alpha$$

$$+ i\bar{N}_\alpha \gamma^\mu \partial_\mu N_\alpha$$

$\mathcal{L}_{Yuk}$

$$+ y_{\alpha\beta}^L \bar{L}_\alpha E_\beta H + y_{\alpha\beta}^N \bar{L}_\alpha N_\beta$$

$$+ y_{\alpha\beta}^N \bar{L}_\alpha N_\beta$$

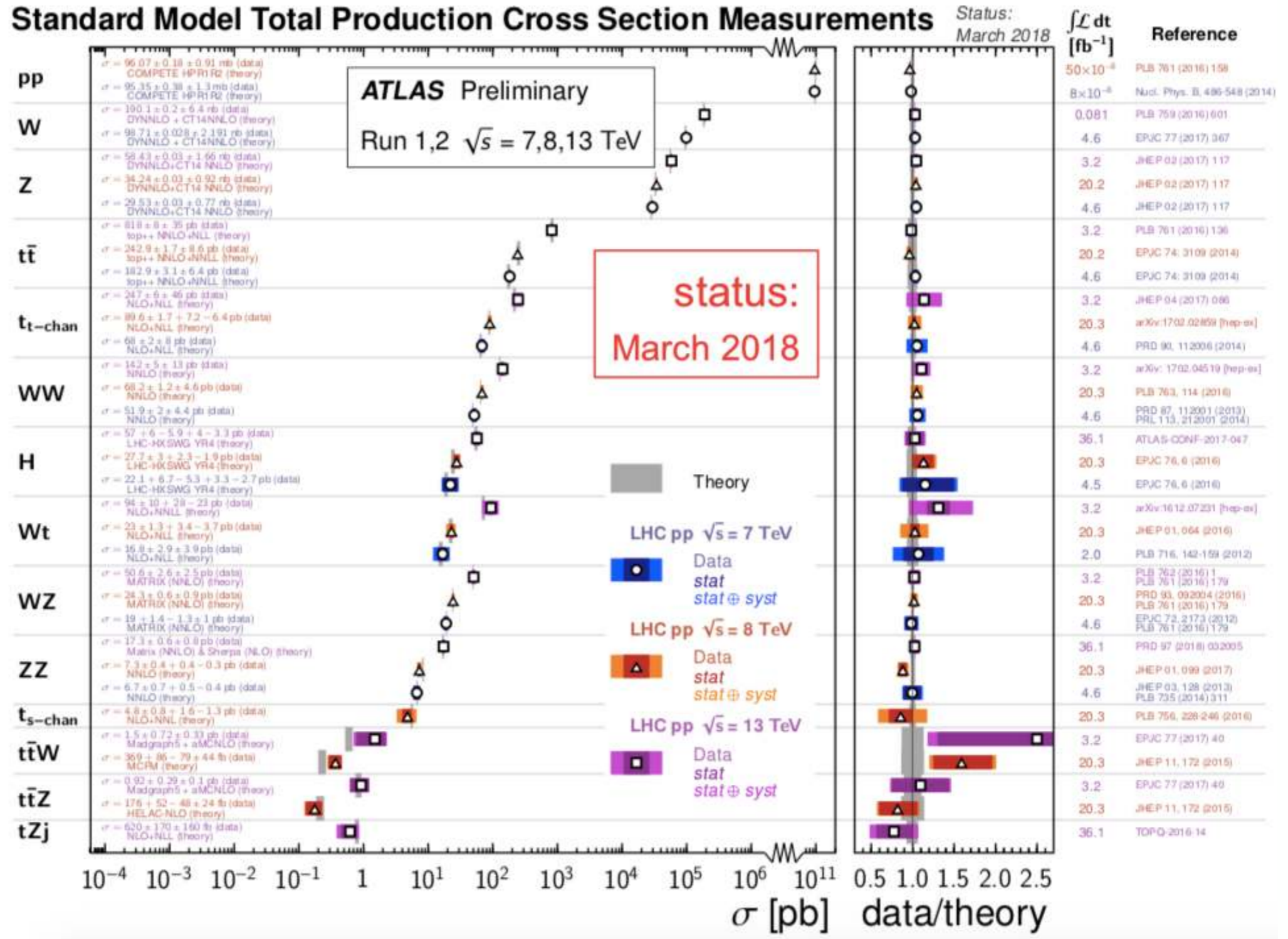
$$\mathcal{L}_{Higgs} = -V =$$

$$\begin{aligned} \mathcal{L}_{SM} = & -\frac{1}{2} \partial_\nu g_\mu^a \partial_\nu g_\mu^a - g_s f^{abc} \partial_\mu g_\nu^a g_\mu^b g_\nu^c - \frac{1}{4} g_s^2 f^{abc} f^{ade} g_\mu^b g_\nu^c g_\mu^d g_\nu^e - \partial_\nu W_\mu^+ \partial_\nu W_\mu^- - \\ & M^2 W_\mu^+ W_\mu^- - \frac{1}{2} \partial_\nu Z_\mu^0 \partial_\nu Z_\mu^0 - \frac{1}{2c_w^2} M^2 Z_\mu^0 Z_\mu^0 - \frac{1}{2} \partial_\mu A_\nu \partial_\mu A_\nu - igc_w (\partial_\nu Z_\mu^0 (W_\mu^+ W_\nu^- - \\ & W_\nu^+ W_\mu^-) - Z_\nu^0 (W_\mu^+ \partial_\nu W_\mu^- - W_\mu^- \partial_\nu W_\mu^+) + Z_\mu^0 (W_\nu^+ \partial_\nu W_\mu^- - W_\nu^- \partial_\nu W_\mu^+)) - \\ & igs_w (\partial_\nu A_\mu (W_\mu^+ W_\nu^- - W_\nu^+ W_\mu^-) - A_\nu (W_\mu^+ \partial_\nu W_\mu^- - W_\mu^- \partial_\nu W_\mu^+) + A_\mu (W_\nu^+ \partial_\nu W_\mu^- - \\ & W_\nu^- \partial_\nu W_\mu^+)) - \frac{1}{2} g^2 W_\mu^+ W_\mu^- W_\nu^+ W_\nu^- + \frac{1}{2} g^2 W_\mu^+ W_\nu^- W_\mu^- W_\nu^+ + g^2 c_w^2 (Z_\mu^0 W_\mu^+ Z_\nu^0 W_\nu^- - \\ & Z_\mu^0 Z_\nu^0 W_\mu^+ W_\nu^-) + g^2 s_w^2 (A_\mu W_\mu^+ A_\nu W_\nu^- - A_\mu A_\nu W_\mu^+ W_\nu^-) + g^2 s_w c_w (A_\mu Z_\nu^0 (W_\mu^+ W_\nu^- - \\ & W_\nu^+ W_\mu^-) - 2A_\mu Z_\nu^0 W_\mu^+ W_\nu^-) - \frac{1}{2} \partial_\mu H \partial_\mu H - 2M^2 \alpha_h H^2 - \partial_\mu \phi^+ \partial_\mu \phi^- - \frac{1}{2} \partial_\mu \phi^0 \partial_\mu \phi^0 - \\ & \beta_h \left( \frac{2M^2}{g^2} + \frac{2M}{g} H + \frac{1}{2} (H^2 + \phi^0 \phi^0 + 2\phi^+ \phi^-) \right) + \frac{2M^4}{g^2} \alpha_h - \\ & g\alpha_h M (H^3 + H\phi^0 \phi^0 + 2H\phi^+ \phi^-) - \\ & \frac{1}{8} g^2 \alpha_h (H^4 + (\phi^0)^4 + 4(\phi^+ \phi^-)^2 + 4(\phi^0)^2 \phi^+ \phi^- + 4H^2 \phi^+ \phi^- + 2(\phi^0)^2 H^2) - \\ & gMW_\mu^+ W_\mu^- H - \frac{1}{2} g \frac{M}{c_w^2} Z_\mu^0 Z_\mu^0 H - \\ & \frac{1}{2} ig (W_\mu^+ (\phi^0 \partial_\mu \phi^- - \phi^- \partial_\mu \phi^0) - W_\mu^- (\phi^0 \partial_\mu \phi^+ - \phi^+ \partial_\mu \phi^0)) + \\ & \frac{1}{2} g (W_\mu^+ (H \partial_\mu \phi^- - \phi^- \partial_\mu H) + W_\mu^- (H \partial_\mu \phi^+ - \phi^+ \partial_\mu H)) + \frac{1}{2} g \frac{1}{c_w} (Z_\mu^0 (H \partial_\mu \phi^0 - \phi^0 \partial_\mu H) + \\ & M (\frac{1}{c_w} Z_\mu^0 \partial_\mu \phi^0 + W_\mu^+ \partial_\mu \phi^- + W_\mu^- \partial_\mu \phi^+) - ig \frac{s_w^2}{c_w} M Z_\mu^0 (W_\mu^+ \phi^- - W_\mu^- \phi^+) + igs_w M A_\mu (W_\mu^+ \phi^- - \\ & W_\mu^- \phi^+) - ig \frac{1-2c_w^2}{2c_w} Z_\mu^0 (\phi^+ \partial_\mu \phi^- - \phi^- \partial_\mu \phi^+) + igs_w A_\mu (\phi^+ \partial_\mu \phi^- - \phi^- \partial_\mu \phi^+) - \\ & \frac{1}{4} g^2 W_\mu^+ W_\mu^- (H^2 + (\phi^0)^2 + 2\phi^+ \phi^-) - \frac{1}{8} g^2 \frac{1}{c_w^2} Z_\mu^0 Z_\mu^0 (H^2 + (\phi^0)^2 + 2(2s_w^2 - 1)^2 \phi^+ \phi^-) - \\ & \frac{1}{2} g^2 \frac{s_w^2}{c_w} Z_\mu^0 \phi^0 (W_\mu^+ \phi^- + W_\mu^- \phi^+) - \frac{1}{2} ig^2 \frac{s_w^2}{c_w} Z_\mu^0 H (W_\mu^+ \phi^- - W_\mu^- \phi^+) + \frac{1}{2} g^2 s_w A_\mu \phi^0 (W_\mu^+ \phi^- + \\ & W_\mu^- \phi^+) + \frac{1}{2} ig^2 s_w A_\mu H (W_\mu^+ \phi^- - W_\mu^- \phi^+) - g^2 \frac{s_w}{c_w} (2c_w^2 - 1) Z_\mu^0 A_\mu \phi^+ \phi^- - \\ & g^2 s_w^2 A_\mu A_\mu \phi^+ \phi^- + \frac{1}{2} ig_s \lambda_{ij}^a (\bar{q}_i^\alpha \gamma^\mu q_j^\alpha) g_\mu^a - \bar{e}^\lambda (\gamma \partial + m_e^\lambda) e^\lambda - \bar{\nu}^\lambda (\gamma \partial + m_\nu^\lambda) \nu^\lambda - \bar{u}_j^\lambda (\gamma \partial + \\ & m_u^\lambda) u_j^\lambda - \bar{d}_j^\lambda (\gamma \partial + m_d^\lambda) d_j^\lambda + igs_w A_\mu (-\bar{e}^\lambda \gamma^\mu e^\lambda + \frac{2}{3} (\bar{u}_j^\lambda \gamma^\mu u_j^\lambda) - \frac{1}{3} (\bar{d}_j^\lambda \gamma^\mu d_j^\lambda)) + \\ & \frac{ig}{4c_w} Z_\mu^0 \{ (\bar{\nu}^\lambda \gamma^\mu (1 + \gamma^5) \nu^\lambda) + (\bar{e}^\lambda \gamma^\mu (4s_w^2 - 1 - \gamma^5) e^\lambda) + (\bar{d}_j^\lambda \gamma^\mu (\frac{4}{3}s_w^2 - 1 - \gamma^5) d_j^\lambda) + \\ & (\bar{u}_j^\lambda \gamma^\mu (1 - \frac{8}{3}s_w^2 + \gamma^5) u_j^\lambda) \} + \frac{ig}{2\sqrt{2}} W_\mu^+ ((\bar{\nu}^\lambda \gamma^\mu (1 + \gamma^5) U^{lep}{}_{\lambda\kappa} e^\kappa) + (\bar{u}_j^\lambda \gamma^\mu (1 + \gamma^5) C_{\lambda\kappa} d_j^\kappa)) + \\ & \frac{ig}{2\sqrt{2}} W_\mu^- ((\bar{e}^\kappa U^{lep}{}_{\kappa\lambda} \gamma^\mu (1 + \gamma^5) \nu^\lambda) + (\bar{d}_j^\kappa C_{\kappa\lambda}^\dagger \gamma^\mu (1 + \gamma^5) u_j^\lambda)) + \\ & \frac{ig}{2M\sqrt{2}} \phi^+ (-m_e^\kappa (\bar{\nu}^\lambda U^{lep}{}_{\lambda\kappa} (1 - \gamma^5) e^\kappa) + m_\nu^\lambda (\bar{\nu}^\lambda U^{lep}{}_{\lambda\kappa} (1 + \gamma^5) e^\kappa) + \\ & \frac{ig}{2M\sqrt{2}} \phi^- (m_e^\lambda (\bar{e}^\lambda U^{lep}{}_{\lambda\kappa} (1 + \gamma^5) \nu^\kappa) - m_\nu^\kappa (\bar{e}^\lambda U^{lep}{}_{\lambda\kappa} (1 - \gamma^5) \nu^\kappa) - \frac{g}{2} \frac{m_\nu^\lambda}{M} H (\bar{\nu}^\lambda \nu^\lambda) - \\ & \frac{g}{2} \frac{m_\nu^\lambda}{M} H (\bar{e}^\lambda e^\lambda) + \frac{ig}{2} \frac{m_\nu^\lambda}{M} \phi^0 (\bar{\nu}^\lambda \gamma^5 \nu^\lambda) - \frac{ig}{2} \frac{m_\nu^\lambda}{M} \phi^0 (\bar{e}^\lambda \gamma^5 e^\lambda) - \frac{1}{4} \bar{\nu}_\lambda M_{\lambda\kappa}^R (1 - \gamma_5) \bar{\nu}_\kappa - \\ & \frac{1}{4} \bar{\nu}_\lambda M_{\lambda\kappa}^R (1 - \gamma_5) \bar{\nu}_\kappa + \frac{ig}{2M\sqrt{2}} \phi^+ (-m_d^\kappa (\bar{u}_j^\lambda C_{\lambda\kappa} (1 - \gamma^5) d_j^\kappa) + m_u^\lambda (\bar{u}_j^\lambda C_{\lambda\kappa} (1 + \gamma^5) d_j^\kappa) + \\ & \frac{ig}{2M\sqrt{2}} \phi^- (m_d^\lambda (\bar{d}_j^\lambda C_{\lambda\kappa}^\dagger (1 + \gamma^5) u_j^\kappa) - m_u^\kappa (\bar{d}_j^\lambda C_{\lambda\kappa}^\dagger (1 - \gamma^5) u_j^\kappa) - \frac{g}{2} \frac{m_u^\lambda}{M} H (\bar{u}_j^\lambda u_j^\lambda) - \\ & \frac{g}{2} \frac{m_u^\lambda}{M} H (\bar{d}_j^\lambda d_j^\lambda) + \frac{ig}{2} \frac{m_u^\lambda}{M} \phi^0 (\bar{u}_j^\lambda \gamma^5 u_j^\lambda) - \frac{ig}{2} \frac{m_u^\lambda}{M} \phi^0 (\bar{d}_j^\lambda \gamma^5 d_j^\lambda) + \bar{G}^a \partial^2 G^a + g_s f^{abc} \partial_\mu \bar{G}^a G^b g_\mu^c + \\ & \bar{X}^+ (\partial^2 - M^2) X^+ + \bar{X}^- (\partial^2 - M^2) X^- + \bar{X}^0 (\partial^2 - \frac{M^2}{c_w^2}) X^0 + \bar{Y} \partial^2 Y + igc_w W_\mu^+ (\partial_\mu \bar{X}^0 X^- - \\ & \partial_\mu \bar{X}^+ X^0) + igs_w W_\mu^+ (\partial_\mu \bar{Y} X^- - \partial_\mu \bar{X}^+ Y) + igc_w W_\mu^- (\partial_\mu \bar{X}^- X^0 - \\ & \partial_\mu \bar{X}^0 X^+) + igs_w W_\mu^- (\partial_\mu \bar{X}^- Y - \partial_\mu \bar{Y} X^+) + igc_w Z_\mu^0 (\partial_\mu \bar{X}^+ X^+ - \\ & \partial_\mu \bar{X}^- X^-) + igs_w A_\mu (\partial_\mu \bar{X}^+ X^+ - \\ & \partial_\mu \bar{X}^- X^-) - \frac{1}{2} g M (\bar{X}^+ X^+ H + \bar{X}^- X^- H + \frac{1}{c_w^2} \bar{X}^0 X^0 H) + \frac{1-2c_w^2}{2c_w} ig M (\bar{X}^+ X^0 \phi^+ - \bar{X}^- X^0 \phi^-) + \\ & \frac{1}{2c_w} ig M (\bar{X}^0 X^- \phi^+ - \bar{X}^0 X^+ \phi^-) + ig M s_w (\bar{X}^0 X^- \phi^+ - \bar{X}^0 X^+ \phi^-) + \\ & \frac{1}{2} ig M (\bar{X}^+ X^+ \phi^0 - \bar{X}^- X^- \phi^0) . \end{aligned}$$

All these parameters are not predicted



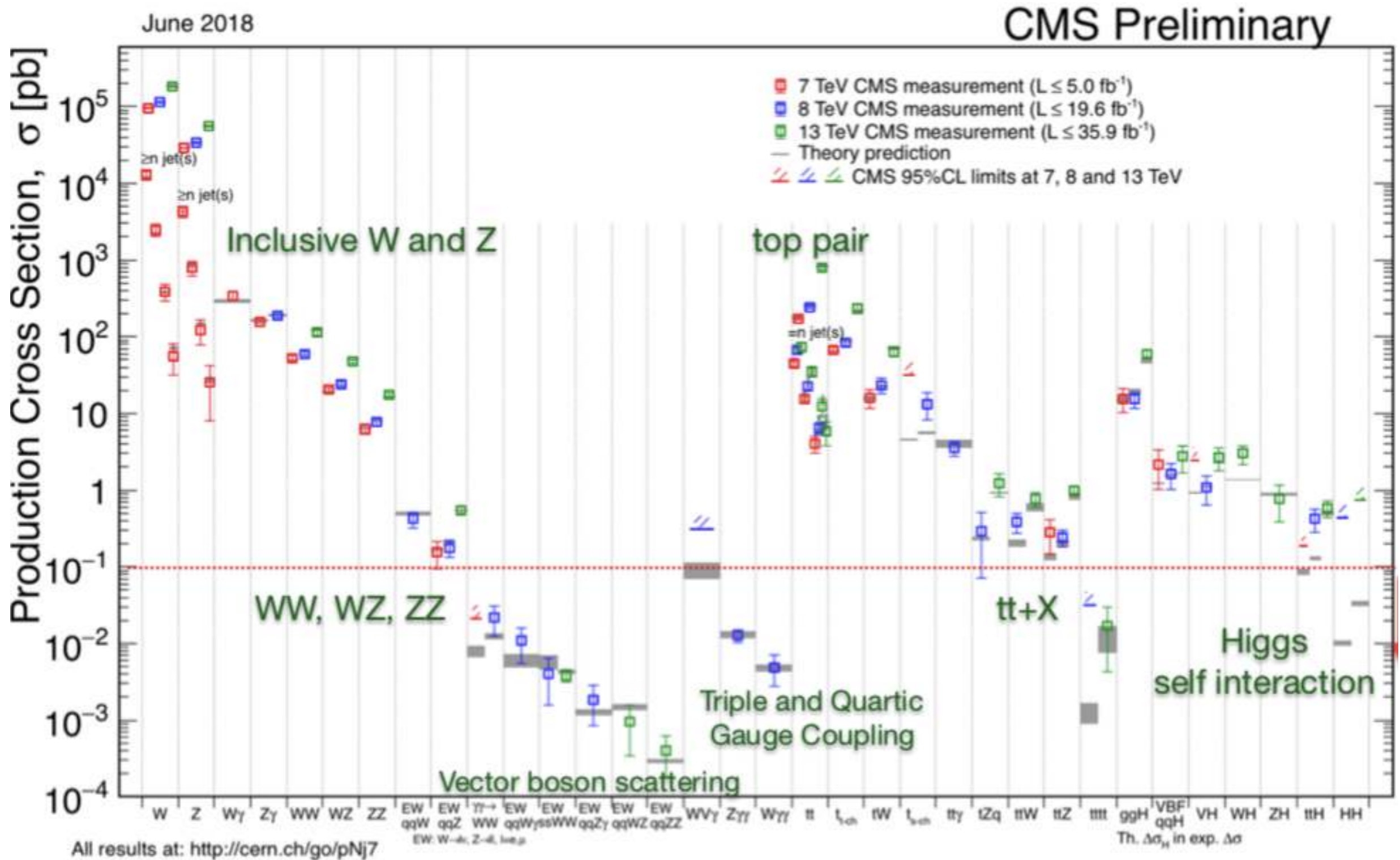
C. Gwenlan  
ICHEP2018



Extraordinary agreement between measurements and SM predictions

S. Rathatlou  
ICHEP2018

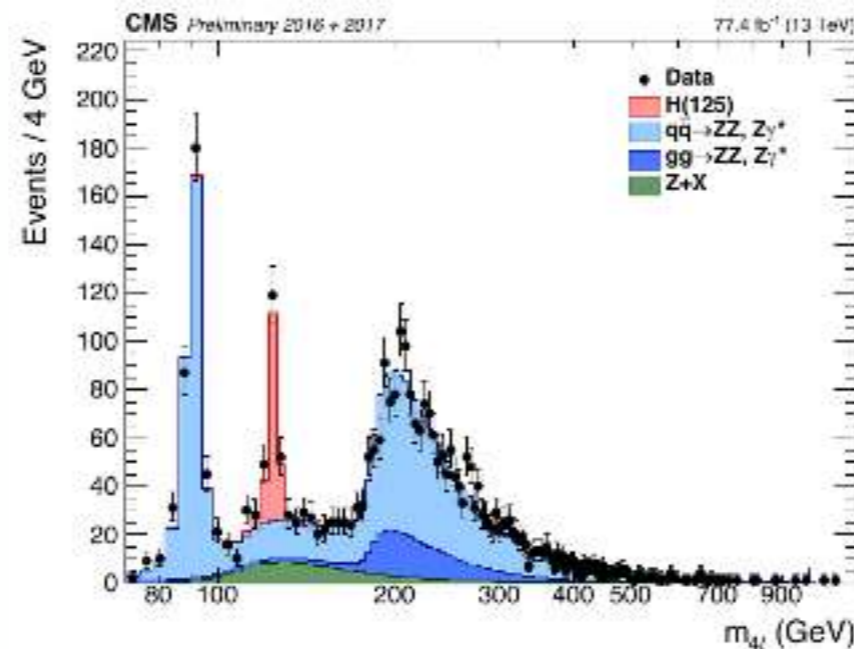
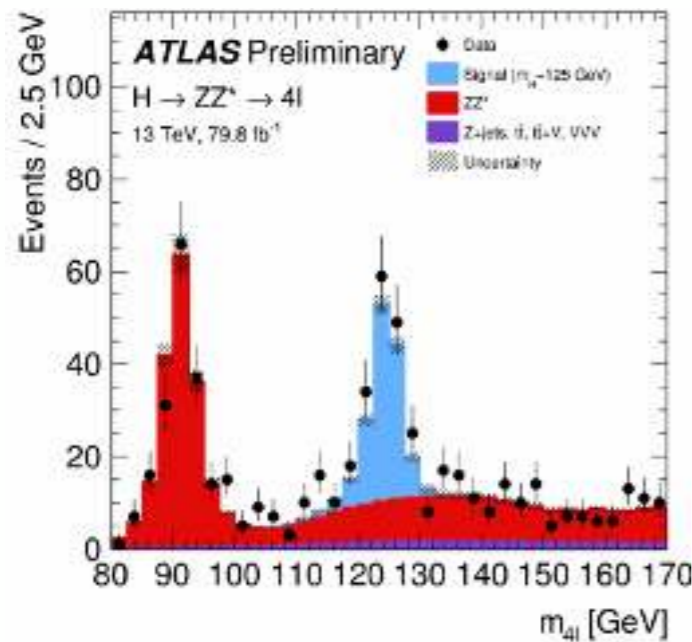
## New physics through precision





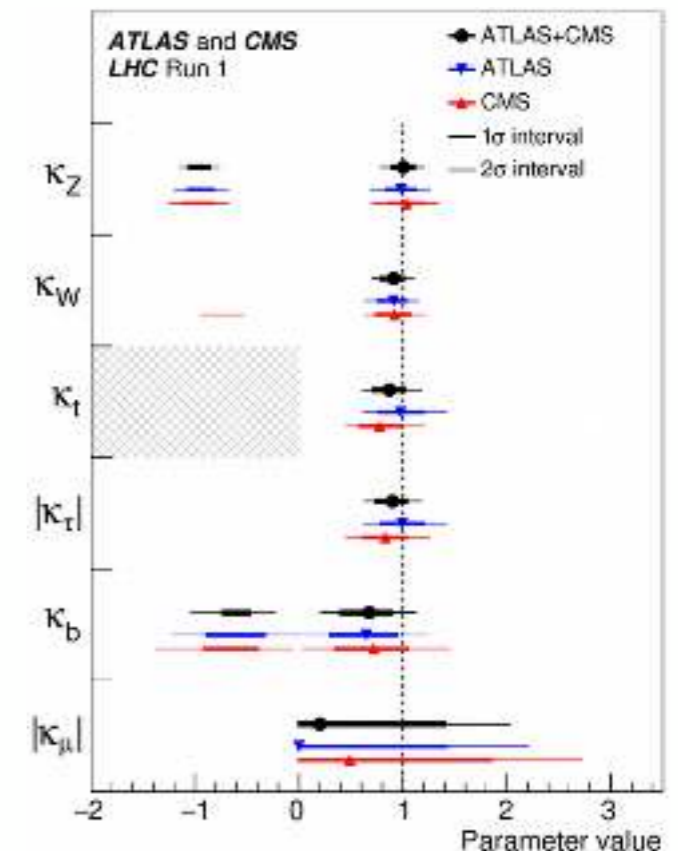
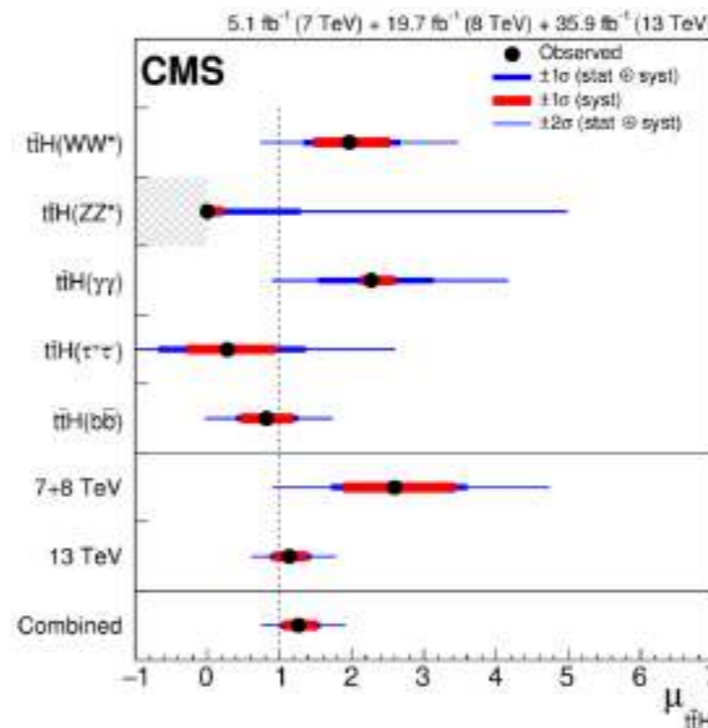
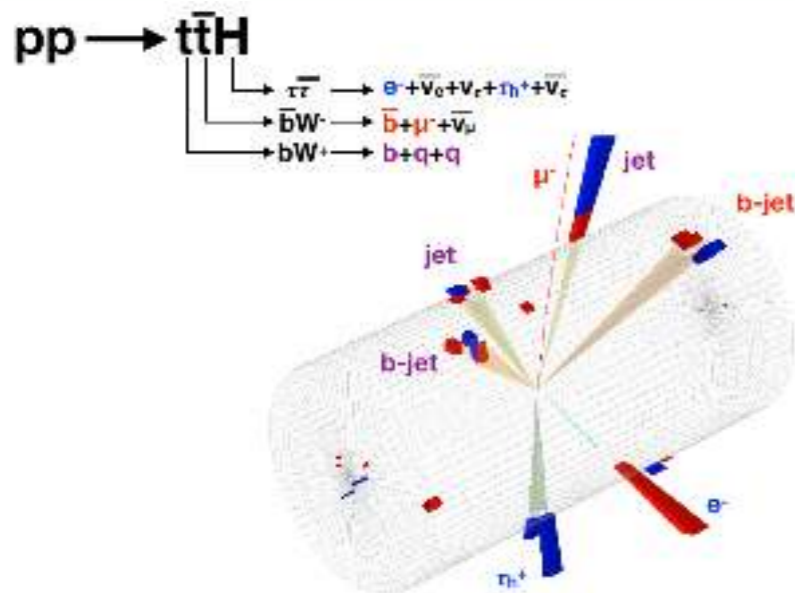
# Higgs bosons - entering precision era

Run-2 analyses with  $80 \text{ fb}^{-1}$  for the first time – higher precision is coming!



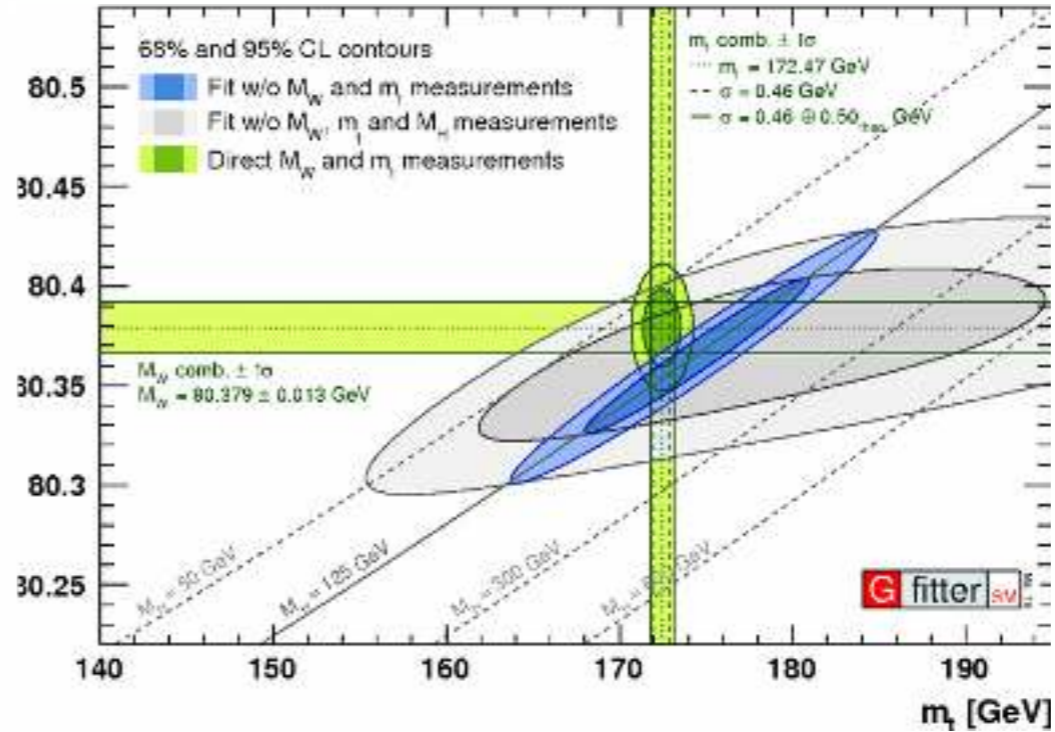
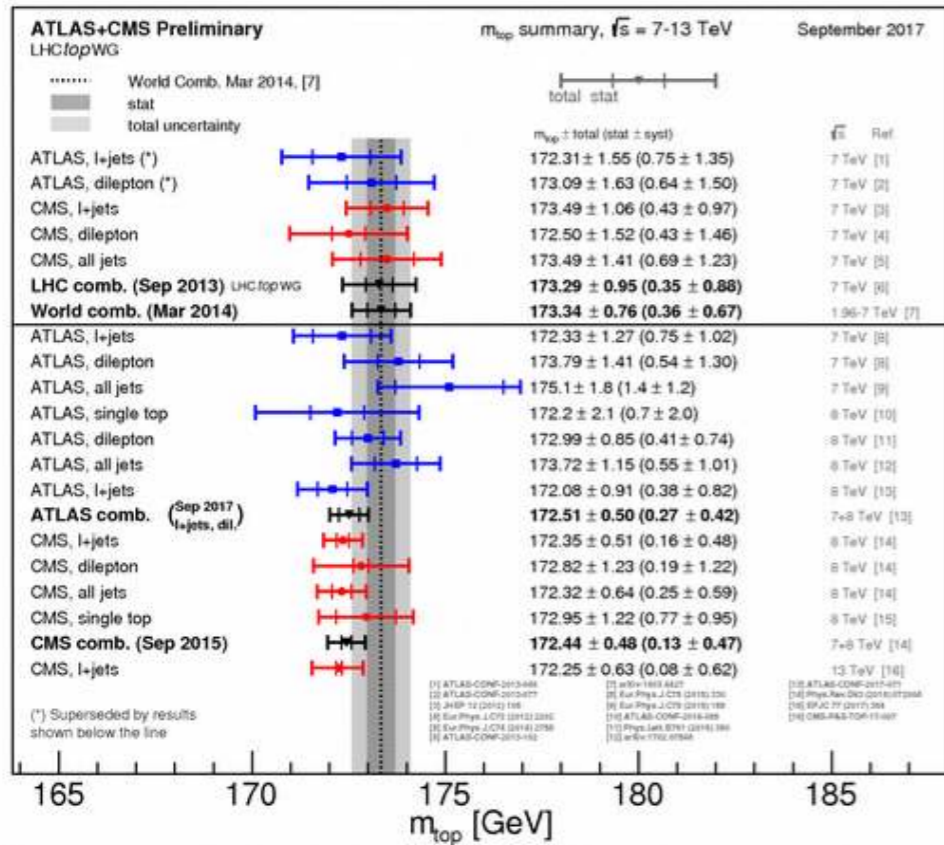
D. Charlton  
 LHCp2018

## ttH observation



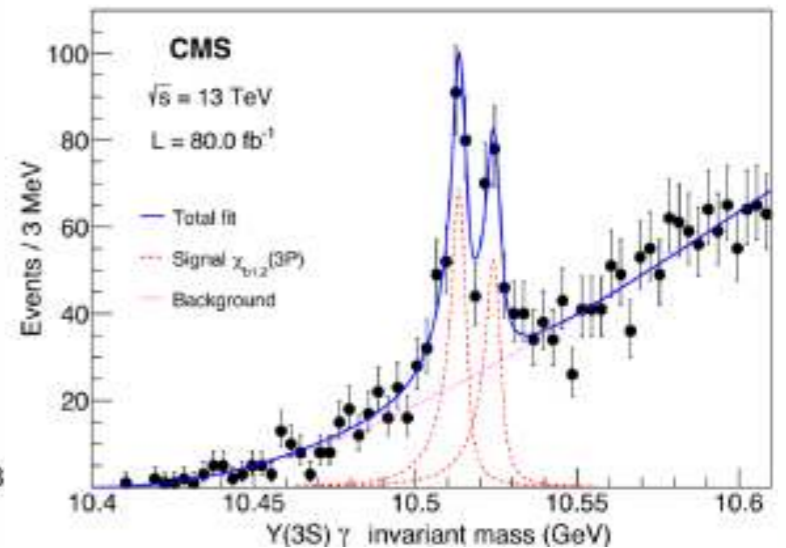
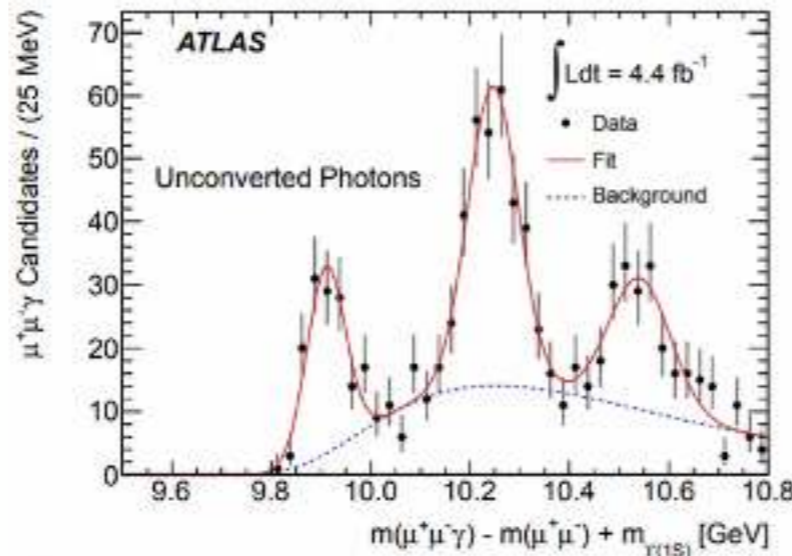
Precision EW mass measurements

D. Charlton  
LHCp2018



Precision spectroscopy!

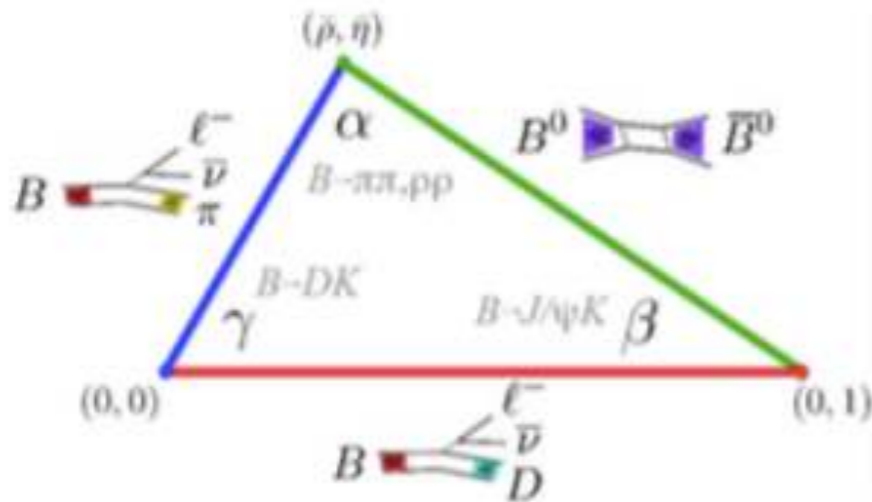
$$m(\chi_{b2}(3P)) - m(\chi_{b2}(3P)) = 10.60 \pm 0.64(\text{stat}) \pm 0.17(\text{syst}) \text{ MeV}$$





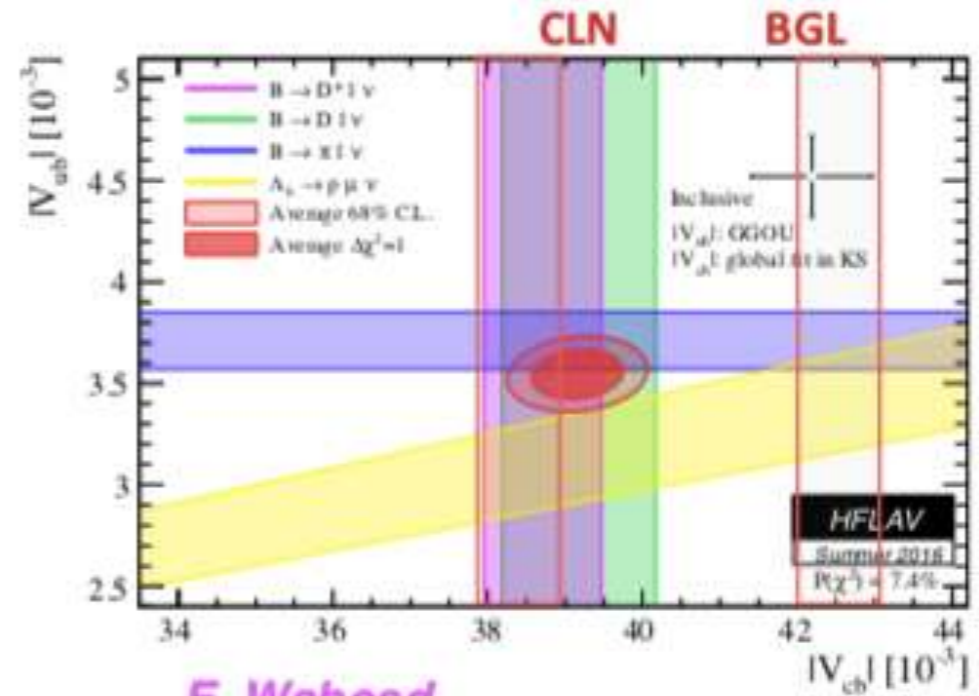
# Flavor Physics

- CKM and CPV
  - $|V_{cb}|$  puzzle resolved (not  $|V_{ub}|$ )
  - new  $\gamma$  ( $\varphi_3$ ) from LHCb

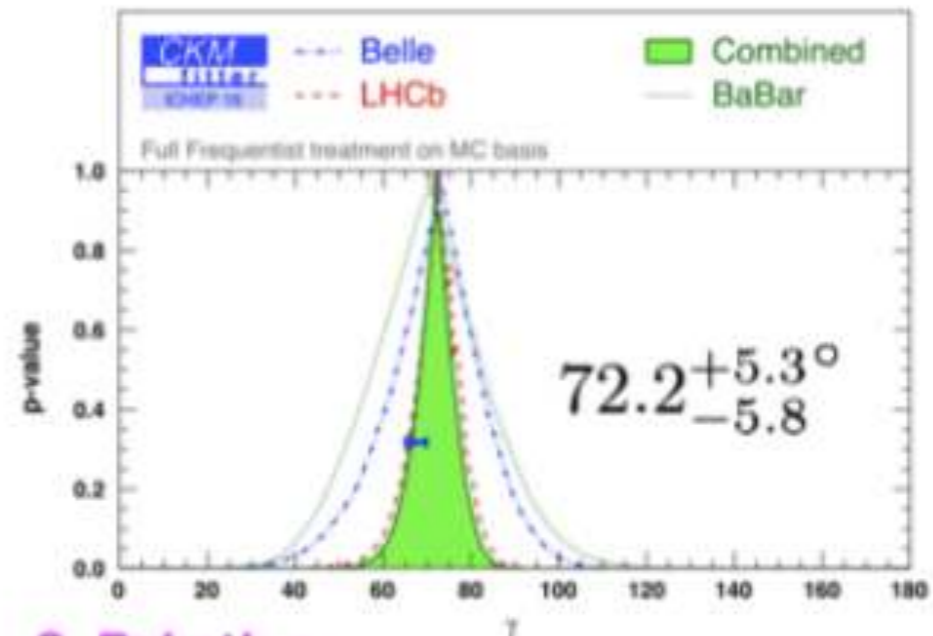


P. Urquijo

$B \rightarrow \pi\pi, \rho\rho$	$\alpha / \Phi_2$
$B \rightarrow D^{(*)} K^{(*)}$	$\gamma / \Phi_3$
$B \rightarrow J/\psi K_S$	$\beta / \Phi_1$
$B_s \rightarrow J/\psi \Phi$	$\beta_s$
$K \rightarrow \pi \nu \text{ anti-}\nu$	$\rho, \eta$



E. Waheed



S. Rahatlou

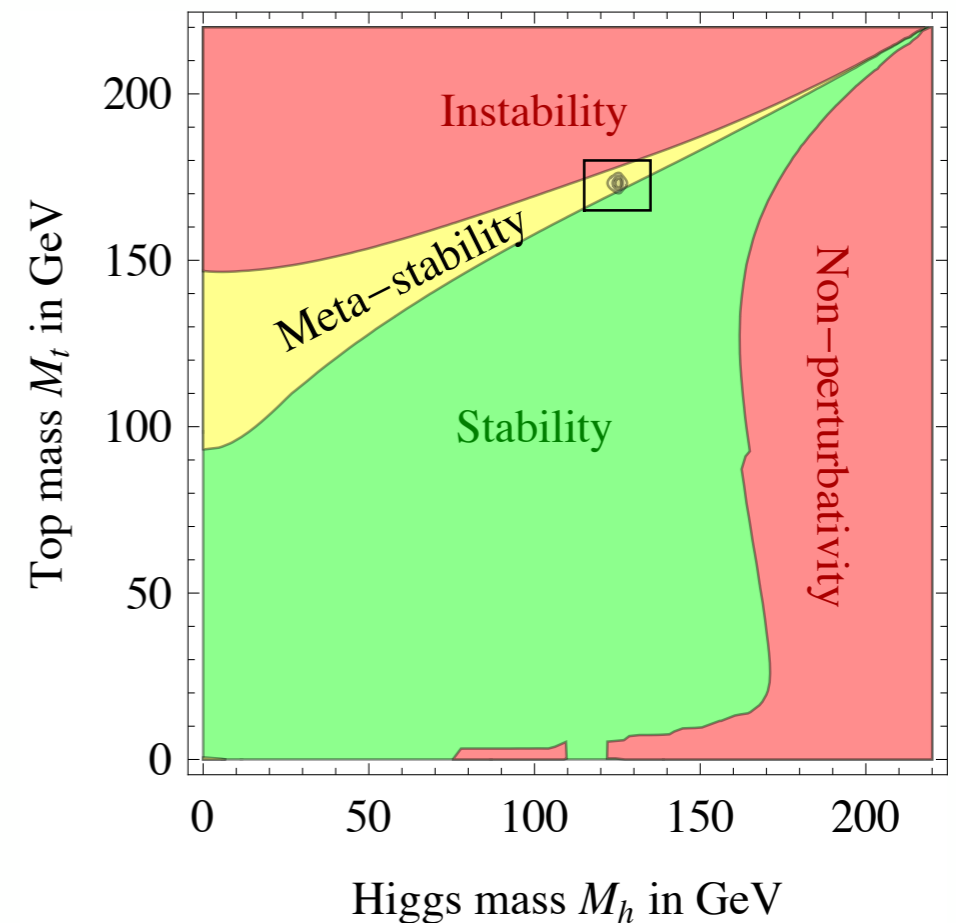
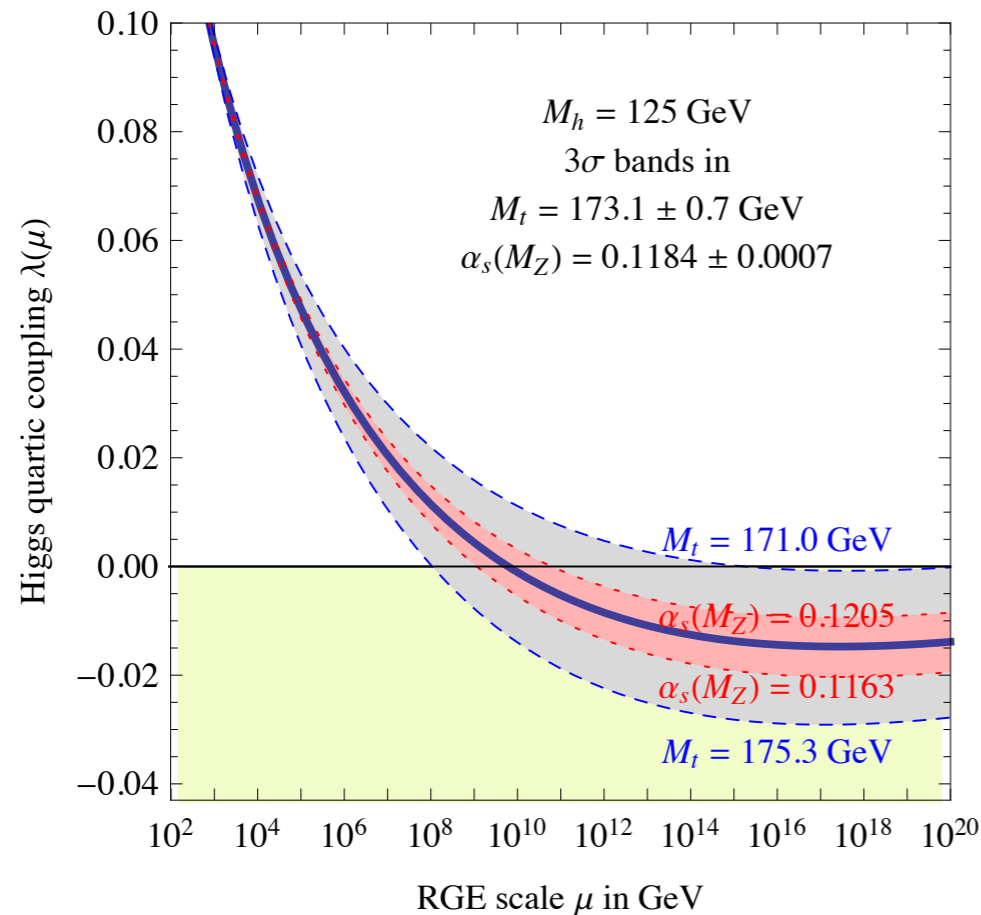
Paul Langacker (IAS)



# THE STANDARD MODEL: THE STATUS REPORT AND OPEN QUESTIONS

The electroweak vacuum is unstable under radiative corrections

The whole construction of the SM may be in trouble being metastable or even unstable



the situation crucially depends on the top and Higgs mass values and requires severe fine-tuning and high accuracy of calculations (3 loops)

## Muon anomalous magnetic moment

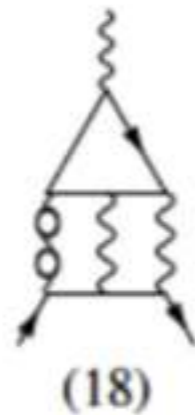
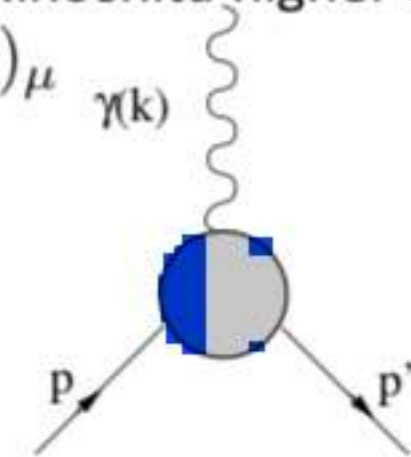
$$ie\bar{u}_\ell(p') \left[ \gamma^\mu - \frac{a_\ell}{2m_\ell} i\sigma^{\mu\nu} q_\nu \right] u_\ell(p) \epsilon_\mu^*$$

$$q_\mu = (p - p')_\mu \quad \gamma(k)$$

(Schwinger  $\alpha/\pi$ ,  
Kinoshita higher orders in  $\alpha$ )

Dirac equation predicts  $g=2$       $a = (g - 2)/2$

For electron  $a_e$  theory and experiment agrees!



$$a_\mu^{th} - a_\mu^{exp} = -(3.06 \pm 0.76) \times 10^{-8} \quad 4\sigma$$

Theory: uncertainty in hadronic contributions to the muon  $g - 2$ , (Jägerlehner, 1802.08019 ).  
Lattice QCD great progress light-by-light study (RBC & UKQCD, 1801.07224).

Fermilab and J-Park experiments are expected to clarify existing discrepancy!

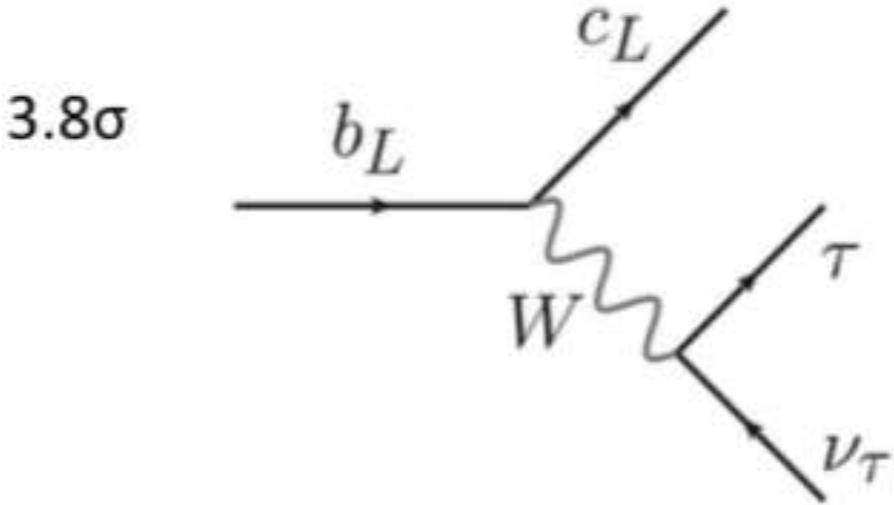


# THE STANDARD MODEL: THE STATUS REPORT AND OPEN QUESTIONS

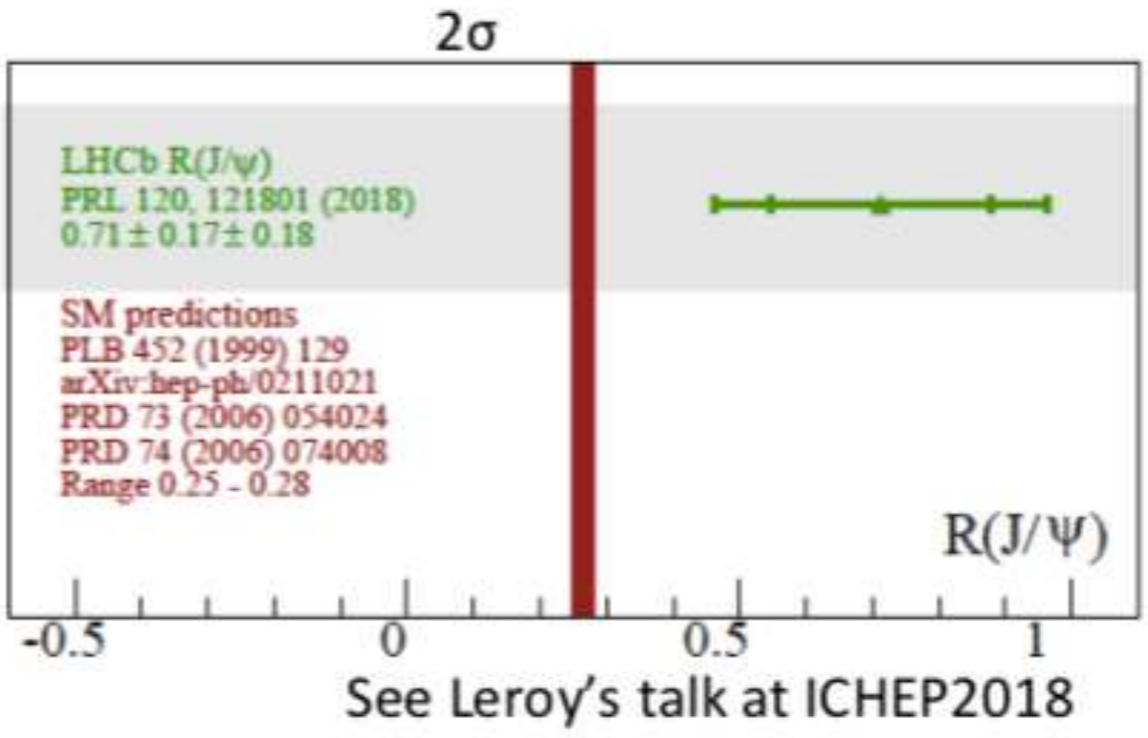
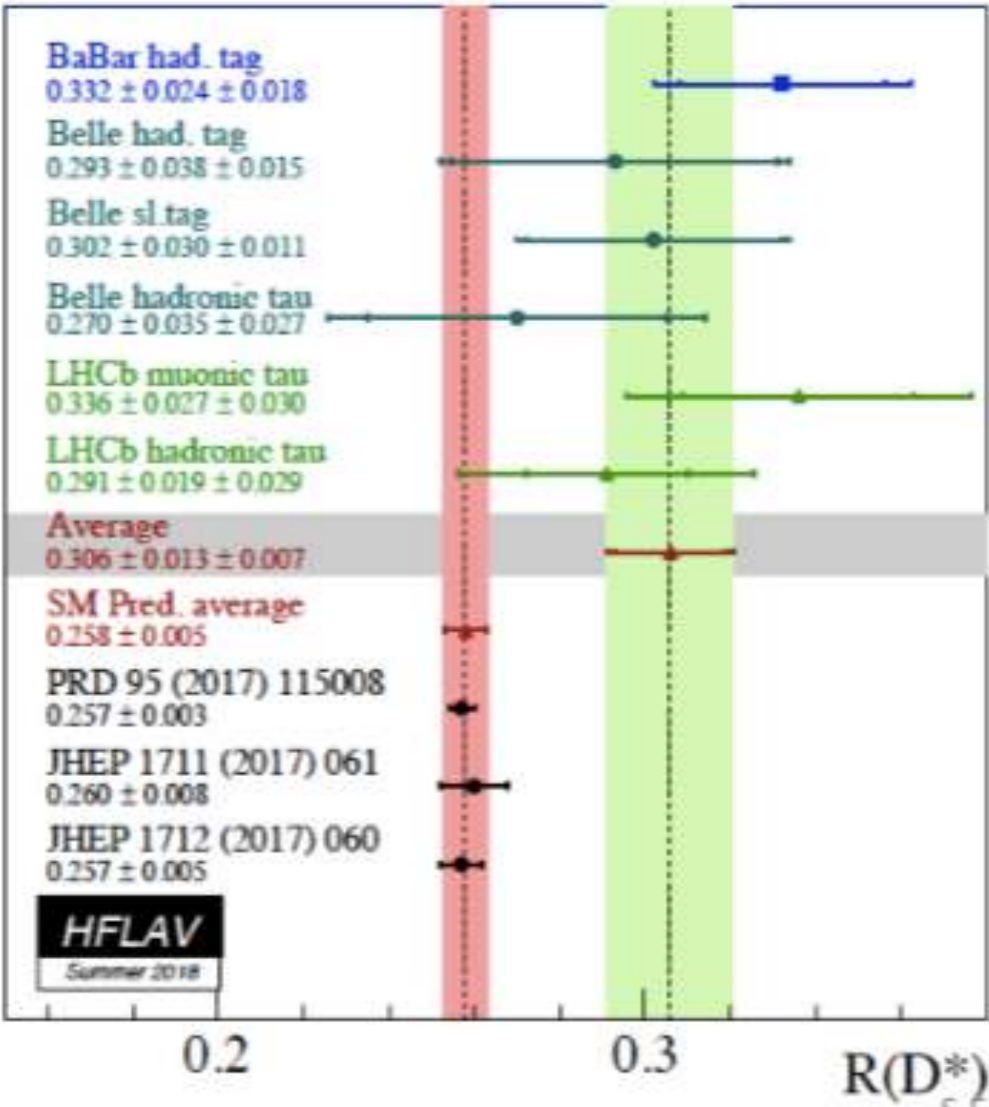
B physics anomalies: experimental results  $\neq$  SM predictions!

charged current (SM tree level)

$$R_{D^{(*)}} = \frac{BR(B \rightarrow D^{(*)} \tau \nu_\tau)}{BR(B \rightarrow D^{(*)} \mu \nu_\mu)}$$

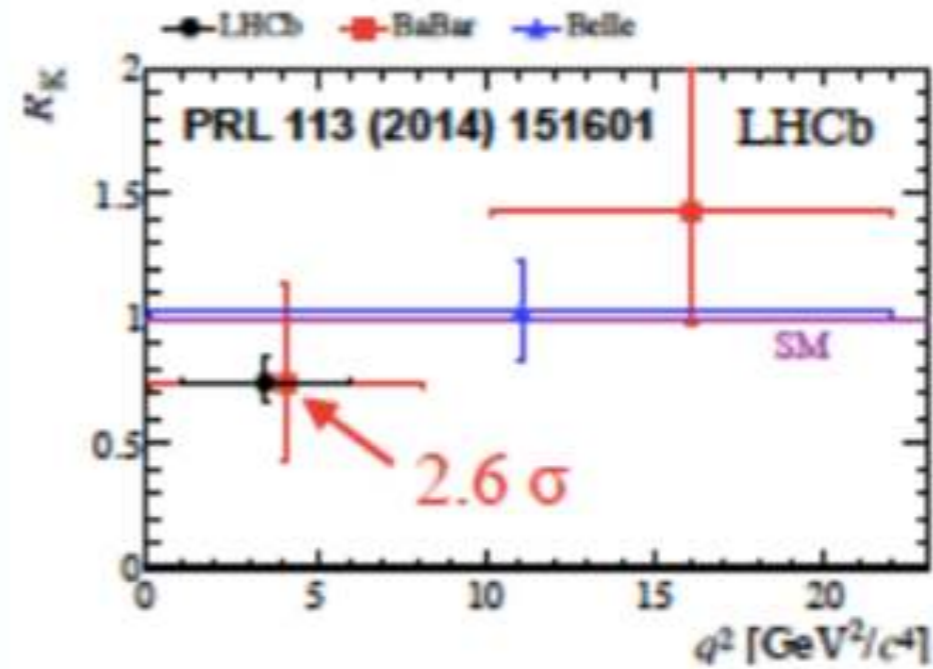
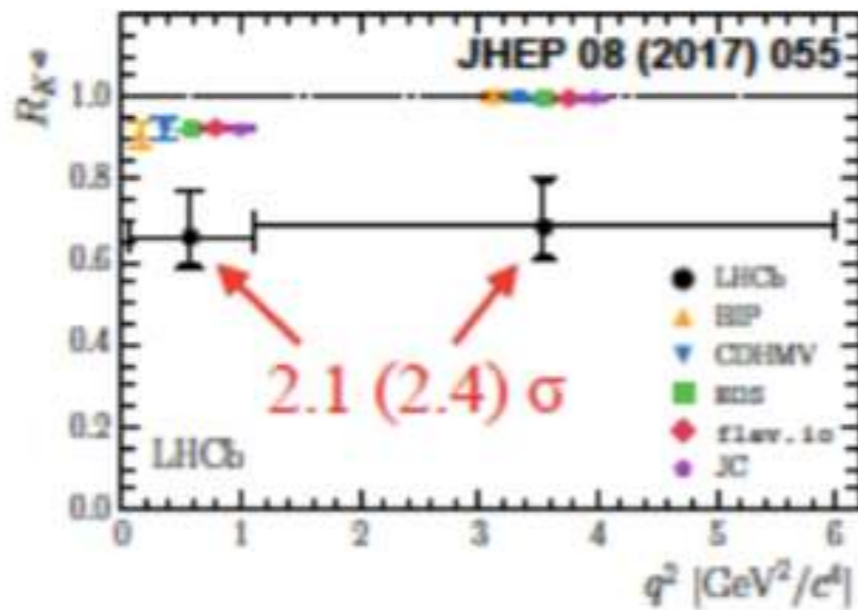


3.8 $\sigma$



## FCNC - SM loop process: $R_{K^{(*)}}$ anomaly

$$R_{K^{(*)}} = \frac{BR(B \rightarrow K^{(*)} \mu\mu)}{BR(B \rightarrow K^{(*)} ee)} \Bigg|_{q^2 \in [q_{min}^2, q_{max}^2]}$$



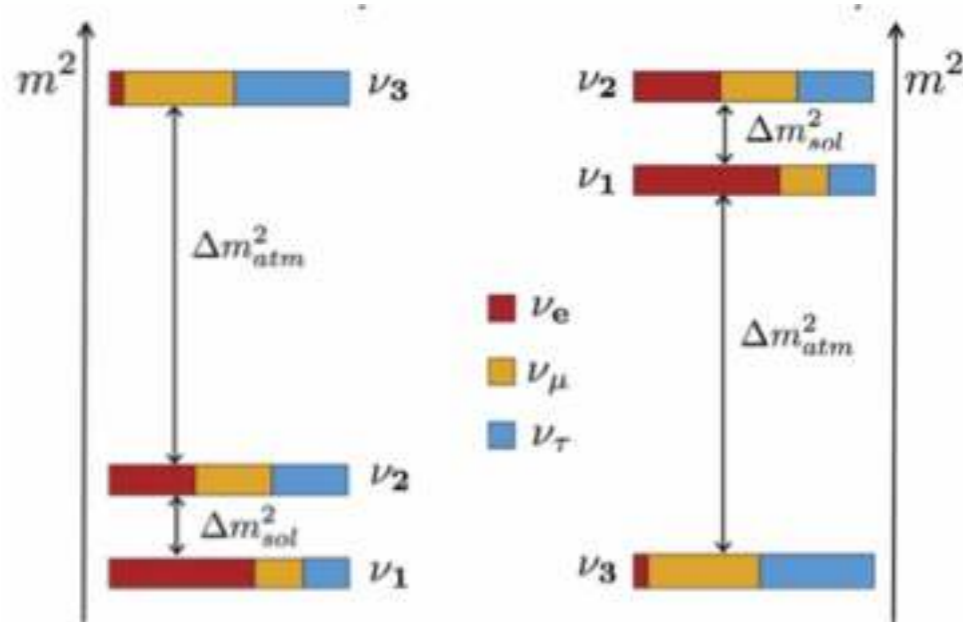
$P_5'$  in  $B \rightarrow K^* \mu^+ \mu^-$  (angular distribution functions)  $3\sigma$

LHCb: the discrepancy present in  $B_s \rightarrow \phi \mu\mu$  and  $\Lambda_b \rightarrow \Lambda \mu\mu$

(see Capriotti talk at ICHEP2018)



# Neutrino Physics



- Absolute value of neutrino masses ?
- Mass hierarchy?
- Dirac or Majorana?
- Fourth sterile neutrino?
- Neutrino dark matter?

$$0.06 \text{ eV} < \sum m_\nu < 0.12 \text{ eV}$$

↑
↑  
**ν-OSC**
**CMB**

PMNS-matrix parameters are measured with high accuracy of few %

- Normal hierarchy favoured at  $3.1 \sigma$
- Nonzero CP phase favoured
- Upper octant favoured

parameter	best fit $\pm 1\sigma$	$3\sigma$ range
$\Delta m_{21}^2$ [ $10^{-5} \text{eV}^2$ ]	$7.55^{+0.20}_{-0.16}$	7.05–8.14
$ \Delta m_{31}^2 $ [ $10^{-3} \text{eV}^2$ ] (NO)	$2.50 \pm 0.03$	2.41–2.60
$ \Delta m_{31}^2 $ [ $10^{-3} \text{eV}^2$ ] (IO)	$2.42^{+0.03}_{-0.04}$	2.31–2.51
$\sin^2 \theta_{12} / 10^{-1}$	$3.20^{+0.20}_{-0.16}$	2.73–3.79
$\sin^2 \theta_{23} / 10^{-1}$ (NO)	$5.47^{+0.20}_{-0.30}$	4.45–5.99
$\sin^2 \theta_{23} / 10^{-1}$ (IO)	$5.51^{+0.18}_{-0.30}$	4.53–5.98
$\sin^2 \theta_{13} / 10^{-2}$ (NO)	$2.160^{+0.083}_{-0.069}$	1.96–2.41
$\sin^2 \theta_{13} / 10^{-2}$ (IO)	$2.220^{+0.074}_{-0.076}$	1.99–2.44
$\delta / \pi$ (NO)	$1.32^{+0.21}_{-0.15}$	0.87–1.94
$\delta / \pi$ (IO)	$1.56^{+0.13}_{-0.15}$	1.12–1.94

## Is it just the SM or requires New physics?

### Three Types of Seesaw Mechanisms

Require the existence of new degrees of freedom (particles) beyond those present in the SM

Type I seesaw mechanism:  $\nu_{IR}$  - RH  $\nu$ s' (heavy).

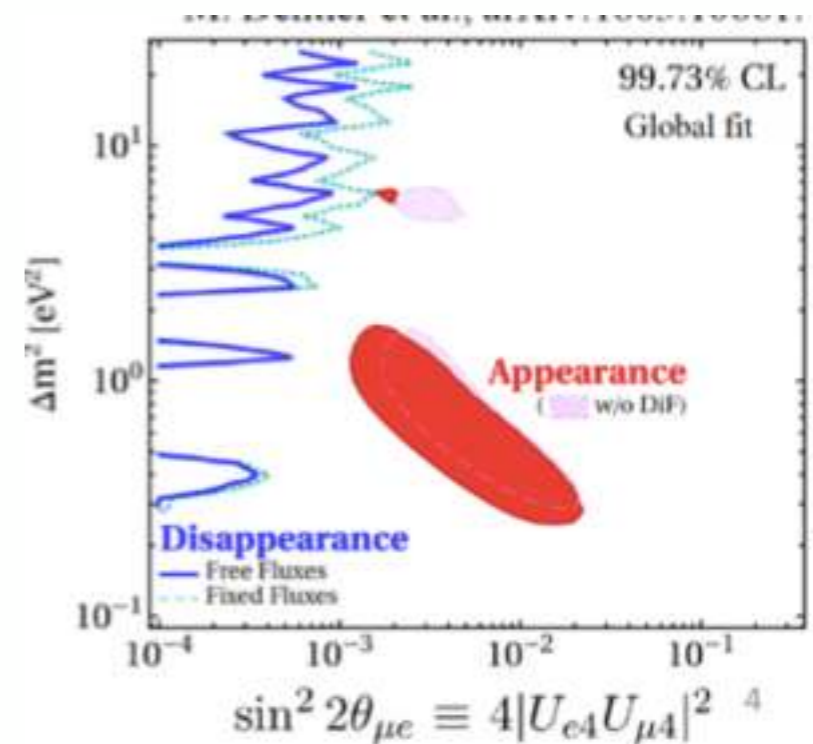
Type II seesaw mechanism:  $H(x)$  - a triplet of  $H^0, H^-, H^{--}$  Higgs fields.

Type III seesaw mechanism:  $T(x)$  - a triplet of fermion fields.

M. Weber ICHEP2018

### • Possible Sterile Neutrino?

- **New MiniBooNE consistent with LSND (but low energy excess?)**
- **Reactor anomaly questioned by Daya Bay/RENO time dependence**
- **New SBL and source experiments**
- **Conflict with  $\nu_\mu$  disappearance**





Major problem: 85% of matter is dark and remains invisible!

Is this compatible with the SM?

Does it requires modification of the SM or addition of gravity?

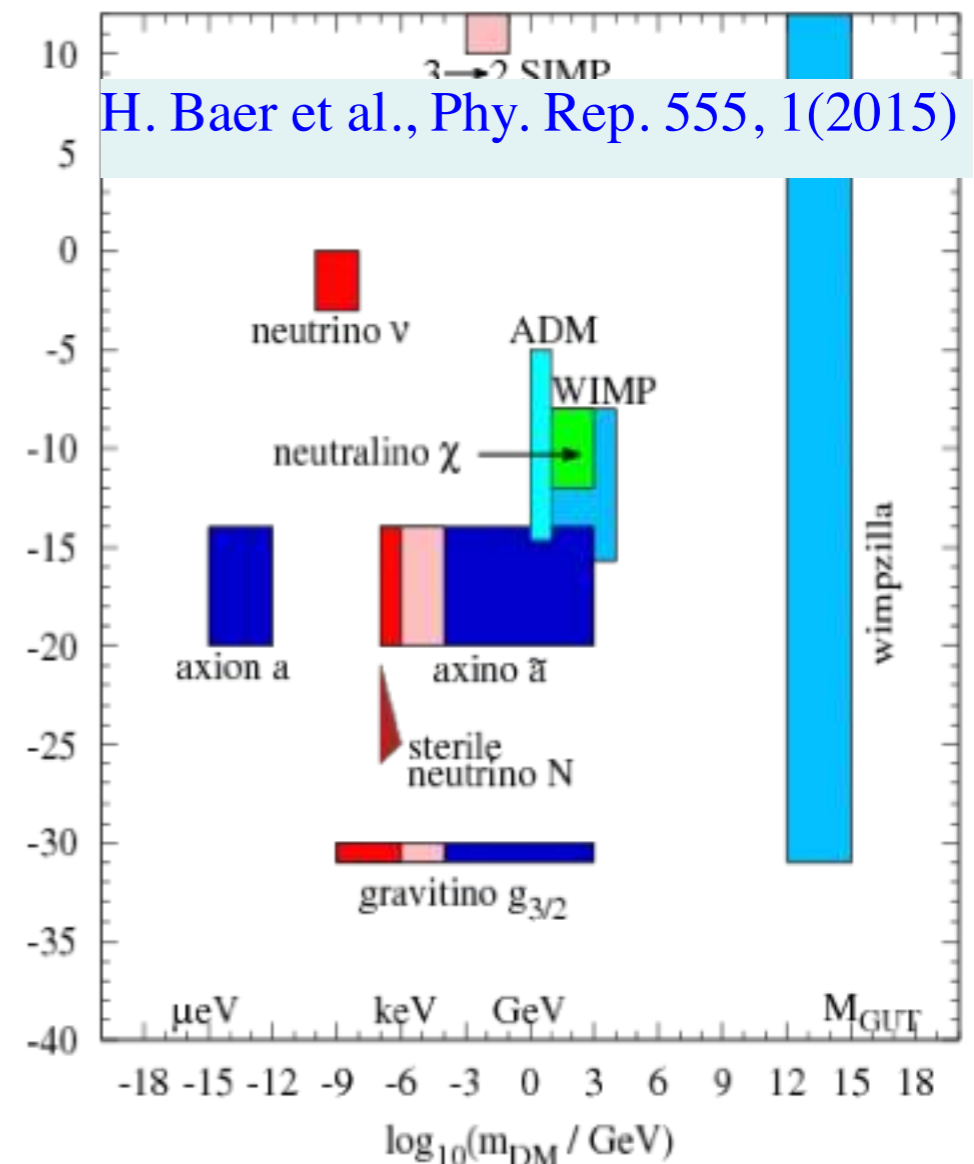
• Many candidates in many orders of magnitude of mass:

- **MOND** (Problems: large scales, Bullet cluster)
- **Primordial black holes** (LIGO, but constraints)
- **Fuzzy** (very light bosons)
- **Warm** (KeV sterile)
- **WIMP**
- **Axions/ALPs**
- **Dark sector**
- **Gravitinos**
- **Moduli**
- **Wimpzillas**



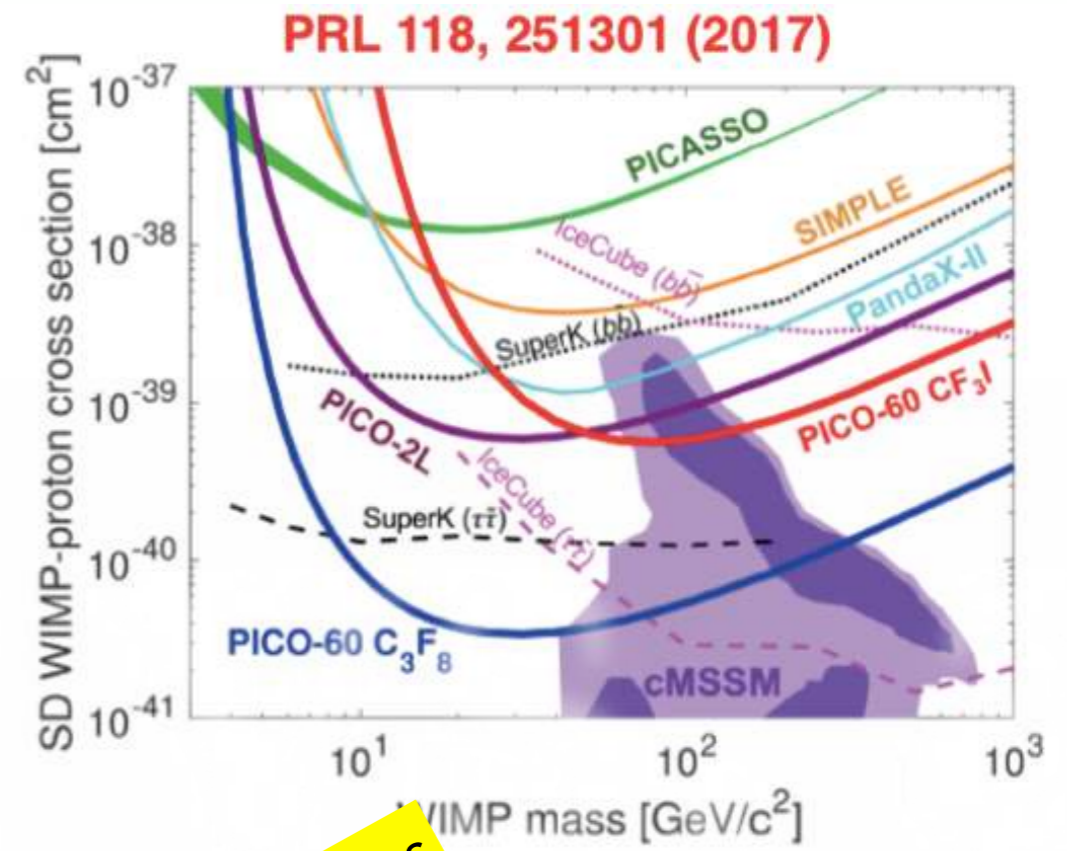
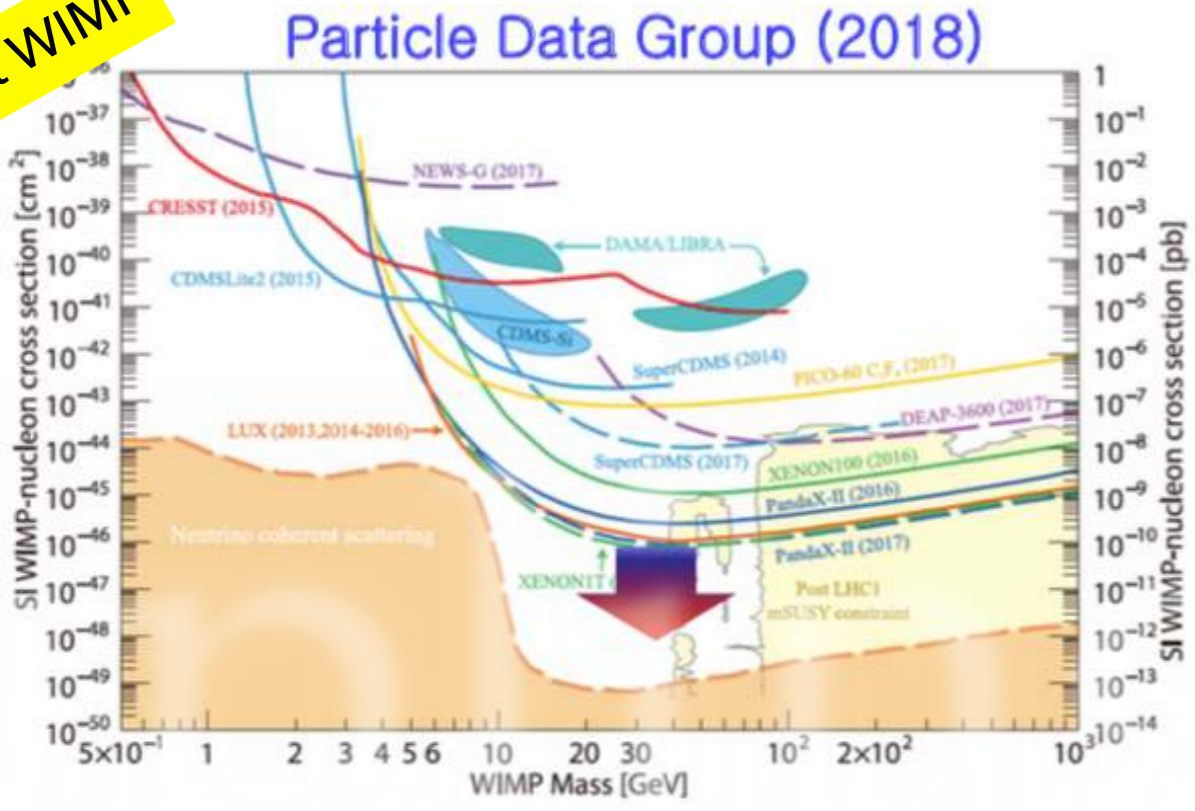
M. Drees

• Direct, indirect, collider

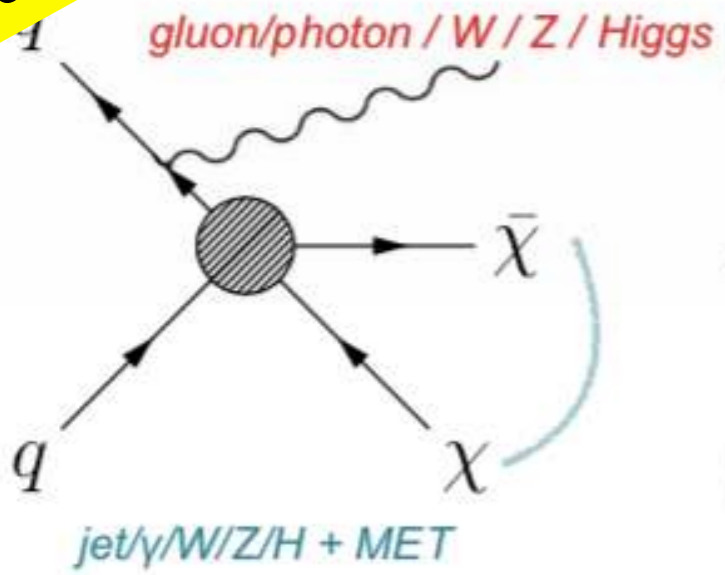


# BEYOND THE STANDARD MODEL: DARK MATTER SEARCHES

Direct WIMP



Colliders WIMP

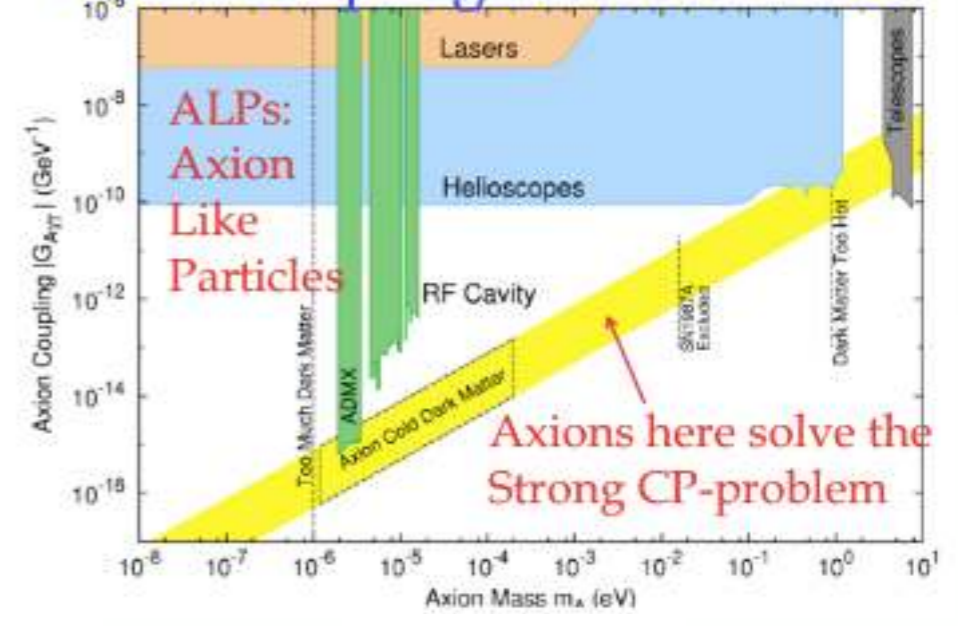


- **mono-jet**
  - most general signature, constraints on many models
- **mono-photon**
  - more challenging for background estimation
  - less powerful: EW vs. strong interaction
- **mono-W/Z leptonic**
  - clean signature and simple trigger
  - penalized by W/Z branching fraction
- **mono-W/Z hadronic**
  - larger statistics with larger background
- **tt+MET/bb+MET and mono-top**
  - more complicated experimentally
  - powerful in some scenarios
- **mono-Higgs**
  - powerful in some scenarios

D. del Re

Axion-likes

## Axion coupling vs. axion mass



Y. Semertzidis



# BEYOND THE STANDARD MODEL: THE MASS SPECTRUM AND MIXINGS

- Mass spectrum?

$$m_{quark} = y_{quark} \cdot v$$

$$m_{lepton} = y_{lepton} \cdot v$$

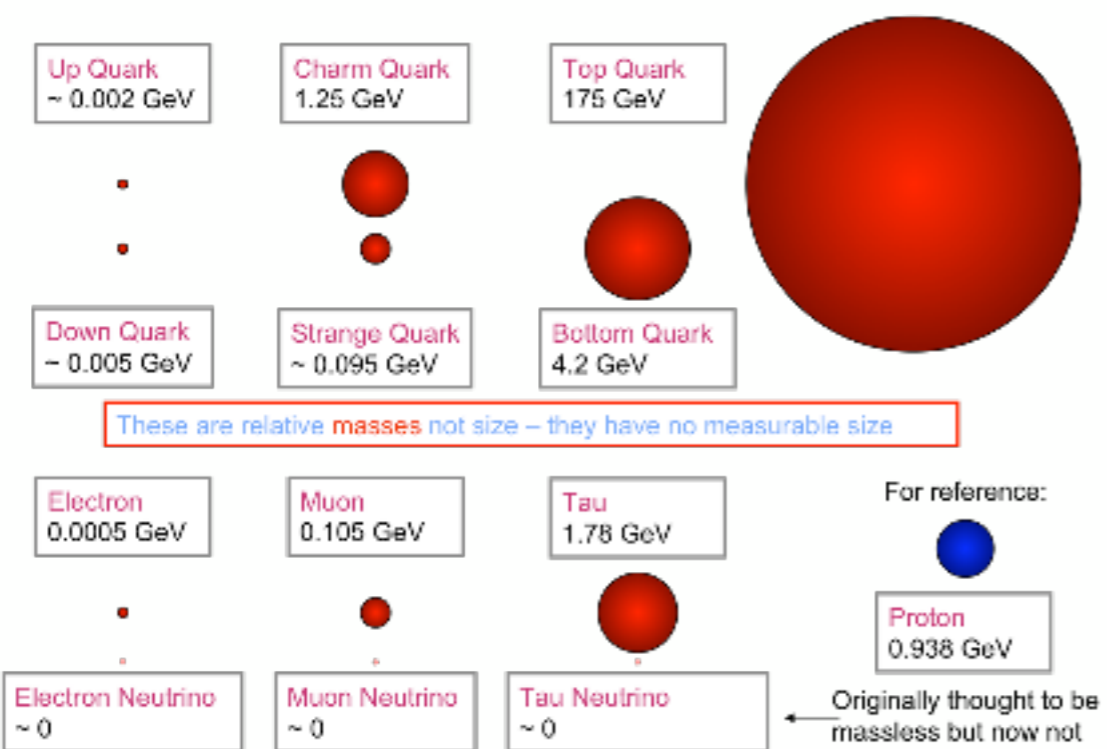
$$m_W = g/\sqrt{2} \cdot v$$

$$m_Z = \sqrt{g^2 + g'^2}/\sqrt{2} \cdot v$$

$$m_H = \sqrt{\lambda} \cdot v$$

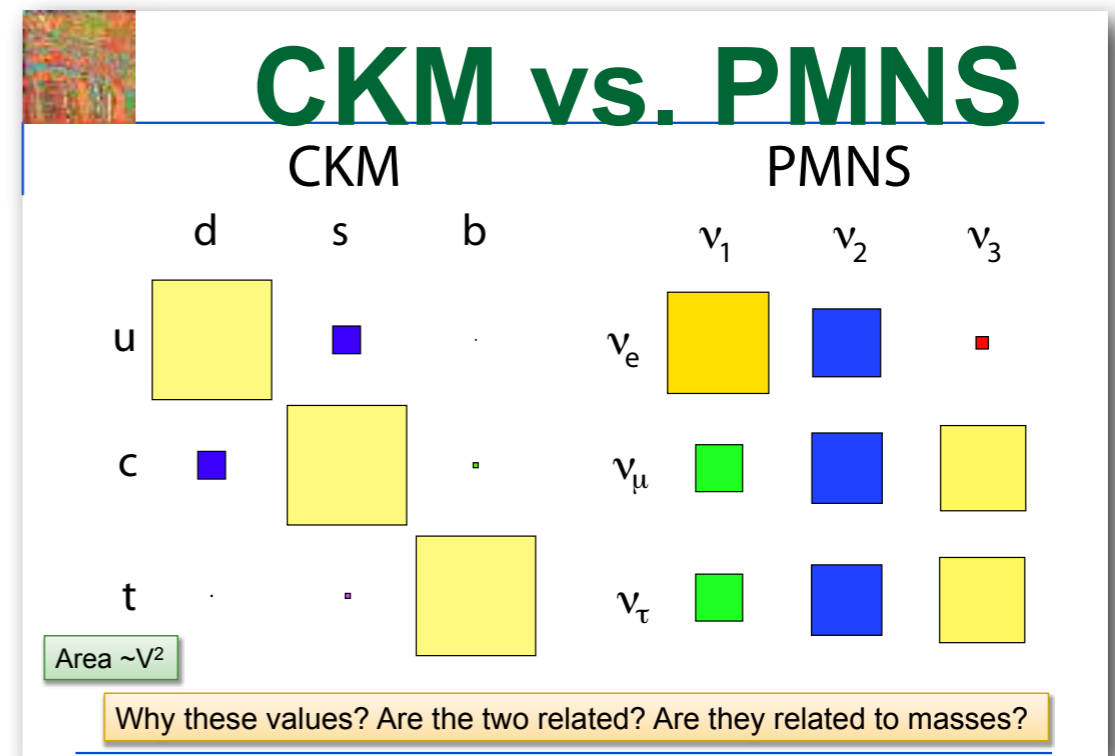
**SM**  $m_\gamma = 0$

$$m_{gluon} = 0$$



- Mixing Matrices?

- Quark-Lepton Symmetry
- Strong difference in parameters



- What are the CKM and PMNS phases?
- Where lies the source of CP violation: in quark or lepton sector?

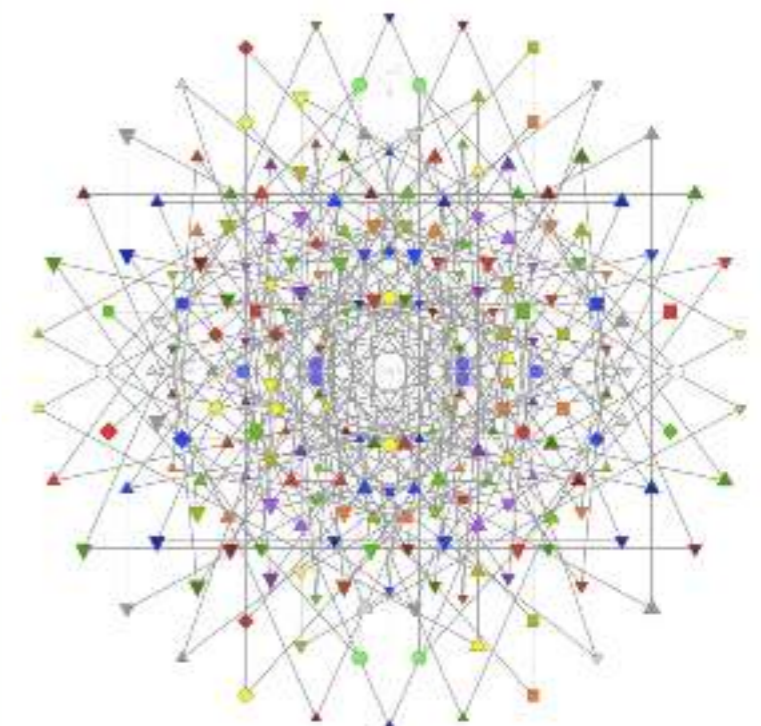
$$J_{CP} = \frac{1}{8} \sin 2\theta_{12} \sin 2\theta_{23} \sin 2\theta_{13} \cos \theta_{13} \sin \delta$$

## Neutrino Mixing: New Symmetry?

- $\theta_{12} = \theta_{\odot} \simeq \frac{\pi}{5.4}$ ,  $\theta_{23} = \theta_{\text{atm}} \simeq \frac{\pi}{4}(?)$ ,  $\theta_{13} \simeq \frac{\pi}{20}$

Follows the attempts in quark sector with 30 years delay: so far unsuccessful

$$U_{\text{TBM}} = \begin{pmatrix} \sqrt{\frac{2}{3}} & \sqrt{\frac{1}{3}} & 0 \\ -\sqrt{\frac{1}{6}} & \sqrt{\frac{1}{3}} & -\sqrt{\frac{1}{2}} \\ -\sqrt{\frac{1}{6}} & \sqrt{\frac{1}{3}} & \frac{1}{\sqrt{2}} \end{pmatrix}, \quad U_{\text{HGM}} = \begin{pmatrix} \frac{\sqrt{3}}{2} & \frac{1}{2} & 0 \\ -\frac{1}{2\sqrt{2}} & \frac{\sqrt{3}}{2\sqrt{2}} & -\frac{1}{\sqrt{2}} \\ -\frac{1}{2\sqrt{2}} & \frac{\sqrt{3}}{2\sqrt{2}} & \frac{1}{\sqrt{2}} \end{pmatrix}$$



Symmetry might be tricky

E8 roots



# THE STANDARD MODEL: CONCEPTUAL PROBLEMS

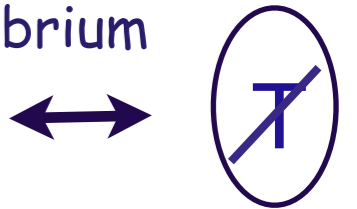
- Baryon asymmetry of the Universe

$$\frac{N(B) - N(\bar{B})}{N_\gamma} \sim (6.19 \pm 0.14) \times 10^{-10}$$

- still not explained
- three conditions (A.D.Sakharov)

1. Violation of a thermal equilibrium in early Universe

A possible scenario in the early Universe when particles drop from thermal equilibrium violations T invariance

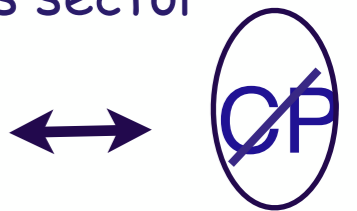


2. Violation of baryon number  $\longleftrightarrow$   $B = \frac{N_q - N_{\bar{q}}}{3}$

Baryon number is conserved in the SM with exponential accuracy

Violation of baryon number occurs in Grand Unified Theories and in Lepton=fourth color models (Pati-Salam model) } New particles = Leptoquarks, Extended Higgs sector

3. Violation of CP invariance (requires larger CP than in the SM)



In the SM achieved via phase factors in the CKM and PMNS mixing matrices

The presence of new phase factors in extended models (2HDM, SUSY, etc)

CPT is exact symmetry of Nature

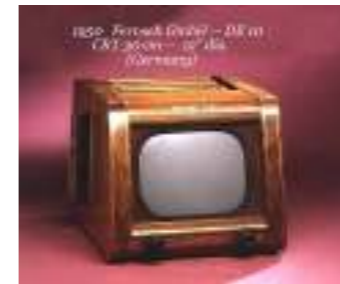
### WHAT MAKES US THINK THAT THERE IS PHYSICS BEYOND THE STANDARD MODEL?

- Small discrepancy with experimental data
- Possible new ingredients in neutrino sector (majorana neutrino)
- Instability of electroweak vacuum
- Inability to describe the Dark matter (unless it has pure gravitational nature)
- Baryon asymmetry of the Universe is a fundamental problem (Baryon and Lepton genesis might require new ingredients)
- Lack of understanding of flavor structure of the SM calls for explanation at higher level
- New era in gravity due to discovery of gravitational waves and black holes might change the landscape

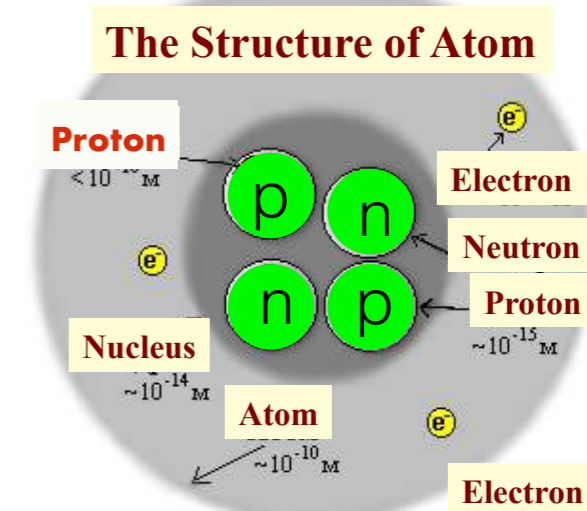
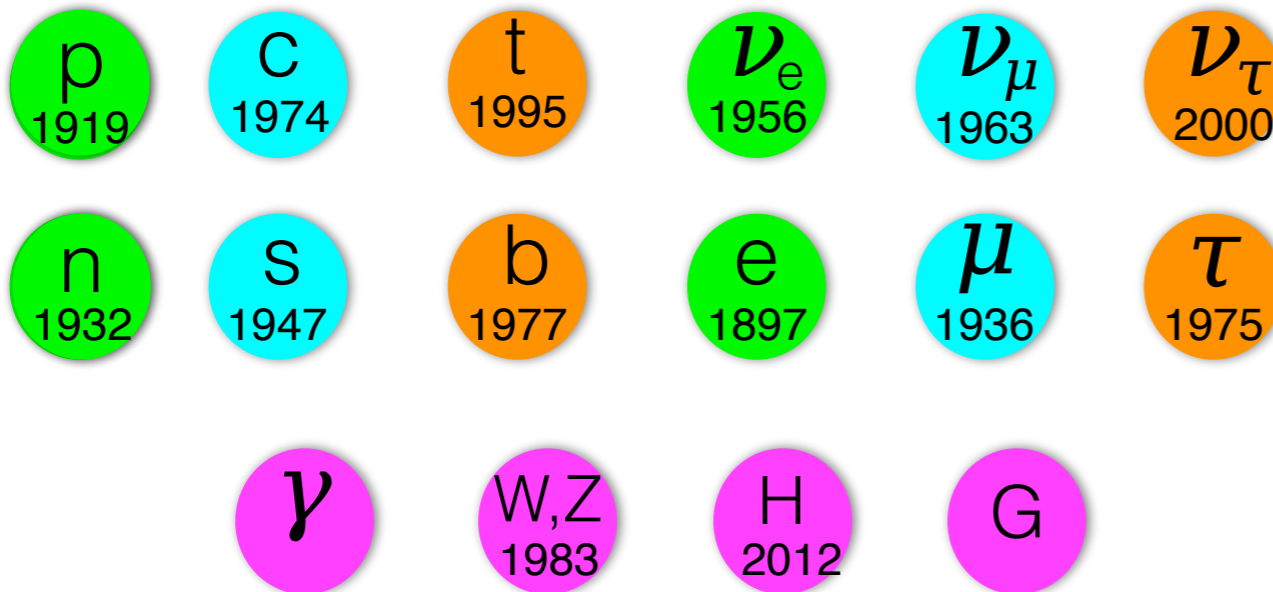


# Physics with a single generation

Back to the middle of the XX century



All the world around us is made of the 1st generation



- Muon - heavy electron - 2nd generation ?
- K-meson - strangeness ?
- Quark model (OK again?)
- GIM Mechanism, J/Psi - charm -2nd generation
- CP-violation: where it comes from?

- Who expected new physics to come?
- What scale of NP?

## Astrophysics & Cosmology challenge

- Baryon asymmetry of the Universe
- Description of the Dark Matter

## ***Ideas*** (conventional and not)

- **Symmetries**
  - Supersymmetry, family, ...
- **Compositeness**
  - Higgs, fermions, ...
- **Extra dimensions**
  - large, warped, ...
- **Dark or hidden sectors**
  - Dark, SUSY-breaking, random, ...
- **Unification**
  - GUT, string, ...
- **New dynamical ideas**
  - Relaxion, unnaturalness, clockwork, string instantons, ...
- **Random or environmental**
  - multiverse
- **String remnants**
  - (need not solve SM problem)
  - $Z'$ , vector fermions, extended Higgs, dark, moduli, axions, ...



# BEYOND THE STANDARD MODEL: CONCLUSIONS

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***How Will We Make Progress?***

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