

New Developments in QFT: D=4 Conformal Field Theories and AdS/CFT Correspondence

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N=4 Super Yang-Mills Theory

- $\mathcal{N} = 4$ Super Yang-Mills theory is the most supersymmetric theory possible without gravity
- Field content: 1 massless gauge boson, 4 massless (Majorana) spin 1/2 fermions, 6 real (or 3 complex) massless spin 0 bosons
All fields are in adjoint representation of the gauge group (Take $SU(N_c)$)
- The theory is exactly scale invariant, conformal field theory at quantum level, i.e. the β function identically vanishes at all orders of PT

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AdS/CFT Correspondence

- $N_c \rightarrow \infty$ (planar limit) is expected to be integrable and solvable
- Maldacena's conjecture: Planar Limit of N=4 SYM at strong coupling is dual to weakly coupled type II b supergravity in 10 dimensional $AdS_5 * S_5$ space.
- How might PT series be organized to produce simple strong coupling result?
- The amplitudes on shell possess IR singularities which should cancel in observables. What are the infrared-safe observables?

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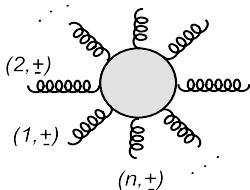
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Gluon scattering amplitudes



All outgoing gluons with helicity + or - on mass shell

In the leading N_c order (planar limit)

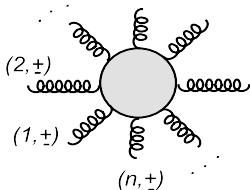
- Colour decomposition of amplitudes in N=4 SYM theory for $N_c \rightarrow \infty$

$$\mathcal{A}_n^{(l)} = g^{n-2} \left(\frac{g^2 N_c}{16\pi^2} \right)^l \sum_{perm} \text{Tr}(T^{a_{\sigma(1)}}, \dots, T^{a_{\sigma(n)}}) A_n^{(l)}(a_{\sigma(1)}, \dots, a_{\sigma(n)}),$$

where \mathcal{A}_n - physical amplitude, A_n - partial amplitude, a_i - is color index of i -th external "gluon"

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N=4 SYM Theory: Perturbation Theory

- Bern, Dixon & Smirnov's conjecture: $M_n^{(L)}(\varepsilon) \equiv A_n^{(L)}/A_n^{(0)}$

$$\mathcal{M}_n \equiv 1 + \sum_{L=1}^{\infty} \left(\frac{g^2 N_c}{16\pi^2} \right)^L M_n^{(L)}(\varepsilon) = \exp \left[\sum_{l=1}^{\infty} \left(\frac{g^2 N_c}{16\pi^2} \right)^l \left(f^{(l)}(\varepsilon) M_n^{(1)}(l\varepsilon) + C^{(l)} + E_n^{(l)}(\varepsilon) \right) \right]$$

$$f^{(l)}(\varepsilon) = f_0^{(l)}(\varepsilon) + \varepsilon f_1^{(l)}(\varepsilon) + \varepsilon^2 f_2^{(l)}(\varepsilon)$$

$$\mathcal{M}_n(\varepsilon) = \exp \left[-\frac{1}{8} \sum_{l=1}^{\infty} \left(\frac{g^2 N_c}{16\pi^2} \right)^l \left(\frac{\gamma_K^{(l)}}{(l\varepsilon)^2} + \frac{2G_0^{(l)}}{l\varepsilon} \right) \sum_{i=1}^n \left(\frac{\mu^2}{-s_{i,i+1}} \right)^{l\varepsilon} + \frac{1}{4} \sum_{l=1}^{\infty} \left(\frac{g^2 N_c}{16\pi^2} \right)^l \gamma_K^{(l)} F_n^{(1)}(0) \right]$$

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Cusp anomalous dimension

- **Cusp anomalous dimension** appears in RG eq. for the expectation value of a Wilson line with a cusp

Loop expansion $\gamma_K = \sum_{l=1}^{\infty} \left(\frac{g^2 N_c}{16\pi^2} \right)^l \gamma_K^{(l)}$

$$\gamma_K^{(1)} = 8, \quad \gamma_K^{(2)} = -16\zeta_2, \quad \gamma_K^{(3)} = 176\zeta_4, \dots$$

- It also controls the large spin limit of anomalous dimension of leading-twist operators

$$O_j \equiv \bar{q}(\gamma_+ \mathcal{D}^+)^j q$$

$$\gamma_j = \frac{1}{2} \gamma_K(\alpha) \log j + \mathcal{O}(j^0), \quad j \rightarrow \infty$$

- and large x limit of the DGLAP kernel for p.d.f.

$$P_{gg} = \frac{1}{2} \frac{\gamma_K(\alpha)}{(1-x)_+} + \dots, \quad x \rightarrow 1, \quad \gamma(j) = - \int_0^1 dx x^{j-1} P(x)$$

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Strong coupling expansion/AdS

- Classical solution (Alday & Maldacena) for the scattering amplitude

$$\mathcal{M}_4(\varepsilon) = \exp[-S_{cl}^E]$$

$$S_{cl}^E = \frac{1}{\varepsilon^2} \frac{\sqrt{g^2 N_c}}{\pi} \left[\left(\frac{\mu_{IR}^2}{-s} \right)^{\varepsilon/2} + \left(\frac{\mu_{IR}^2}{-t} \right)^{\varepsilon/2} \right]$$

$$+ \frac{1}{\varepsilon} \frac{\sqrt{g^2 N_c}}{2\pi} (1-\log 2) \left[\left(\frac{\mu_{IR}^2}{-s} \right)^{\varepsilon/2} + \left(\frac{\mu_{IR}^2}{-t} \right)^{\varepsilon/2} \right] - \frac{\sqrt{g^2 N_c}}{8\pi} \left[\log^2\left(\frac{s}{t}\right) + c \right] + \mathcal{O}(\varepsilon)$$

- $\gamma_K(g^2) \sim \frac{\sqrt{g^2 N_c}}{\pi}$, $G_0(g^2) \sim \sqrt{g^2 N_c} \frac{1-\log 2}{2\pi}$, for $g^2 N_c \rightarrow \infty$
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Violation of BDS ansatz

- For $n = 4, 5$ the BDS ansatz goes through all checks, namely the amplitudes were calculated up to three loops for four gluons and up to two loops for five gluons.
- However, starting from $n = 6$ it fails.
 - ▶ In the strong coupling calculation in the limit $n \rightarrow \infty$ discrepancy with the BDS formula was found.
 - ▶ Starting from $n = 6$ the Regge limit factorization of the amplitude in some physical regions fails. This was also shown by explicit two-loop calculation.
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Virtual Correction (MHV)

- Born Term

$$\left(\frac{d\sigma}{d\Omega}\right)_0^{-++} = \frac{\alpha^2 N_c^2}{8E^2} \frac{s^4 (s^2 + t^2 + u^2)}{s^2 t^2 u^2} \left(\frac{\mu^2}{s}\right)^\epsilon = \frac{\alpha^2 N_c^2}{E^2} \left(\frac{\mu^2}{s}\right)^\epsilon \frac{3 + c^2}{(1 - c^2)^2}$$

$$c \equiv \cos \theta$$

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- Virtual Correction

$$\begin{aligned} \left(\frac{d\sigma}{d\Omega}\right)_{virt}^{-+} &= \frac{\alpha^2 N_c^2}{8E^2} \left(\frac{\mu^2}{s}\right)^\epsilon \left\{ \frac{\alpha N_c}{2\pi} \frac{s^4}{s^2 t^2 u^2} \left[-\frac{8}{\epsilon^2} \left(\left(\frac{\mu^2}{-t}\right)^\epsilon + \left(\frac{\mu^2}{-u}\right)^\epsilon \right) s^2 \right. \right. \\ &\quad \left. \left. + \left(\left(\frac{\mu^2}{s}\right)^\epsilon + \left(\frac{\mu^2}{-t}\right)^\epsilon \right) u^2 + \left(\left(\frac{\mu^2}{s}\right)^\epsilon + \left(\frac{\mu^2}{-u}\right)^\epsilon \right) t^2 \right] \right\} \\ &\quad \left. + \frac{16}{3} \pi^2 (s^2+t^2+u^2) + 4(u^2 \log^2\left(\frac{-s}{t}\right) + t^2 \log^2\left(\frac{-s}{u}\right) + s^2 \log^2\left(\frac{t}{u}\right)) \right\} \\ &= \frac{\alpha^2 N_c^2}{E^2} \left(\frac{\mu^2}{s}\right)^{2\epsilon} \left\{ \frac{\alpha N_c}{2\pi} \left[-\frac{16}{\epsilon^2} \frac{3+c^2}{(1-c^2)^2} + \frac{4}{\epsilon} \left(\frac{5+2c+c^2}{(1-c^2)^2} \log\left(\frac{1-c}{2}\right) \right. \right. \right. \\ &\quad \left. \left. + (c \leftrightarrow -c) \right) + \frac{16(3+c^2)\pi^2}{3(1-c^2)^2} - \frac{16}{(1-c^2)^2} \log\left(\frac{1-c}{2}\right) \log\left(\frac{1+c}{2}\right) \right] \right\} \end{aligned}$$

Virtual Correction (MHV)

- Born Term

$$\left(\frac{d\sigma}{d\Omega}\right)_0^{--++} = \frac{\alpha^2 N_c^2}{8E^2} \frac{s^4(s^2+t^2+u^2)}{s^2 t^2 u^2} \left(\frac{\mu^2}{s}\right)^\epsilon = \frac{\alpha^2 N_c^2}{E^2} \left(\frac{\mu^2}{s}\right)^\epsilon \frac{3+c^2}{(1-c^2)^2}$$

$$c \equiv \cos \theta$$

- Virtual Correction

$$\begin{aligned} \left(\frac{d\sigma}{d\Omega}\right)_{virt}^{--++} &= \frac{\alpha^2 N_c^2}{8E^2} \left(\frac{\mu^2}{s}\right)^\epsilon \left\{ \frac{\alpha N_c}{2\pi} \frac{s^4}{s^2 t^2 u^2} \left[-\frac{8}{\epsilon^2} \left(\left(\frac{\mu^2}{-t}\right)^\epsilon + \left(\frac{\mu^2}{-u}\right)^\epsilon \right) s^2 \right. \right. \\ &\quad \left. \left. + \left(\left(\frac{\mu^2}{s}\right)^\epsilon + \left(\frac{\mu^2}{-t}\right)^\epsilon \right) u^2 + \left(\left(\frac{\mu^2}{s}\right)^\epsilon + \left(\frac{\mu^2}{-u}\right)^\epsilon \right) t^2 \right] \right. \\ &\quad \left. + \frac{16}{3} \pi^2 (s^2+t^2+u^2) + 4(u^2 \log^2\left(\frac{-s}{t}\right) + t^2 \log^2\left(\frac{-s}{u}\right) + s^2 \log^2\left(\frac{t}{u}\right)) \right\} \\ &= \frac{\alpha^2 N_c^2}{E^2} \left(\frac{\mu^2}{s}\right)^{2\epsilon} \left\{ \frac{\alpha N_c}{2\pi} \left[-\frac{16}{\epsilon^2} \frac{3+c^2}{(1-c^2)^2} + \frac{4}{\epsilon} \left(\frac{5+2c+c^2}{(1-c^2)^2} \log\left(\frac{1-c}{2}\right) \right. \right. \right. \\ &\quad \left. \left. + (c \leftrightarrow -c) \right) + \frac{16(3+c^2)\pi^2}{3(1-c^2)^2} - \frac{16}{(1-c^2)^2} \log\left(\frac{1-c}{2}\right) \log\left(\frac{1+c}{2}\right) \right] \right\} \end{aligned}$$

Real Emission

- MHV

$$\left(\frac{d\sigma}{d\Omega_{14}}\right)_{Born}^{(----++)} = \frac{\alpha^2 N_c^2}{E^2} \left(\frac{\mu^2}{s}\right)^{2\epsilon} \frac{\alpha N_c}{2\pi} \left\{ \frac{1}{\epsilon^2} \frac{8(3+c^2)}{(1-c^2)^2} \right. \\ \left. + \frac{1}{\epsilon} \left[\frac{2}{(1+c)^2} \log\left(\frac{1-c}{2}\right) + \frac{2}{(1-c)^2} \log\left(\frac{1+c}{2}\right) + \frac{16\delta(2\delta-3)}{(1-c^2)^2(1-\delta)^2} \right. \right. \\ \left. \left. + \frac{12(3+c^2)(\log(1-\delta) - \log(\delta))}{(1-c^2)^2} \right] + \text{Finite part} \right\}$$

- Anti MHV

$$\left(\frac{d\sigma}{d\Omega_{14}}\right)_{Born}^{(----+-)} = \frac{\alpha^2 N_c^2}{E^2} \left(\frac{\mu^2}{s}\right)^{2\epsilon} \frac{\alpha N_c}{2\pi} \left\{ \frac{1}{\epsilon^2} \frac{8(3+c^2)}{(1-c^2)^2} \right. \\ \left. + \frac{2}{\epsilon} \left[\frac{64}{3(1-c^2)^3} + \delta \frac{(6\delta^2 - 3\delta + 30)c^2 + (10\delta^2 - 57\delta + 66)}{3(1-c^2)^2} - \frac{6(c^2 + 3) \log \delta}{(1-c^2)^2} + \right. \right. \\ \left. \left(\frac{3c^2 - 24c + 85}{(1-c)(1+c)^3} \log \frac{1-c}{2} - \frac{4(c^2 - 6c + 21)}{(1-c)(1+c)^3} \log \frac{1+\delta - (1-\delta)c}{2} \right. \right. \\ \left. \left. + \frac{16\delta(5-c)}{(1-c^2)^2(1+\delta - (1-\delta)c)} - \frac{32(1-c)}{3(1+c)^3(1+\delta - (1-\delta)c)^3} + (c \leftrightarrow -c) \right] \right. \\ \left. + \text{Finite part} \right\}$$

Real Emission

- MHV

$$\left(\frac{d\sigma}{d\Omega_{14}}\right)_{Born}^{(----++)} = \frac{\alpha^2 N_c^2}{E^2} \left(\frac{\mu^2}{s}\right)^{2\epsilon} \frac{\alpha N_c}{2\pi} \left\{ \frac{1}{\epsilon^2} \frac{8(3+c^2)}{(1-c^2)^2} \right.$$

$$+ \frac{1}{\epsilon} \left[\frac{2}{(1+c)^2} \log\left(\frac{1-c}{2}\right) + \frac{2}{(1-c)^2} \log\left(\frac{1+c}{2}\right) + \frac{16\delta(2\delta-3)}{(1-c^2)^2(1-\delta)^2} \right.$$

$$\left. \left. + \frac{12(3+c^2)(\log(1-\delta) - \log(\delta))}{(1-c^2)^2} \right] + \text{Finite part} \right\}$$

- Anti MHV

$$\left(\frac{d\sigma}{d\Omega_{14}}\right)_{Born}^{(----++)} = \frac{\alpha^2 N_c^2}{E^2} \left(\frac{\mu^2}{s}\right)^{2\epsilon} \frac{\alpha N_c}{2\pi} \left\{ \frac{1}{\epsilon^2} \frac{8(3+c^2)}{(1-c^2)^2} \right.$$

$$+ \frac{2}{\epsilon} \left[\frac{64}{3(1-c^2)^3} + \delta \frac{(6\delta^2 - 3\delta + 30)c^2 + (10\delta^2 - 57\delta + 66)}{3(1-c^2)^2} - \frac{6(c^2+3)\log\delta}{(1-c^2)^2} + \right.$$

$$\left(\frac{3c^2 - 24c + 85}{(1-c)(1+c)^3} \log \frac{1-c}{2} - \frac{4(c^2 - 6c + 21)}{(1-c)(1+c)^3} \log \frac{1+\delta - (1-\delta)c}{2} \right.$$

$$\left. \left. + \frac{16\delta(5-c)}{(1-c^2)^2(1+\delta - (1-\delta)c)} - \frac{32(1-c)}{3(1+c)^3(1+\delta - (1-\delta)c)^3} + (c \leftrightarrow -c) \right] \right.$$

$$\left. + \text{Finite part} \right\}$$

Real Emission

- MHV

$$\left(\frac{d\sigma}{d\Omega_{14}}\right)_{Born}^{(---+++)} = \frac{\alpha^2 N_c^2}{E^2} \left(\frac{\mu^2}{s}\right)^{2\epsilon} \frac{\alpha N_c}{2\pi} \left\{ \frac{1}{\epsilon^2} \frac{8(3+c^2)}{(1-c^2)^2} \right.$$

$$+ \frac{1}{\epsilon} \left[\frac{2}{(1+c)^2} \log\left(\frac{1-c}{2}\right) + \frac{2}{(1-c)^2} \log\left(\frac{1+c}{2}\right) + \frac{16\delta(2\delta-3)}{(1-c^2)^2(1-\delta)^2} \right.$$

$$\left. \left. + \frac{12(3+c^2)(\log(1-\delta) - \log(\delta))}{(1-c^2)^2} \right] + \text{Finite part} \right\}$$

- Anti MHV

$$\left(\frac{d\sigma}{d\Omega_{14}}\right)_{Born}^{(---+-)} = \frac{\alpha^2 N_c^2}{E^2} \left(\frac{\mu^2}{s}\right)^{2\epsilon} \frac{\alpha N_c}{2\pi} \left\{ \frac{1}{\epsilon^2} \frac{8(3+c^2)}{(1-c^2)^2} \right.$$

$$+ \frac{2}{\epsilon} \left[\frac{64}{3(1-c^2)^3} + \delta \frac{(6\delta^2 - 3\delta + 30)c^2 + (10\delta^2 - 57\delta + 66)}{3(1-c^2)^2} - \frac{6(c^2+3)\log\delta}{(1-c^2)^2} + \right.$$

$$\left(\frac{3c^2 - 24c + 85}{(1-c)(1+c)^3} \log \frac{1-c}{2} - \frac{4(c^2 - 6c + 21)}{(1-c)(1+c)^3} \log \frac{1+\delta - (1-\delta)c}{2} \right.$$

$$\left. \left. + \frac{16\delta(5-c)}{(1-c^2)^2(1+\delta - (1-\delta)c)} - \frac{32(1-c)}{3(1+c)^3(1+\delta - (1-\delta)c)^3} + (c \leftrightarrow -c) \right] \right\}$$

$$+ \text{Finite part}$$

Real Emission (Matter)

● Fermions

$$\left(\frac{d\sigma}{d\Omega_{14}}\right)_{Real}^{(--+\bar{q}q)} = \frac{\alpha^2 N_c^2}{E^2} \left(\frac{\mu^2}{s}\right)^{2\epsilon} \frac{\alpha N_c}{2\pi} \left\{ \frac{4}{\epsilon} \left[\frac{32(79 - 25c^2)}{3(1 - c^2)^2} \right. \right. \\ \left. \left. + \frac{64(3 - c)^2}{(1 - c)(1 + c)^3} \log\left(\frac{1 - c}{2}\right) + \frac{64(3 + c)^2}{(1 - c)^3(1 + c)} \log\left(\frac{1 + c}{2}\right) \right] + \text{Finite part} \right\}$$

● Sfermions

$$\left(\frac{d\sigma}{d\Omega_{14}}\right)_{Real}^{(--+\bar{q}\bar{q})} = \frac{\alpha^2 N_c^2}{E^2} \left(\frac{\mu^2}{s}\right)^{2\epsilon} \frac{\alpha N_c}{2\pi} \left\{ \frac{1}{\epsilon} \left[-\frac{128(10 + 7c^2)}{(1 - c^2)^2} \right. \right. \\ \left. \left. - \frac{192(5 - c)}{(1 + c)^3} \log\left(\frac{1 - c}{2}\right) - \frac{192(5 + c)}{(1 - c)^3} \log\left(\frac{1 + c}{2}\right) \right] + \text{Finite part} \right\}$$

Real Emission (Matter)

- Fermions

$$\left(\frac{d\sigma}{d\Omega_{14}}\right)_{Real}^{(--+\bar{q}q)} = \frac{\alpha^2 N_c^2}{E^2} \left(\frac{\mu^2}{s}\right)^{2\epsilon} \frac{\alpha N_c}{2\pi} \left\{ \frac{4}{\epsilon} \left[\frac{32(79 - 25c^2)}{3(1 - c^2)^2} \right. \right. \\ \left. \left. + \frac{64(3 - c)^2}{(1 - c)(1 + c)^3} \log\left(\frac{1 - c}{2}\right) + \frac{64(3 + c)^2}{(1 - c)^3(1 + c)} \log\left(\frac{1 + c}{2}\right) \right] + \text{Finite part} \right\}$$

- Sfermions

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Real Emission (Matter)

- Fermions

$$\left(\frac{d\sigma}{d\Omega_{14}}\right)_{Real}^{(--+\bar{q}q)} = \frac{\alpha^2 N_c^2}{E^2} \left(\frac{\mu^2}{s}\right)^{2\epsilon} \frac{\alpha N_c}{2\pi} \left\{ \frac{4}{\epsilon} \left[\frac{32(79 - 25c^2)}{3(1 - c^2)^2} \right. \right. \\ \left. \left. + \frac{64(3 - c)^2}{(1 - c)(1 + c)^3} \log\left(\frac{1 - c}{2}\right) + \frac{64(3 + c)^2}{(1 - c)^3(1 + c)} \log\left(\frac{1 + c}{2}\right) \right] + \text{Finite part} \right\}$$

- Sfermions

$$\left(\frac{d\sigma}{d\Omega_{14}}\right)_{Real}^{(--+\bar{q}\bar{q})} = \frac{\alpha^2 N_c^2}{E^2} \left(\frac{\mu^2}{s}\right)^{2\epsilon} \frac{\alpha N_c}{2\pi} \left\{ \frac{1}{\epsilon} \left[-\frac{128(10 + 7c^2)}{(1 - c^2)^2} \right. \right. \\ \left. \left. - \frac{192(5 - c)}{(1 + c)^3} \log\left(\frac{1 - c}{2}\right) - \frac{192(5 + c)}{(1 - c)^3} \log\left(\frac{1 + c}{2}\right) \right] + \text{Finite part} \right\}$$

Splitting of Massless states

- There are NO isolated massless asymptotic states!
- Massless particle can split into two (or more) indistinguishable collinear particles.
- One has to consider coherent states of parallel massless particles.

$$d\sigma_{2 \rightarrow 2}^{spl, init} = \frac{\alpha}{2\pi} \frac{1}{\epsilon} \left(\frac{\mu^2}{Q_f^2} \right)^\epsilon \int_0^1 dz P_{gg}(z) \sum_{i,j=1,2; i \neq j} d\sigma_{2 \rightarrow 2}(z p_i, p_j, p_3, p_4) S_2^{spl, init}(z)$$

$$d\sigma_{2 \rightarrow 2}^{spl, fin} = \frac{\alpha}{2\pi} \frac{1}{\epsilon} \left(\frac{\mu^2}{Q_f^2} \right)^\epsilon d\sigma_{2 \rightarrow 2}(p_1, p_2, p_3, p_4) \int_0^1 dz \sum_{l=g,q,\Lambda} P_{gl}(z) S_2^{spl, fin}(z)$$

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Initial and final state splitting (MHV)

- Initial

$$\left(\frac{d\sigma}{d\Omega_{14}} \right)_{InSplit}^{(--++++)} = \frac{\alpha^2 N_c^2}{E^2} \left(\frac{\mu^2}{s} \right)^\epsilon \left(\frac{\mu^2}{Q_f^2} \right)^\epsilon$$

$$\frac{\alpha N_c}{2\pi} \left\{ \frac{1}{\epsilon} \left[-\frac{4(c^2+3)}{(1-c^2)^2} \left(\log \frac{1-c}{2} + \log \frac{1+c}{2} \right) - \frac{8(c^2+3)}{(1-c^2)^2} \log \frac{1-\delta}{\delta} \right. \right.$$

$$\left. \left. - \frac{16\delta(2\delta-3)}{(1-\delta^2)(1-c^2)^2} \right] + \text{Finite part} \right\}$$

- Final

$$\left(\frac{d\sigma}{d\Omega_{14}} \right)_{FnSplit}^{(--++++)} = \frac{\alpha^2 N_c^2}{E^2} \left(\frac{\mu^2}{s} \right)^\epsilon \left(\frac{\mu^2}{Q_f^2} \right)^\epsilon$$

$$\frac{\alpha N_c}{2\pi} \left\{ -\frac{1}{\epsilon} \frac{4(c^2+3)}{(1-c^2)^2} \log \frac{1-\delta}{\delta} + \text{Finite part} \right\}$$

Initial and final state splitting (MHV)

- Initial

$$\left(\frac{d\sigma}{d\Omega_{14}} \right)_{InSplit}^{(--++++)} = \frac{\alpha^2 N_c^2}{E^2} \left(\frac{\mu^2}{s} \right)^\epsilon \left(\frac{\mu^2}{Q_f^2} \right)^\epsilon$$

$$\frac{\alpha N_c}{2\pi} \left\{ \frac{1}{\epsilon} \left[-\frac{4(c^2 + 3)}{(1 - c^2)^2} \left(\log \frac{1 - c}{2} + \log \frac{1 + c}{2} \right) - \frac{8(c^2 + 3)}{(1 - c^2)^2} \log \frac{1 - \delta}{\delta} \right. \right.$$

$$\left. \left. - \frac{16\delta(2\delta - 3)}{(1 - \delta^2)(1 - c^2)^2} \right] + \text{Finite part} \right\}$$

- Final

$$\left(\frac{d\sigma}{d\Omega_{14}} \right)_{FnSplit}^{(--++++)} = \frac{\alpha^2 N_c^2}{E^2} \left(\frac{\mu^2}{s} \right)^\epsilon \left(\frac{\mu^2}{Q_f^2} \right)^\epsilon$$

$$\frac{\alpha N_c}{2\pi} \left\{ -\frac{1}{\epsilon} \frac{4(c^2 + 3)}{(1 - c^2)^2} \log \frac{1 - \delta}{\delta} + \text{Finite part} \right\}$$

Initial and final state splitting (MHV)

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$$\left(\frac{d\sigma}{d\Omega_{14}} \right)_{InSplit}^{(--++++)} = \frac{\alpha^2 N_c^2}{E^2} \left(\frac{\mu^2}{s} \right)^\epsilon \left(\frac{\mu^2}{Q_f^2} \right)^\epsilon$$

$$\frac{\alpha N_c}{2\pi} \left\{ \frac{1}{\epsilon} \left[-\frac{4(c^2 + 3)}{(1 - c^2)^2} \left(\log \frac{1 - c}{2} + \log \frac{1 + c}{2} \right) - \frac{8(c^2 + 3)}{(1 - c^2)^2} \log \frac{1 - \delta}{\delta} \right. \right.$$

$$\left. \left. - \frac{16\delta(2\delta - 3)}{(1 - \delta^2)(1 - c^2)^2} \right] + \text{Finite part} \right\}$$

- Final

$$\left(\frac{d\sigma}{d\Omega_{14}} \right)_{FnSplit}^{(---++)} = \frac{\alpha^2 N_c^2}{E^2} \left(\frac{\mu^2}{s} \right)^\epsilon \left(\frac{\mu^2}{Q_f^2} \right)^\epsilon$$

$$\frac{\alpha N_c}{2\pi} \left\{ -\frac{1}{\epsilon} \frac{4(c^2 + 3)}{(1 - c^2)^2} \log \frac{1 - \delta}{\delta} + \text{Finite part} \right\}$$

Infrared-free sets

$$A^{MHV} = \frac{1}{2} \left(\frac{d\sigma}{d\Omega_{14}} \right)_{Virt}^{(----)} + \left(\frac{d\sigma}{d\Omega_{14}} \right)_{Real}^{(----)} + \left(\frac{d\sigma}{d\Omega_{14}} \right)_{InSplit}^{(----)} + \left(\frac{d\sigma}{d\Omega_{14}} \right)_{FnSplit}^{(----)}$$

$$B^{AntiMHV} = \frac{1}{2} \left(\frac{d\sigma}{d\Omega_{14}} \right)_{Virt}^{(----)} + \left(\frac{d\sigma}{d\Omega_{14}} \right)_{Real}^{(----)} + \left(\frac{d\sigma}{d\Omega_{14}} \right)_{InSplit}^{(----)} + \left(\frac{d\sigma}{d\Omega_{14}} \right)_{FnSplit}^{(----)}$$

$$C^{Matter} = \left(\frac{d\sigma}{d\Omega_{14}} \right)_{Real}^{(---+\bar{q}q, \bar{q}\bar{q})} + \left(\frac{d\sigma}{d\Omega_{14}} \right)_{InSplit}^{(---+\bar{q}q, \bar{q}\bar{q})} + \left(\frac{d\sigma}{d\Omega_{14}} \right)_{FnSplit}^{(---+\bar{q}q, \bar{q}\bar{q})}$$

Infrared-free sets

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$$B^{AntiMHV} = \frac{1}{2} \left(\frac{d\sigma}{d\Omega_{14}} \right)_{Virt}^{(----)} + \left(\frac{d\sigma}{d\Omega_{14}} \right)_{Real}^{(----)} + \left(\frac{d\sigma}{d\Omega_{14}} \right)_{InSplit}^{(----)} + \left(\frac{d\sigma}{d\Omega_{14}} \right)_{FnSplit}^{(----)}$$

$$C^{Matter} = \left(\frac{d\sigma}{d\Omega_{14}} \right)_{Real}^{(---+\bar{q}q, \bar{q}\bar{q})} + \left(\frac{d\sigma}{d\Omega_{14}} \right)_{InSplit}^{(---+\bar{q}q, \bar{q}\bar{q})} + \left(\frac{d\sigma}{d\Omega_{14}} \right)_{FnSplit}^{(---+\bar{q}q, \bar{q}\bar{q})}$$

Infrared-free sets

- $$A^{MHV} = \frac{1}{2} \left(\frac{d\sigma}{d\Omega_{14}} \right)_{Virt}^{(----)} + \left(\frac{d\sigma}{d\Omega_{14}} \right)_{Real}^{(----)} + \left(\frac{d\sigma}{d\Omega_{14}} \right)_{InSplit}^{(----)} + \left(\frac{d\sigma}{d\Omega_{14}} \right)_{FnSplit}^{(----)}$$

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Infrared-free sets

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Infrared-free observables

- Registration of two fastest gluons of positive chirality

$$A^{MHV} \Big|_{\delta=1/3} + B^{AntiMHV} \Big|_{\delta=1}$$

- Registration of one fastest gluon of positive chirality

$$A^{MHV} \Big|_{\delta=1/3} + B^{AntiMHV} \Big|_{\delta=1/3} + C^{Matter} \Big|_{\delta=1}$$

- Anti MHV cross-section

$$B^{AntiMHV} \Big|_{\delta=1} + C^{Matter} \Big|_{\delta=1} \Rightarrow \text{Finite Part}$$

Infrared-free observables

- Registration of two fastest gluons of positive chirality

$$A^{MHV} \Big|_{\delta=1/3} + B^{AntiMHV} \Big|_{\delta=1}$$

- Registration of one fastest gluon of positive chirality

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$$B^{AntiMHV} \Big|_{\delta=1} + C^{Matter} \Big|_{\delta=1} \Rightarrow \text{Finite Part}$$

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The simplest IR finite answer so far ($Q_f = E$): **N=4 SYM Anti MHV**

L.Bork, D.K., G.Vartanov, A.Zhiboedov

$$\left(\frac{d\sigma}{d\Omega_{14}} \right)_{AntiMHV} = \frac{\alpha^2 N_c^2}{E^2} \left\{ \frac{3 + c^2}{(1 - c^2)^2} - \frac{\alpha N_c}{2\pi} \left[2 \frac{(c^4 + 2c^3 + 4c^2 + 6c + 19) \log^2\left(\frac{1-c}{2}\right)}{(1-c)^2(1+c)^4} + 2 \frac{(c^4 - 2c^3 + 4c^2 - 6c + 19) \log^2\left(\frac{1+c}{2}\right)}{(1-c)^4(1+c)^2} - 8 \frac{(c^2 + 1) \log\left(\frac{1+c}{2}\right) \log\left(\frac{1-c}{2}\right)}{(1-c^2)^2} - \frac{6\pi^2(c^2 - 1) + 5(61c^2 + 99)}{9(1-c^2)^2} + 2 \frac{(11c^3 + 31c^2 - 47c + 133) \log\left(\frac{1+c}{2}\right)}{3(1-c)^3(1+c)^2} - 2 \frac{(11c^3 - 31c^2 - 47c - 133) \log\left(\frac{1-c}{2}\right)}{3(1+c)^3(1-c)^2} \right] \right\}$$

Summary For Part I

- In observable cross-sections the IR divergences do cancel in accordance with Kinoshita-Lee-Nauenberg theorem

$$d\sigma_{obs}^{incl} = \sum_{n=2}^{\infty} \int_0^1 dz_1 q_1(z_1, \frac{Q_f^2}{\mu^2}) \int_0^1 dz_2 q_2(z_2, \frac{Q_f^2}{\mu^2}) \prod_{i=1}^n \int_0^1 dx_i q_i(x_i, \frac{Q_f^2}{\mu^2}) \times$$

$$\times d\sigma^{2 \rightarrow n}(z_1 p_1, z_2 p_2, \dots) S_n(\{z\}, \{x\}) = g^4 \sum_{L=0}^{\infty} \left(\frac{g^2}{16\pi^2} \right)^L d\sigma_L^{Finite}(s, t, u, Q_f^2)$$

- The simple structure of the MHV amplitude DOES NOT reveal at the level of IR finite cross-sections;
- In some cases the cancellation of complicated functions occurs, though not always;
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UV Divergences in Gravity

Power counting in gravity and supergravity in D spacetime dimensions

$$\omega = (D - 2)L + 2 \quad D = 4 \rightarrow \omega = 2L + 2$$

Restrict analysis to ON-SHELL amplitudes only!

- Pure Gravity in D=4: Finite in 1 loop
 $R_{ij}^2, R^2, \rightarrow 0, R_{ijkl}^2 \rightarrow$ Gauss – Bonnet term $\rightarrow 0$
- Gravity with matter in D=4: Divergence in 1 loop
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UV Divergences in Supergravity

- Local supersymmetry implies that the pure curvature part of such a D=4, 3-loop divergent structure must be built from the square of the Bel-Robinson tensor

$$\int \sqrt{-g} T_{\mu\nu\rho\sigma} T^{\mu\nu\rho\sigma}, \quad T_{\mu\nu\rho\sigma} = R_{\mu\nu}^{\alpha\beta} R_{\rho\alpha\sigma\beta} + {}^* R_{\mu\nu}^{\alpha\beta} {}^* R_{\rho\alpha\sigma\beta}$$

- This is directly related to the α'^3 corrections in the superstring effective action, except that in the string context such contributions occur with finite coefficients.
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Non-Renormalization Theorems

- Powerful SUSY "non-renormalization theorems" exclude infinite renormalization within D=4, N=1 supersymmetry of chiral invariants, given in N=1 superspace by integrals over half the superspace

$$\int d^2\theta W(\Phi(x, \theta, \bar{\theta})), \quad \bar{D}\Phi = 0$$

The counterterms have to be written as full superspace integrals $\int d^{4M}\theta$ and be local.

- So, in D=4, N=1 supersymmetry, full superspace integrals like (or "D terms") are allowed, but chiral integrals like (or "F terms") are not.

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The First Divergent Loop Orders

- The strength of a given supersymmetric non-renormalization theorem depends on the extent of linearly realizable, or "off-shell" supersymmetry.
- Assuming that 1/2 supersymmetry is linearly realizable and requiring gauge and supersymmetry invariances, predictions were derived for the first divergent loop orders in maximal ($N = 4 < - > 16$ supercharges) SYM and ($N = 8 < - > 32$ sc.) SUGRA

P.Howe,K.Stelle,P.Townsend

Max. SYM first divergences,
assuming half SUSY off-shell
(8 supercharges)

Dimension D	10	8	7	6	5	4
Loop order L	1	1	2	3	4	∞
Gen. form	$\partial^2 F^4$	F^4	$\partial^2 F^4$	$\partial^2 F^4$	F^4	finite

Max. SUGRA first divergences,
assuming half SUSY off-shell
(16 supercharges)

Dimension D	11	10	8	7	6	5	4
Loop order L	2	2	1	2	3	2	3
Gen. form	$\partial^{12} R^4$	$\partial^{10} R^4$	R^4	$\partial^6 R^4$	$\partial^6 R^4$	R^4	R^4

Unitarity-based calculations

Bern, Carrasco, Dixon, Dunbar, Johansson,

Kosower, Perelstein, Roiban, Rozowsky et al.

- Within the last decade, there have been significant advances in the computation of loop corrections in quantum field theory.
- These developments include the organization of amplitudes into a new kind of perturbation theory starting with maximal helicity violating amplitudes (MHV), then next-to-MHV (NMHV), etc.
- They also incorporate a specific use of dimensional regularization together with a clever use of unitarity cutting rules.

Cut Construction Procedure

- Normally, one thinks of unitarity relations such as the optical theorem as giving information only about the imaginary parts of amplitudes. However, if one keeps all orders in an expansion in $\varepsilon = 4 - D$, then loop integrals like require integrands to have an additional momentum dependence

$$f(s) \rightarrow f(s)s^{-\varepsilon/2}$$

- Then, since

$$s^{-\varepsilon/2} = 1 - \varepsilon/2 \log(s)$$

and

$$\log(s) = \log(|s|) + i\pi\Theta(s),$$

one can learn about the real parts of an amplitude by retaining imaginary terms at order ε . This gives rise to a procedure for the *cut construction* of higher-loop diagrams.

Kawai-Lewellen-Tai relation

- For maximal supergravity, another relation allowing amplitudes to be evaluated is the Kawai-Lewellen-Tai (KLT) relation between open- and closed-string amplitudes. This gives rise to tree-level relations between field-theoretic max. SUGRA and max. SYM field-theory amplitudes, e.g.

$$M_4^{tree}(1, 2, 3, 4) = -is_{12} A_4^{tree}(1, 2, 3, 4) A_4^{tree}(1, 2, 4, 3)$$

- Combining this with unitarity-based calculations, in which all amplitudes are ultimately reduced to integrals over products of tree amplitudes, one has a way to obtain higher-loop supergravity amplitudes from SYM amplitudes.

On-shell Simplification

- Another remarkable aspect of the unitarity-based methods is the simplification of vertices. The off-shell 3- graviton vertex has the form (with about 100 terms)

$$\begin{aligned}
 G_{3,\mu\alpha,\nu\beta,\sigma\gamma}(k_1, k_2, k_3) = & \kappa \text{Sym}[\\
 & -\frac{1}{2}P_3(k_1 \cdot k_2 \eta_{\mu\alpha} \eta_{\nu\beta} \eta_{\sigma\gamma}) - \frac{1}{2}P_6(k_{1\nu} k_{1\beta} \eta_{\mu\alpha} \eta_{\sigma\gamma}) + \frac{1}{2}P_3(k_1 \cdot k_2 \eta_{\mu\nu} \eta_{\alpha\beta} \eta_{\sigma\gamma}) \\
 & + P_6(k_1 \cdot k_2 \eta_{\mu\alpha} \eta_{\nu\sigma} \eta_{\beta\gamma}) + 2P_3(k_{1\nu} k_{2\gamma} \eta_{\mu\alpha} \eta_{\beta\sigma}) - P_3(k_{1\beta} k_{2\mu} \eta_{\alpha\nu} \eta_{\sigma\gamma}) \\
 & + P_3(k_{1\sigma} k_{2\gamma} \eta_{\mu\nu} \eta_{\alpha\beta}) + P_6(k_{1\sigma} k_{1\gamma} \eta_{\mu\nu} \eta_{\alpha\beta}) + 2P_6(k_{1\nu} k_{2\gamma} \eta_{\beta\mu} \eta_{\alpha\sigma}) \\
 & + 2P_3(k_{1\nu} k_{2\mu} \eta_{\beta\sigma} \eta_{\gamma\alpha}) - 2P_3(k_1 \cdot k_2 \eta_{\alpha\nu} \eta_{\beta\sigma} \eta_{\gamma\mu})]
 \end{aligned}$$

- Putting this vertex on-shell for tree diagrams simplifies it to

$$i\kappa(\eta_{\mu\nu}(k_1 - k_2)_\rho + \text{cyclic}) \times (\eta_{\alpha\beta}(k_1 - k_2)_\gamma + \text{cyclic})$$

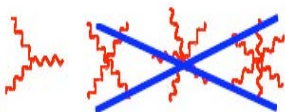
which is just the square of the colour-stripped version of the SYM amplitude,

$$-gf^{abc}(\eta_{\mu\nu}(k_1 - k_2)_\rho + \text{cyclic})$$

in agreement with KLT.

Simplifies Feynman Rules

- More remarkable still is the fact that on-shell tree amplitudes can be built exclusively using 3-point vertices: contributions from the infinite numbers of higher-point vertices all cancel out:



- This is also a reflection of a simplification in SYM, where the 4-point vertex does not contribute to on-shell tree amplitude calculations:



- Using these on-shell simplifications inside general tree diagrams requires also a technique of making complex shifts of external momenta.

New expectations for UV divergences

- In this way, a different set of anticipated first loop orders for ultraviolet divergences arose using the unitarity-based approach:

Max. SYM first divergences,
early unitarity-based
predictions

Dimension D	10	8	7	6	5	4
Loop order L	1	1	2	3	6	∞
Gen. form	$\partial^2 F^4$	F^4	$\partial^2 F^4$	$\partial^2 F^4$	$\partial^2 F^4$	finite

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Gen. form	$\partial^{12} R^4$	$\partial^{10} R^4$	R^4	$\partial^6 R^4$	$\partial^6 R^4$	$\partial^6 R^4$	$\partial^4 R^4$

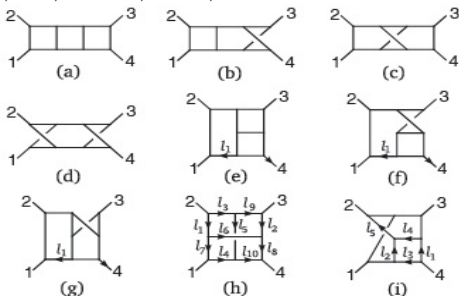
- These anticipations were based on iterated 2-particle cuts, however. Full calculations can reveal different behavior, however.

Explicit Calculations

- An important development is the subsequent completion of the 3-loop calculation:

Bern, Carrasco, Dixon, Johansson, Kosower, Roiban.

Normal Feynman
diagram calculation
of these would
involve about 10^{20}
terms

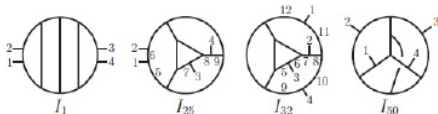


The result is finite at $L=3$ in $D=4$.

Explicit Calculations II

- Meanwhile, the 4-loop calculation has now been done (May 2009).

Bern, Carrasco, Dixon, Johansson & Roiban



+ 46 more topologies

The result

$$M_4^{4-loop} = \left(\frac{\kappa}{2}\right)^{10} stu M_4^{tree} \sum_{S_4} \sum_{i=1}^{50} c_i I_i$$

is ultraviolet finite in D=4 (as expected) and in D=5 (unexpected).

IR Divergences in QG

L.Bork, D.K., G.Vartanov, A.Zhiboedov

- For the MHV amplitudes

$$A_{Gravity}^{MHV} = \frac{1}{2} \left(\frac{d\sigma}{d\Omega_{14}} \right)_{Virt}^{(---++)} + \left(\frac{d\sigma}{d\Omega_{14}} \right)_{Real}^{(---++)} + \left(\frac{d\sigma}{d\Omega_{14}} \right)_{InSplit}^{(---++)} + \left(\frac{d\sigma}{d\Omega_{14}} \right)_{FnSplit}^{(---++)}$$

All the IR and Collinear divergences CANCEL like in SYM theory!

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- There is new hope for the finiteness of maximally supersymmetric Quantum gravity

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Conclusion

- It may well be that maximally supersymmetric N=4 SYM theory and N=8 supergravity are finite and integrable.
- One may hope to obtain exact solution of the models in all orders of PT and beyond
- Any possible solution seems to be related by duality to a proper string model in AdS space.