



BEYOND THE STANDARD MODEL' 06

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ICHEP'06, Moscow, August 1, 2006

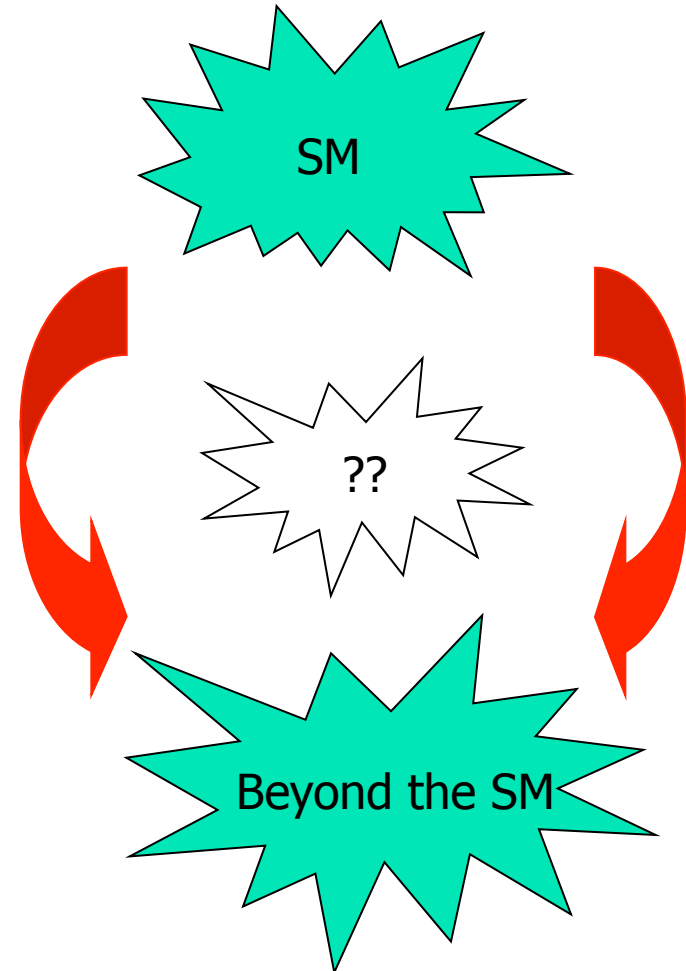
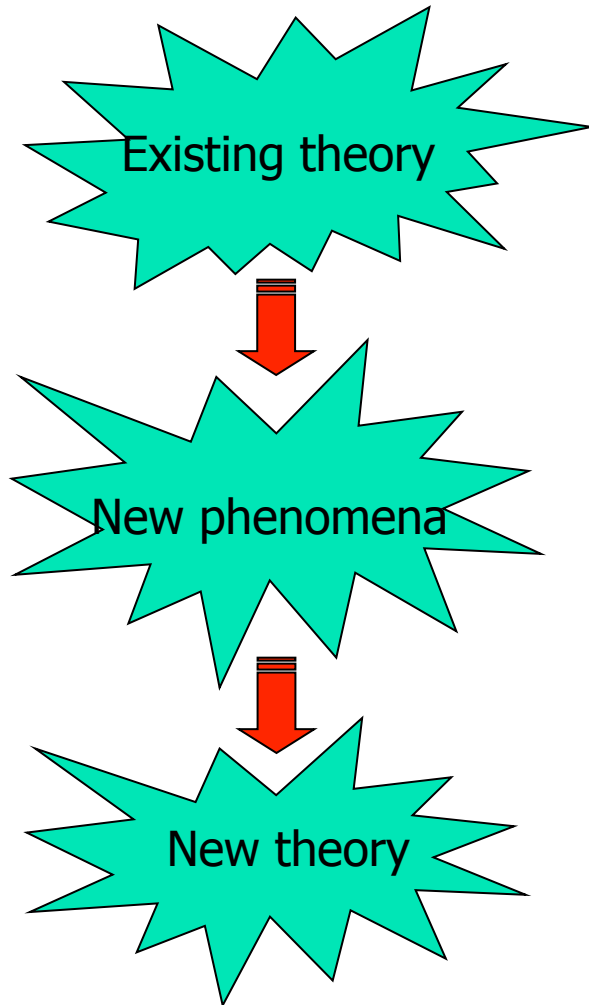
The Standard Model of Fundamental Interactions



Something one may be proud of

HEP Paradox

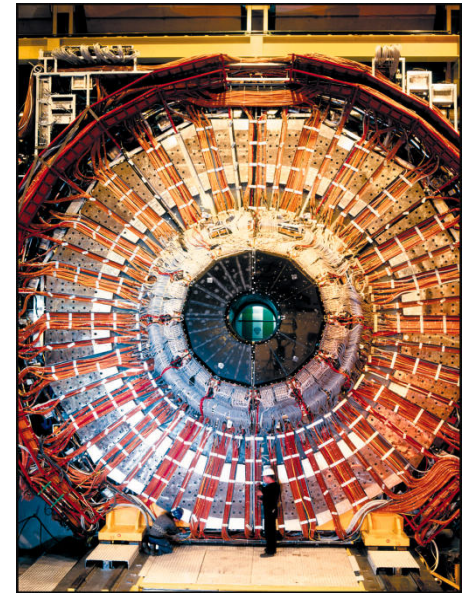
HEP Paradox



HEP Paradox

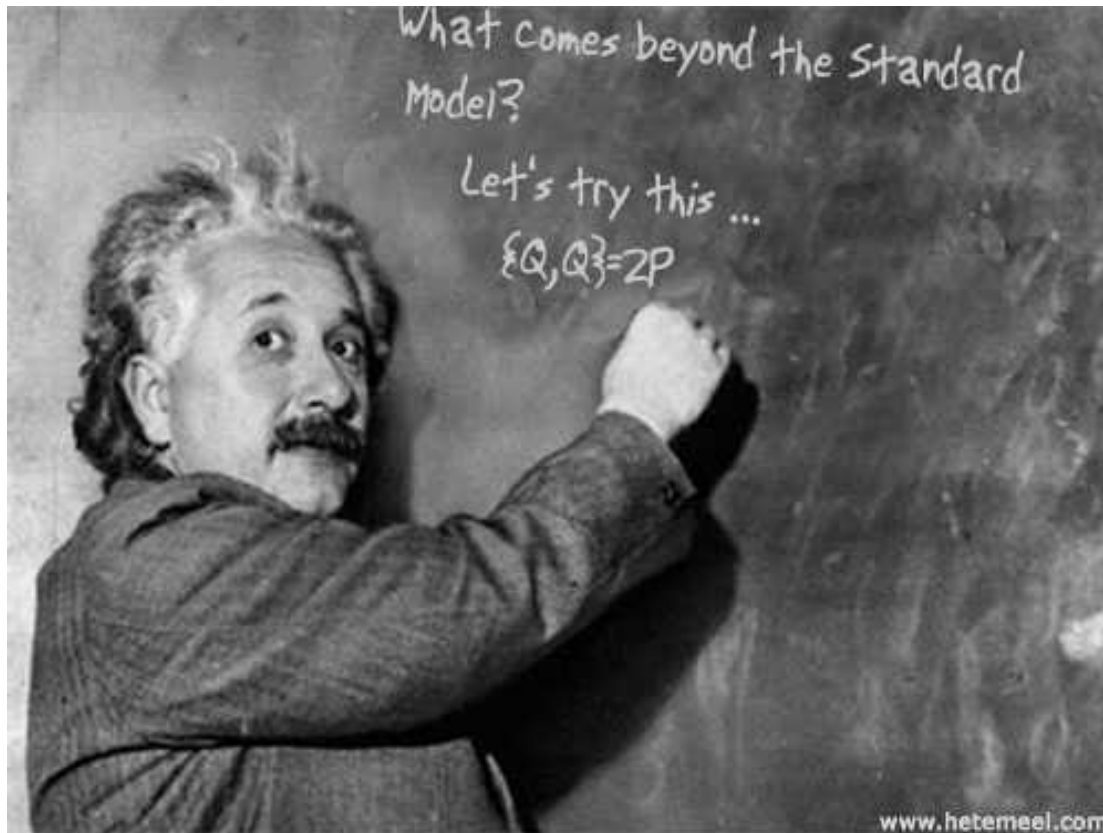
Physics beyond the SM

- Low Energy Supersymmetry
- Extra gauge bosons
- Axions
- Extra dimensions
- Deviation from Unitarity triangle
- Modification of Newton law
- Free quarks
- New forces / particles
- Violation of Baryon number
- Violation of Lepton number
- Monopoles
- Violation of Lorentz invariance
- Compositeness

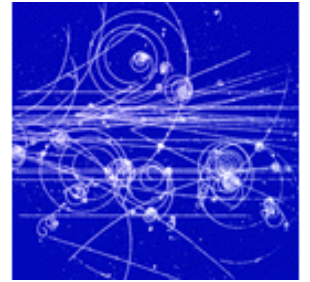


Not found so far ...

Physics beyond the SM

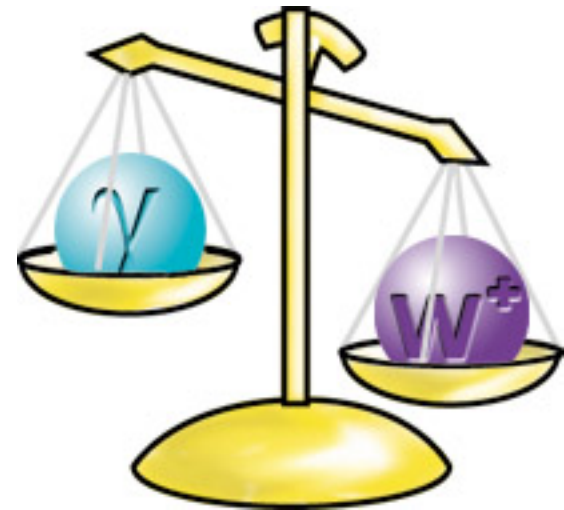


Problem # 1



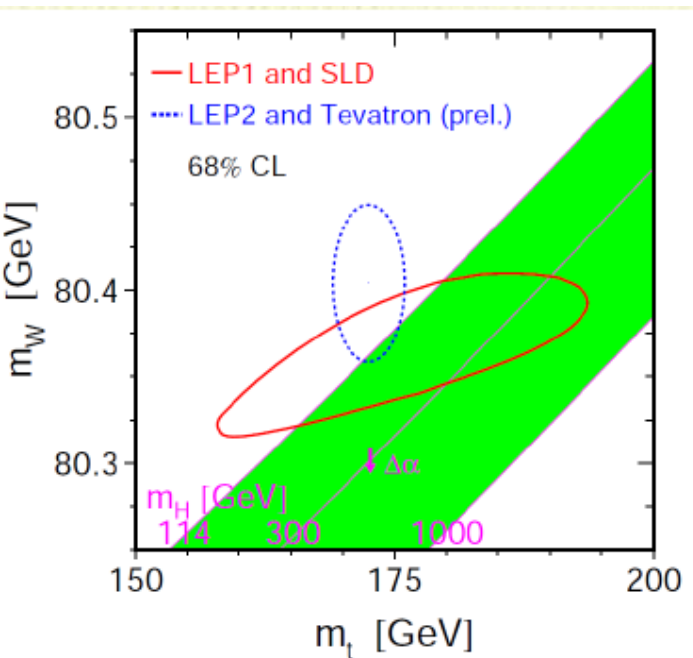
Mechanism of Electroweak Symmetry Breaking:

- The Higgs mechanism
- Alternatives



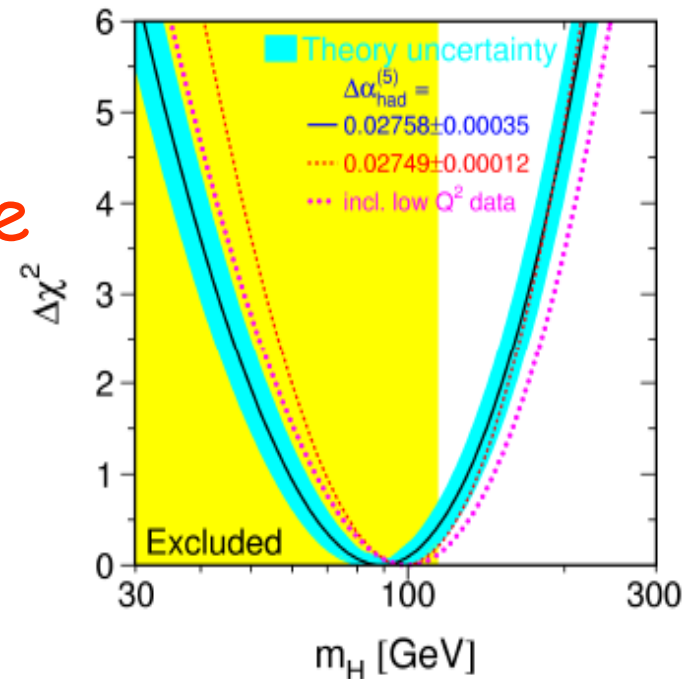
The SM Higgs Boson

- Indirect limit from radiative corrections
- Direct limit from Higgs non observation at LEP II (CERN)



If it is there we may see it soon

χ^2 versus M_H for SM Fit
 $\pm M_H = 89 + 42 - 30$ @68%CL
 $\pm M_H < 175$ GeV @95%CL
 for $m_{\text{top}} = 172.5$ GeV

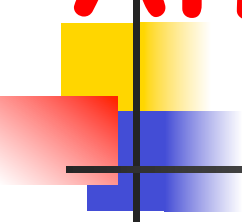


Alternatives to SM Higgs



- Two-Higgs Doublet Models
- Inert Higgs Model
- Little Higgs Models
- Twin Higgs Model
- Gauge-Higgs Unification Models
- Higgsless Models

Alternatives to SM Higgs



■ Two-Higgs Doublet Model

Hundreds of papers ..

- More Higgses -> more freedom in parameters
- Multiphase case -> new source of CP violation
- Making the lightest Higgs almost sterile
(very weak coupling to $Z \sim \sin(\alpha-\beta)$) allows Higgs boson mass below 100 GeV

■ Inert Higgs Model

R.Barbieri, L.Hall,V.Rychkov

- Inert Higgs doublet that has neither a vev nor couplings to quarks and leptons,
- It is the lightest particle and may compose the Dark Matter
- The usual Higgs boson is heavy (> 400 GeV)
- Does not contradict the precision EW tests



Alternatives to HM

■ Little Higgs Model

Arkani-Hamed, Cohen and Georgi

- Higgs bosons as pseudo-goldstones \rightarrow quadratically divergent one-loop contribution to the Higgs mass is not generated \rightarrow Log hierarchy
- Needs larger gauge group and “collective” symmetry breaking
- New heavy states with masses around 1 TeV
- LHC signatures are similar to SUSY albeit different angular dependence due to spin structure
- T-parity (similar to R-parity) \rightarrow stable particle \rightarrow DM

Perelstein

Birkedal, Noble, Spray

Martin

■ Twin Higgs Model

Chacko, Goh, Hornik, Su

- Similar to Little Higgs, Higgses are pseudo-goldstones
- Has discrete symmetry – twin symmetry (mirror or L-R)
- SUSY generalization

Chang, Hall, Weiner;

Fialkowski, Pokorski, Schmaltz



Alternatives to HM

■ Gauge-Higgs Unification Models

- Higgs is a gauge scalar of XD
- Discrete symmetry protects Higgs mass
- Orbifold compactification to get chiral matter and fundamental rep Higgs
- Leads to KK excitations of W and Z at 500 GeV – 1 TeV and extra scalars

Manton; Fairlie;

Candelas, Horowitz, Strominger, Witten;
Hosotani; Csaki, Grojean, Murayama, ...

■ Higgsless Models

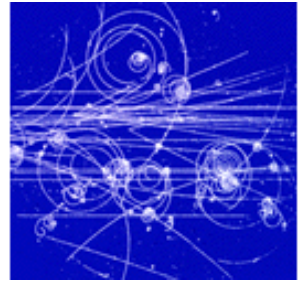
- Extra dimensions/Non-flat geometry (warped)
- EWSB from boundary conditions \rightarrow W & Z are KK states \rightarrow unitarity is preserved
- No scalars at low energies
- Fermions in a bulk, mass at IR vector-like brane
- New (KK) spin1 states, slightly different couplings to SM gauge bosons
- Not to contradict precision EW tests needs heavy scale
- All these theories are effective (non-renormalizable)

Csaki, Grojean, Murayama;

Cacciapaglia, Lillie, Terning, Hubisz, Meade,
Shirman, Barbieri, Pomarol, Rattazzi,
Hebecker, March-Russell, ...



Problem # 2



New physics at the TeV scale:

- Supersymmetry
- Extra Dimensions
- Unknown ?

TeV Scale SUSY

- MSSM (gravity mediation)
- MSSM (gauge mediation)
- MSSM (anomaly med, etc)
- MSSM (non-universality)
- NMSSM (Singlet extensions)
- R-parity violation

SUSY'06
Review talks

M.Nojiri

H.Baer

V.Barger

A.Dedes

ICHEP'06

T.Hebbeker
K.Klein

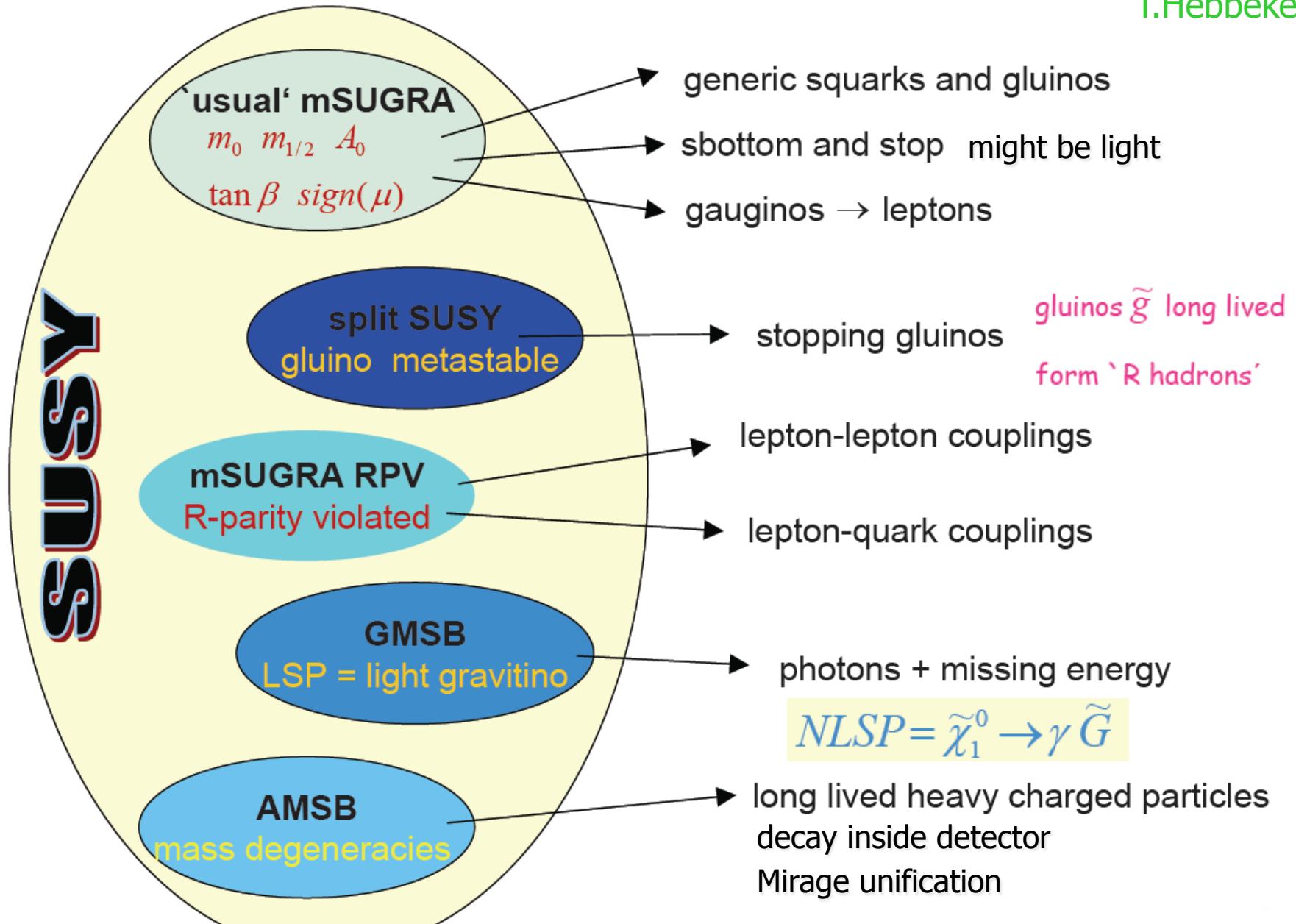
T.Krupovnickas

R.Godbole

T.Hebbeker

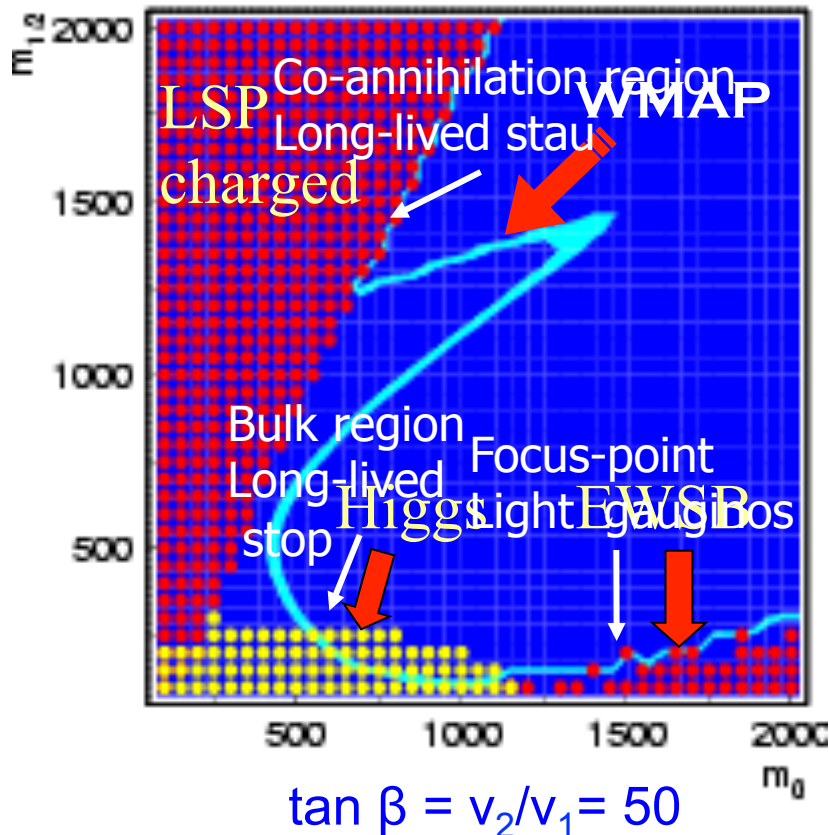
SUSY Models and Signatures

T.Hebbeker



Allowed regions after WMAP (mSUGRA)

m_0 common spin 0 mass
 $m_{1/2}$ common spin $1/2$ mass

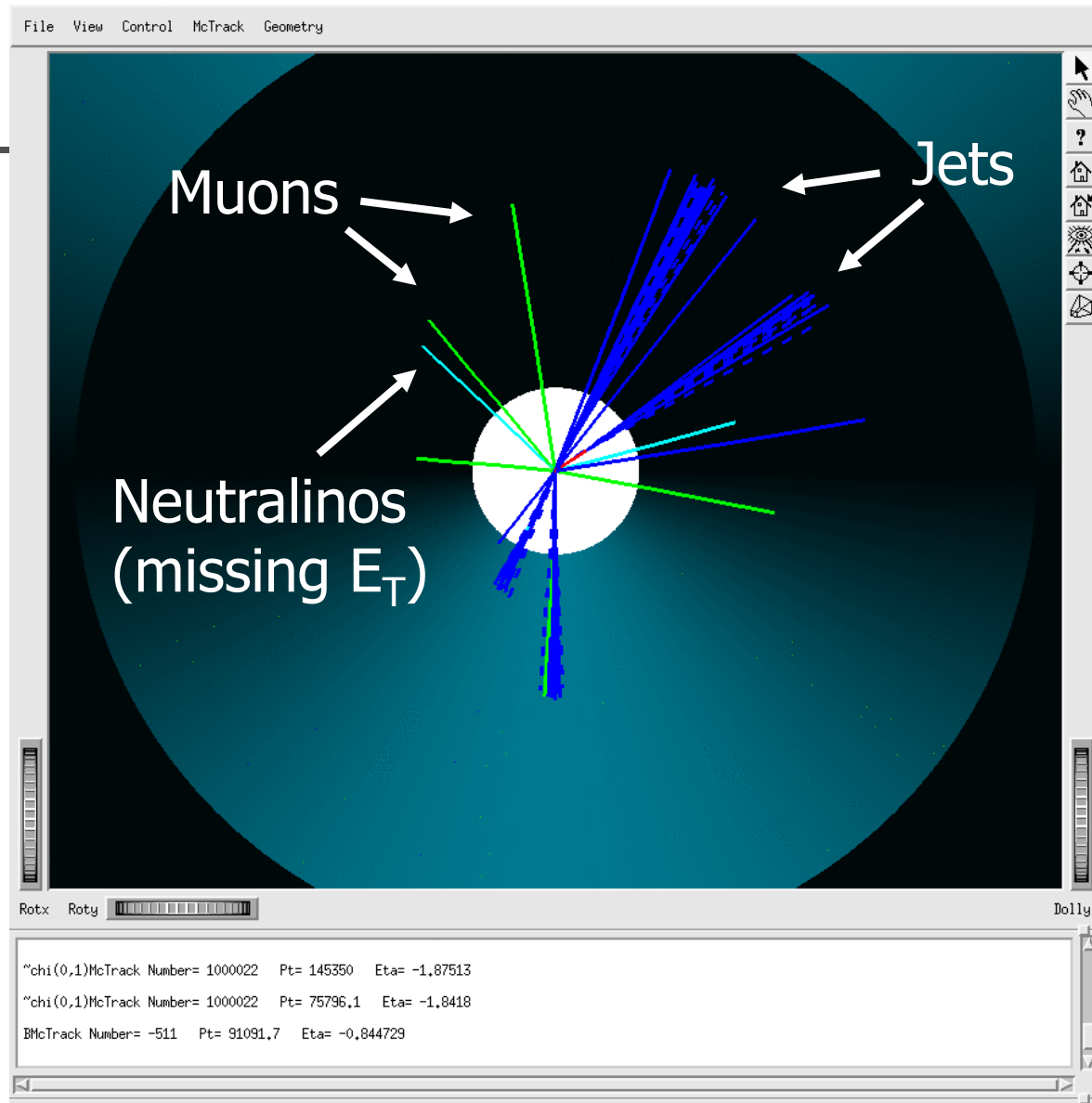


In allowed region one fulfills all the constraints simultaneously and has the suitable amount of the dark matter

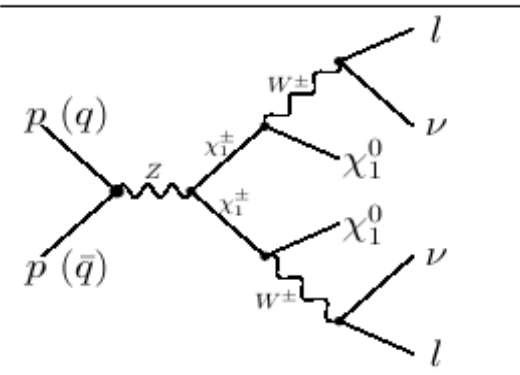
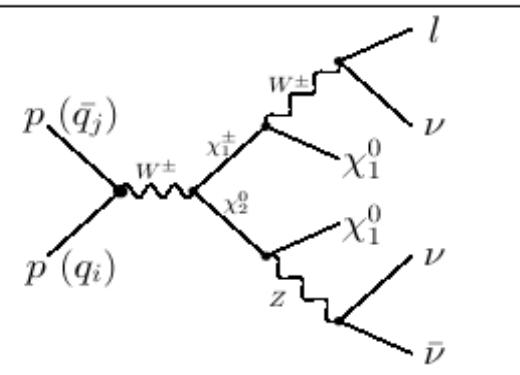
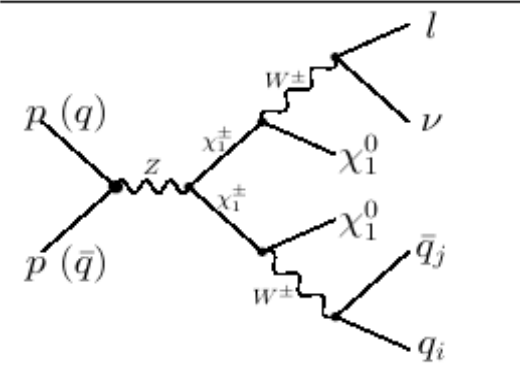
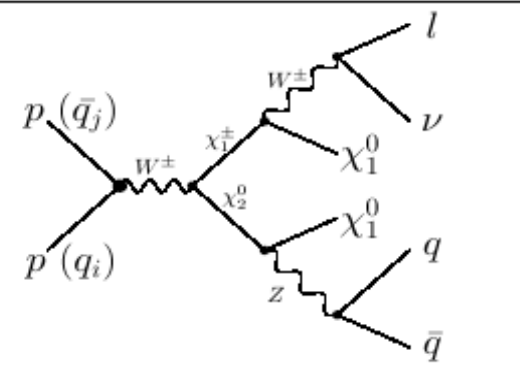
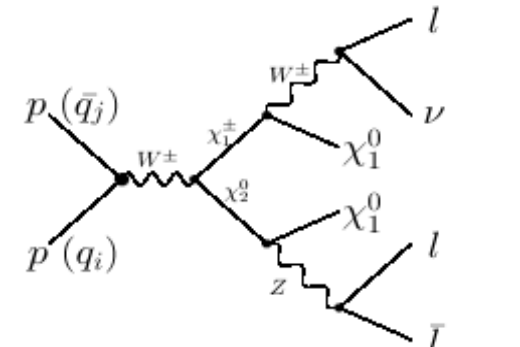
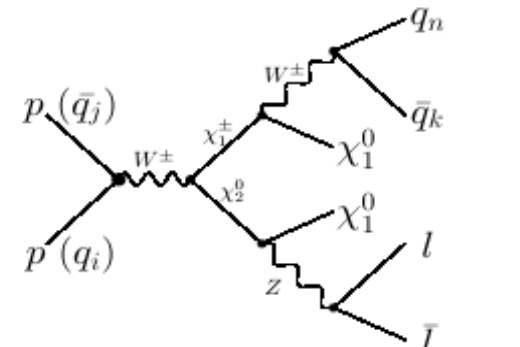
Narrow allowed region enables one to predict the particle spectra and the main decay patterns

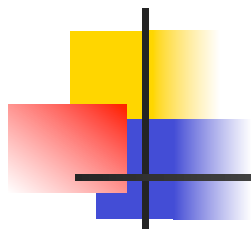
Phenomenology essentially depends on the region of parameter space and has direct influence on the strategy of SUSY searches

SUSY Production at LHC



Cascade Processes

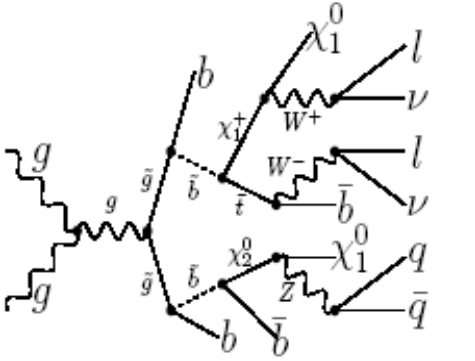
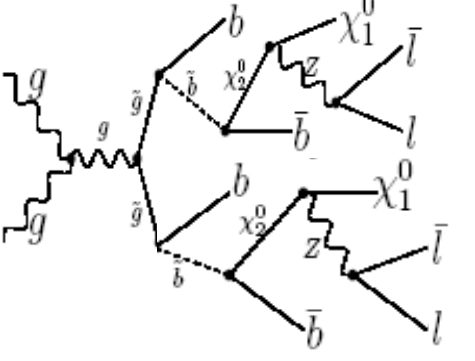
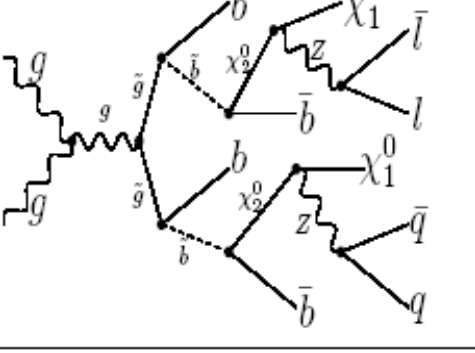
process	final states	process	final states
	2ℓ 2ν $\cancel{H^0}$		ℓ 3ν $\cancel{H^0}$
	1ℓ $2j$ ν $\cancel{H^0}$		ℓ ν $2j$ $\cancel{H^0}$
	3ℓ ν $\cancel{H^0}$		2ℓ $2j$ $\cancel{H^0}$

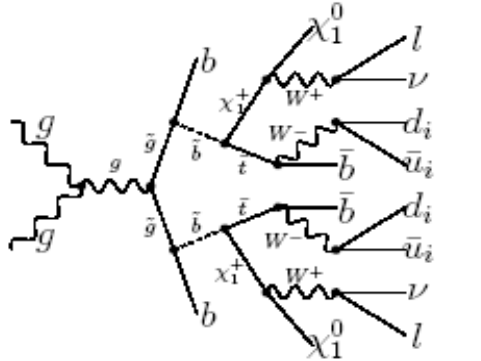
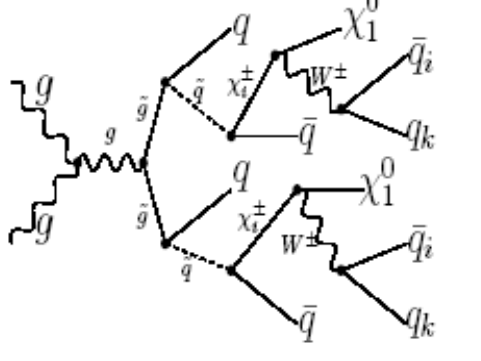
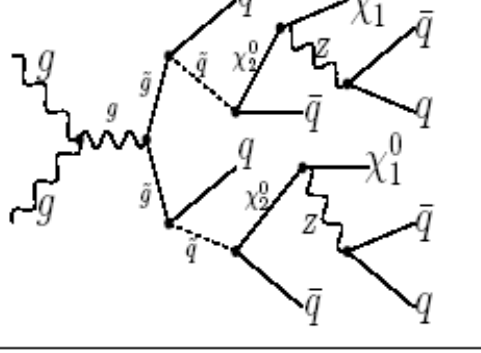


W
Z
H
A
S
I

Cascade Processes

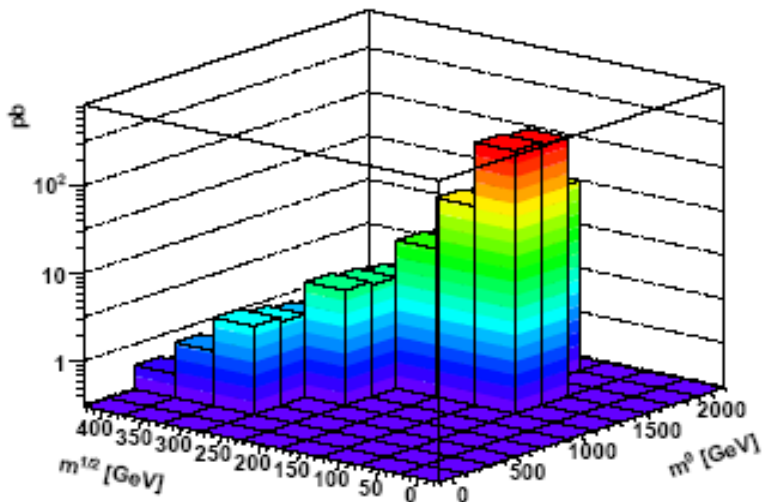
Strong Interactions

process	final states
	$2l$ 2ν $6j$ $\cancel{F_T}$
	$4l$ $4j$ $\cancel{F_T}$
	$2l$ $6j$ $\cancel{F_T}$

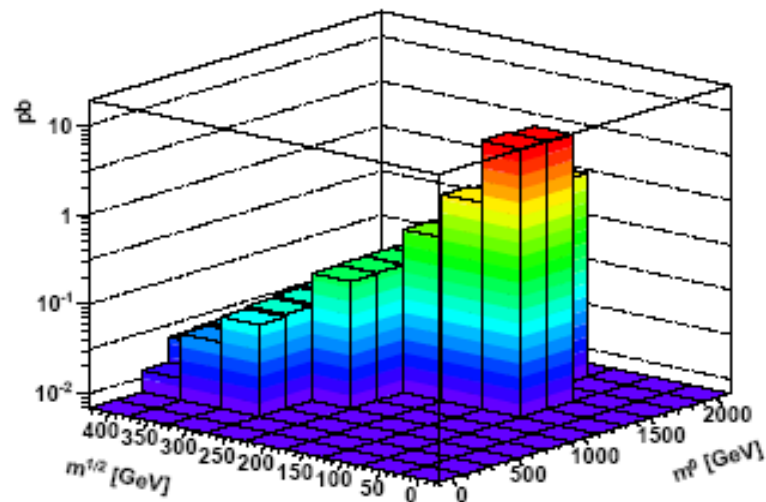
process	final states
	$2l$ 2ν $8j$ $\cancel{F_T}$
	$8j$ $\cancel{F_T}$
	$8j$ $\cancel{F_T}$

Cross-Sections at LHC

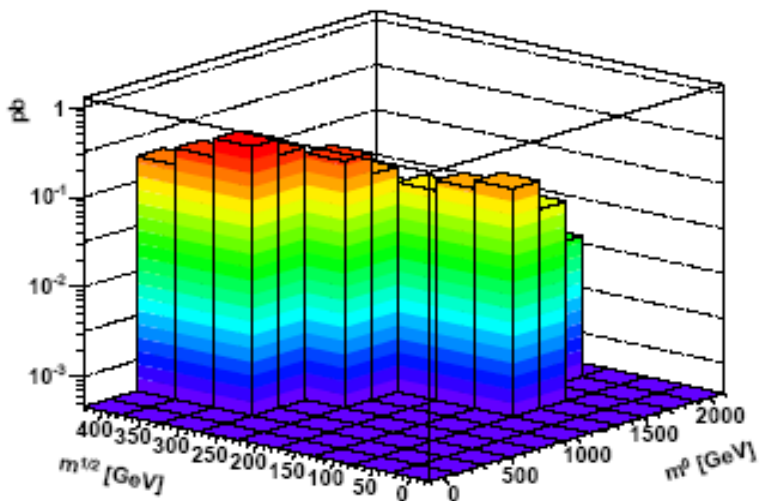
cross section p-p to $\tilde{g}\tilde{g}$



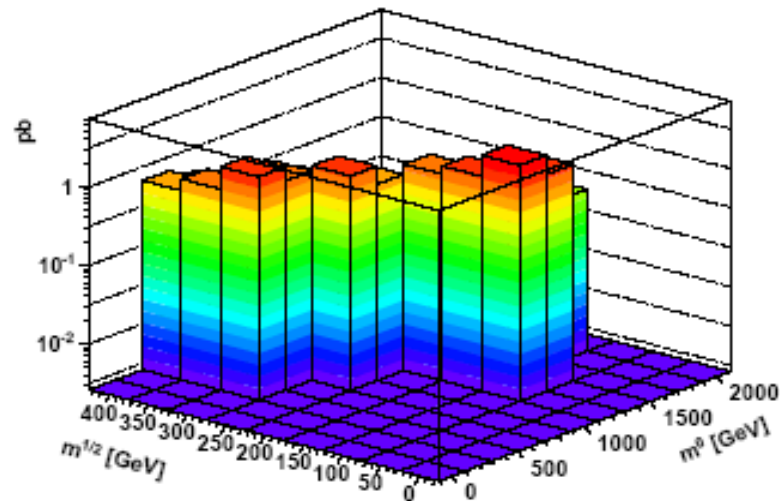
cross section p-p to $\tilde{\chi}_1^0\tilde{\chi}_2^0$



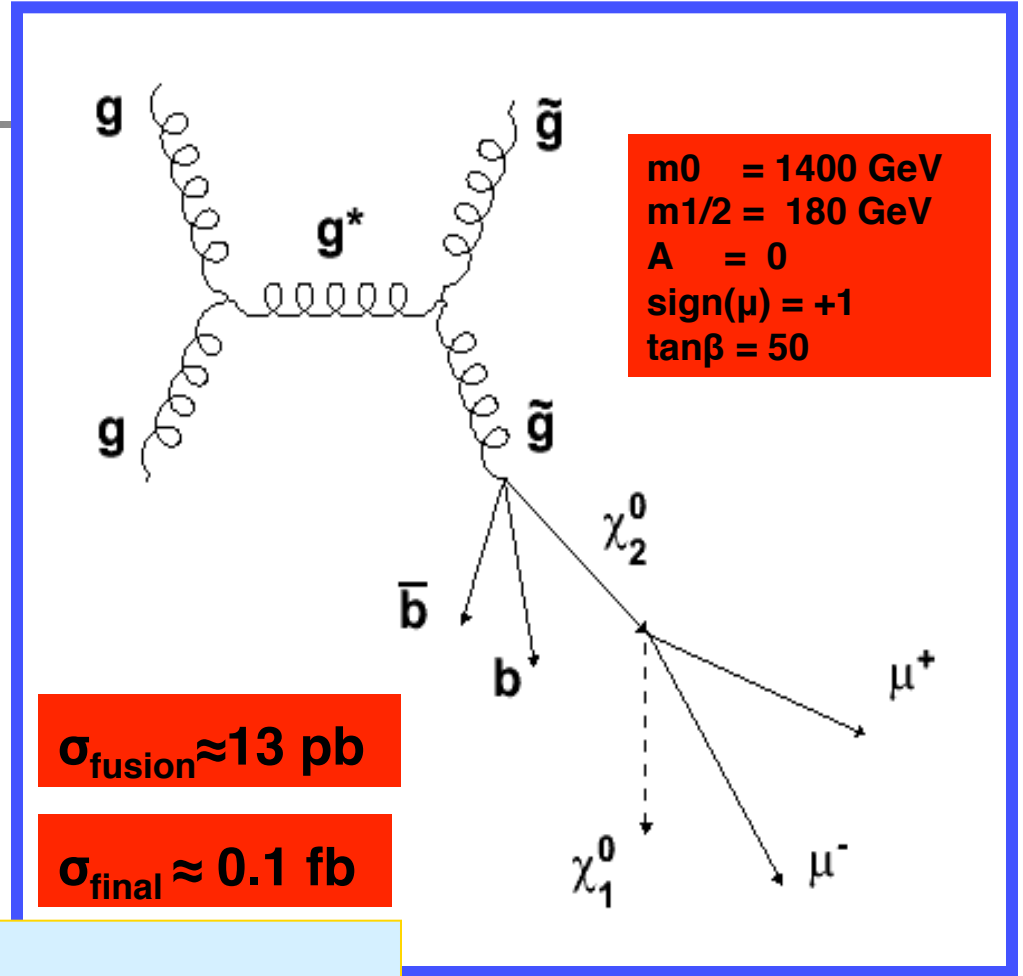
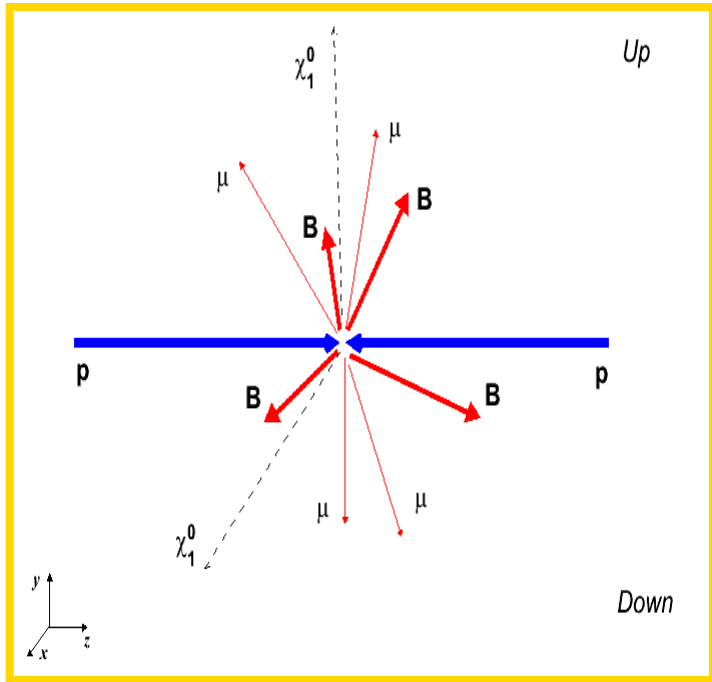
cross section p-p to $u\tilde{L}u\tilde{R}$



cross section p-p to $u\tilde{L}\tilde{g}$



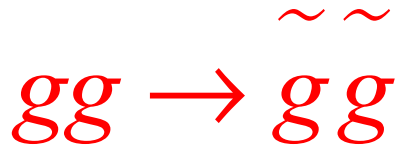
SUSY Production AT LHC



SIGNATURE:
4 b-jets + 4 muons + E_T^{miss}

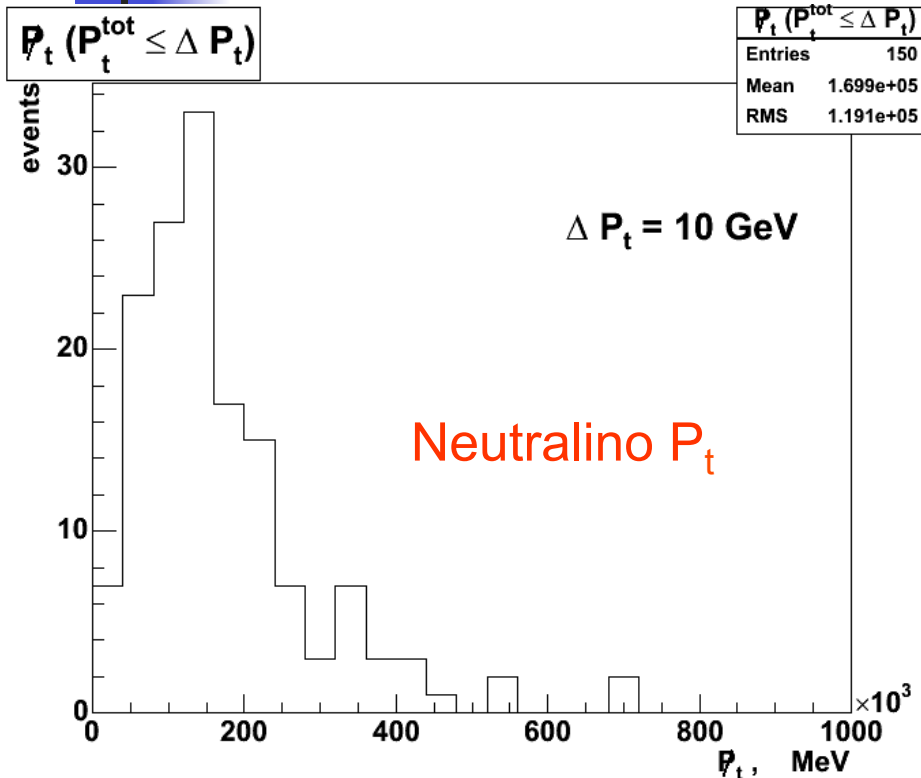
LARGE!

SUSY Event in ATLAS

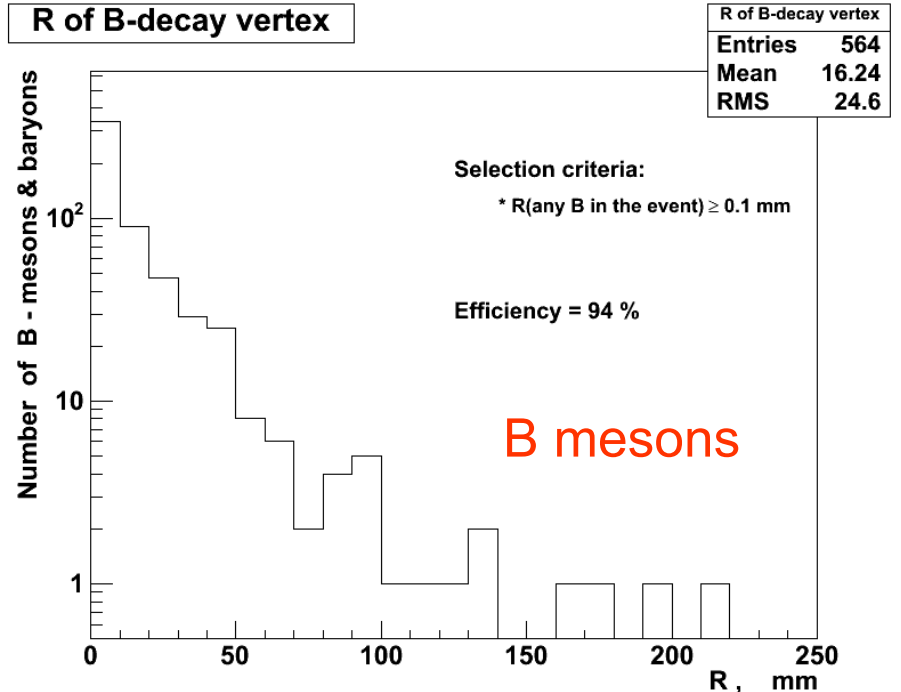


JINR(Dubna) ATLAS Group

V. Bednyakov, Y. Budagov, G. Khorauli, J. Khubua



Pythia within ATHENA,
B-vertex tagging

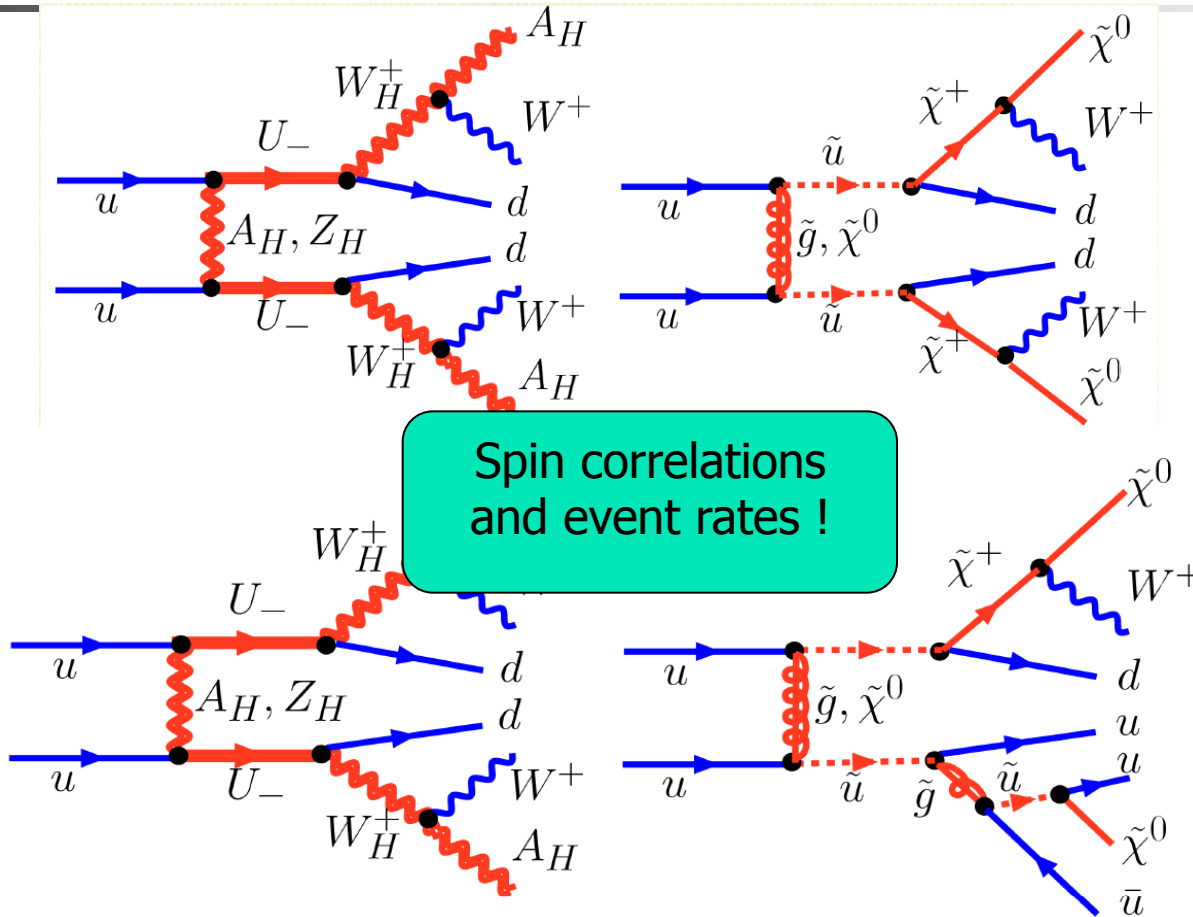


$$\sum P_t^{B,B,\mu,\mu} (\text{down}) - \sum P_t^{B,B,\mu,\mu} (\text{up}) = P_t \equiv E_t$$

SUSY Versus Little Higgs

A.Belyaev

Little Higgs

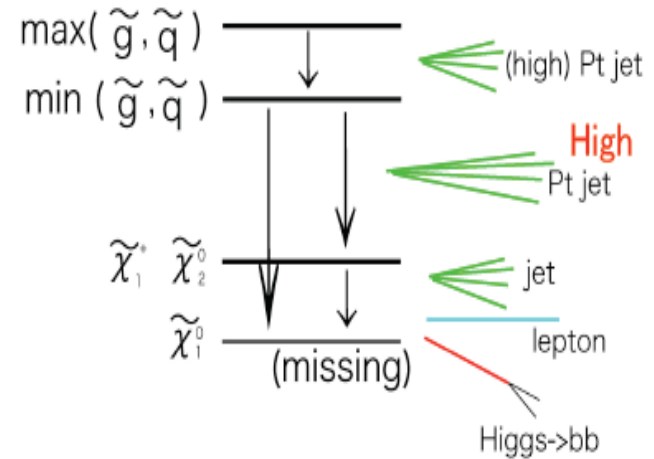
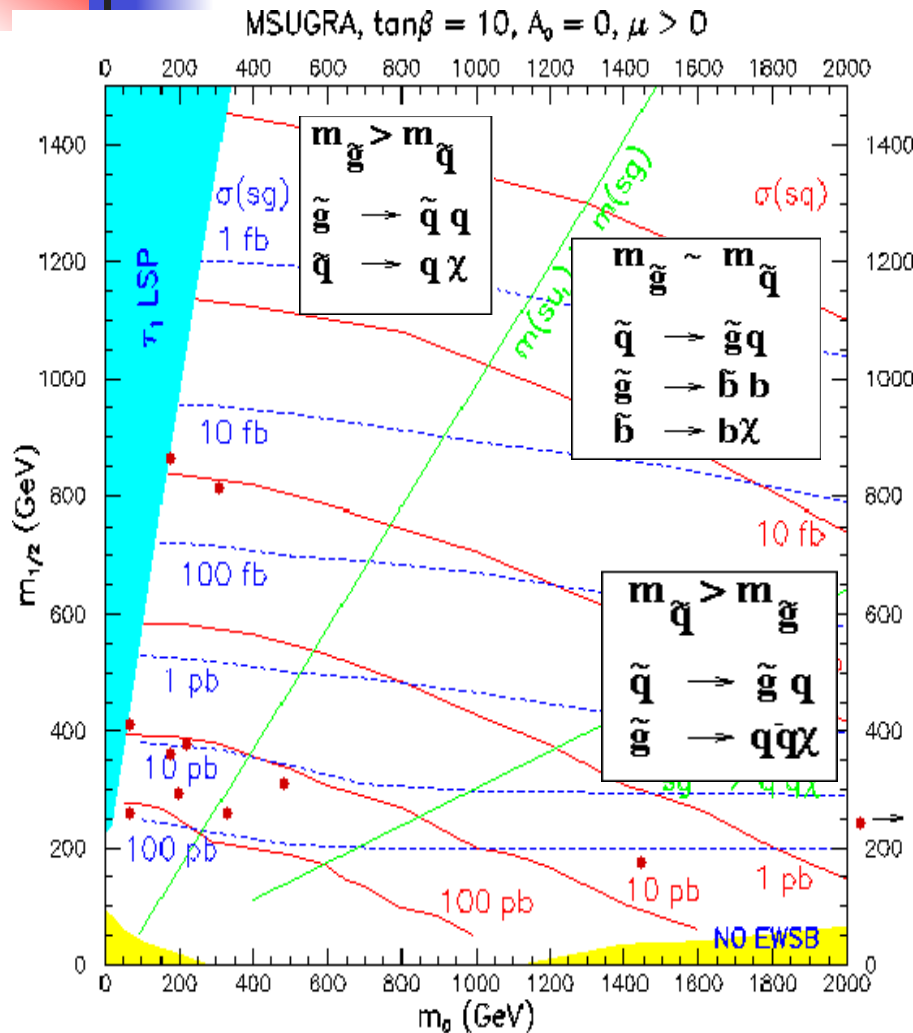


SUSY

Study of spin correlations is quite a challenge for LHC

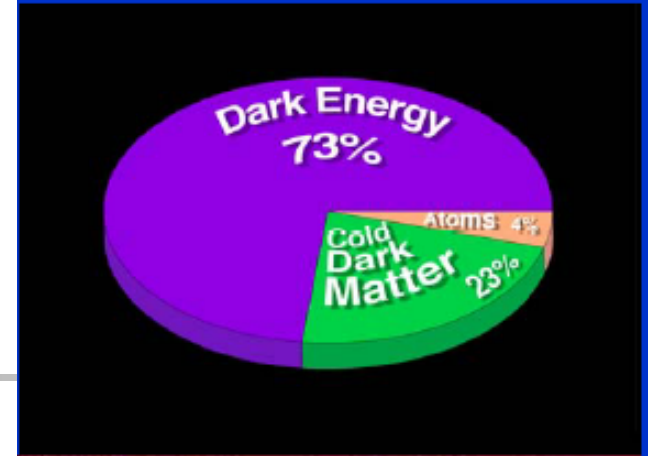
SUSY Searches at LHC

V.Zhukov



Main final state in the cascade decays:
MET+Jets+Leptons

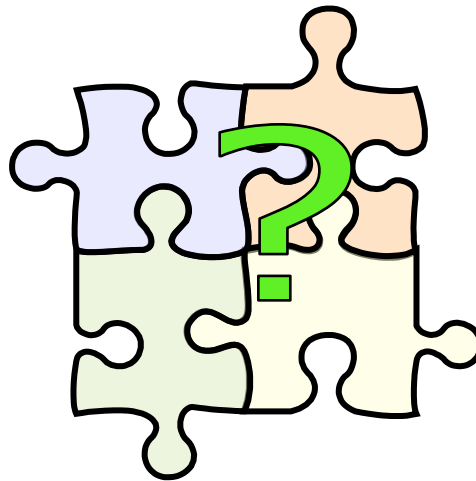
Problem # 3



What is Dark Matter ?



DARK



TRANSPARENT



INVISIBLE

What is it made of ?

The Origin of Dark Matter

The Dark Matter is made of:

- Macro objects – **Not seen**
- New particles – axion (axino)

Belram, Garcia-Bellido, Lesgourges; Kaplan, Zurek
Chun, Kim Jedamzik, Lemoike, Moutaka
Steffen; Panotopoulos

- neutralino

mSUGRA

- sneutrino

- right neutrino

Asaka, Blanchet, Shapohsnikov

- gravitino

Gauge Mediation

- heavy photon

Little Higgs Models

- heavy pseudo-goldstone

Inert Higgs Model

- light sterile higgs

Non from the SM

DM Detection

Direct detection

DAMA, Zeplin,
CDMS, Edelweiss

No convincing evidence so far
Hope for new results soon

Indirect detection

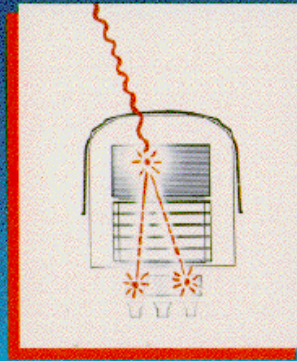
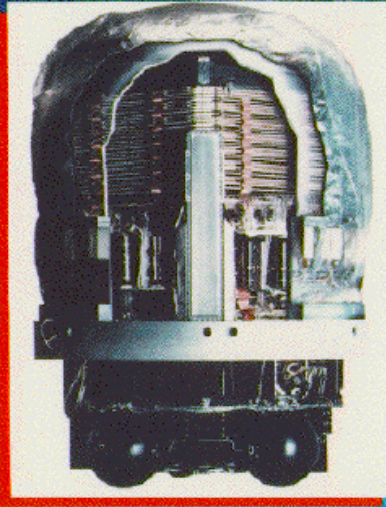
- EGRET -> GLAST
Diffuse Gamma Rays
- HEAT, AMS01 -> PAMELA
Positrons in Cosmic Rays
- BESS -> AMS02
Antiprotons in Cosmic Rays

E.Bloom
A.Morselli
P.Picozza
C.Goy

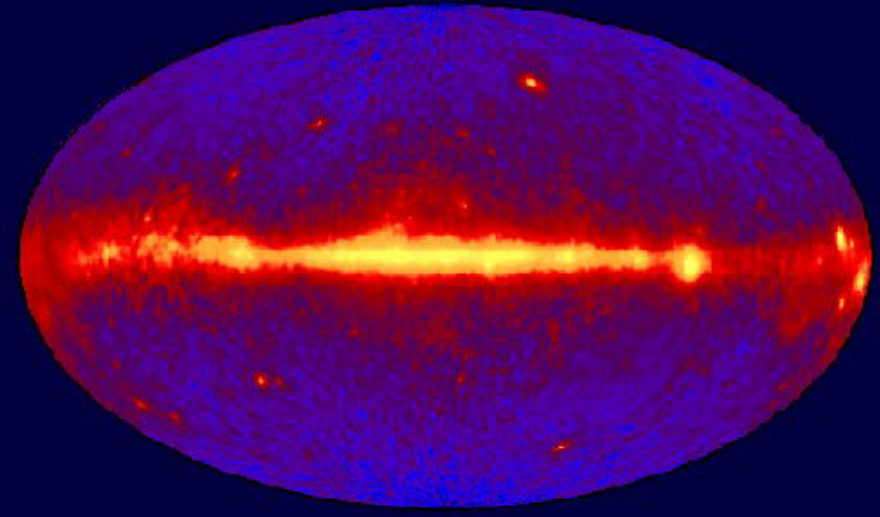
First Evidence of DM annihilation!

Diffuse Gamma Rays from the Sky

Energetic Gamma Ray Experiment Telescope (EGRET)



EGRET All-Sky Gamma-Ray Survey Above 100 MeV



Instrumental parameters:

Energy range: 0.02-30 GeV

Energy resolution: $\sim 20\%$

Effective area: 1500 cm²

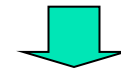
Angular resol.: $< 0.5^\circ$

Data taking: 1991-2000

Main EGRET results:

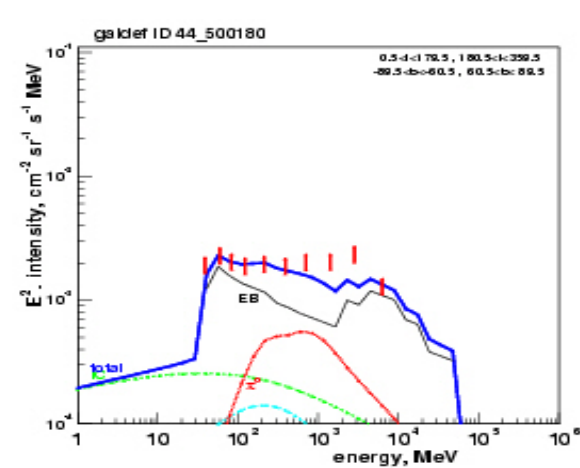
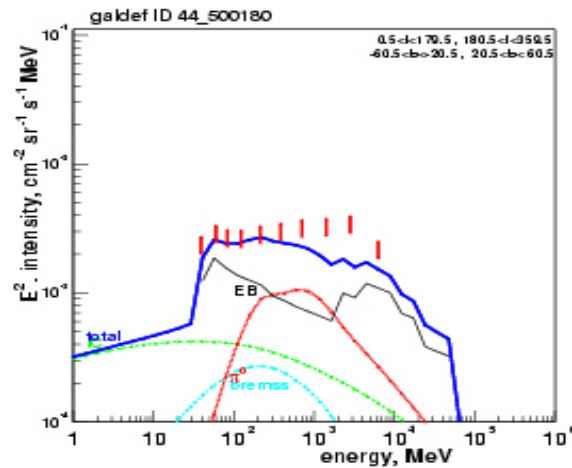
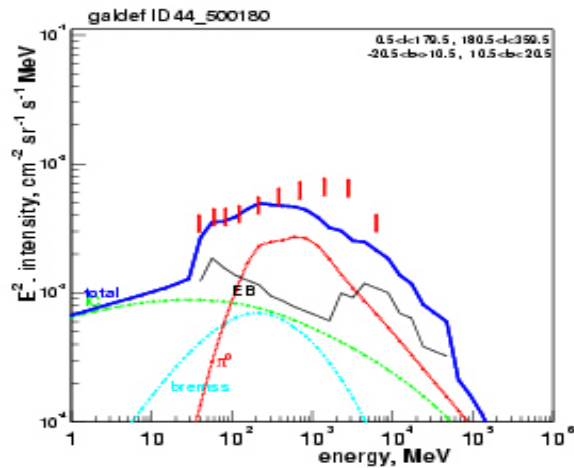
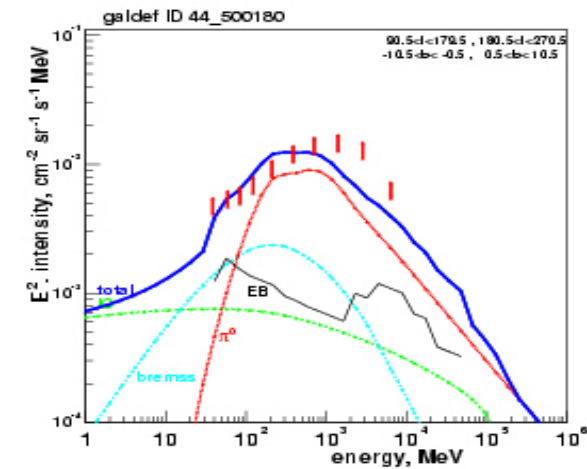
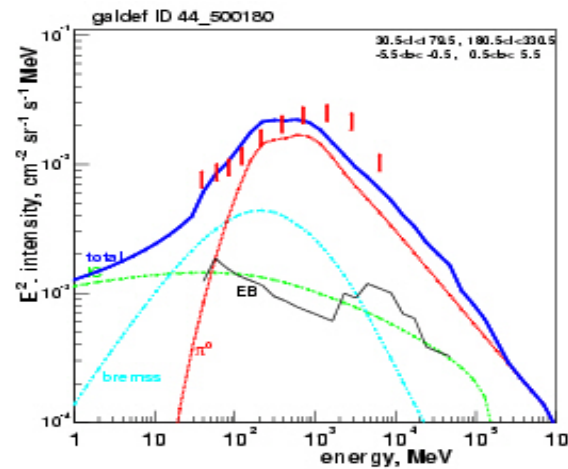
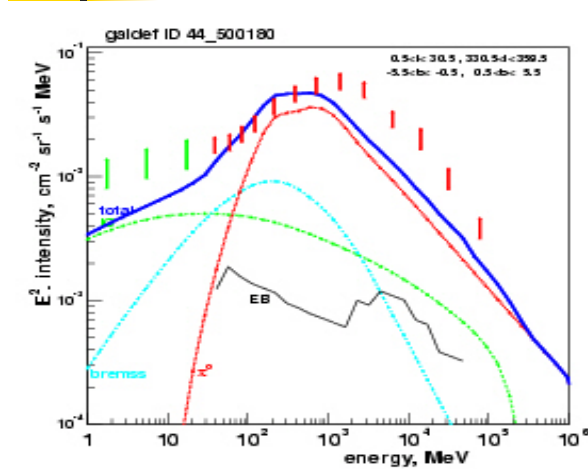
Catalogue of point sources

Excess in diffuse gamma rays



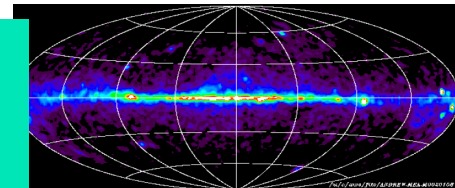
First indirect evidence for DM annihilation

EXCESS OF DIFFUSE GAMMAY RAYS ABOVE 1 GEV



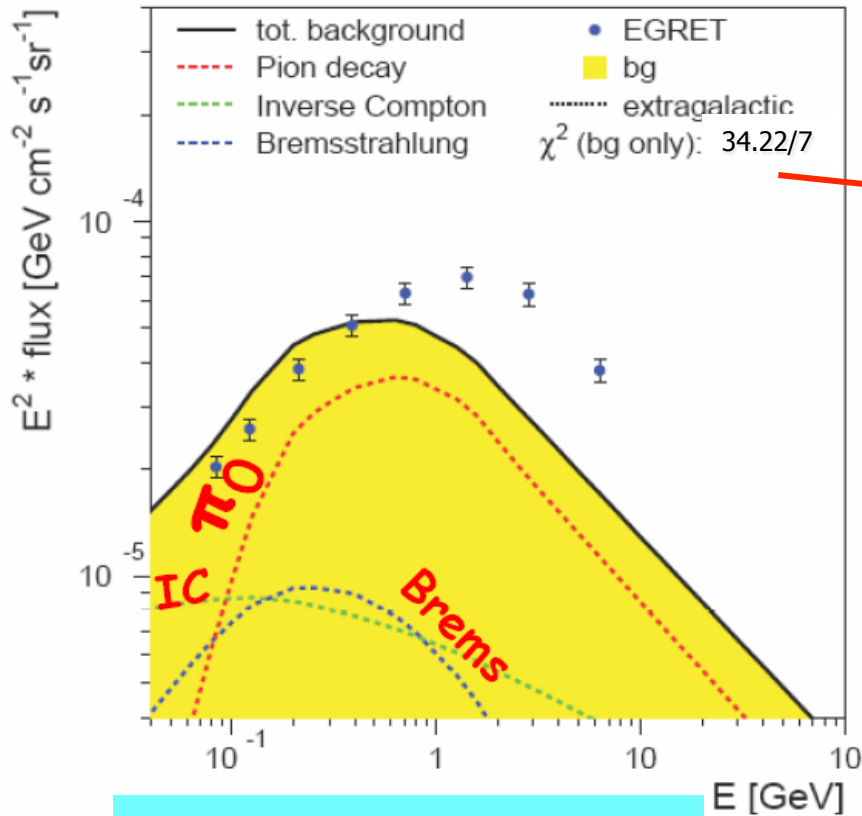
A: inner Galaxy ($l = \pm 30^\circ$, $|b| < 5^\circ$)
B: Galactic plane avoiding A
C: Outer Galaxy

D: low latitude ($10-20^\circ$)
E: intermediate lat. ($20-60^\circ$)
F: Galactic poles ($60-90^\circ$)

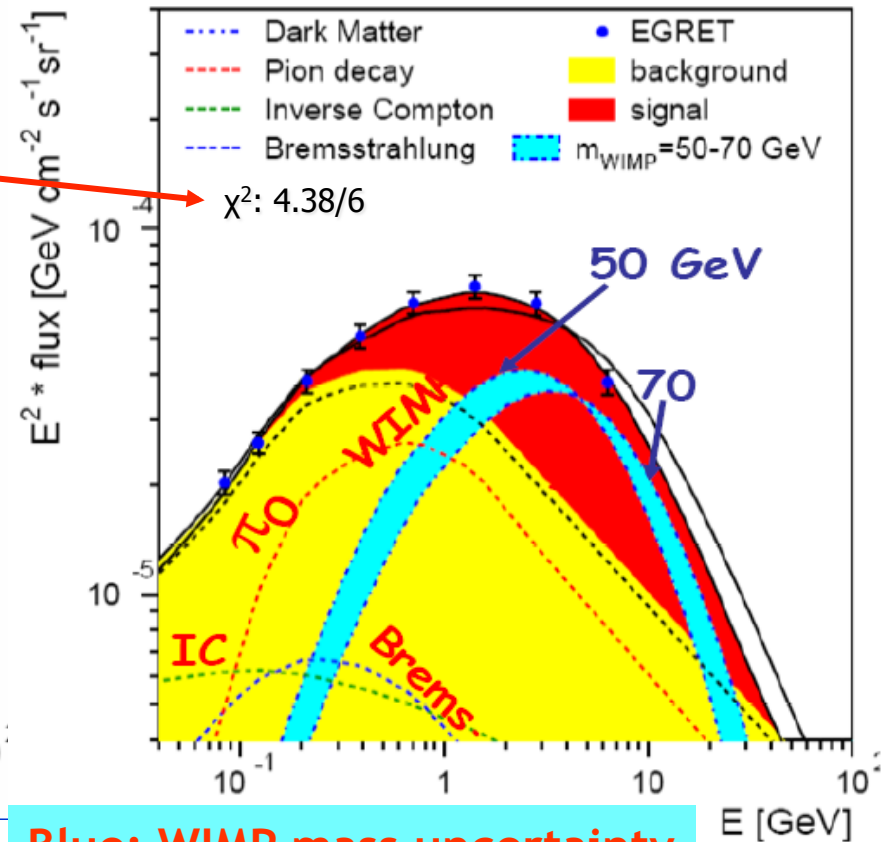


Excess of Diffuse Gamma Rays with and without DM Annihilation

W.De Boer, C.Sanders, V.Zhukov, A.Gladyshev and D.Kazakov



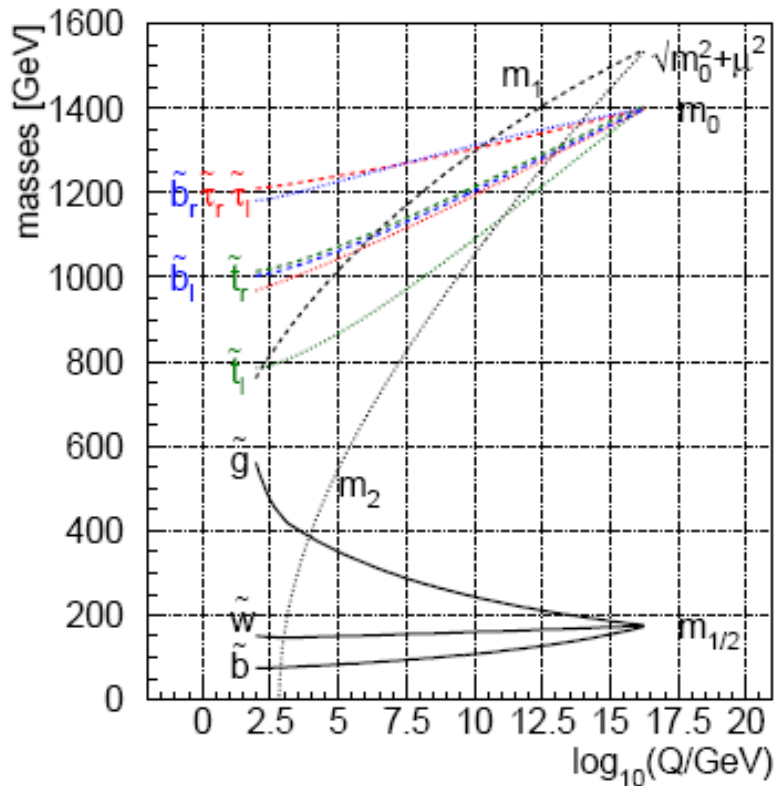
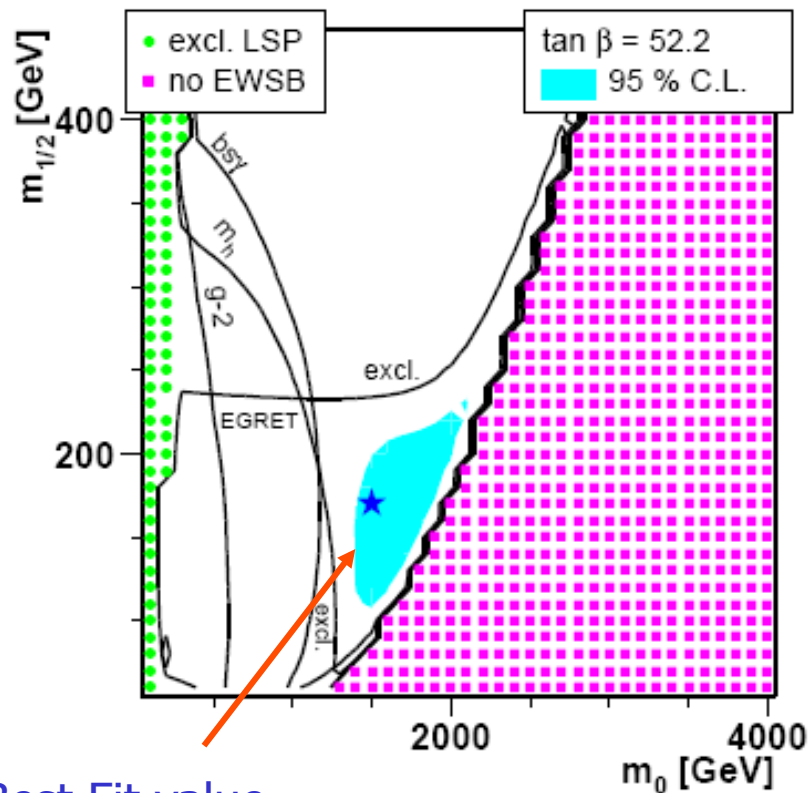
Without DM annihilation



Blue: WIMP mass uncertainty

Fit to the shape of spectrum + intensity from rotation curve of Milky Way

EGRET Point and Mass Spectrum



Best Fit value

Split SUSY: light gauginos & heavy matter

WIMP = LSP (Neutralino) – superpartner of CMB

SUSY MASS SPECTRUM

Fitted SUSY Parameters

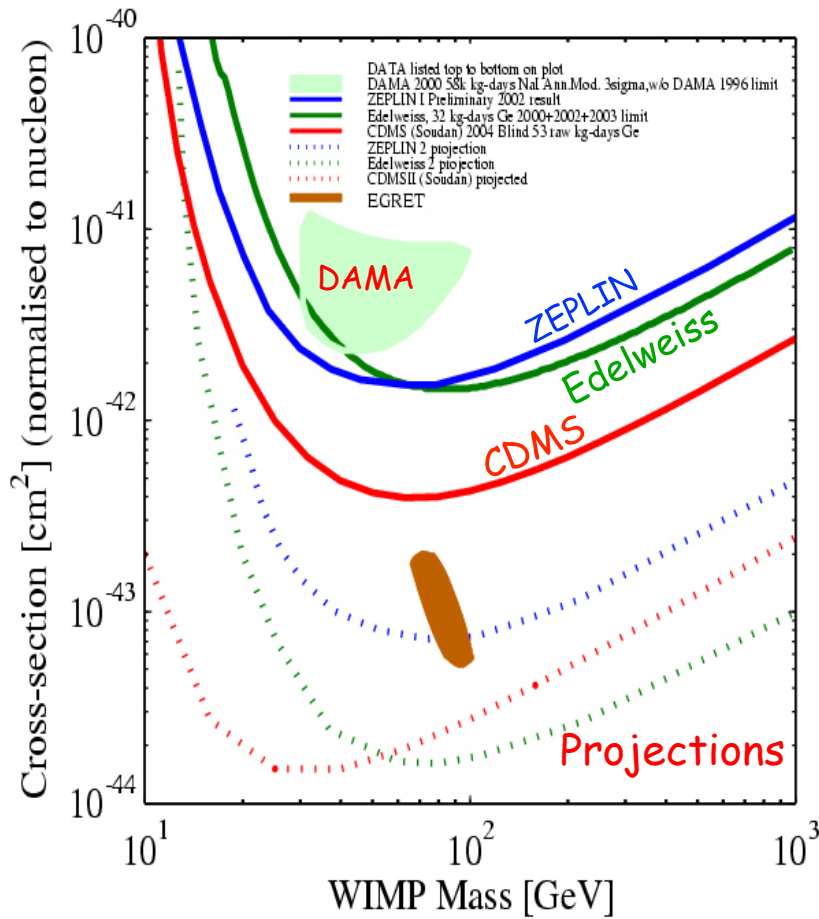
SUSY Masses in GeV

Parameter	Value
Tan β	52.2
m_0	1500
$m_{1/2}$	170
Sign μ	+
$A(0)$	0
$\alpha_s(M_Z)$	0.122
$\alpha_{em}(M_Z)$	0.0078153697
$\text{Sin}^2 \theta_w _{\overline{MS}}$	0.2314
m_t	175 GeV
m_b	4.214 GeV

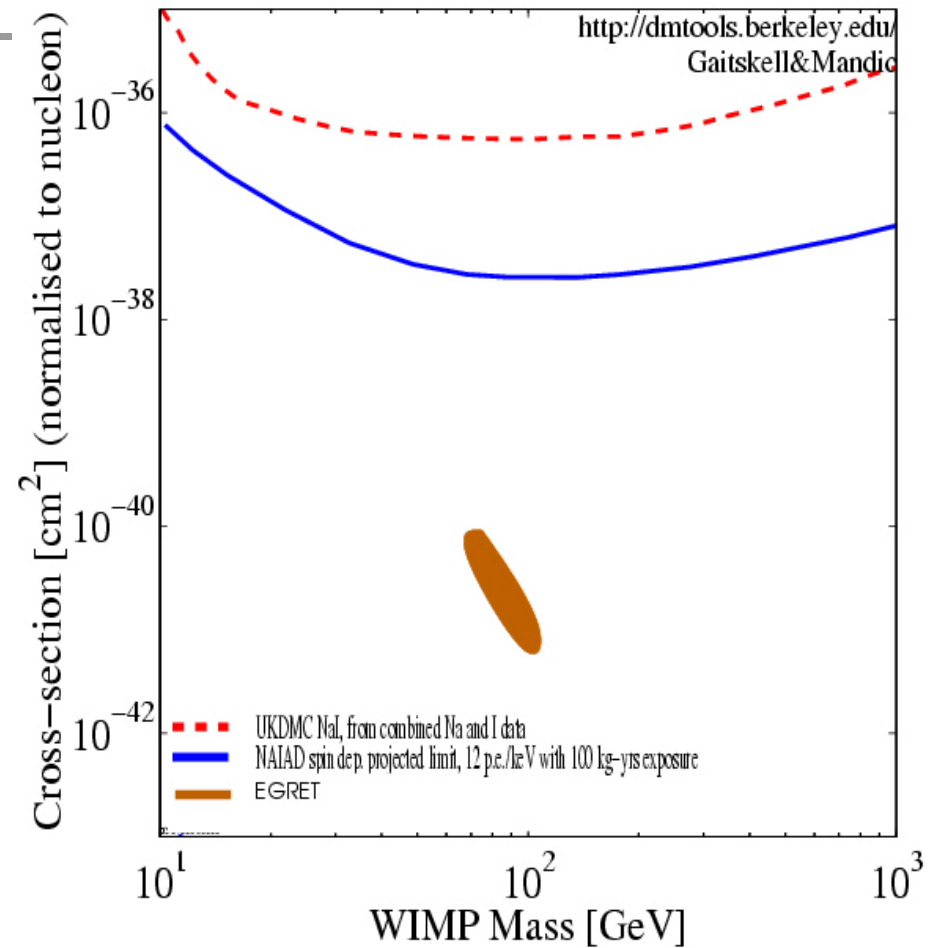
Particle	Mass
$\tilde{\chi}_{1,2,3,4}^0$	→ 64, 113, 194, 229
$\tilde{\chi}_{1,2}^\pm, \tilde{g}$	→ 110, 130, 516
$\tilde{u}_{1,2} = \tilde{c}_{1,2}$	1519, 1523
$\tilde{d}_{1,2} = \tilde{s}_{1,2}$	1522, 1524
$\tilde{t}_{1,2}$	906, 1046
$\tilde{b}_{1,2}$	1309, 1152
$\tilde{e}_{1,2} = \tilde{\mu}_{1,2}$	1497, 1499
$\tilde{\tau}_{1,2}$	1305, 1288
$\tilde{\nu}_e, \tilde{\nu}_\mu, \tilde{\nu}_\tau$	1495, 1495, 1286
h, H, A, H^\pm	→ 115, 372, 372, 383

Direct DM Searches

Spin-independent



Spin-dependent



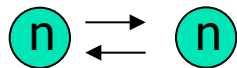
Predictions from EGRET data assuming Supersymmetry

Further Major Problems

- Baryon Asymmetry → Baryo and leptogenesis

- Violation of B-L (required for baryogenesis) Kuzmin, Rubakov, Shaposhnikov

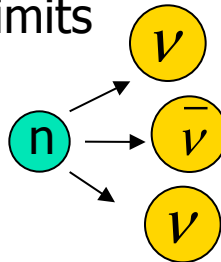
Neutron-antineutron oscillations $\Delta B=2$



B. Dutta, K. Babu et al.
A. Dolgov et al., C. Bambi

Nucleon decays, $\tau \rightarrow$ nucleon decays with $\Delta|B-L|=2$

New $n \rightarrow \nu\nu\nu$ and $n \rightarrow \nu\nu$ limits



Yu. Kamyshev
KamLAND

- Dark Energy

Physics beyond the SM



What the future may bring?



Instead of Conclusion

“Blessed are those who believe
and yet have not seen”

St. John, XX, 29

“Блаженны не видевшие
и уверовавшие”

Св. Иоанн, XX, 29