



## Lattice QCD for Beyond the Standard Model physics

M.P. Lombardo



Beyond the Standard Model:  
HUGE territory!

This talk: Two examples of applications of topics  
discussed here which have a close connections  
to BSM physics

I Dark Matter: limit on the mass for post-inflationary Axions  
from QCD topology

II Higgs mass generation via QCD-like dynamics

## The two faces of the axion



Dark Matter Candidate

QCD Topology,  
Strong CP Problem

Inspiring paper:  
Berkowitz, Buchoff, Rinaldi **Phys.Rev. D92 (2015) no.3, 034507**

$\theta$  term, strong CP problem and topology

$$\mathcal{L}_{QCD}(\theta) = \mathcal{L}_{QCD} + \frac{g^2\theta}{32\pi^2} \epsilon^{\mu\nu\rho\sigma} F_{\mu\nu}^a F_{\rho\sigma}^a.$$

Admitted but  $\theta < 10^{-9}$

$$Q = \int d^4x \frac{g^2}{32\pi^2} \text{tr} F \tilde{F}$$

$$Z_{QCD}(\theta, T) = \int [dA][d\psi][d\bar{\psi}] \exp \left( -T \sum_t d^3x \mathcal{L}_{QCD}(\theta) \right) = \exp[-VF(\theta, T)]$$

$$\frac{\partial^2 F(\theta, T)}{\partial \theta^2} \Big|_{\theta=0} \equiv \chi(T) = (\langle Q^2 \rangle - \langle Q \rangle^2) / V$$

Axions 'must' be there: solution to the strong CP problem

$$\mathcal{L}_{QCD}(\theta) = \mathcal{L}_{QCD} + \frac{g^2 \theta}{32\pi^2} \epsilon^{\mu\nu\rho\sigma} F_{\mu\nu}^a F_{\rho\sigma}^a$$

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Postulate axions, coupled to Q:

$$\mathcal{L}_{\text{axions}} = \frac{1}{2} (\partial_\mu a)^2 + \left( \frac{a}{f_a} + \theta \right) \frac{1}{32\pi^2} \epsilon^{\mu\nu\rho\sigma} F_{\mu\nu} F_{\rho\sigma}$$

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Axion potential

$$m_a^2(T) f_a^2 = \left. \frac{\partial^2 F(\theta, T)}{\partial \theta^2} \right|_{\theta=0} \equiv \chi(T),$$

# Axions 'must' be there: solution to the strong CP problem

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Axion potential

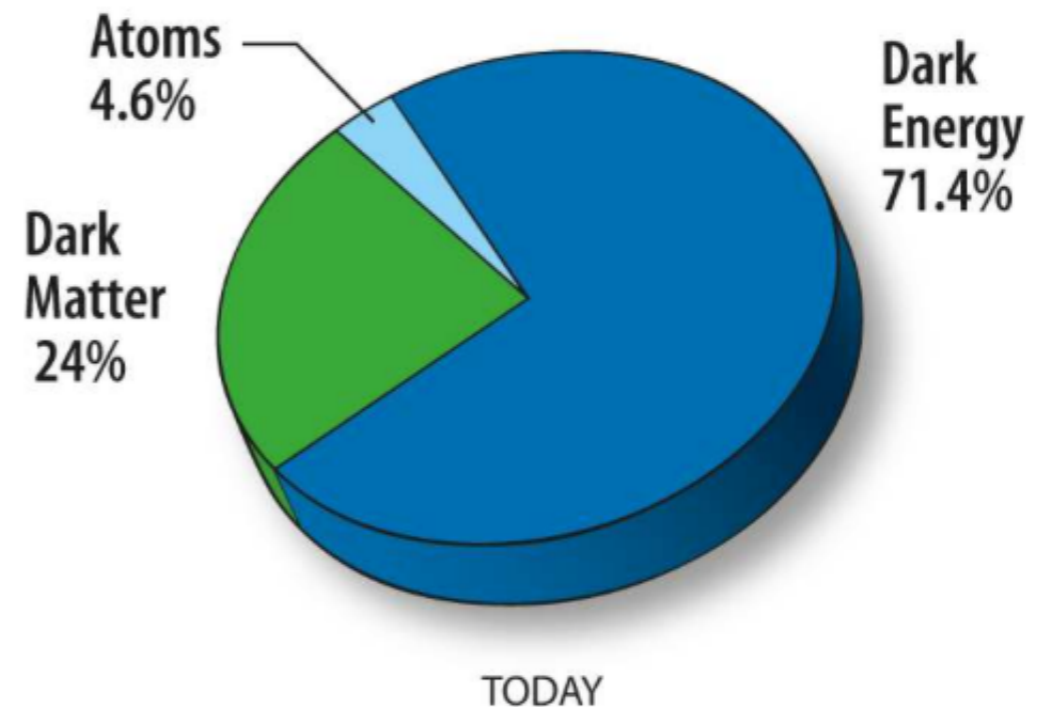
$$m_a^2(T) f_a^2 = \left. \frac{\partial^2 F(\theta, T)}{\partial \theta^2} \right|_{\theta=0} \equiv \chi(T),$$

Axion:

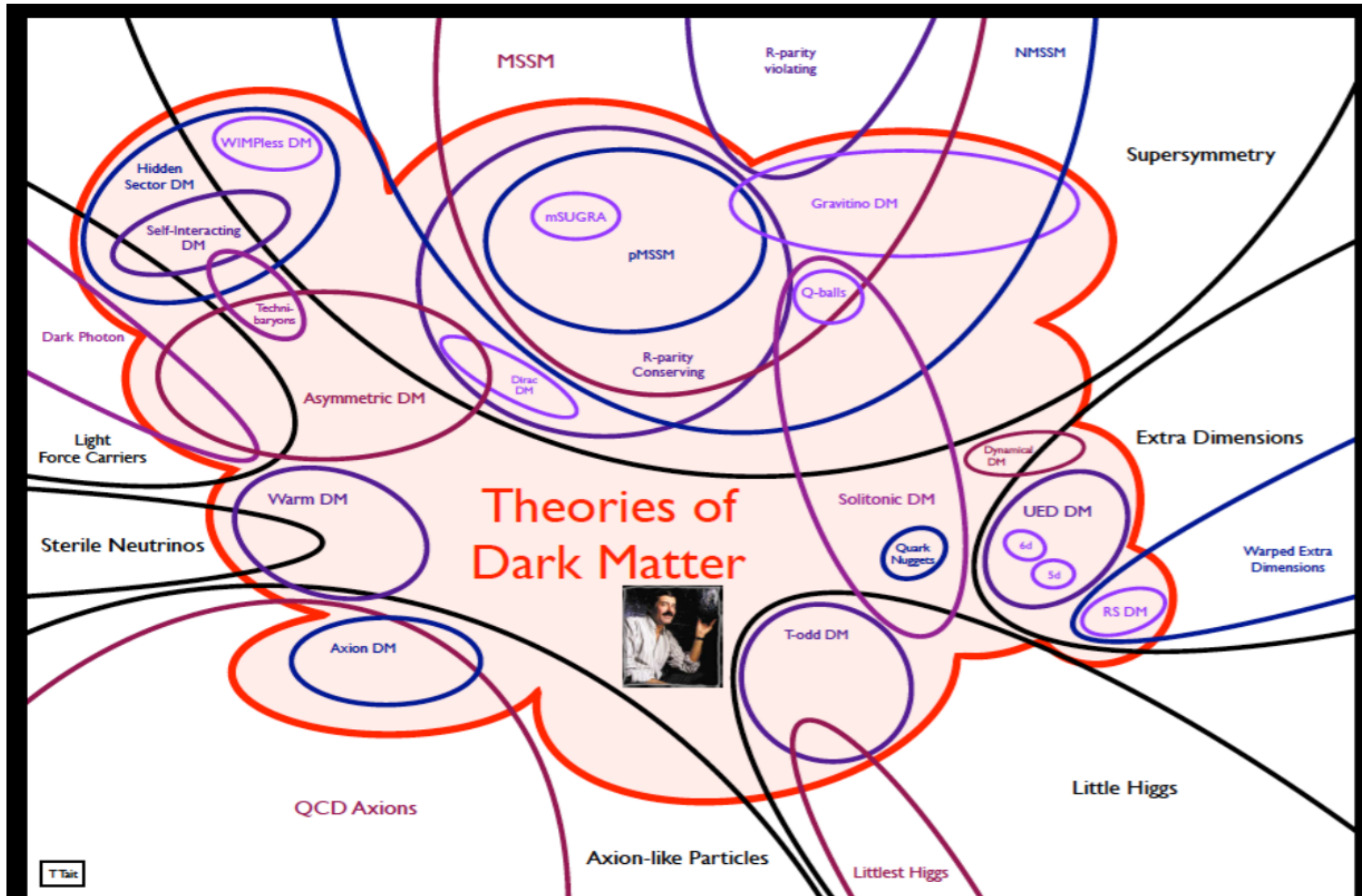
theoretically well motivated

searched and not found in experiments -> weakly coupled

-> Dark Matter Candidate



# Theory landscape (From Tim Tait, Snowmass)





Axion:

theoretically well motivated

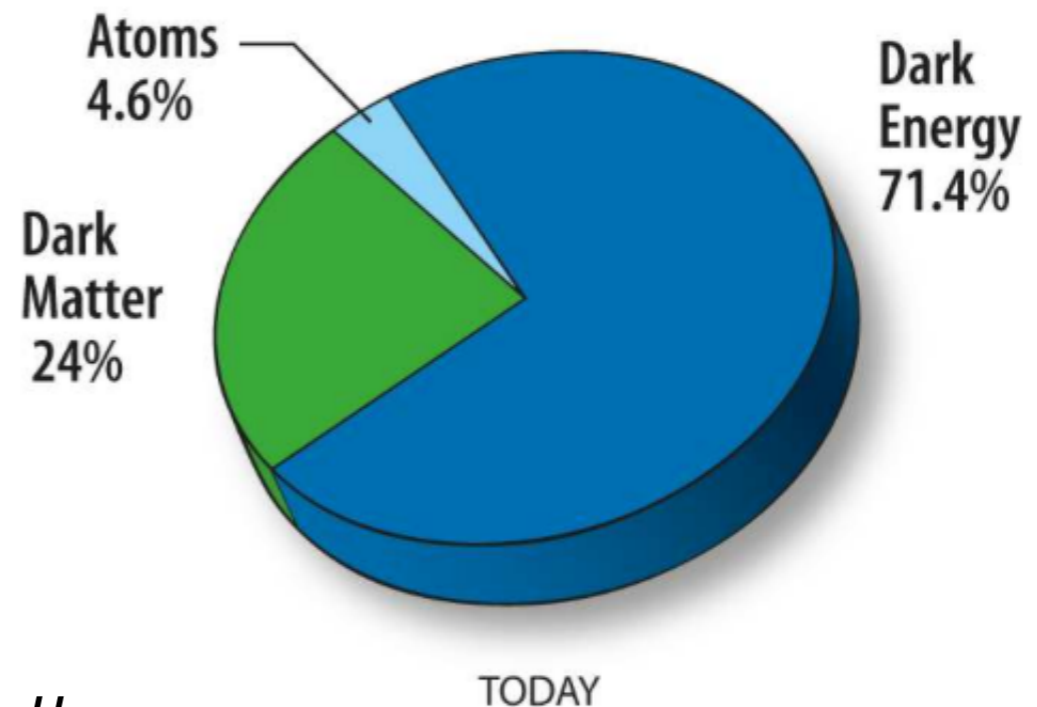
searched and not found in experiments -> weakly coupled

-> Dark Matter Candidate

Rough estimate for axion contribution to DM

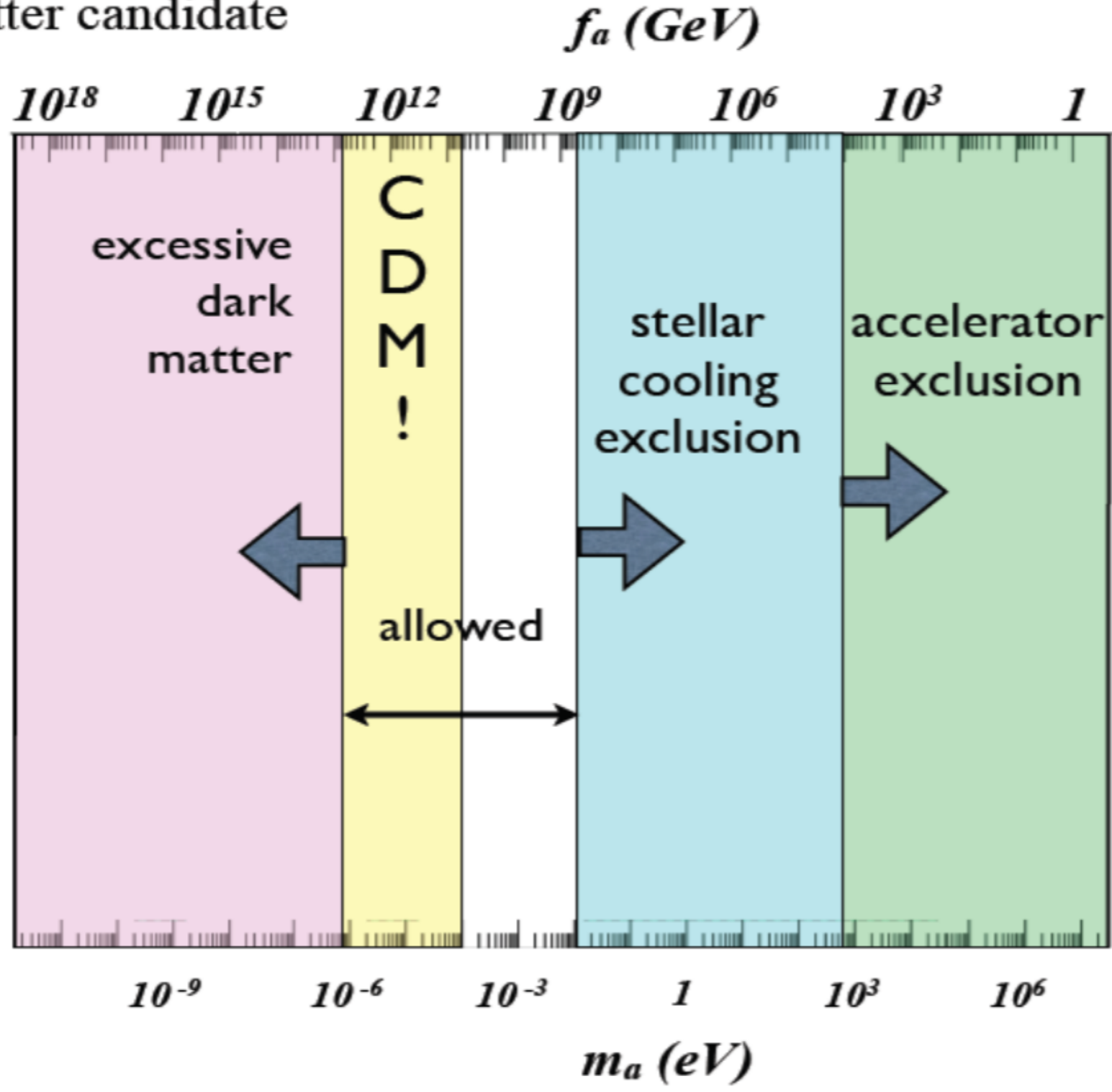
$$\frac{\rho_{\text{axion}}}{\rho_{DM}} \sim \left( \frac{f_a}{10^{11} \text{GeV}} \right)^{\frac{3}{2}}$$

*from 'almost' 1 till 'very small'*



# Allowed axion window

Allowed axion window  
Cold Dark Matter candidate

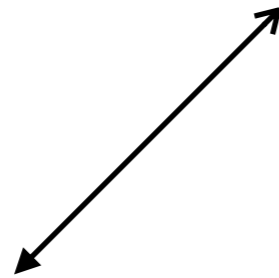


Time from Big Bang



Axions's freezout

$$3H(T) = m_a(T)$$



**Axions' mass and density today**

Temperature

Hubble parameter

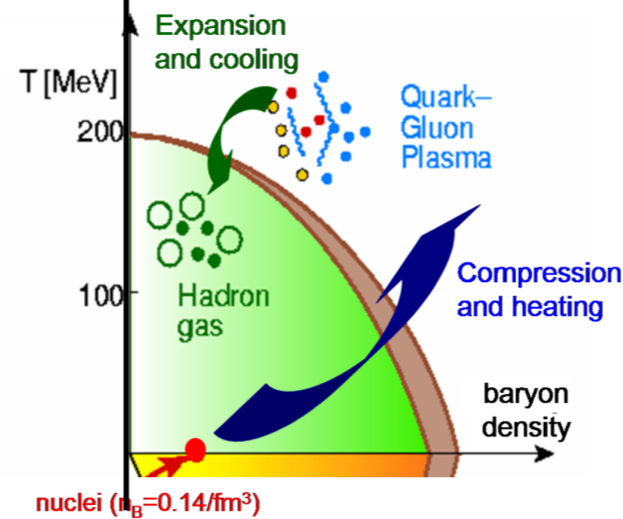
$$H(T) \simeq T^2 / M_P$$

Details

from the QCD eos

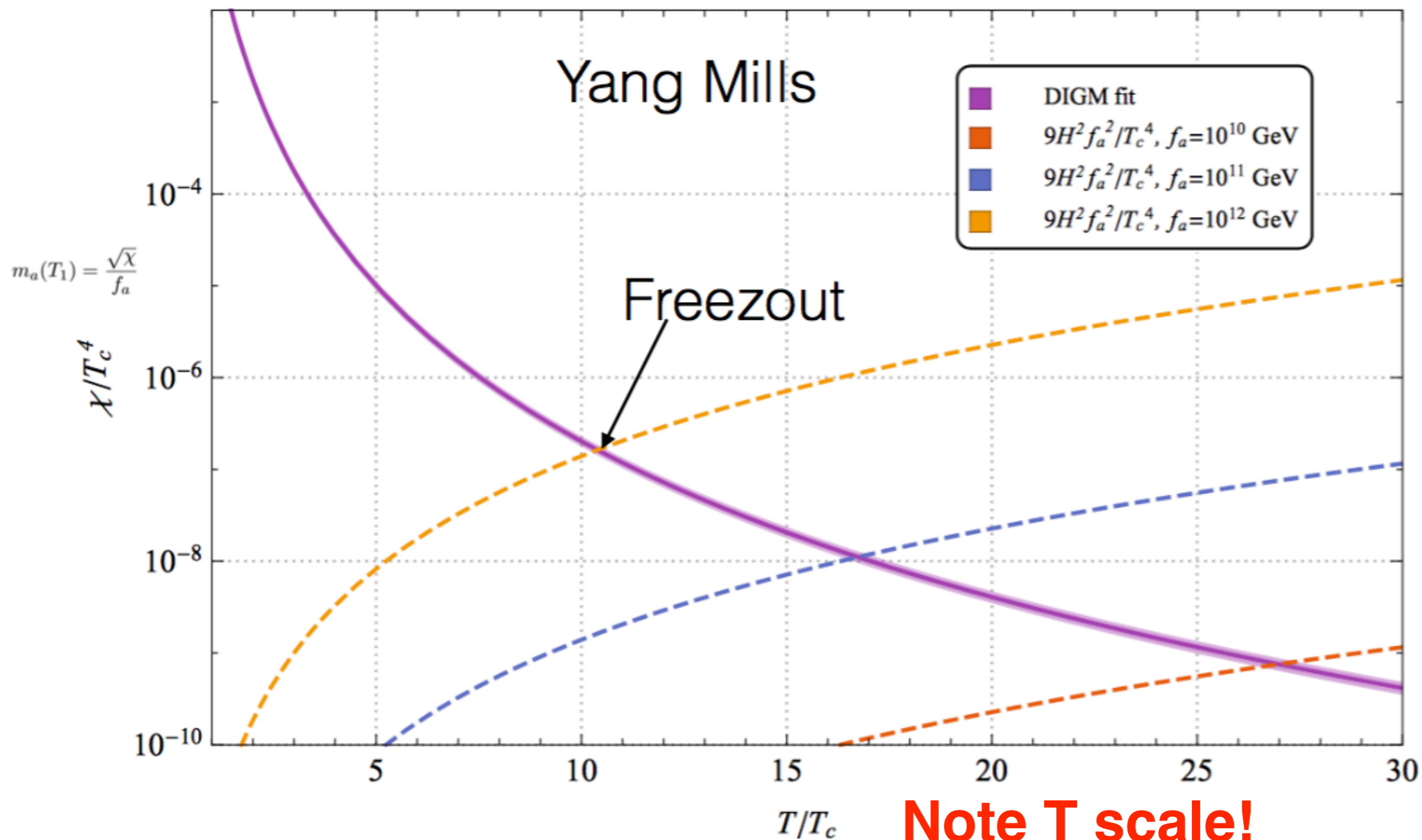
$$m_a(T) = \sqrt{\chi(T)} / f_a$$

**Quark Gluon Plasma:  
Topology**



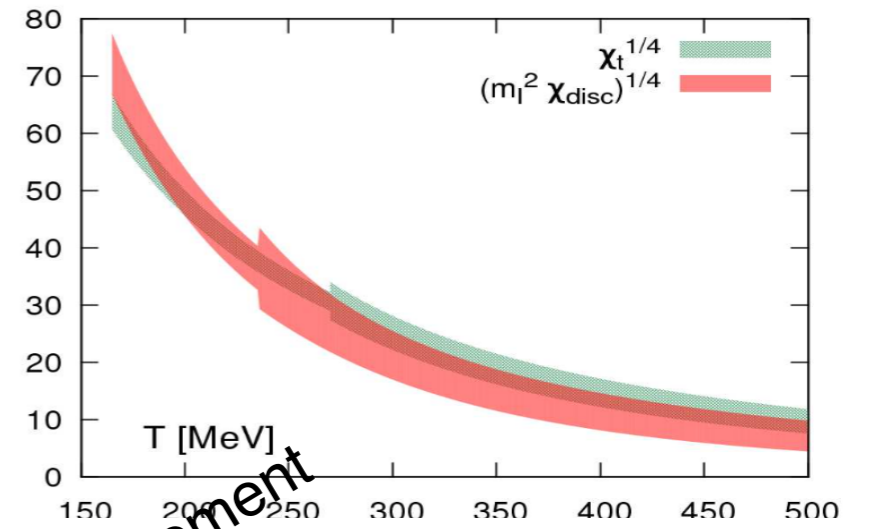
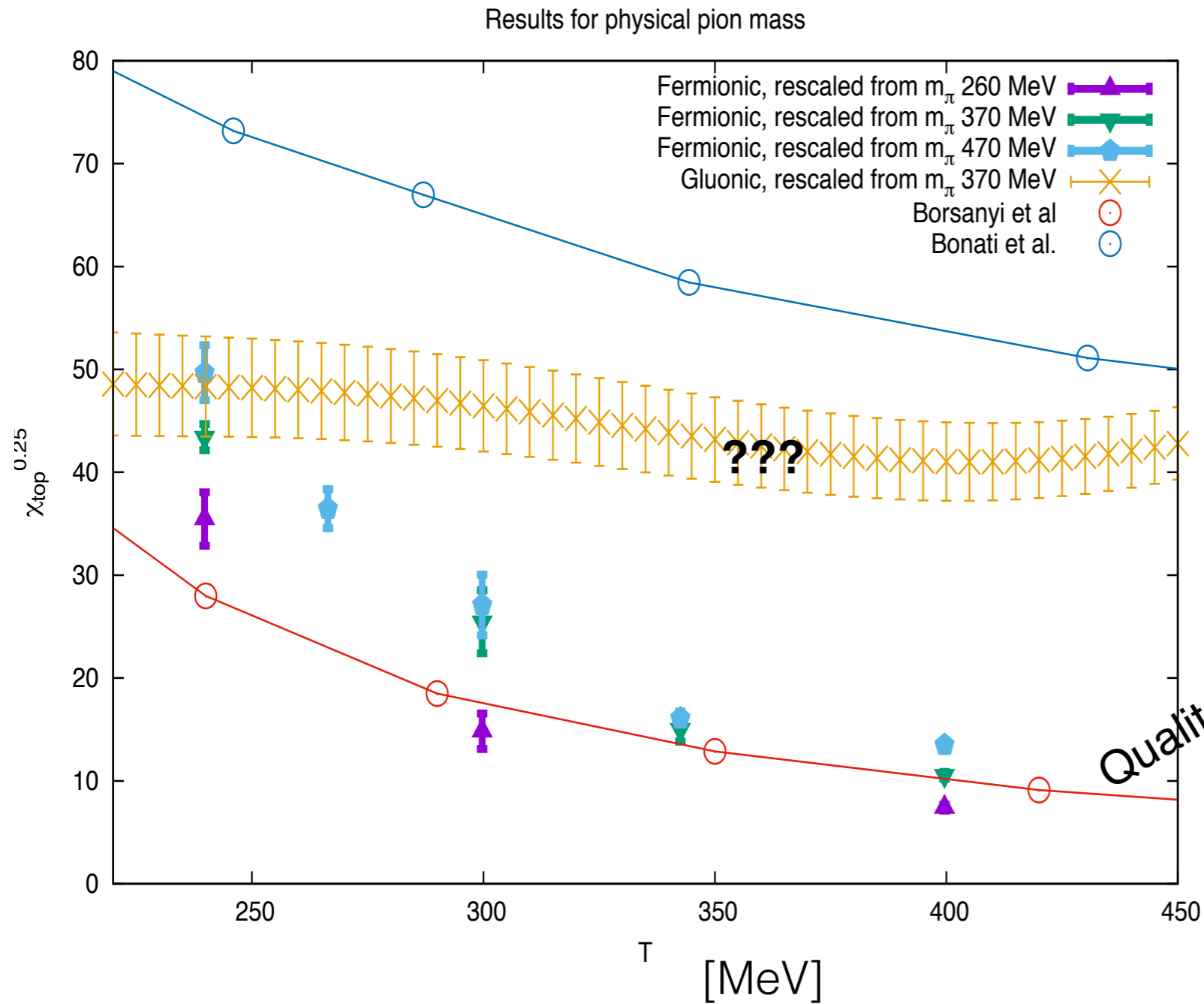
Axion freezout :  $3H(T) = m_a(T) = \sqrt{\chi(T)} / f_a$

Berkowitz Buchhoff Rinaldi 2015



Axion density at freezout controls axion density today

# Needed: Topological susceptibility at 'nearly unreachable' temperatures!



Qualitative agreement

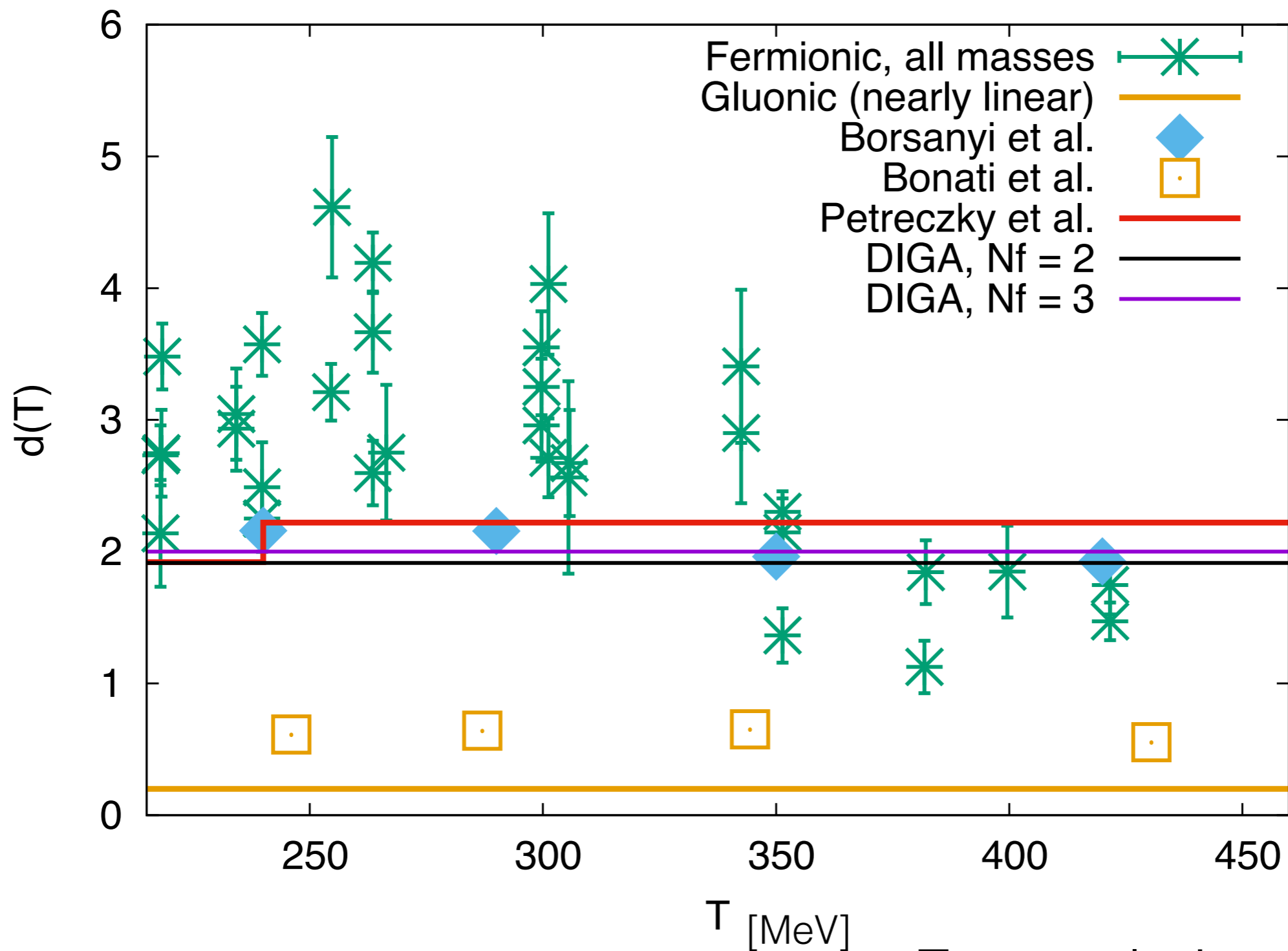
Petreczky et al.

Current results

Effective exponent  $d(T)$ :

Current results

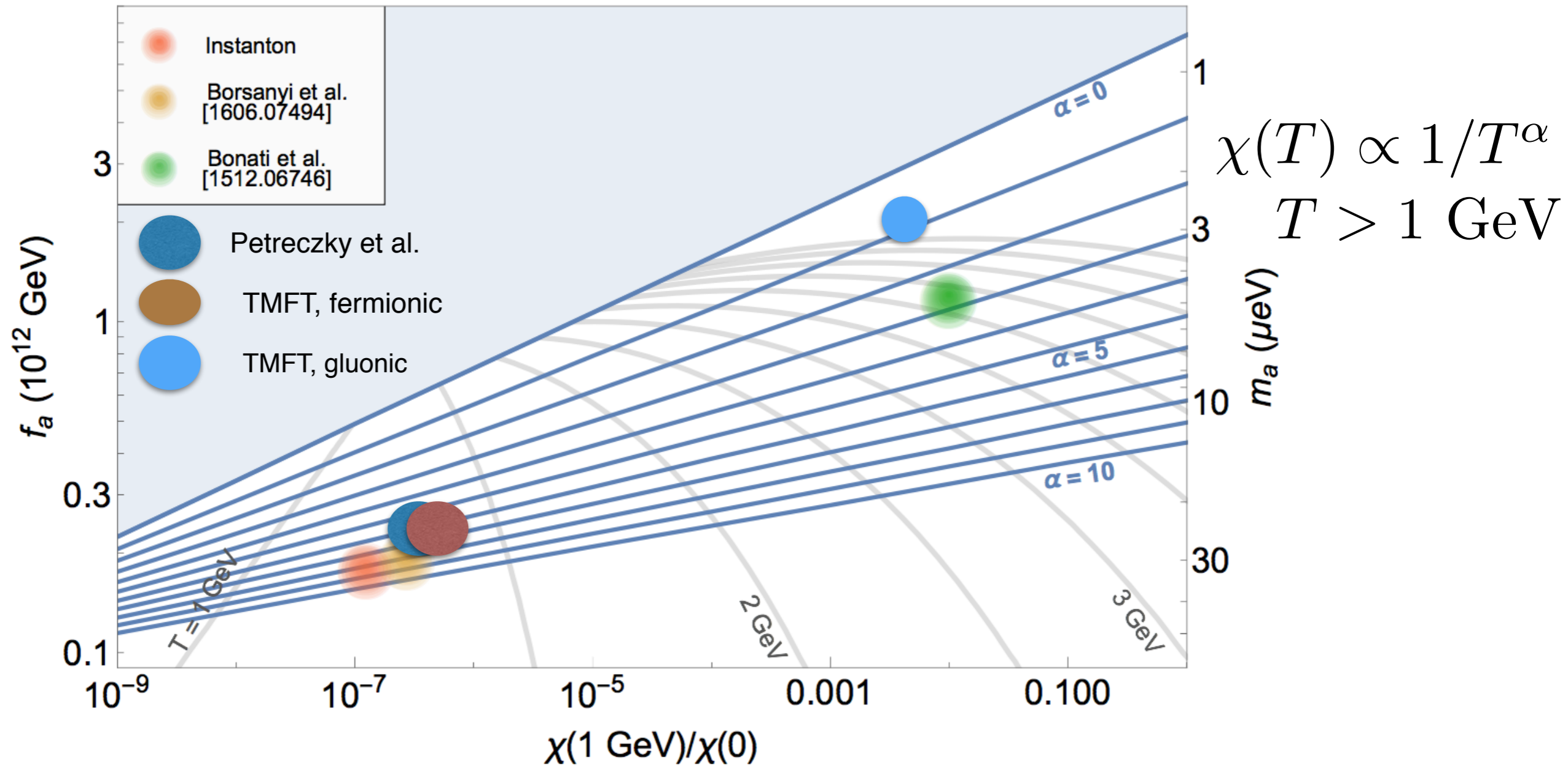
$$\chi_{top}^{1/4} = aT^{-d(T)}$$



Extrapolation justifiable??

Needed assumption on fraction of DM made of axions

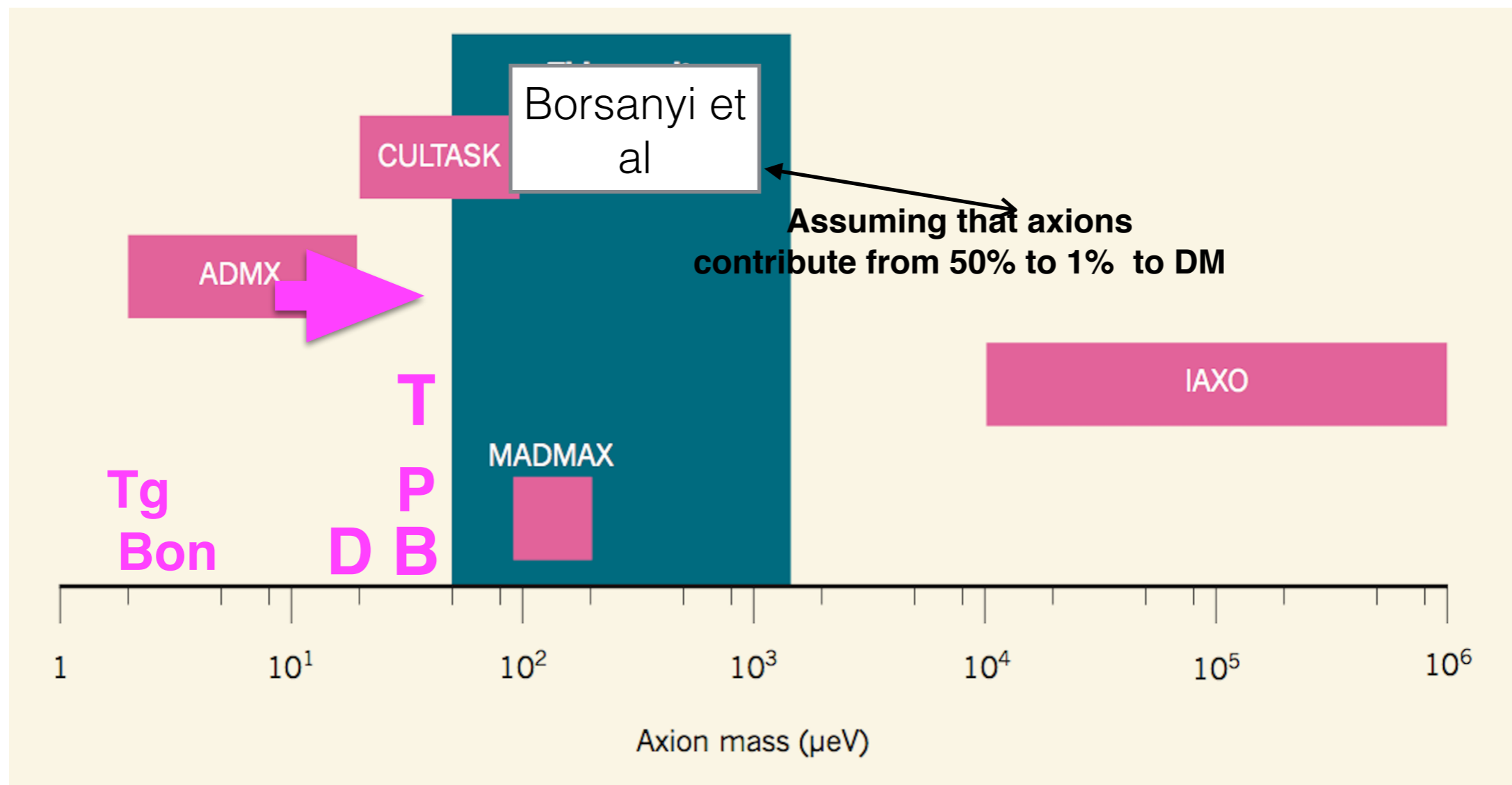
*Assume:* Axions make all of Dark Matter



PhD Thesis, G. Grilli di Cortona, Sissa 2016  
(advisor G. Villadoro)

# Lower limits on the axion mass assuming that axions make 100% of DM:

T<sub>g</sub>: Tmft, gluonic; B<sub>on</sub>: Bonati et al.; D: DIGA, B: Borsanyi et al.,  
P: Petreczky et al., T: Tmft, fermionic



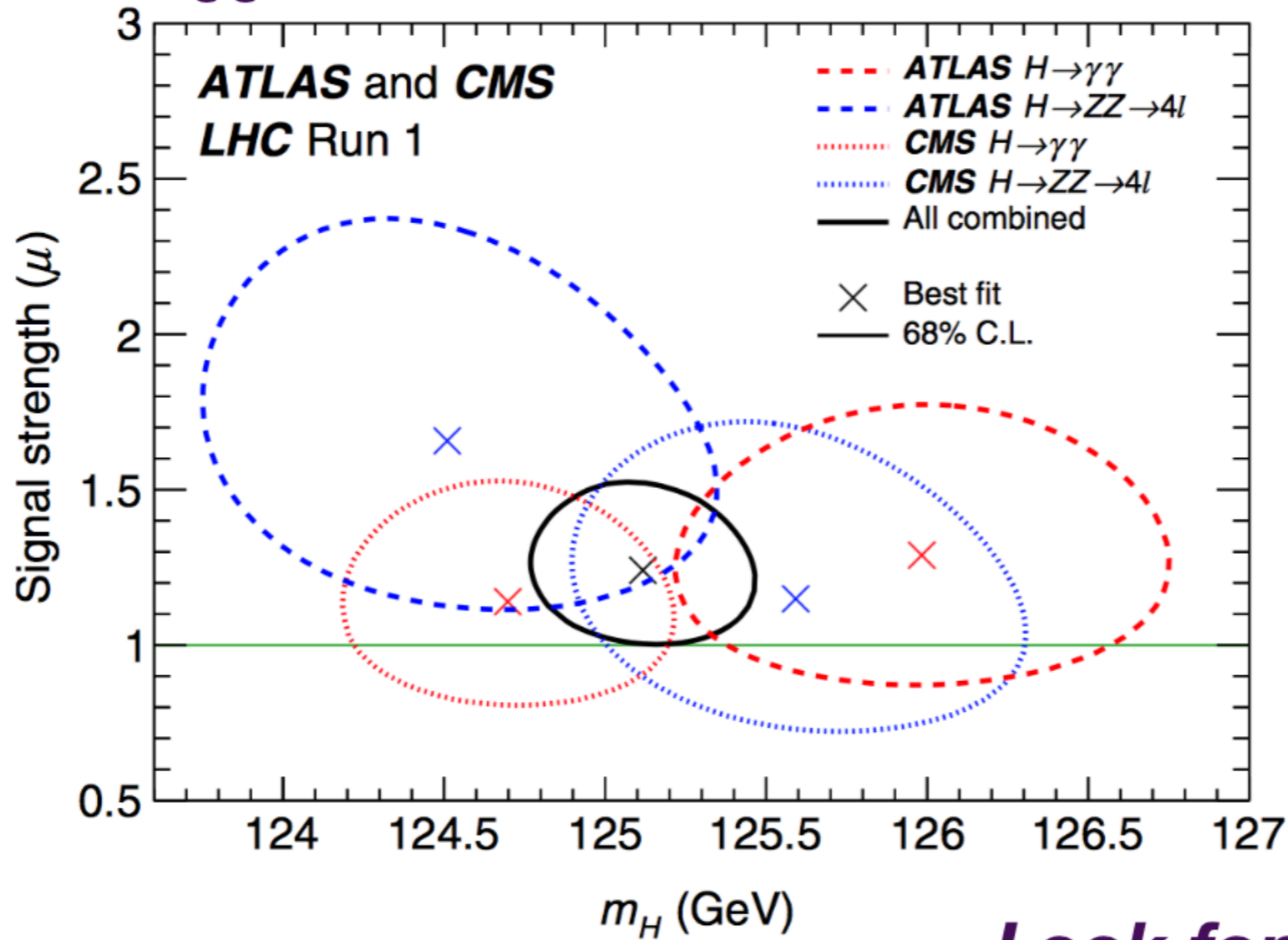
Updated from Nature N&V



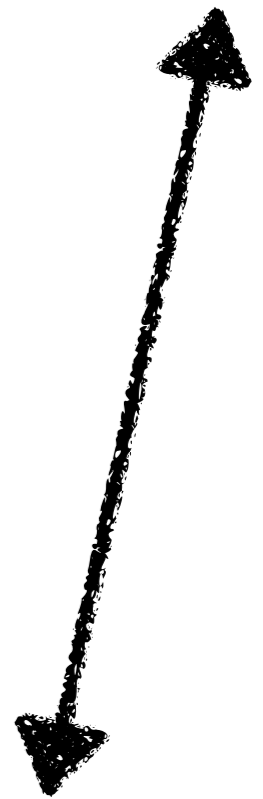
# Phases of QCD and Higgs physics

# Beyond the Standard Model:

Can we find a theory which produces a narrow Higgs-like status?



Maybe  
in the  
preconformal  
region?



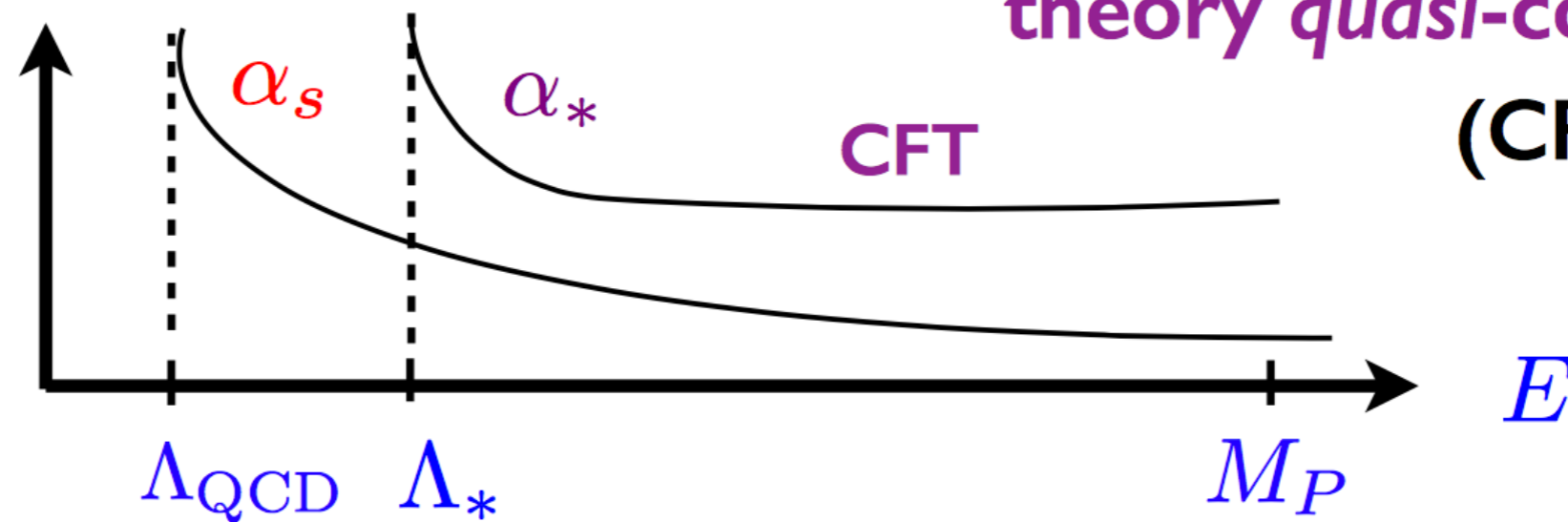
It should be a composite scalar particle, lighter than so-far unobserved composite vector states

**Look for theories with scale separation**

**..as possible BSM candidates**

# Explanation for the smallness of the EW-scale

QCD as an inspiration:



*Slide from A. Pommerol*

New strong dynamics at TeV

It could explain why  $m_H \lesssim \Lambda_* \sim \text{TeV} \ll M_P$

**Composite Higgs**

**Solves the problem in one shot!**

(in supersymmetry we still need strong dynamics to break susy at some low-scale)

# Adding flavours to strong interactions

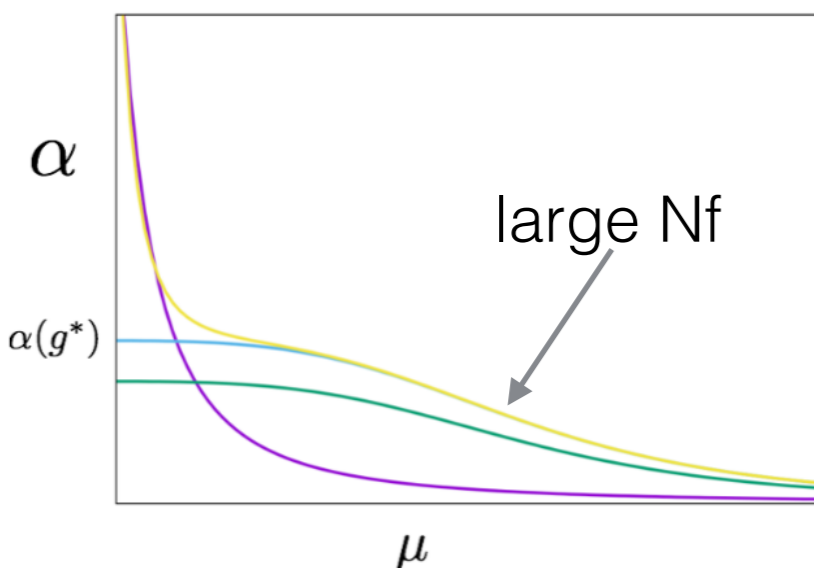


# Adding Flavours to QCD:

## *Another world*

Symmetry  
is  
preserved

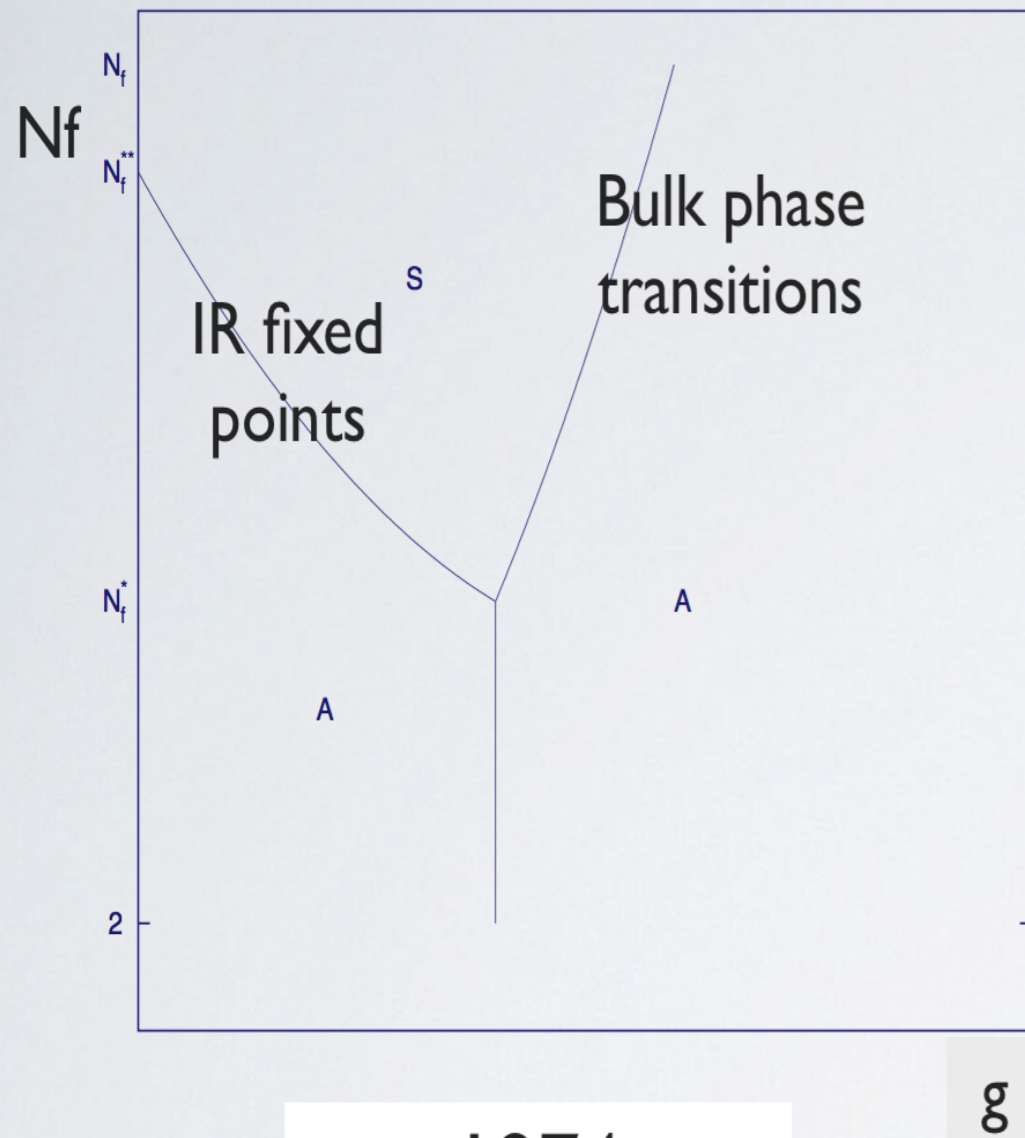
$$SU(N_f) \times SU(N_f) \rightarrow SU(N_f)$$



***Interplay between chiral symmetry and running of the coupling which depends on the flavor content***

For large  $N_f$   $\alpha_s$  is too small for  $\chi$  breaking

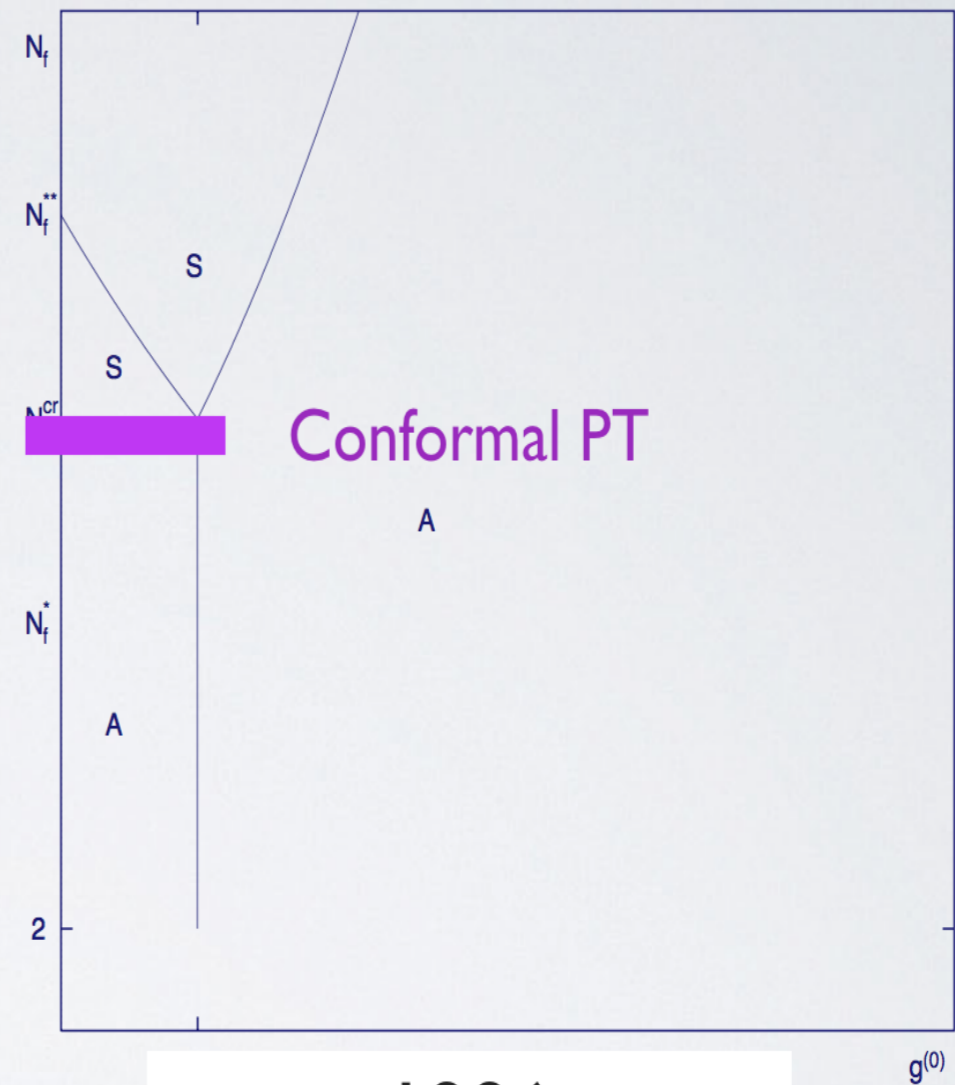
# Historical overview



1974

Banks Zacs

# Discovery of the QCD conformal window



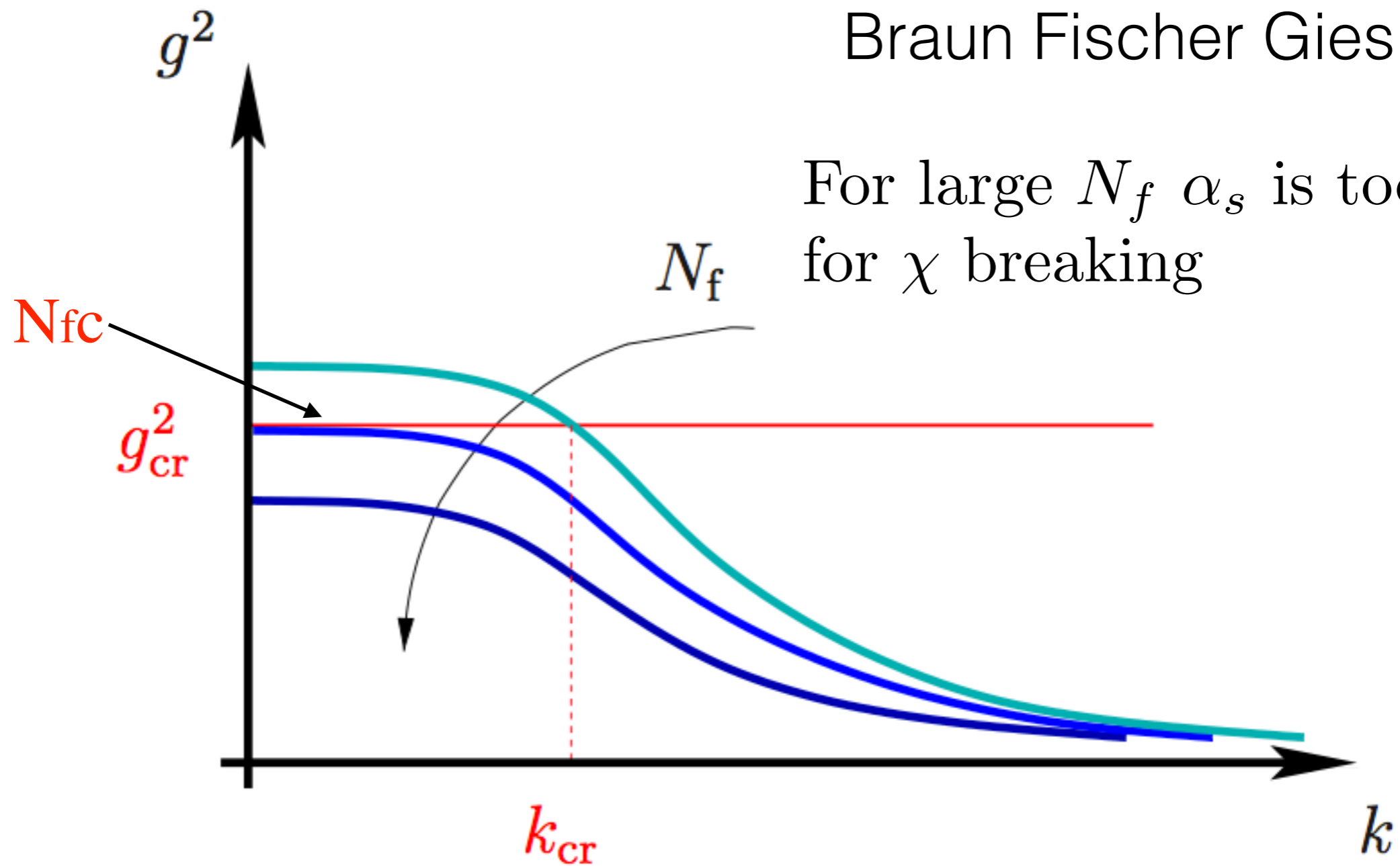
1996

Miransky Yamawaki

# IR running of the gauge coupling for different $N_f$ 's

Braun Fischer Gies 2010

For large  $N_f$   $\alpha_s$  is too small for  $\chi$  breaking



# Phases of Strong Interactions



Which are  
its features?

Can we  
observe it?



# Establishing the conformal window



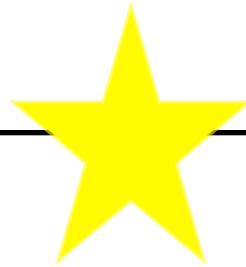
# Similarities and differences between a conformal PT and a 2nd order one

**Conformal transition**

IR IR IR IR IR.....



(X power-law)



Conformal scaling

*Analogies  
in the broken phase*

*Differences  
in the symmetric phase*

**EoS**

IR

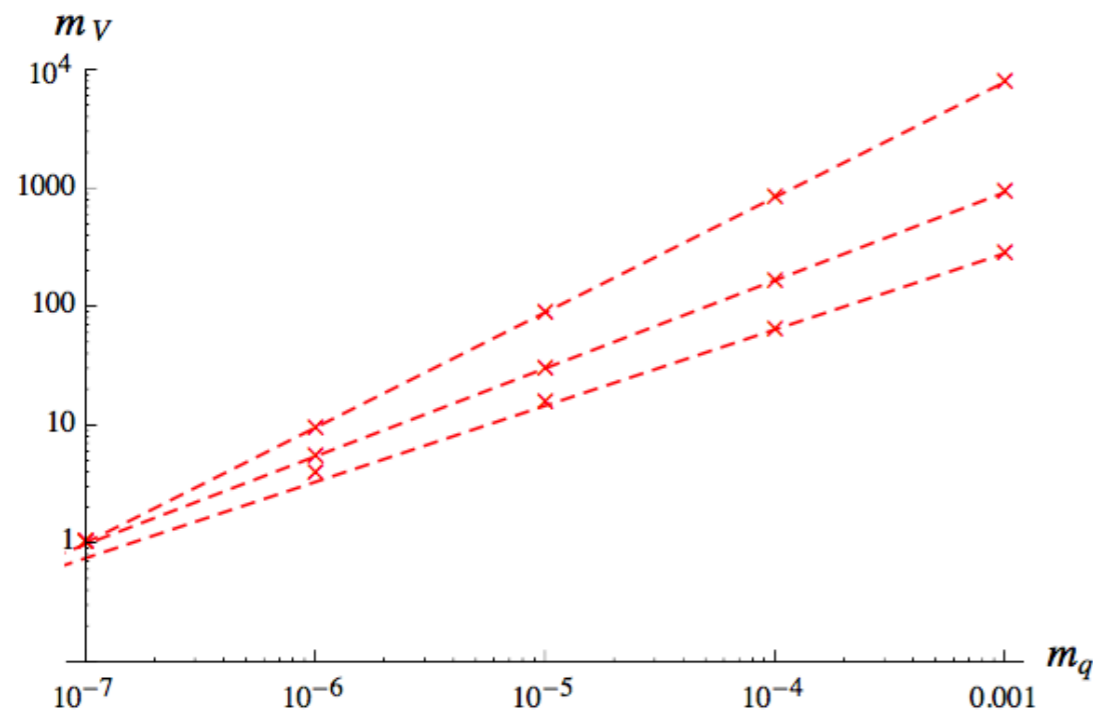


**2nd order transition**



Griffith's analyticity

# Conformal scaling



$$M_H = c_H m^{1/y_h}$$



Power law Scaling with anomalous dimension

Alho Evans Tuominen 2014

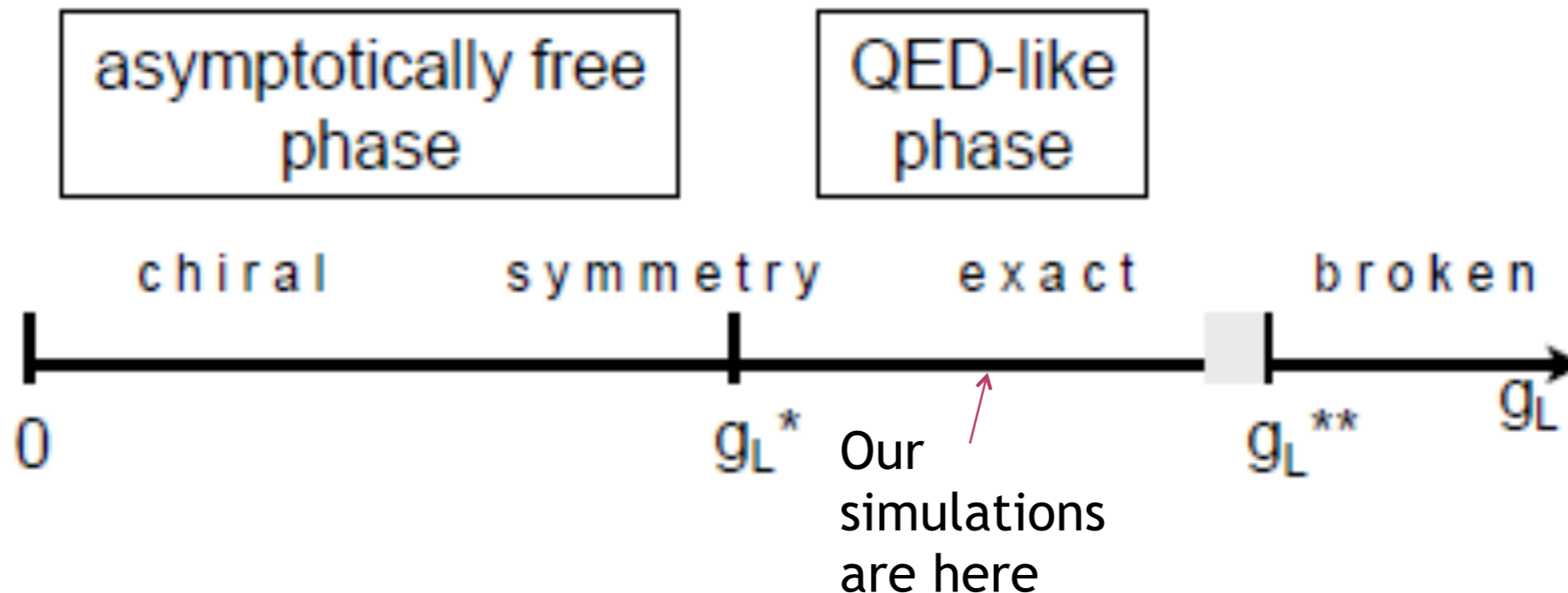
# Lattice corrections to conformal scaling

1: Size  $M_H = L^{-1} f_H(x), \quad x \equiv Lm^{1/y_m}$

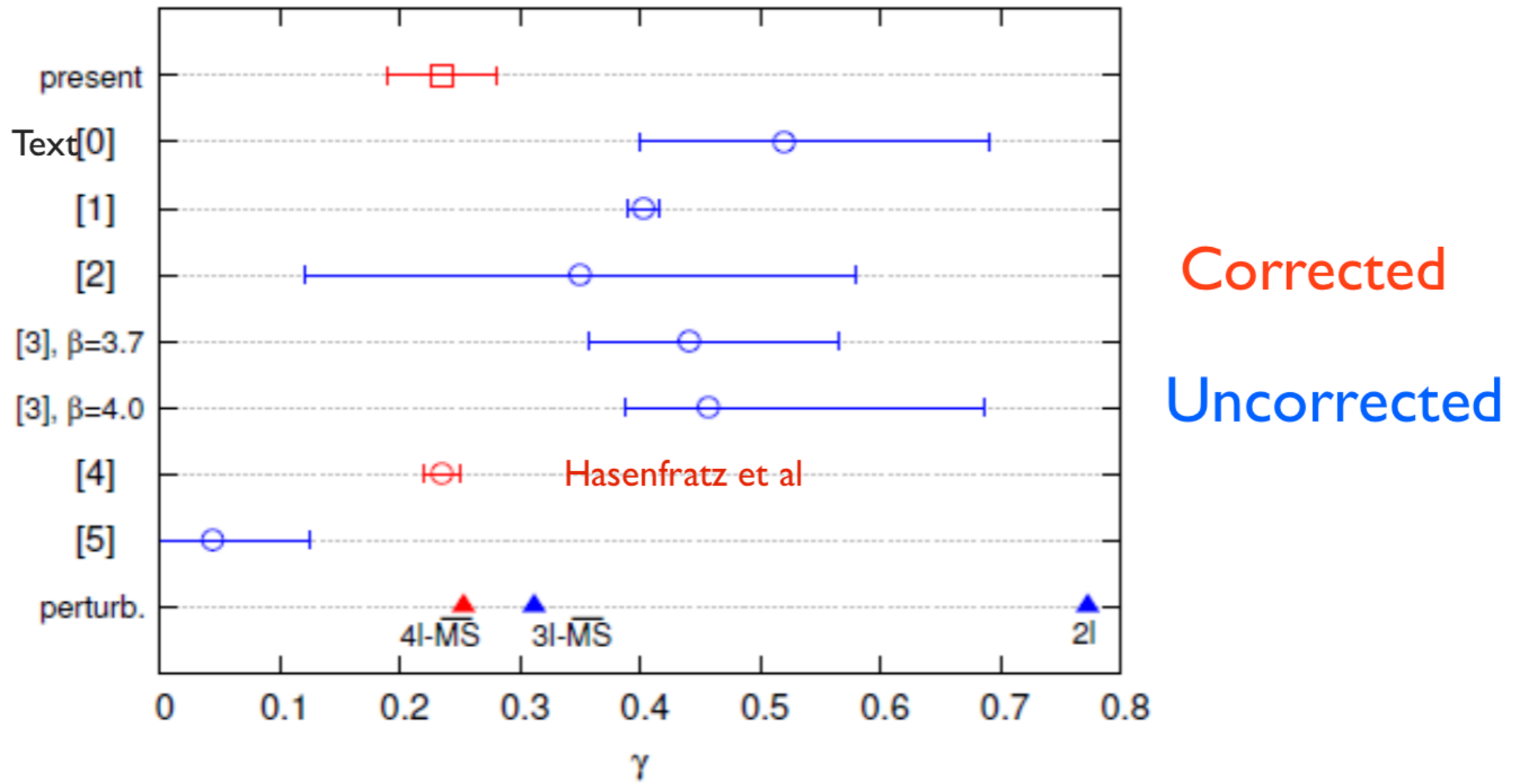
2: Coupling  $M_H = L^{-1} f_H(x, g_0 m^\omega)$

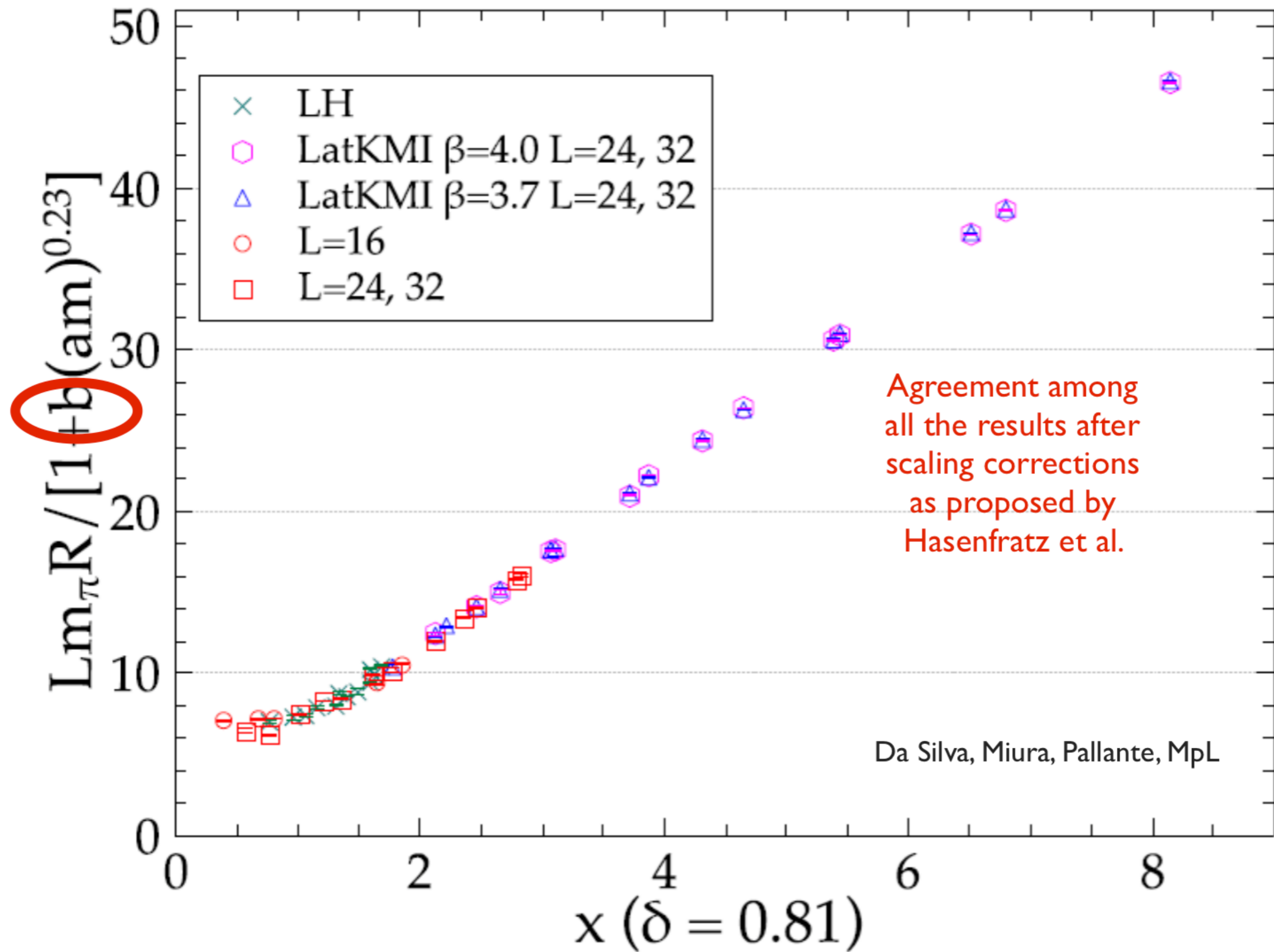
Del Debbio, Zwicky;  
Hasenfratz et al;  
MpL, da Silva, Miura, Pallante

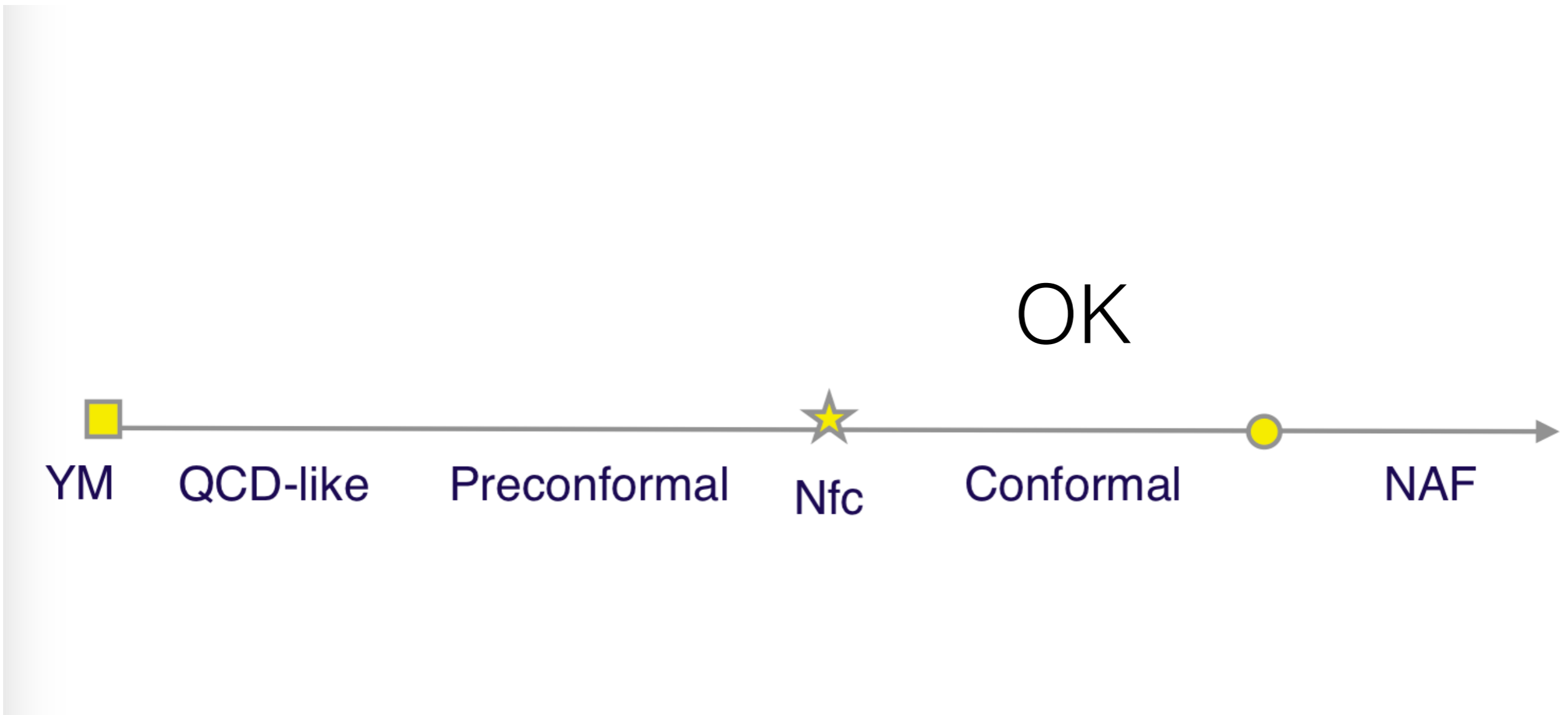
$$LM_H = F_H(x) \left\{ 1 + g_0 m^\omega G_H(x) + \mathcal{O}(g_0^2 m^{2\omega}) \right\}$$



- Anomalous dimension results with and without scaling corrections







YM

QCD-like

Preconformal

Nfc

Conformal

NAF

OK

# Next





# Standard picture of scale separation

$$\Lambda_{\text{IR}}/\Lambda_{\text{UV}} = \mathcal{O}(1).$$

$\Lambda_{\text{UV}}$

In the conformal phase IR scales vanish but UV ones survive

$\Lambda_{\text{IR}}$

$N_{\text{fc}}$

$$\frac{\Lambda_{\text{UV}}}{\Lambda_{\text{IR}}} \sim \exp\left(\frac{\hat{K}}{\sqrt{x_c - x}}\right)$$

$$x = N_f/N_c$$

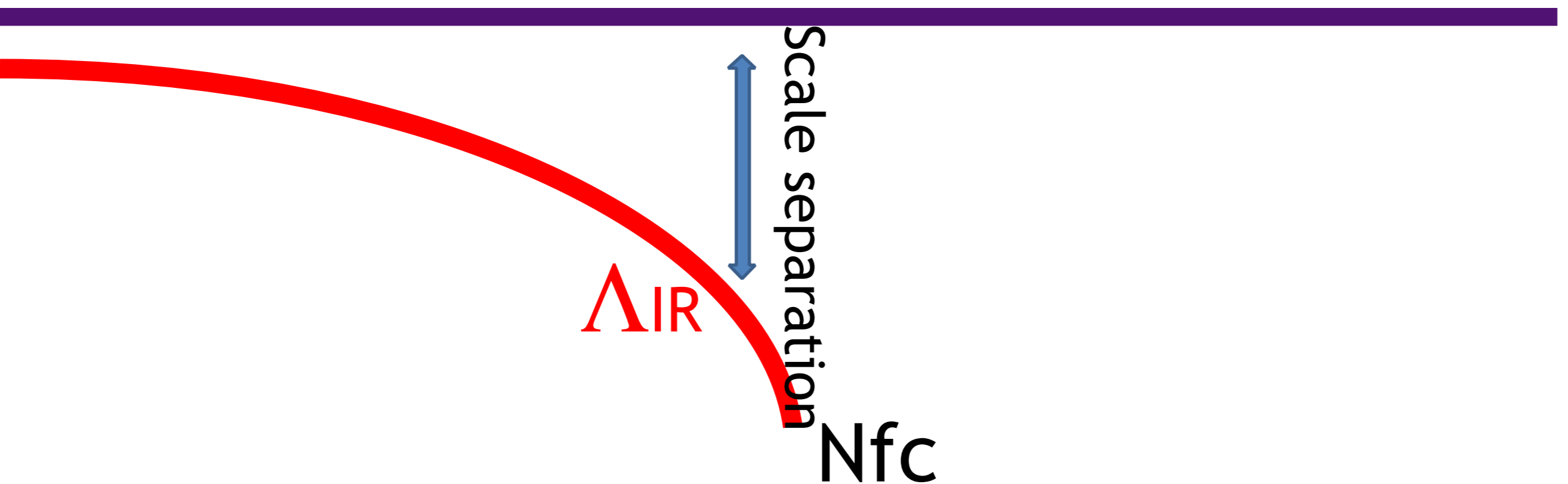
The coupling walks for

$$\Lambda_{\text{UV}}^{-1} \ll r \ll \Lambda_{\text{IR}}^{-1}$$

# Standard picture of scale separation

$$\Lambda_{\text{IR}}/\Lambda_{\text{UV}} = \mathcal{O}(1).$$

$\Lambda_{\text{UV}}$



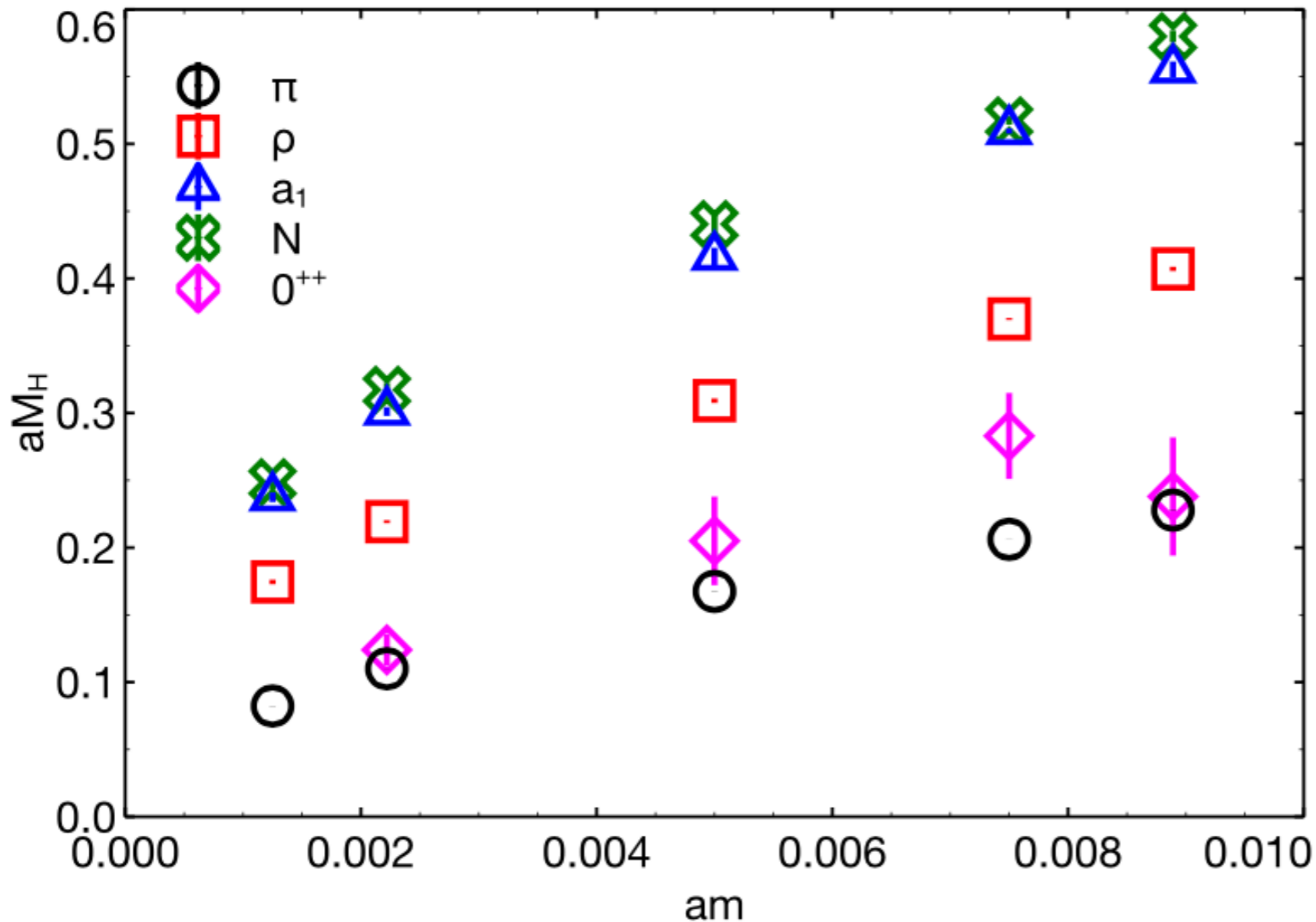
$$\frac{\Lambda_{\text{UV}}}{\Lambda_{\text{IR}}} \sim \exp\left(\frac{\hat{K}}{\sqrt{x_c - x}}\right)$$

# Strongly interacting dynamics and the search for new physics at the LHC

T. Appelquist,<sup>1</sup> R. C. Brower,<sup>2,3</sup> G. T. Fleming,<sup>1,3</sup> A. Hasenfratz,<sup>4,3</sup> X. Y. Jin,<sup>5</sup> J. Kiskis,<sup>6</sup> E. T. Neil,<sup>4,7,3</sup>  
J. C. Osborn,<sup>5,3</sup> C. Rebbi,<sup>2</sup> E. Rinaldi,<sup>8,3</sup> D. Schaich,<sup>9,3,10</sup> P. Vranas,<sup>8</sup> E. Weinberg,<sup>11</sup> and O. Witzel<sup>12,3</sup>

(Lattice Strong Dynamics (LSD) Collaboration)

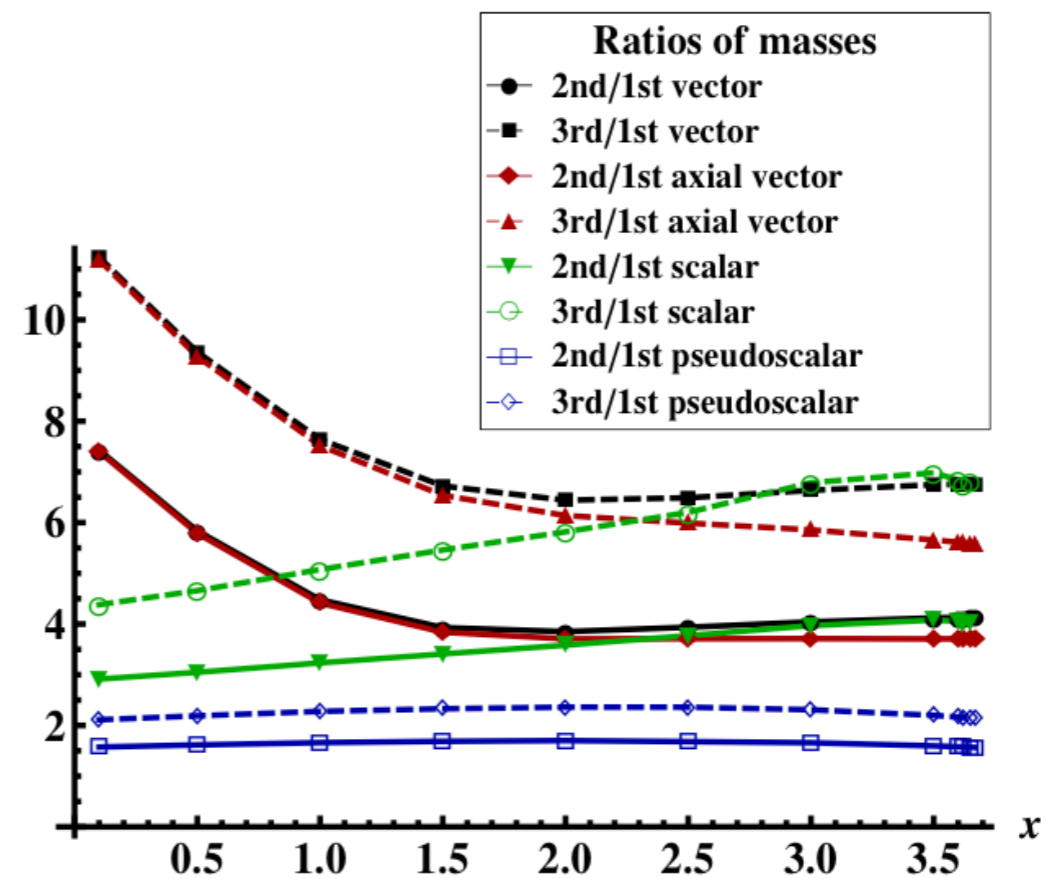
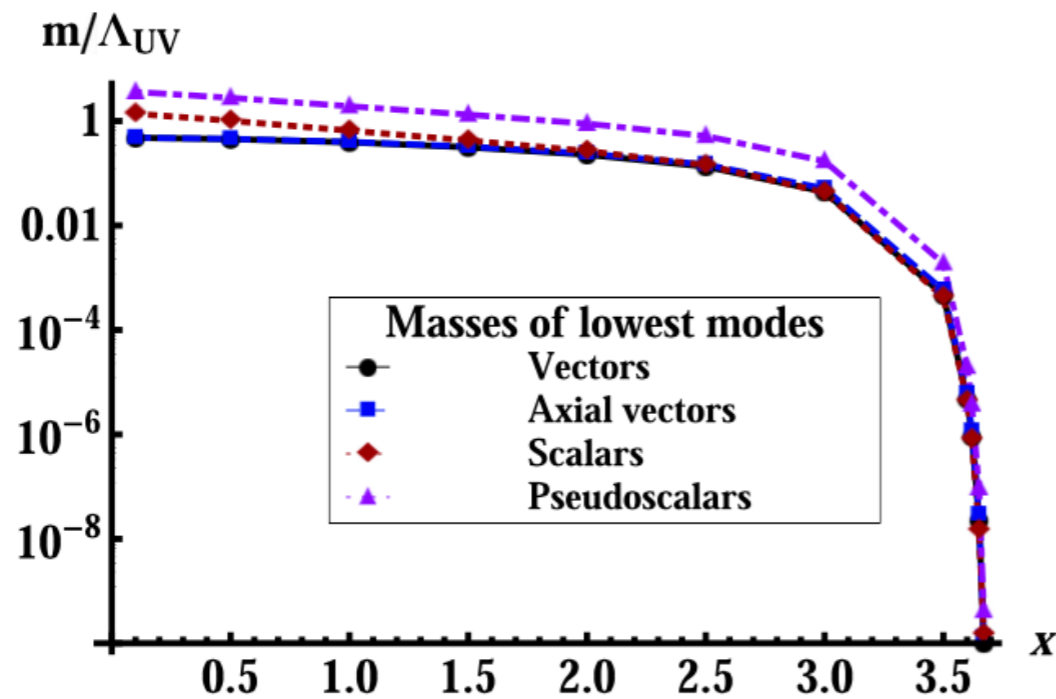
Beyond scale separation:



hierarchy of scales  
in the spectrum:  
light scalar  
needed for  
phenomenology

(Essential)  
singularity in the  
chiral limit and mass ratios:  
example from holographic  
V-QCD

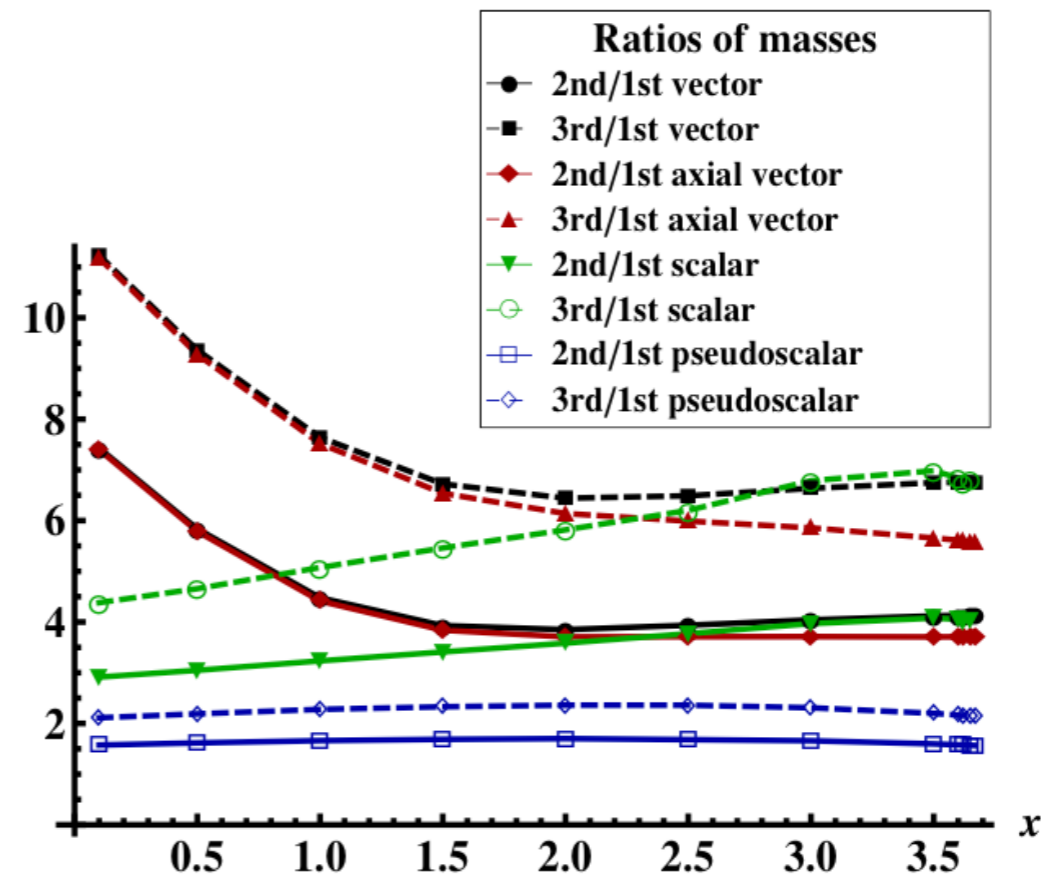
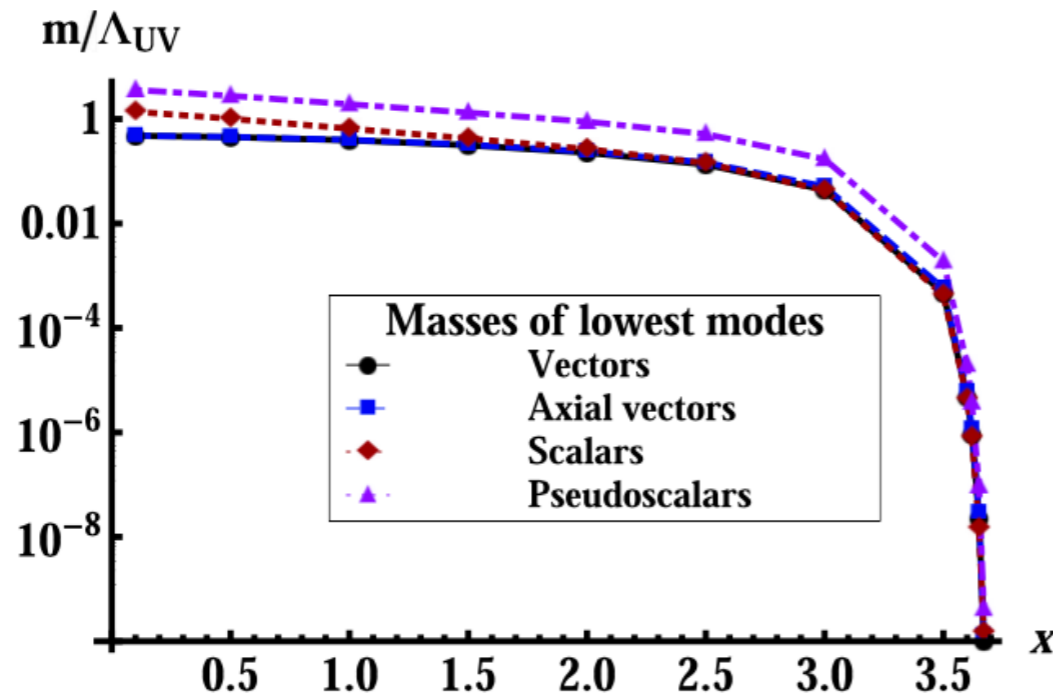
$$\frac{\Lambda_{UV}}{\Lambda_{IR}} \sim \exp\left(\frac{\hat{K}(n+1)}{\sqrt{x_c - x}}\right)$$



(Essential)  
singularity in the  
chiral limit and mass ratios:  
example from holographic  
V-QCD

$$\frac{\Lambda_{UV}}{\Lambda_{IR}} \sim \exp\left(\frac{\hat{K}(n+1)}{\sqrt{x_c - x}}\right)$$

*Not unique*



$\overline{\Lambda_{\text{IR}}}$  not unique:

## Power-law corrections to essential singularity

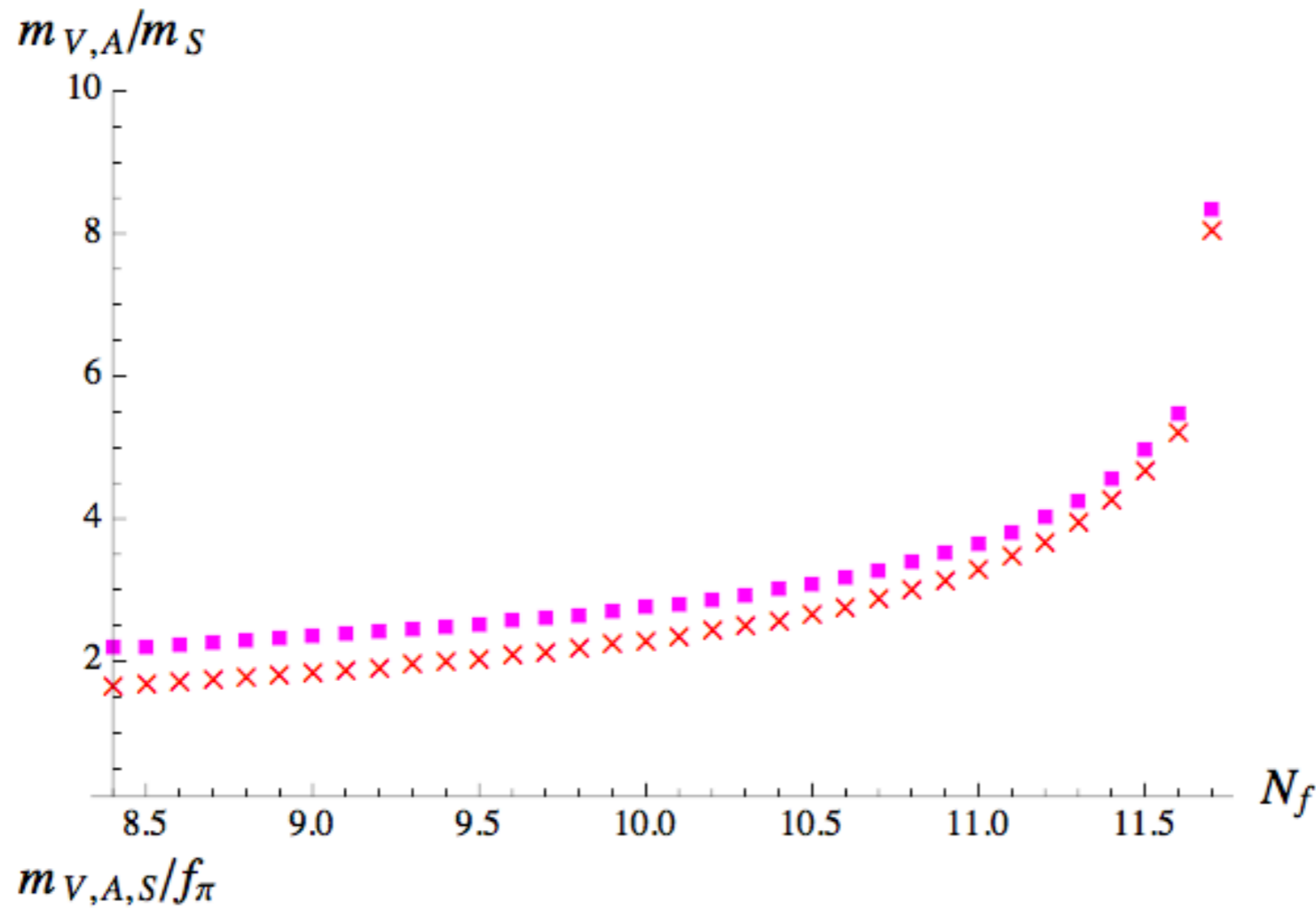
Gies et al. 2013

Alho, Evans, Tuominen 2013

$$O_i = A_i (N_f^c - N_f)^{p_i} \langle \bar{q}q \rangle^{1/3}$$

Power-law X Miranski scaling

May account for hierarchy of scales



# Mass deformed theory I: EoS approach for IR quantities

$$y = f(x)$$

$$y = m / \langle \bar{\psi}\psi \rangle^\delta \quad \delta = \frac{6-\eta}{2-\eta}$$

Second order transition:

$$x = (N_f^c - N_f) / \langle \bar{\psi}\psi \rangle^{\frac{1}{\beta}} \quad \langle \bar{\psi}\psi \rangle = (N_f^c - N_f)^\beta$$

Essential singularity:

Nogawa, Hasegawa, Nemoto, 2012

$$x = e^{\sqrt{(N_f^c - N_f)}} / \langle \bar{\psi}\psi \rangle \quad \langle \bar{\psi}\psi \rangle = e^{\sqrt{(N_f^c - N_f)}}$$

Continuity of  $f(x)$  plus asymptotic forms for  $m \rightarrow 0$  and  $N_f \rightarrow N_f^c$  imply

$$\langle \bar{\psi}\psi \rangle \propto e^{\sqrt{(N_f^c - N_f)}} \text{ for } m \text{ smallish and } (N_f^c - N_f) \text{ largish}$$

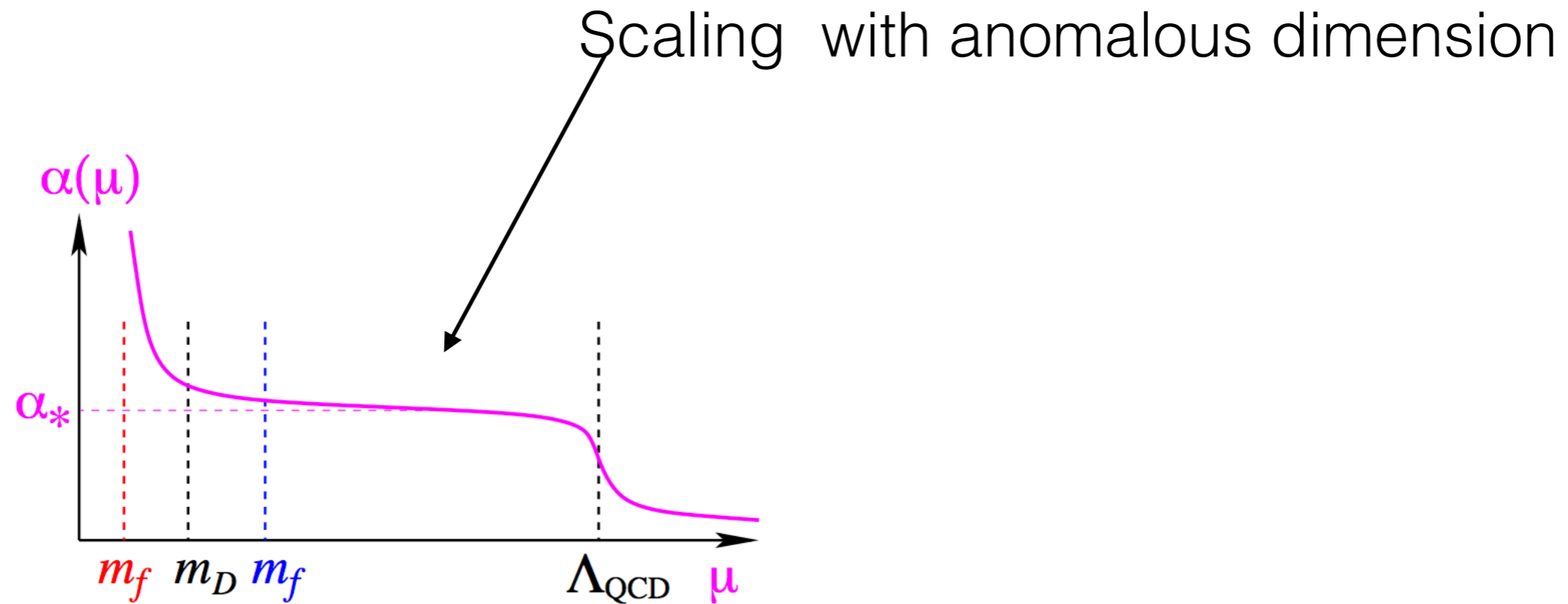
$$\langle \bar{\psi}\psi \rangle \propto m^{1/\delta} \text{ for } m \text{ largish and } (N_f^c - N_f) \text{ smallish}$$

Anomalous dimension appears naturally below  $N_f^c$

Scaling limited by Goldstone singularities in the chiral limit (Wallace Zia)

# Mass deformed theory II: KMI discussion

Mutatis mutandis,  
Eos approach reproduces KMI scenario:

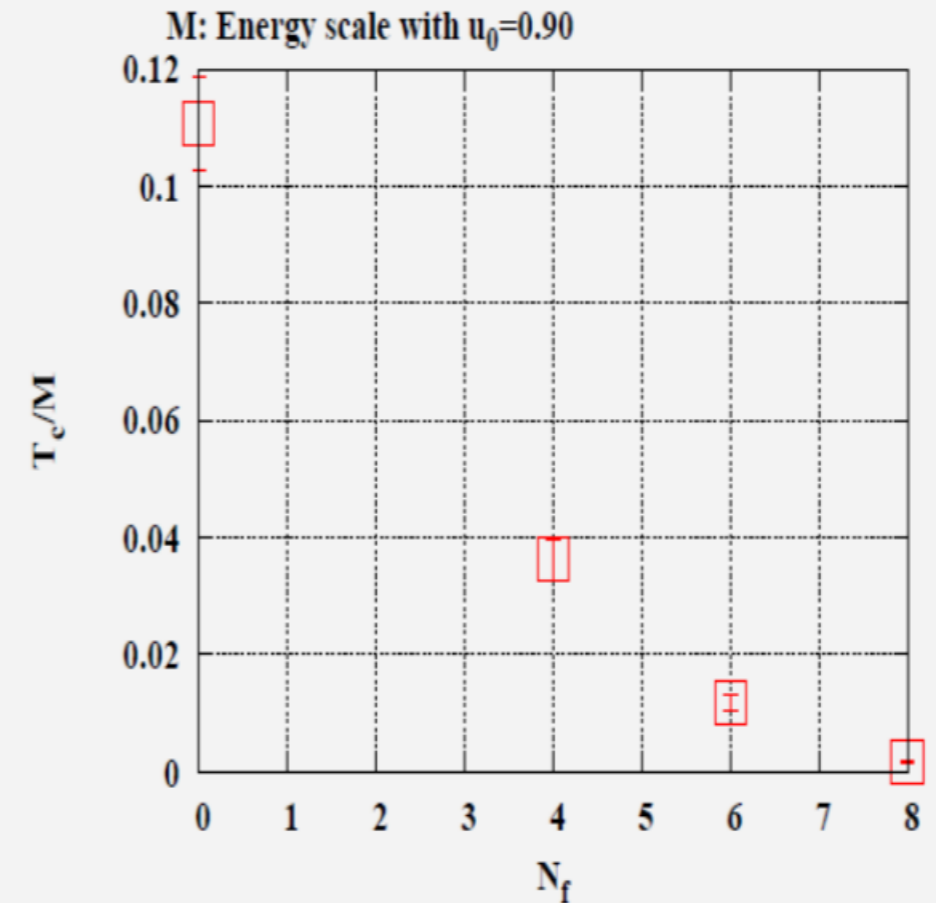
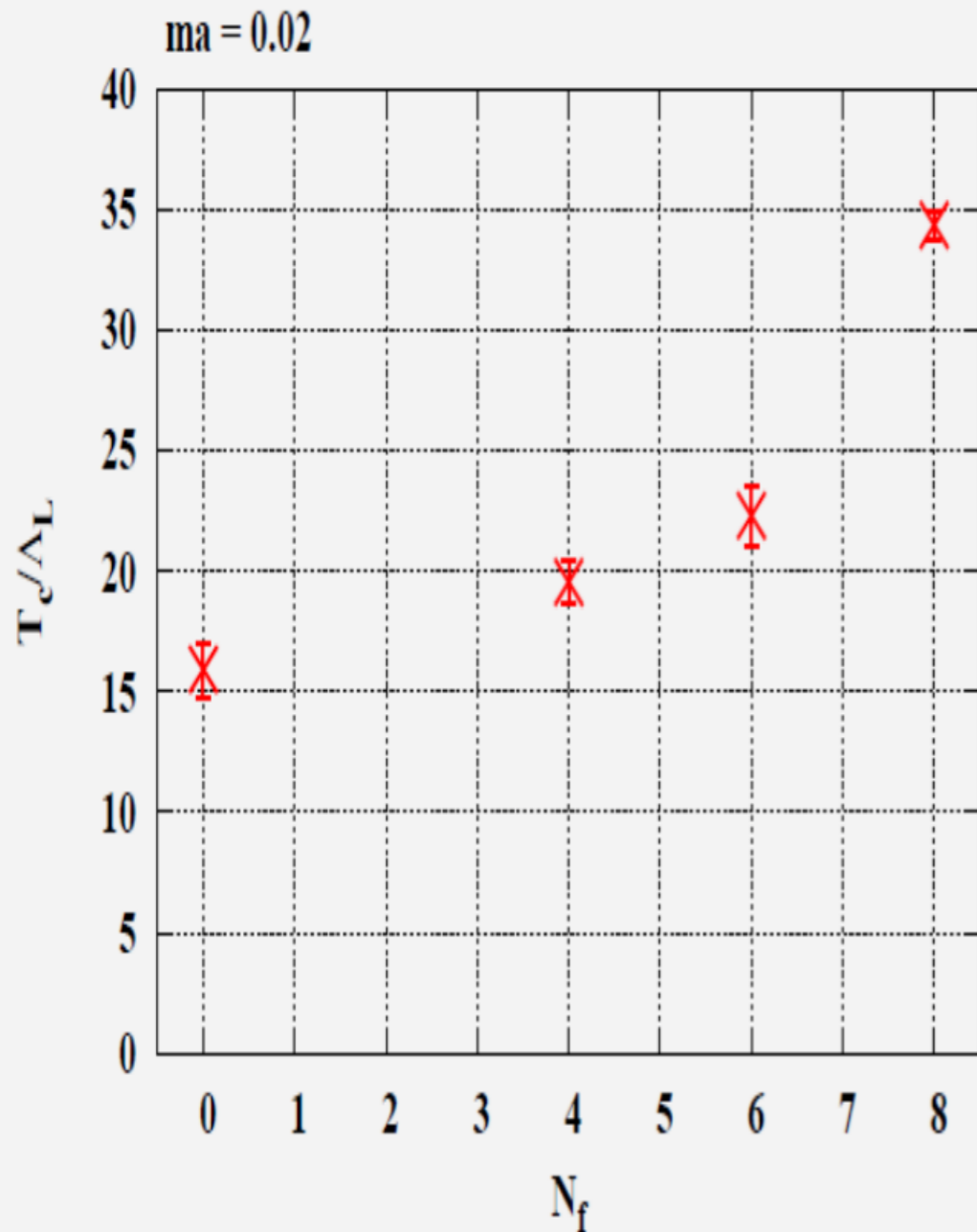


KMI 2013



# Search for scale hierarchy -

Kohtaroh Miura, MpL, Tiago Nunes da Silva, E Pallante

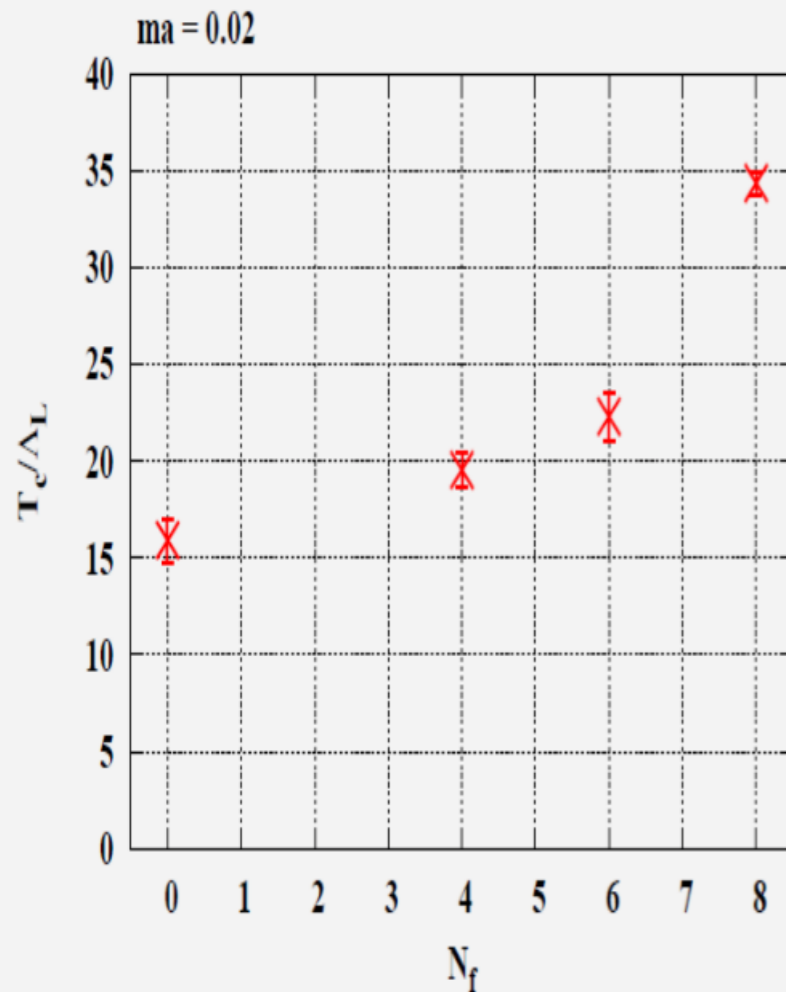


$$\frac{T_c}{M} = \frac{1}{N_t} \exp \left[ \int_{g_{\text{ref}}}^{g_c} \frac{dg}{B(g)} \right].$$

# Towards a quantitative comparison with holography

K. Miura, MpL, E. Pallante, in progress

$$\frac{2\pi T_c}{M_{KK}} = 1 - \frac{1}{126\pi^3} \lambda^2 \frac{N_f}{N_c} \left( 1 + \frac{12\pi^{3/2}}{\Gamma(-\frac{2}{3}) \Gamma(\frac{1}{6})} \right)$$



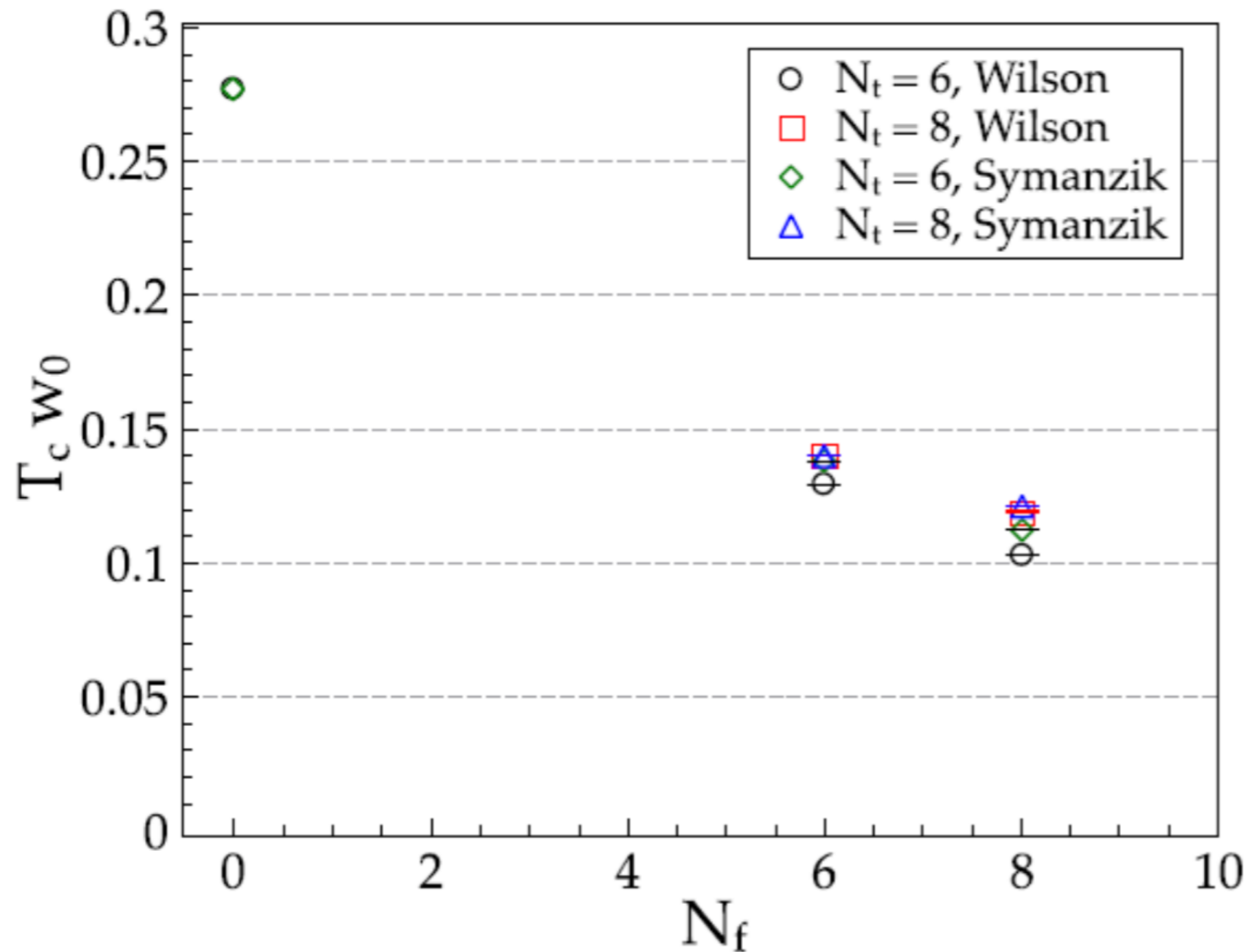
Bigazzi and Cotrone, JHEP 2015



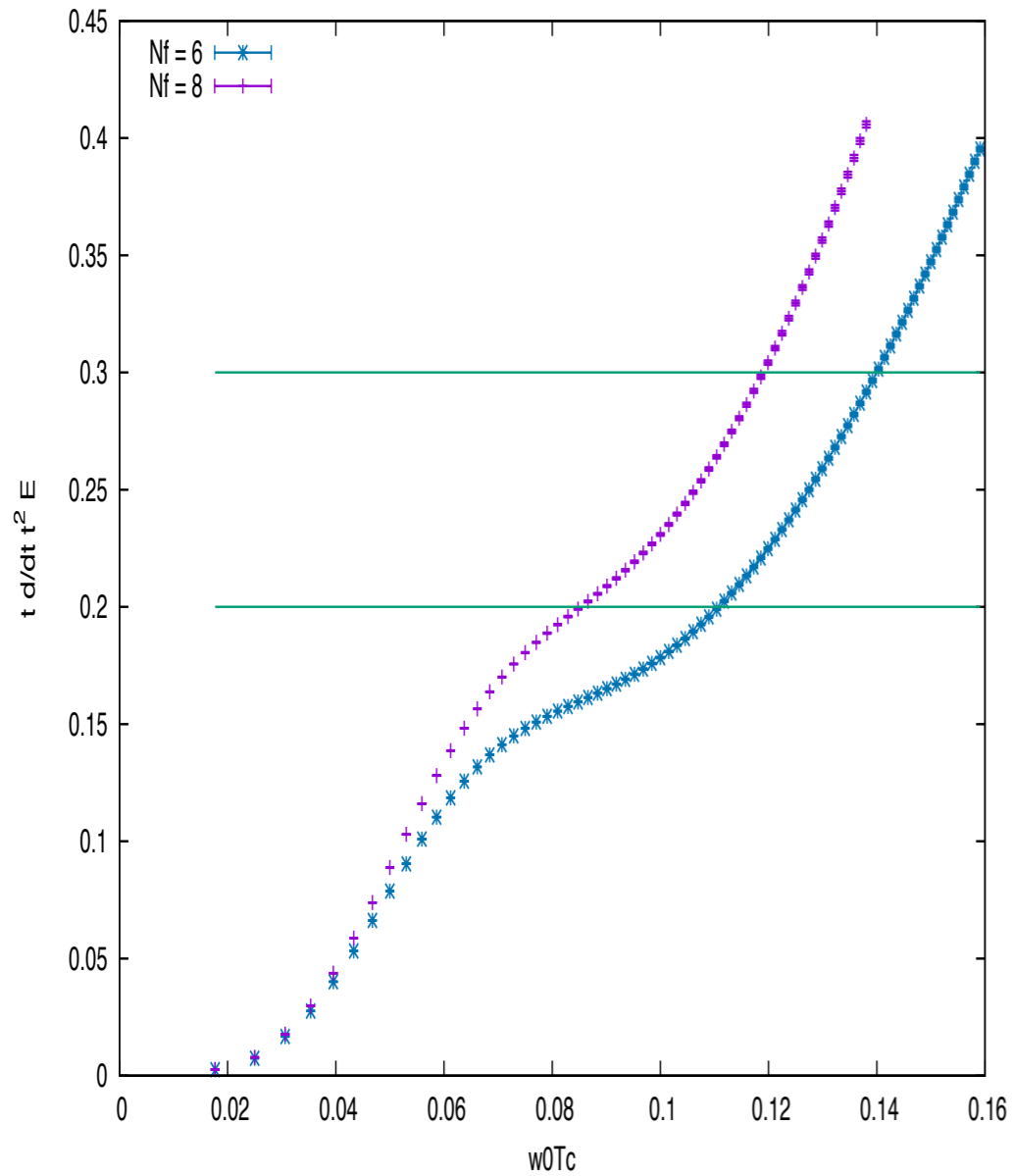
$$\left( 1 + \frac{12\pi^{3/2}}{\Gamma(-\frac{2}{3}) \Gamma(\frac{1}{6})} \right) \approx -1.987$$

T increases with  $N_f$  on the scales used in these two studies

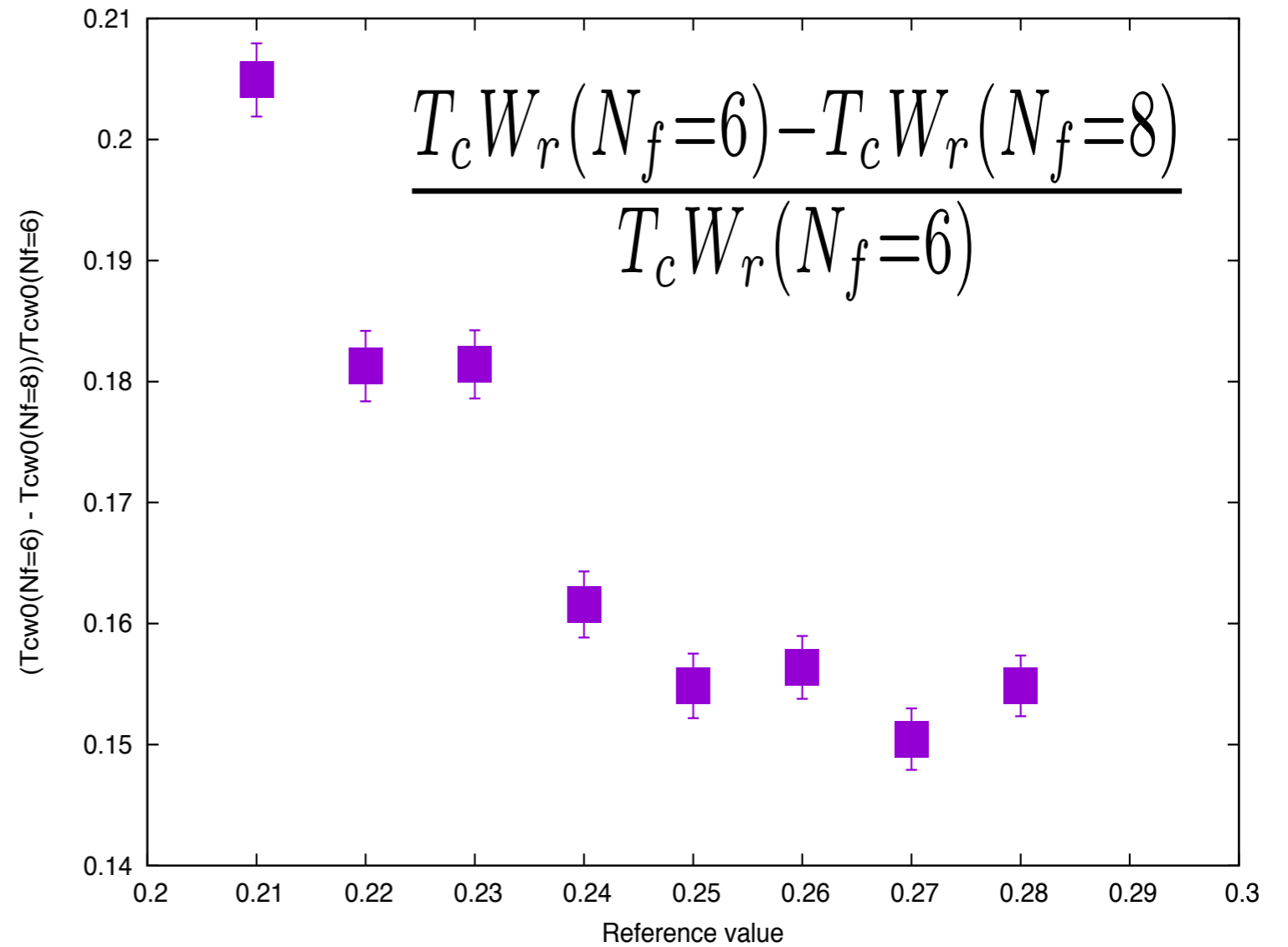
# Tc on the 1/w0 scale



# Moving the scale with Wilson flow



Qualitatively as expected,  
limited by lattice artifacts



UV

# Tc and the string tension

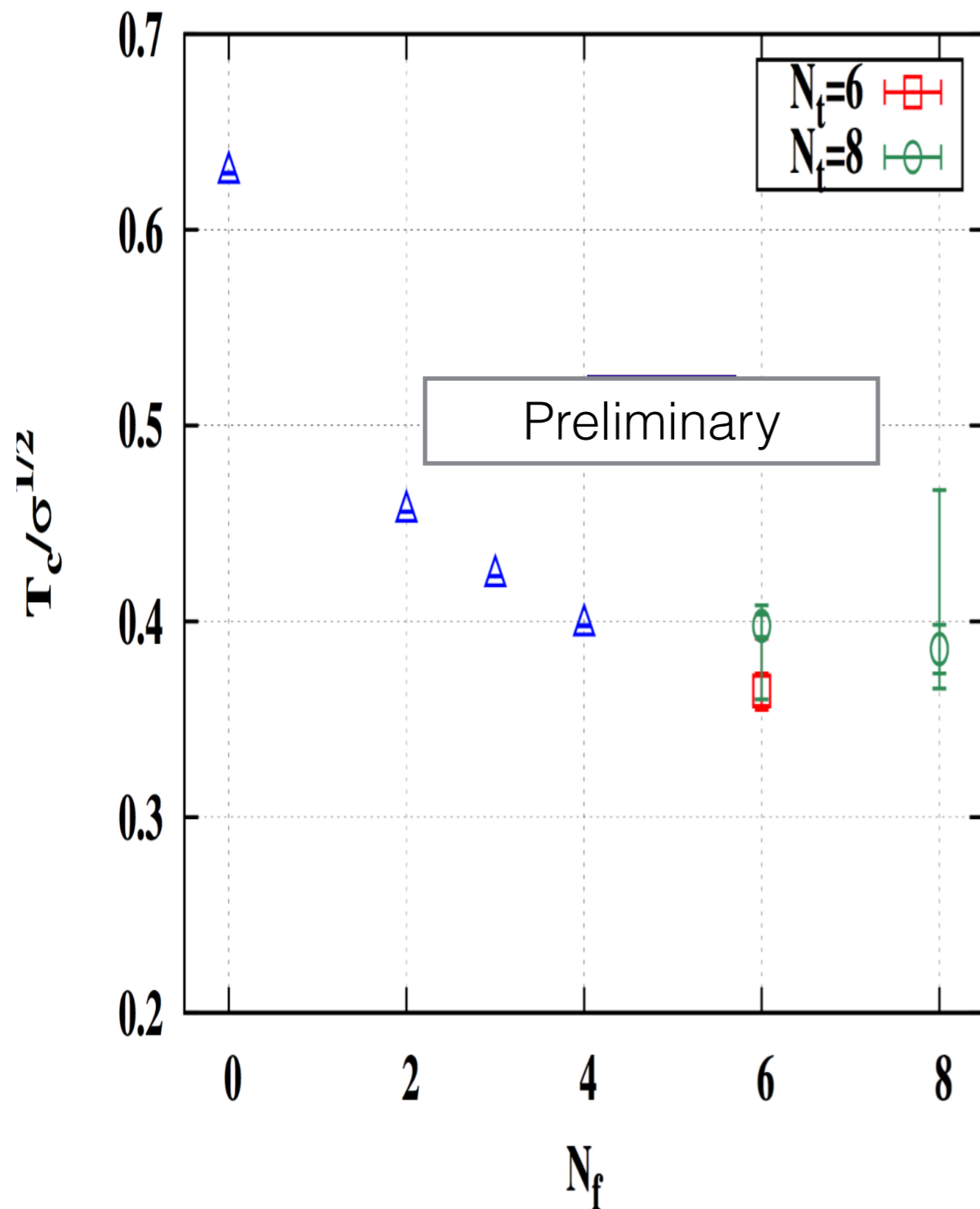
KM, MpL, EP, in progress

Mild decrease, possibly constant as  $N_f \rightarrow N_f^c$

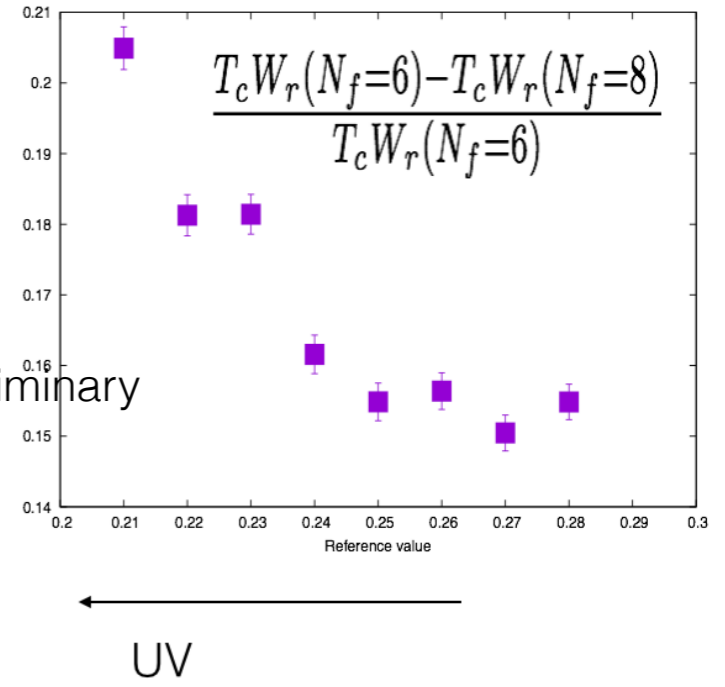
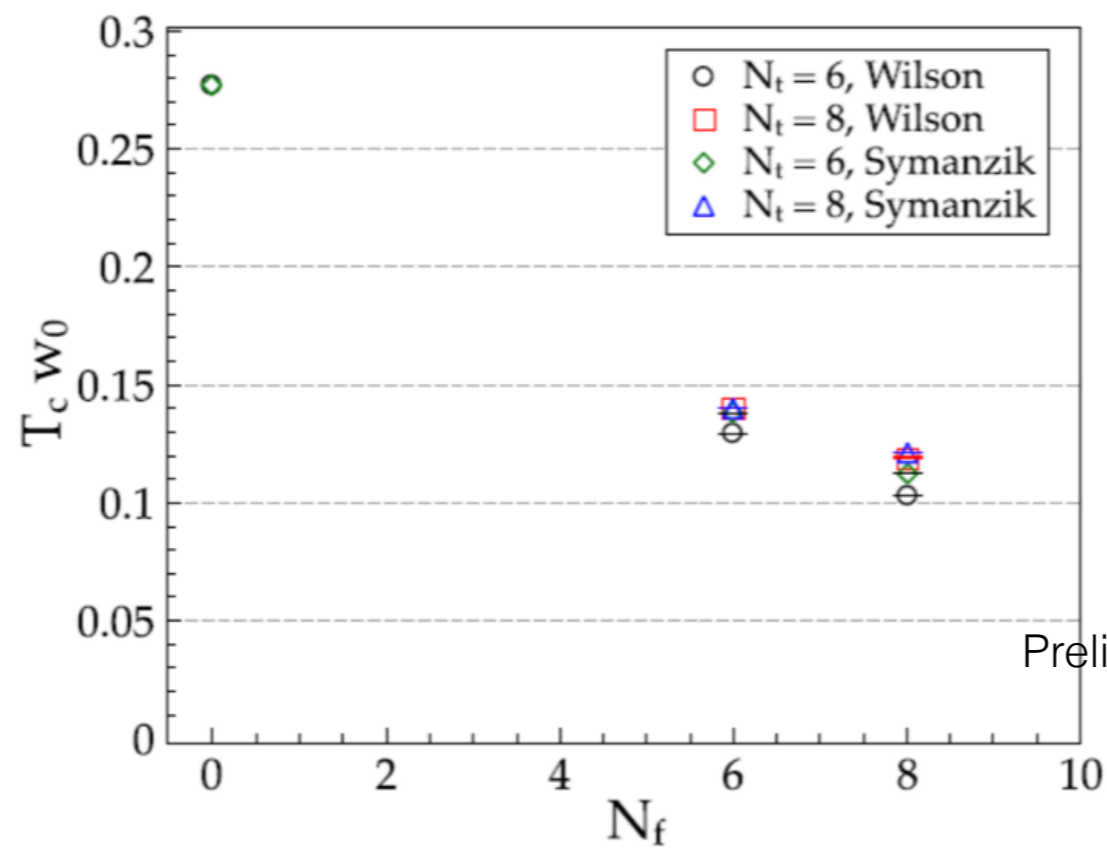
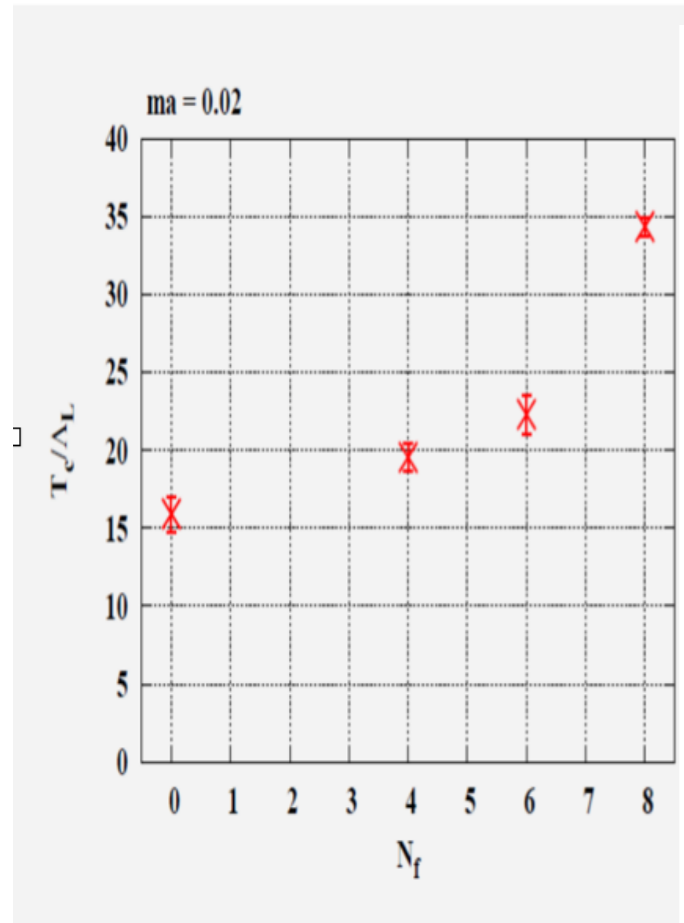
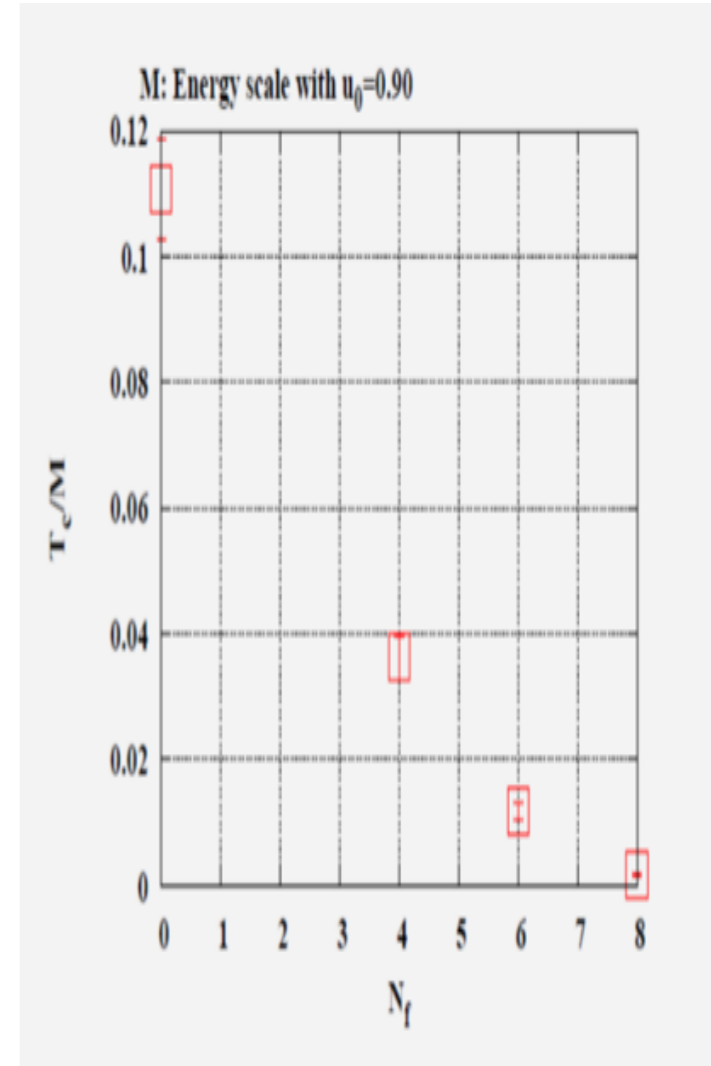
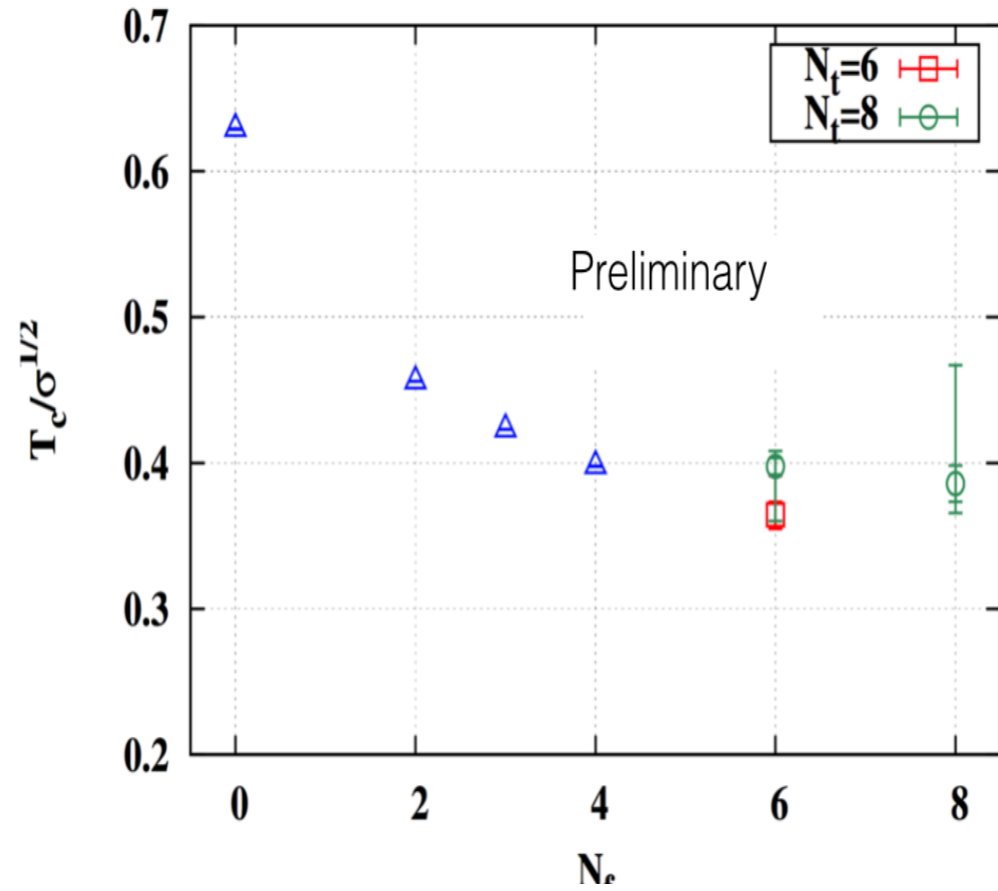
Again similar to the prediction of the WSS model:

$$\frac{T_c}{\sqrt{\sigma}} \propto (1 - \epsilon N_f / N_c)$$

communicated by F. Bigazzi



# Hierarchy of scales in the near conformal phase



# Hierarchy of scales

$M$

$\Lambda_{UV}$

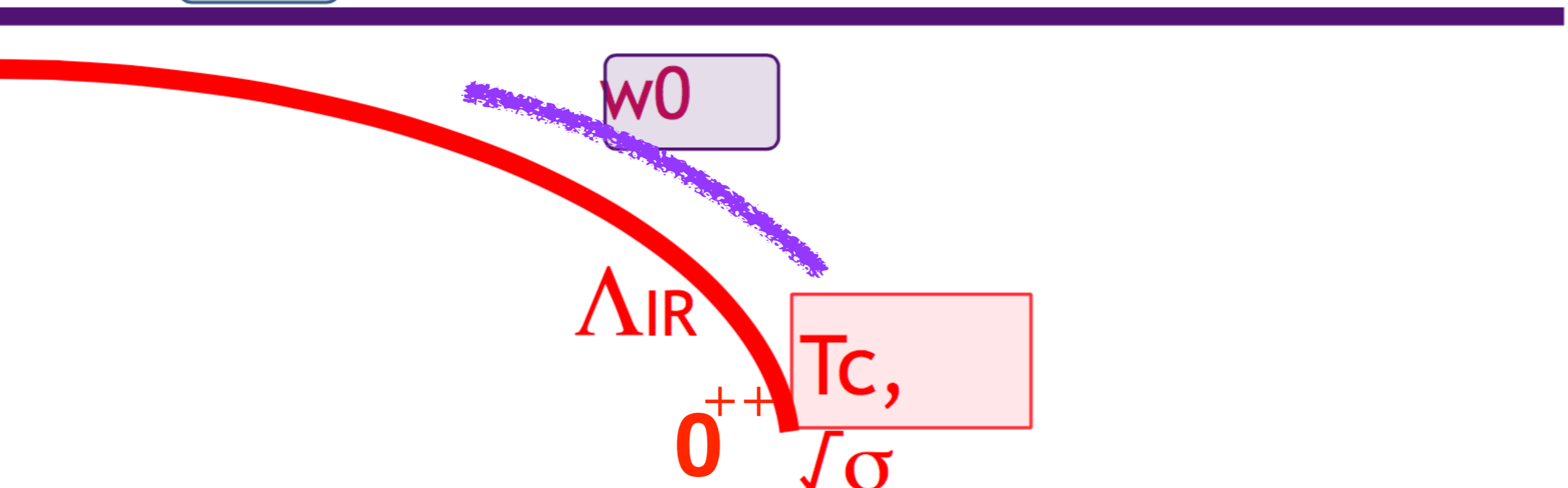
$w_0$

$\Lambda_{IR}$

$T_C,$

$0^{++}$

$\sqrt{\sigma}$



Summary: *Two examples of applications of topics discussed here which have a close connections to BSM physics*

I *Dark Matter: limit of the mass for post-inflationary Axions from QCD topology —>*

calls for more controlled results from topology

II *Higgs mass generation via QCD-like dynamics*

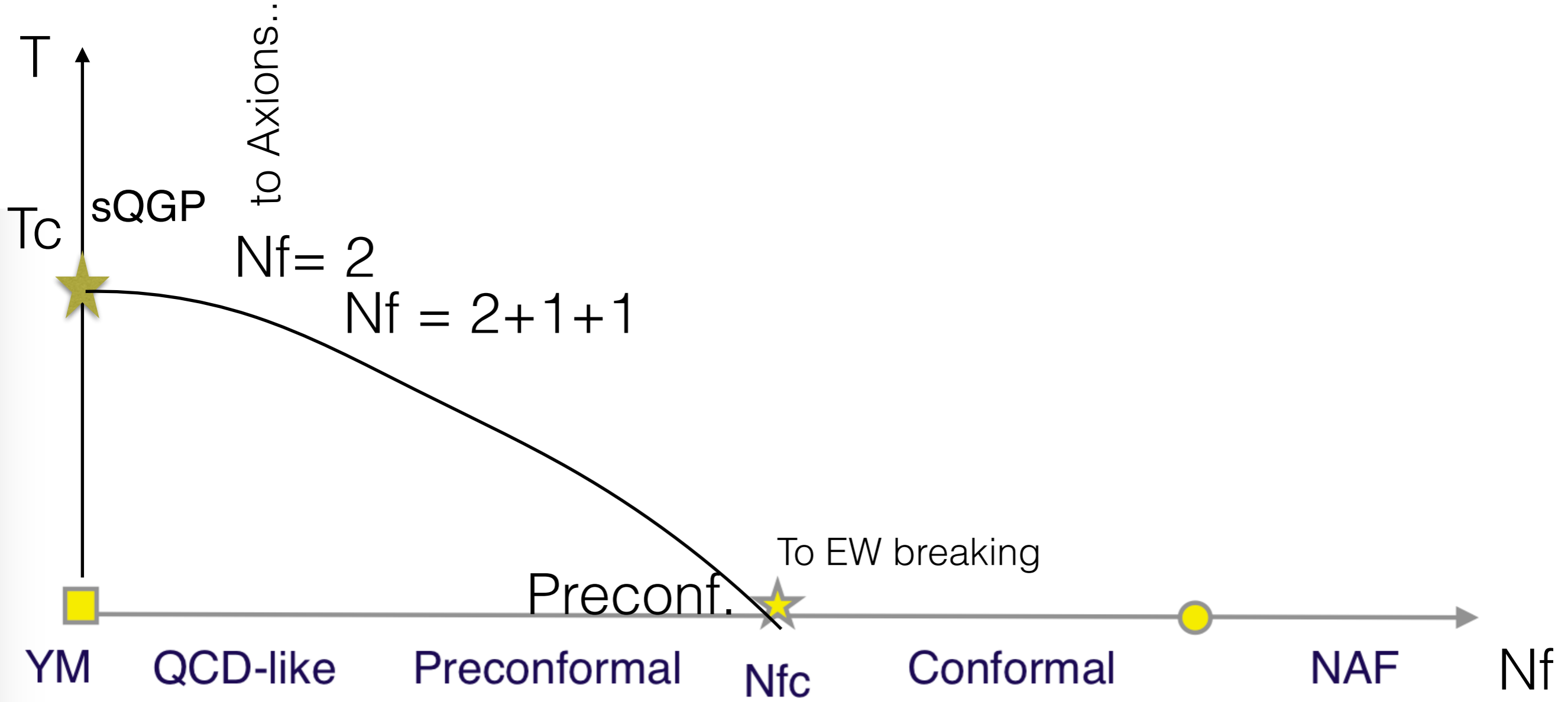
—>observed hierarchy of scales,

special role of the scalar —> further insight from lattice /

FRG?? connection with topology

(as advocated by E. Zmitchinski)





**Comment: interesting to study  
 the critical line in the  $T, N_f$  plane...  
 as a way to strongly coupled near conformal QGP..**

Thank You !

