

Antimatter Experiments at CERN

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Outline

- CPT Invariance and its Tests
- The Antiproton Decelerator at CERN
- ASACUSA: Atomic Spectroscopy And Collisions Using Slow Antiprotons
- ALPHA: Antimatter Laser PHysics Apparatus
- ATRAP: Antimatter trap
- AEGIS: Antimatter gravity
- ACE: Antimatter Cell Experiment
- AMS2: Alpha Magnetic Spectrometer
- Outlook: ELENA

Antimatter mysteries

- Why there is practically no antimatter in our Universe?
At the Big Bang particles and antiparticles should have been produced together. Where did antimatter go?
- Could they be hiding in parts of the Universe inaccessible for us?
- Could there be a tiny difference between particle and antiparticle to cause this asymmetry?
- Are there particles which are their own antiparticles (Majorana particles)? Could the dark matter of the Universe consist of such particles?
- Can antimatter be used for something in everyday life or is it just an expensive curiosity?

CPT Invariance

Charge conjugation: $C|\mathbf{p}(r, t)\rangle = |\bar{\mathbf{p}}(r, t)\rangle$

Space reflection: $P|\mathbf{p}(r, t)\rangle = |\mathbf{p}(-r, t)\rangle$

Time reversal: $T|\mathbf{p}(r, t)\rangle = |\mathbf{p}(r, -t)\rangle$

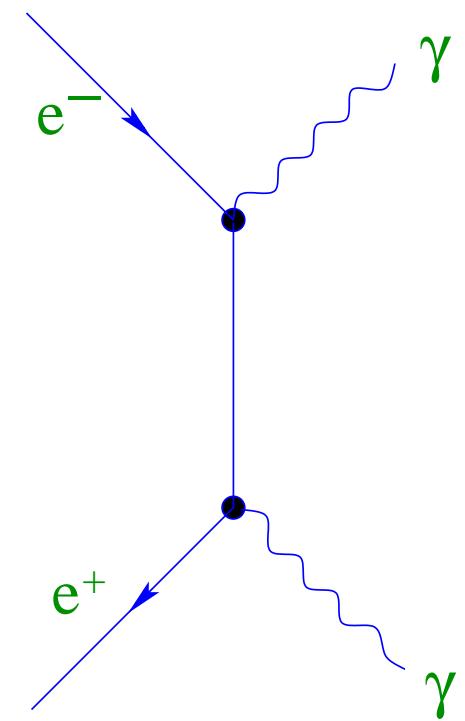
Basic assumption of field theory:

$$CPT|\mathbf{p}(r, t)\rangle = |\bar{\mathbf{p}}(-r, -t)\rangle \sim |\mathbf{p}(r, t)\rangle$$

meaning free antiparticle \sim particle
going backwards in space and time.

Giving up *CPT* one has to give up:

- locality of interactions \Rightarrow causality, or
- unitarity \Rightarrow conservation of matter, information, ... or
- Lorentz invariance



CPT-violating theories

Weak interaction violates P and CP symmetry
Theoreticians in general: CPT cannot be violated

- Standard Model valid up to Planck scale ($\sim 10^{19}$ GeV).
Above Planck scale new physics \Rightarrow
Lorentz violation possible
- Quantum gravity: fluctuations \Rightarrow Lorentz violation
loss of information in black holes \Rightarrow unitarity violation

Motivation for testing CPT at low energy

- Quantitative expression of Lorentz and CPT invariance needs violating theory
- low-energy tests can limit possible high energy violation

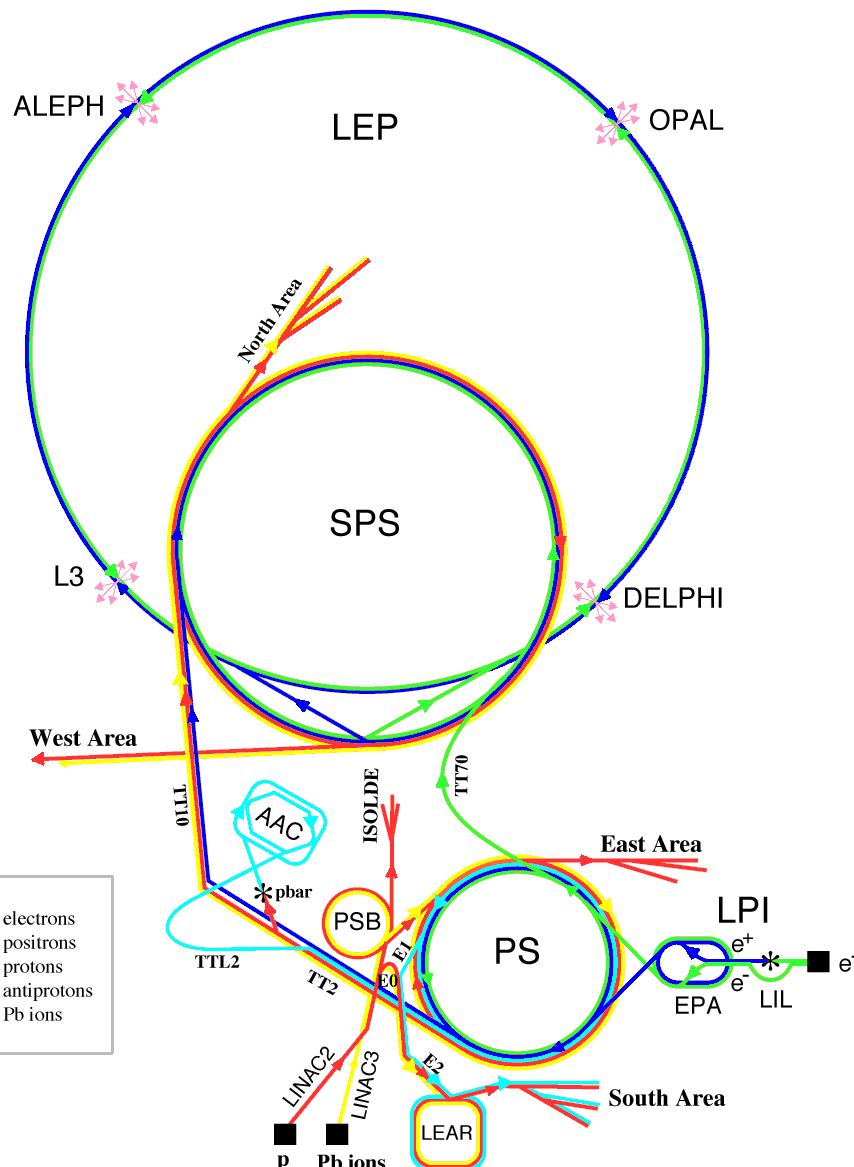
How to test CPT ?

Particle = – antiparticle ?

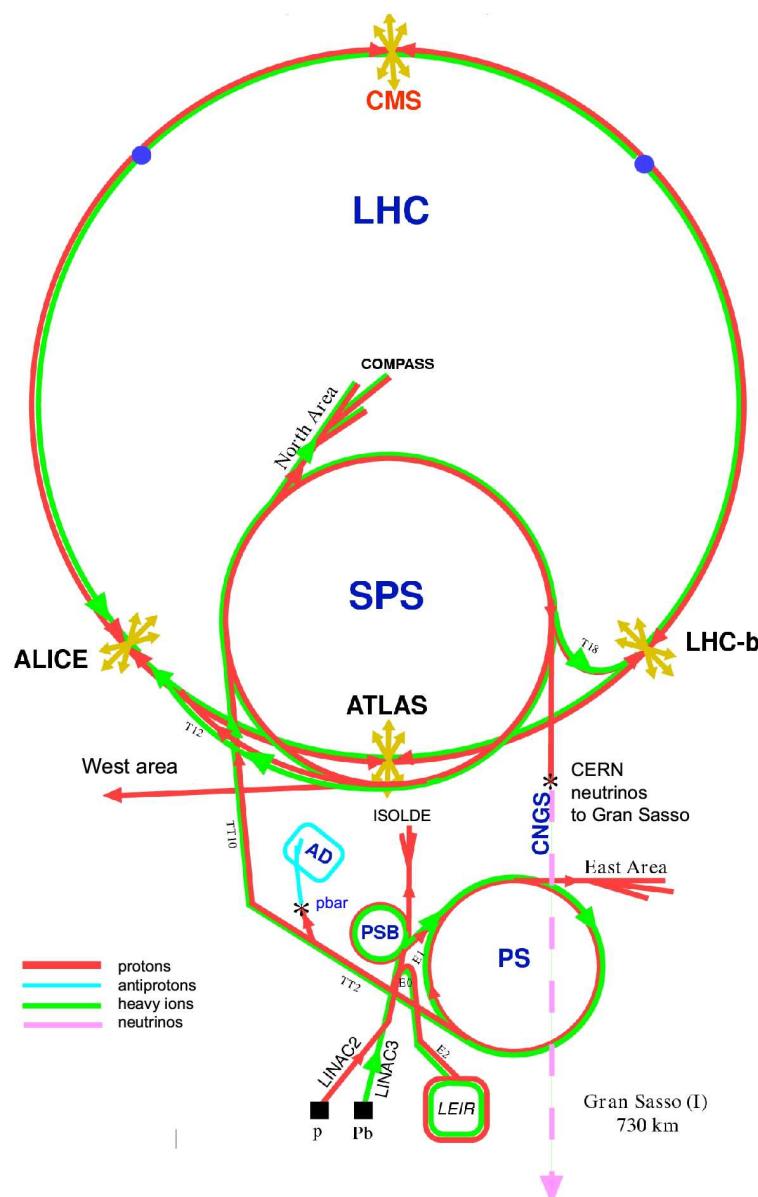
- $[m(K^0) - m(\bar{K}^0)]/m(\text{average}) < 10^{-18}$
- proton ~ antiproton? (compare $m, q, \vec{\mu}$)
- hydrogen ~ antihydrogen? ($2S - 1S$, HFS)

Accelerators at CERN

1989–2000



2009–2025??



The Antiproton Decelerator at CERN



has been built to test *CPT* invariance



Three experiments test CPT:

ATRAP: $q(\bar{p})/m(\bar{p}) \leftrightarrow q(p)/m(p)$

$$\bar{H}(2S - 1S) \leftrightarrow H(2S - 1S)$$

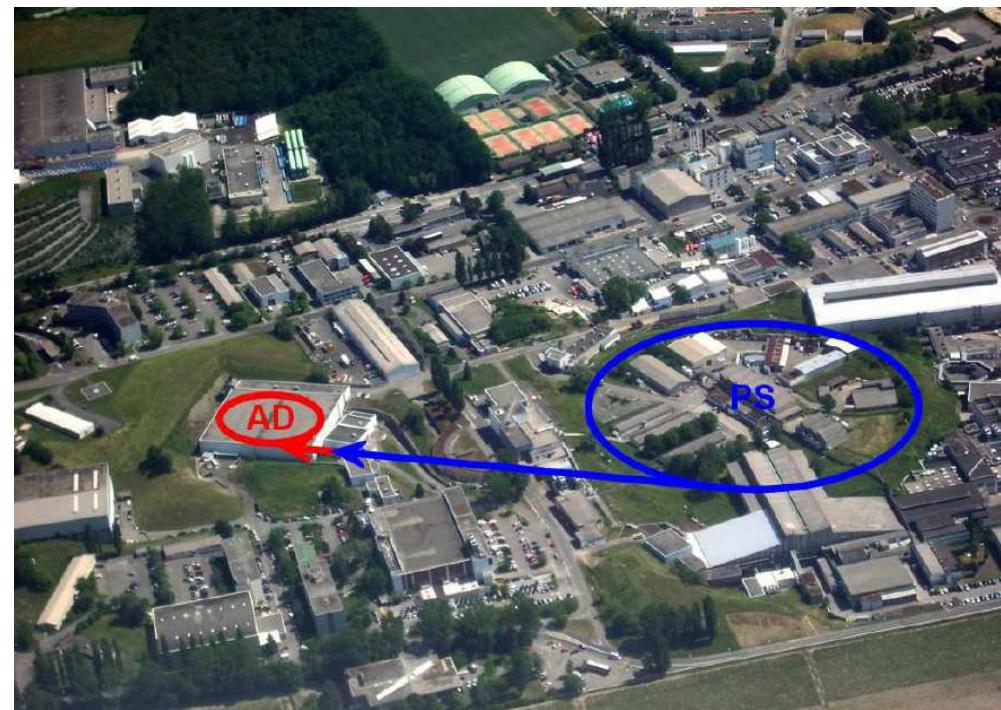
ALPHA: $\bar{H}(2S - 1S) \leftrightarrow H(2S - 1S)$

ASACUSA: $q(\bar{p})^2 m(\bar{p}) \leftrightarrow q(p)^2 m(p)$

$$\mu_\ell(\bar{p}) \leftrightarrow \mu_\ell(p)$$

$$\bar{H} \leftrightarrow H \text{ HF structure}$$

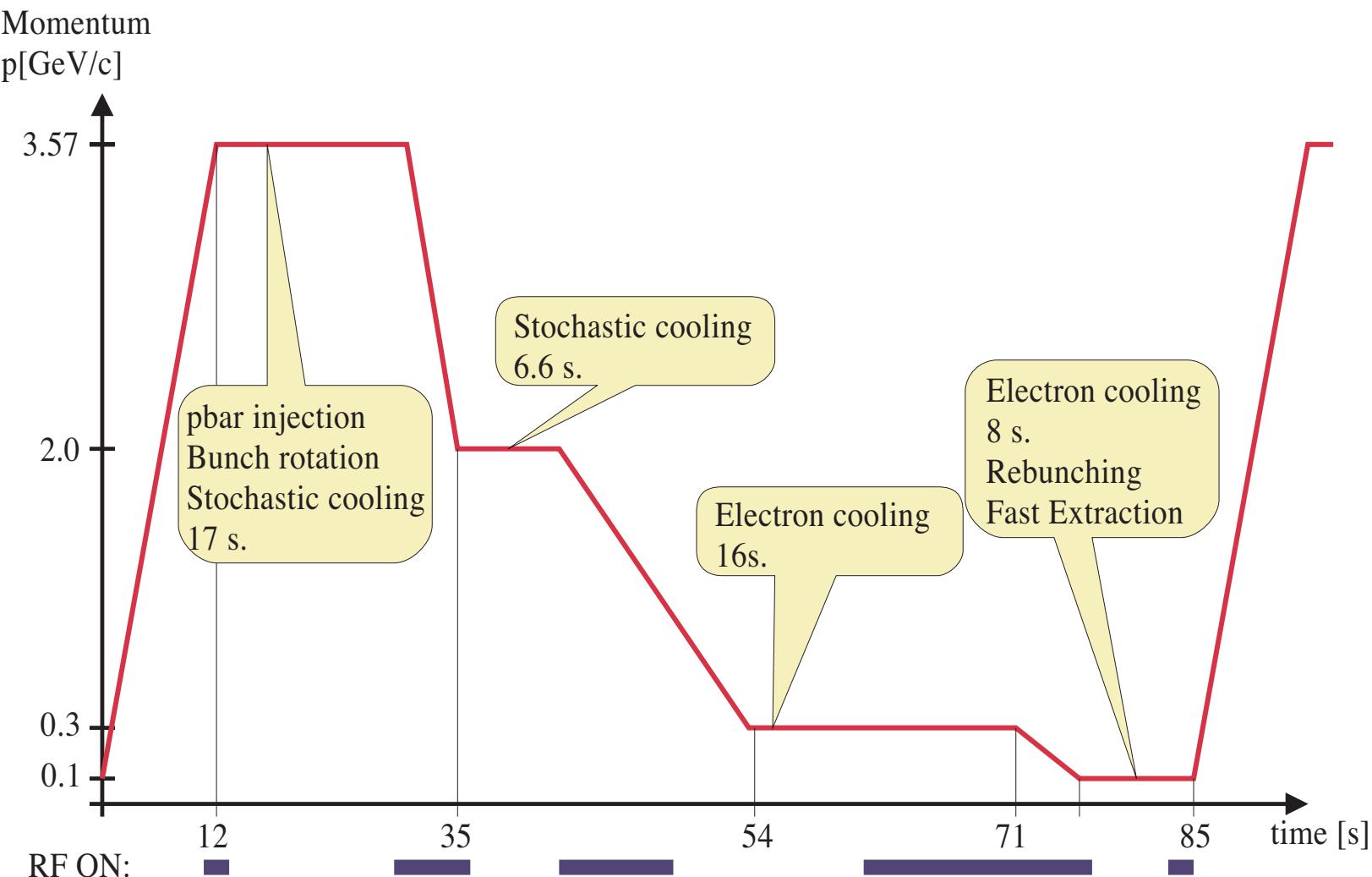
RED: done, GREEN: planned



©Ryugo S. Hayano



The Antiproton Decelerator: cooling

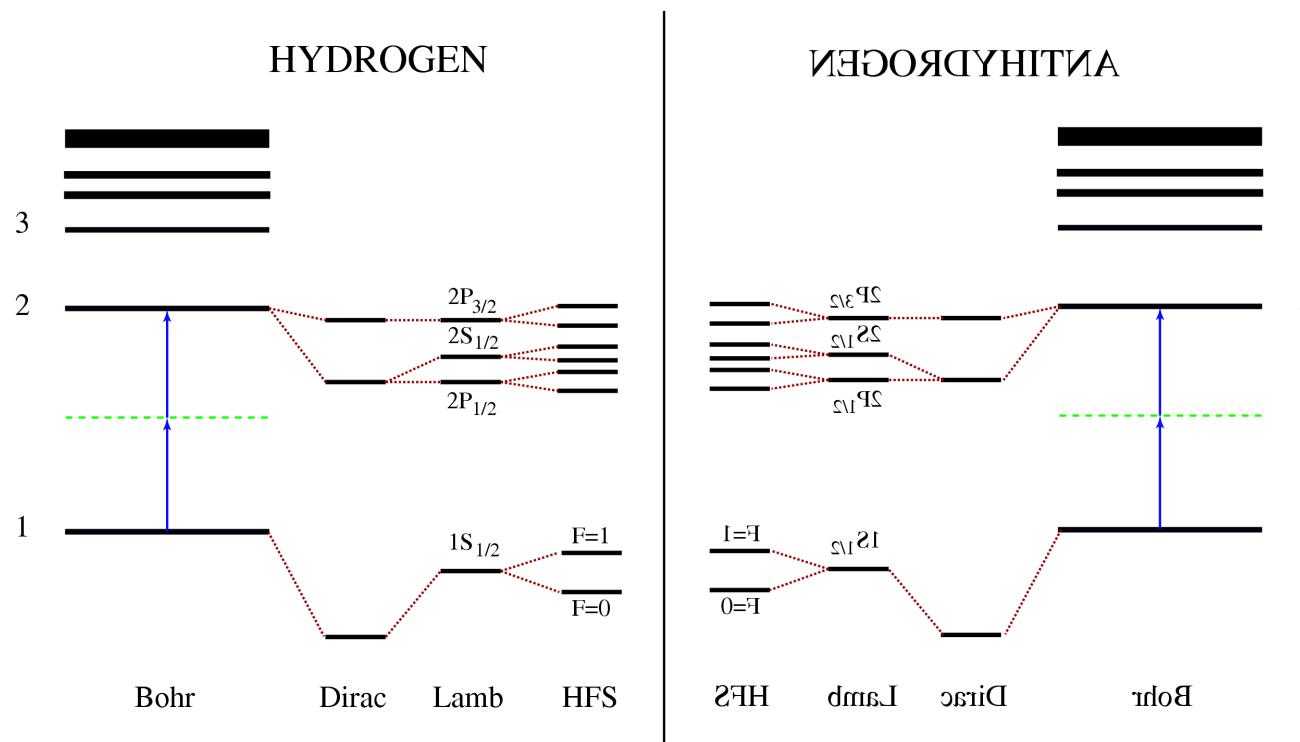


$\sim 4 \times 10^7$ 100 MeV/c antiprotons every 85 s

Pavel Belochitskii: AIP Conf. Proc. 821 (2006) 48



Antihydrogen, $e^+ - \bar{p}$ atom



$2S - 1S$ transition with 2-photons

Long lifetime, narrow transition, Doppler-free spectroscopy

M. Charlton, J. Eades, D. Horváth, R. J. Hughes, C. Zimmermann:
Antihydrogen physics, Physics Reports 241 (1994) 65.

M.H. Holzscheiter, M. Charlton, M.M. Nieto:

The route to ultra-low energy antihydrogen, Physics Reports 402 (2004) 1.



Steps toward \bar{H} spectroscopy

- Putting antiprotons (\bar{p}) in electromagnetic trap
- Trapping and cooling antiprotons
- Cooling slow positrons (e^+ from ^{22}Na) in trap
- Mixing \bar{p} and e^+ \rightarrow recombination
- Trapping antihydrogen, waiting for deexcitation

2014

- Cooling antihydrogen
- Laser spectroscopy on antihydrogen

Future



ASACUSA: Mass of the antiproton

Proton's well (?) known:

$$m(p)/m(e) = 1836.15267245(75)$$

$$q(e) = 1.602176565(35) \times 10^{-19} \text{ C}$$

Precision: $4 \cdot 10^{-10}$ and $2 \cdot 10^{-8}$

Relative measurements: proton vs. antiproton

Cyclotron frequency in trap $\rightarrow q/m$

TRAP \Rightarrow ATRAP collaboration

Harvard, Bonn, München, Seoul

\bar{p} and H^- together $\Rightarrow 10^{-10}$ precision

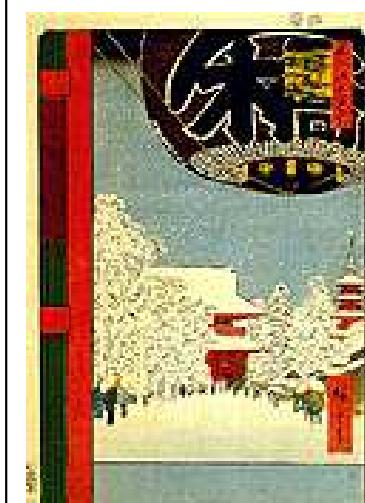
Atomic transitions:

$$E_n \approx -m_{\text{red}} c^2 (Z\alpha)^2 / (2n) \rightarrow m \cdot q^2$$

PS-205 \Rightarrow ASACUSA collaboration

Tokyo, Brescia, Budapest, Debrecen, Munich, Vienna

A
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S
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A
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C
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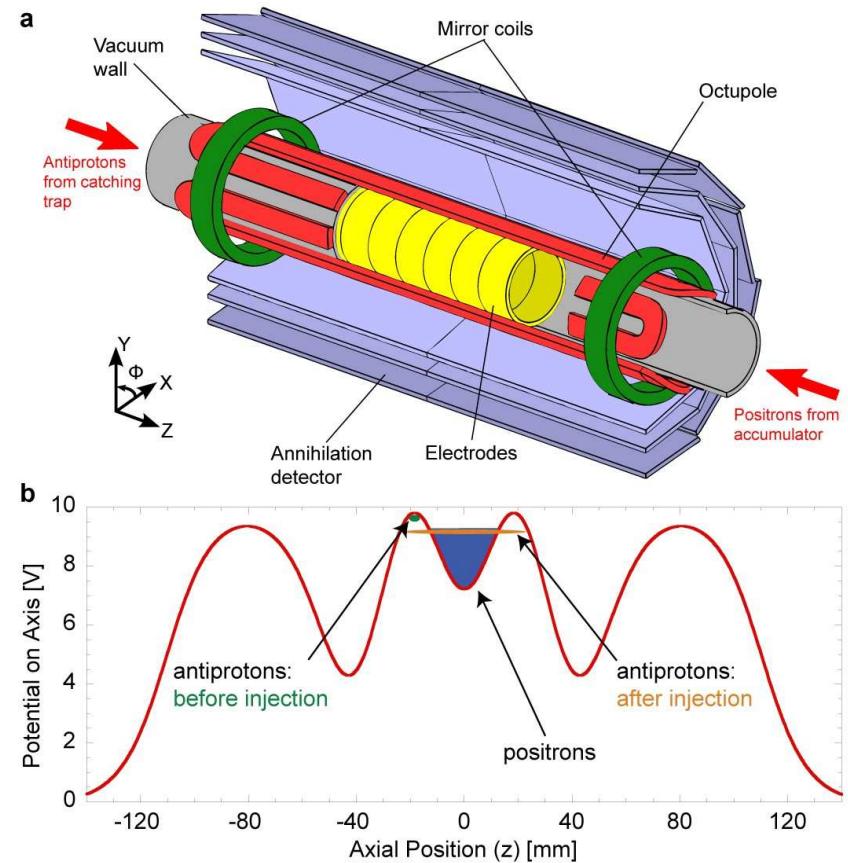
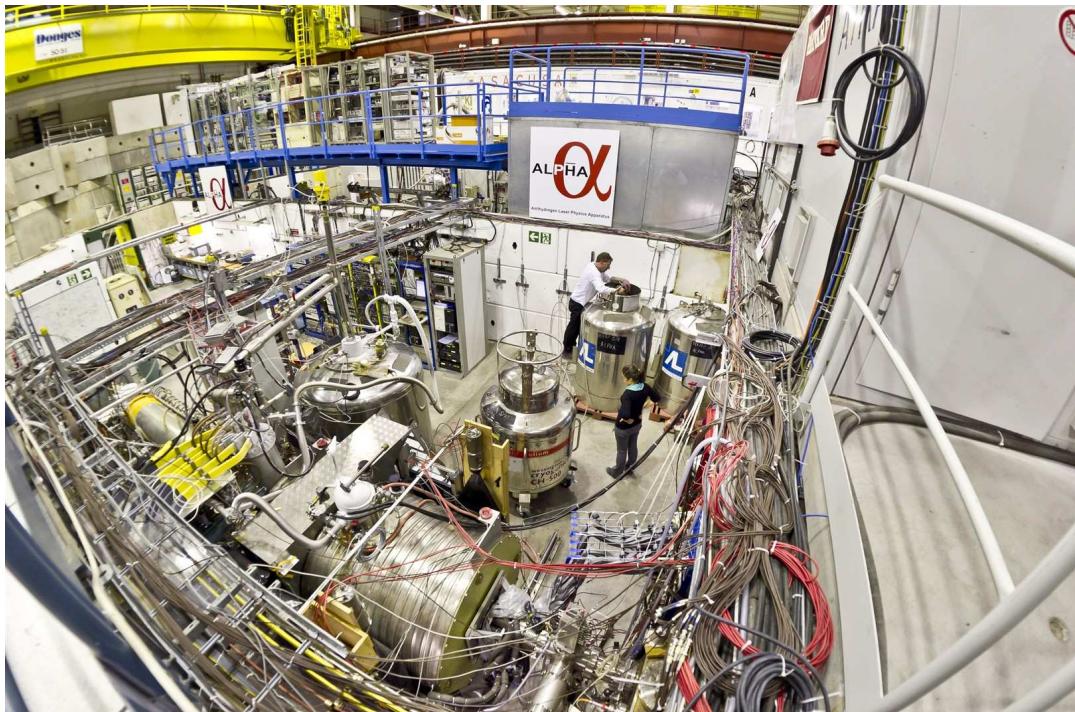


Asakusa, Tokyo



ALPHA

ALPHA: Antimatter Laser PHysics Apparatus



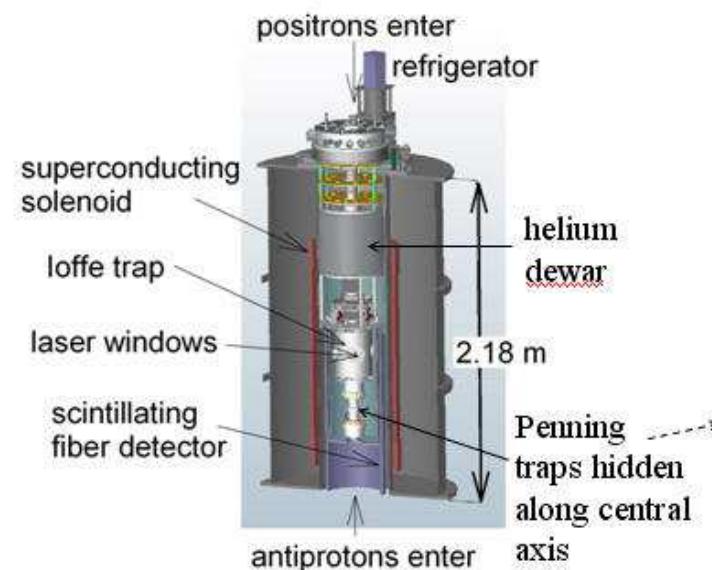
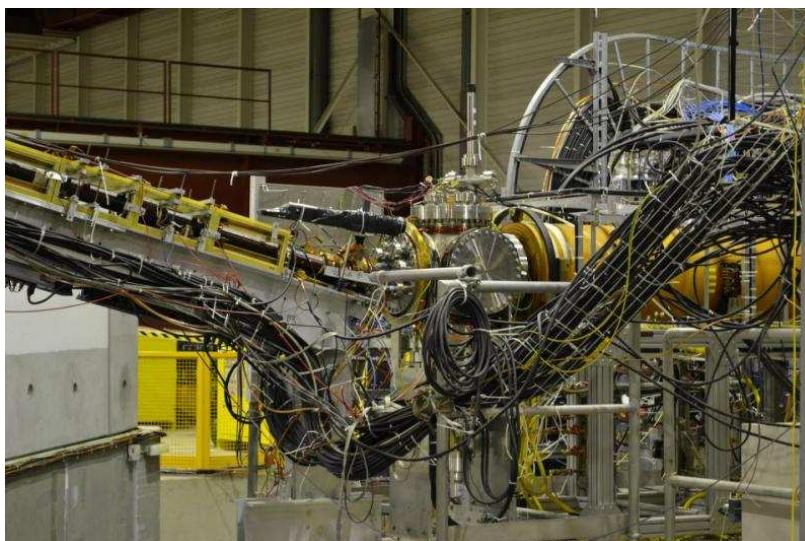
\bar{H} trapped for 1000 s; resonant HF transition induced;
limit on \bar{H} charge; proposal for gravity measurement.



ATRAP: Antimatter trap

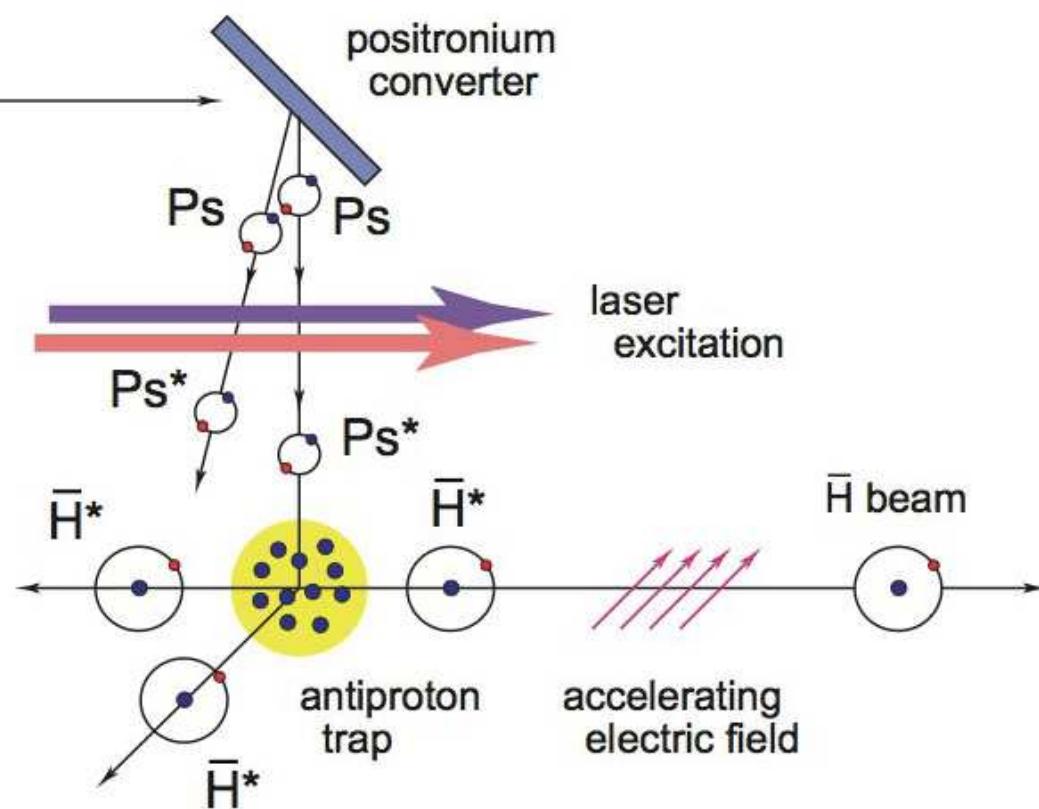


Continuous \bar{H} production

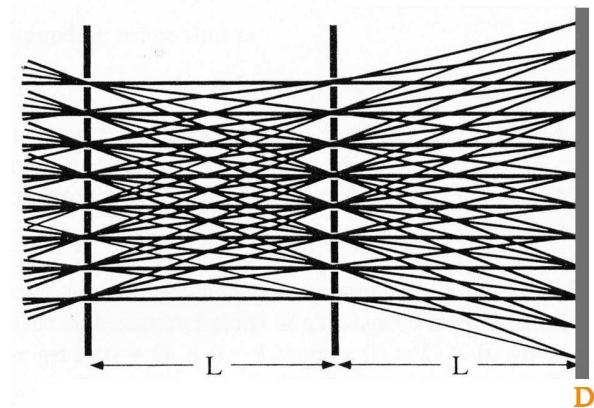


AEGIS: antimatter gravity

Antihydrogen Experiment: Gravity, Interferometry, Spectroscopy

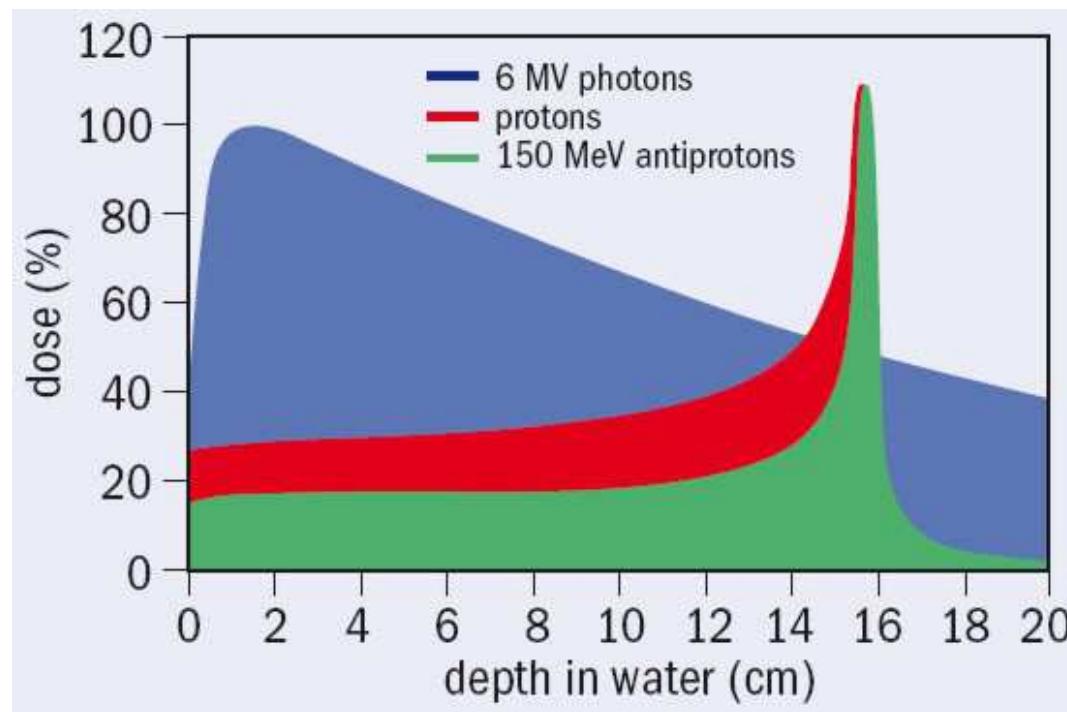


Moiré deflectometry:
gravitational falling of
collimated \bar{H}
as compared to light



ACE: Antimatter Cell Experiment

Cancer therapy research (USA) at AD of CERN



Advantage: Antiprotons lose energy in very small volume, choosing the right energy concentrates damage in tumor.

Disadvantage: Antiprotons are very expensive and annihilation radiation damages as well.



Antihydrogen beam

ASACUSA: MUSASHI



Monoenergetic
Ultra
Slow
Antiproton
Source for
High-precision
Investigations

Musashi Miyamoto self-portrait ~ 1640

5.8 MeV \bar{p} injected into RFQ

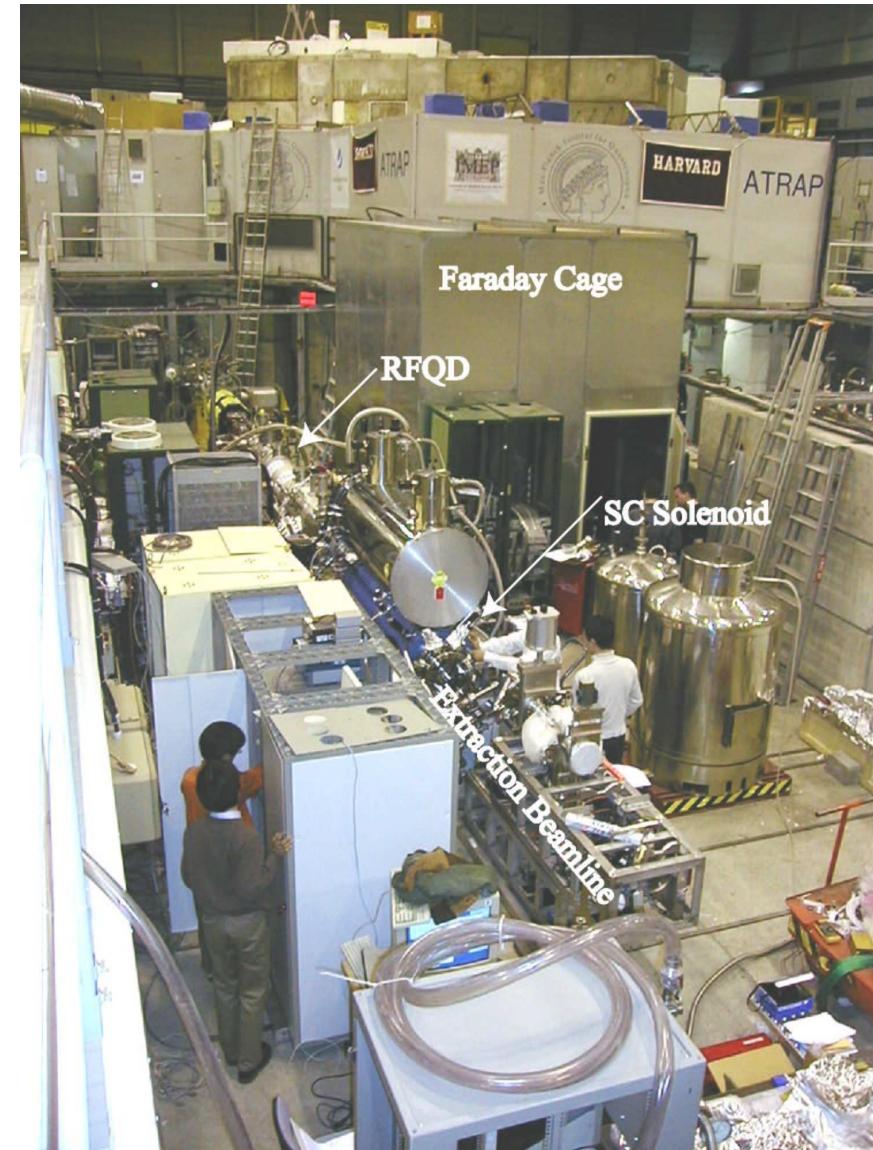
100 keV \bar{p} injected into trap

$10^6 \bar{p}$ trapped and cooled (2002)

~ 350000 slow \bar{p} extracted (2004)

Cold \bar{p} compressed in trap (2008)

$(5 \times 10^5 \bar{p}, E = 0.3 \text{ eV}, R = 0.25 \text{ mm})$

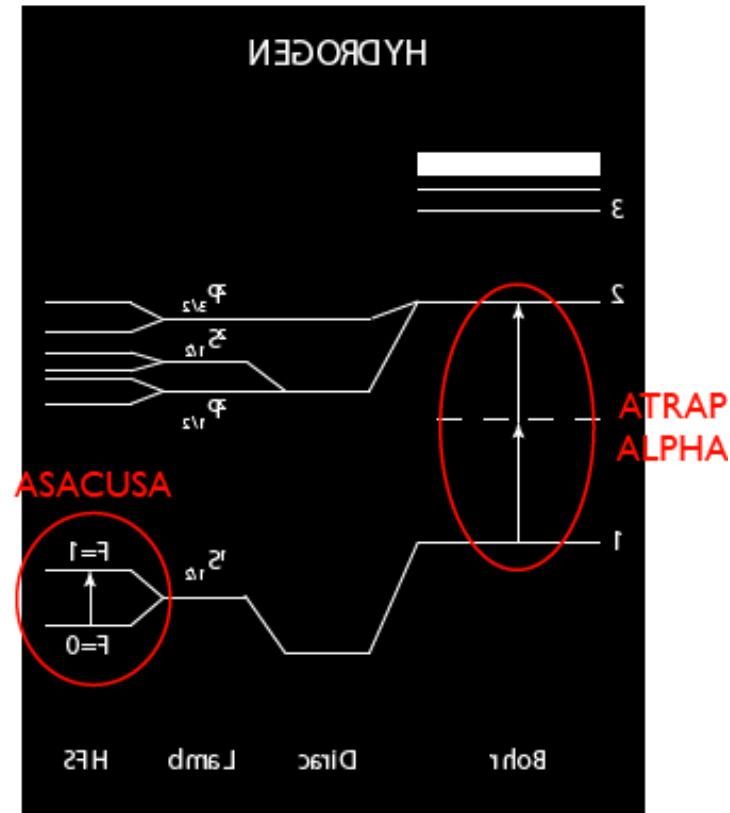
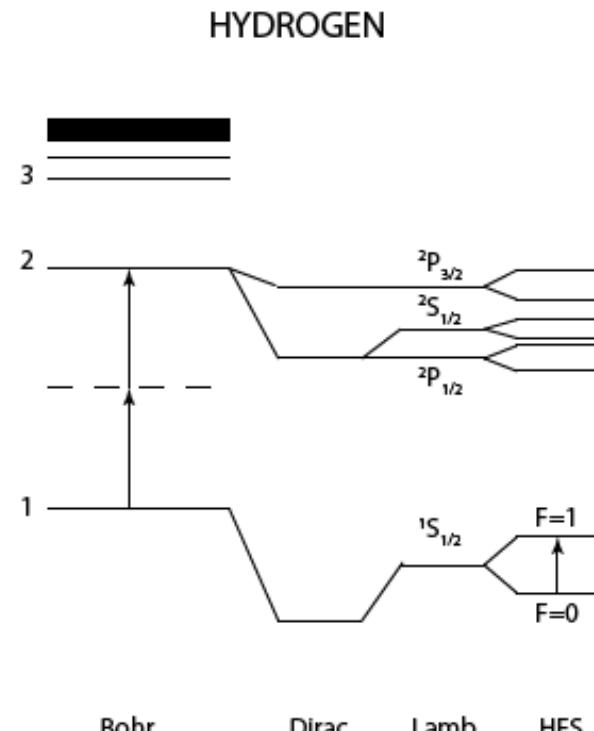
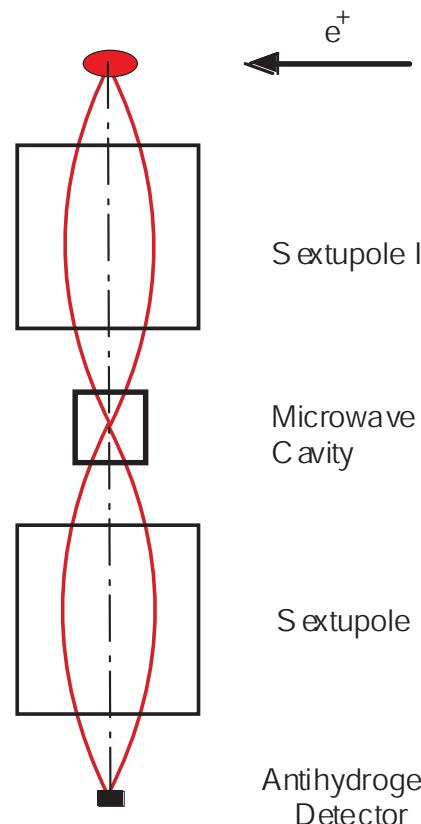


\bar{H} -beam formed: ASACUSA, 2010–2012



Spectroscopy with \bar{H} -beam

antiproton and positron
Trap / Recombination



\bar{H} -beam path: polariser, resonator, analyser
Analogy to polarised light

R.S. Hayano, M. Hori, D. Horváth, E. Widmann, Rep. Progr. Phys. 70 (2007) 1995.



Extra Low ENergy Antiprotons (ELENA)

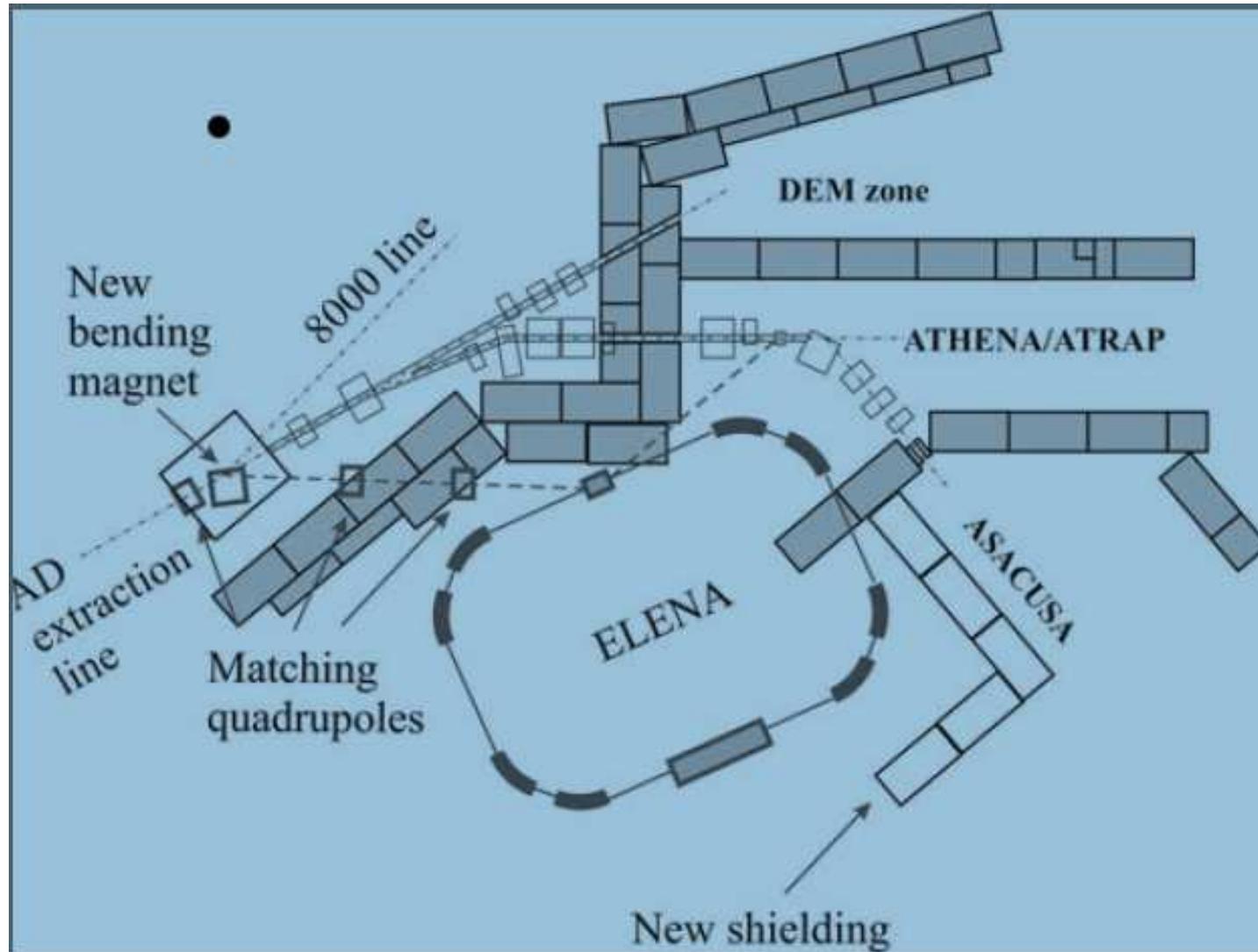
Physics Motivation

- Test the Standard Model and General Relativity for antimatter
- Test SM extensions for antimatter (Lorentz-violation, black holes, new interactions, ...)
- Stringent CPT tests with antihydrogen
- Antimatter gravity measurement (weak equivalence)
- Added precision for physical constants (CODATA) assuming CPT invariance

All existing AD experiments profit, new ones made possible
(gravity, X-rays, nuclear studies)



ELENA at the AD: plan



M.-E.Angoletta et al: *ELENA: A Preliminary Cost And Feasibility Study*,
CERN-AB-2007-079.



Antimatter in Space

AMS-2: Alpha Magnetic Spectrometer
to discover antimatter (anti-helium!) and
dark matter

Mass: 8500 kg,
1200 kg perm. magnet

Father: Sam Ting, cost: 2 G\$

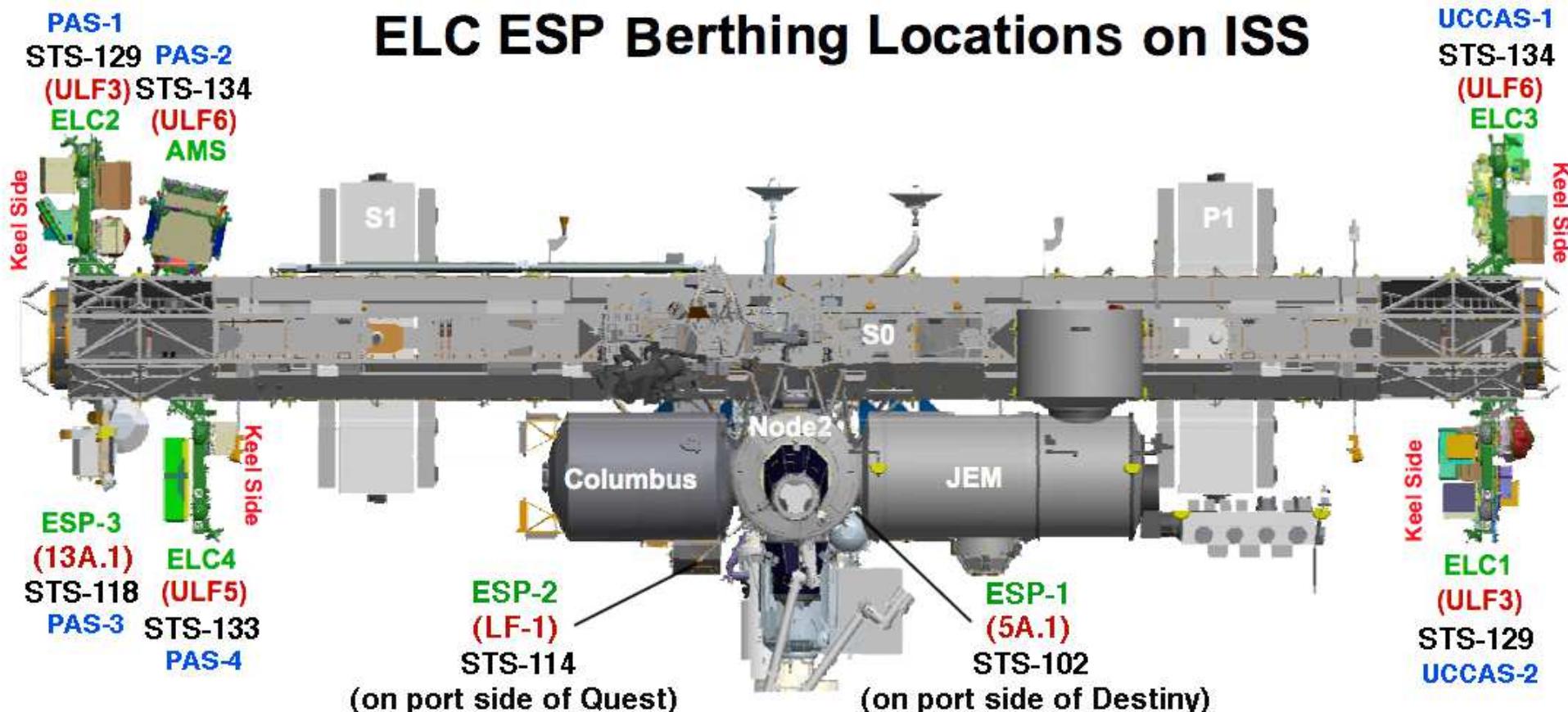
Construction: CERN

Launch: May 2011, USA

Control room: CERN



AMS-2: Alpha Magnetic Spectrometer



First results (2013-14):
No antihelium observed.
High energy positrons everywhere.
Could come from dark matter or pulsars.
AMS2 will collect data for 10–15 years.



Conclusion

ASACUSA at AD

- Two-photon spectroscopy of antiprotonic helium: results agree with 3-body calculations.
- Determined $M_{\bar{p}}/m_e$ ratio to 1.3 ppb. Result agrees with CODATA proton value (0.4 ppb).
- Further improvement partially hindered by theoretical uncertainty.

Future prospects

- Colder atoms, better lasers, better detectors.
- ELENA (colder antiproton beams at 100 keV of higher luminosity).
- Spectroscopy with trapped antihydrogen and with antihydrogen beam.
- AMS2 delivers more and more data on antimatter in space.



Next: ASACUSA, antiprotonic helium

