

 All observables can be predicted in terms of N<sub>par</sub> = 18 (?) parameters

 $O_i(E) = O_i(E, \alpha, \alpha_s, M_Z, M_W, M_H, M_t, \dots)$ 

where E corresponds to some charesterstic energy scale.

- If we meausere at least N<sub>par</sub> different observables we can extract the values of the parameters from the experiment and make predictions....
- Rather naive question:
  - Why can't we extract all the parameters just from ONE observable by choosing different E<sub>i</sub>,i=1,N<sub>par</sub>?

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In princple «YES»:

- Choose appropriate O<sub>i</sub>
- Calculate RADIATIVE corrections as precise as possible (dependence on all the parameters)
- · Get what you want!

 All observables can be predicted in terms of N<sub>par</sub> = 18 (?) parameters

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In practice «NO»! sensitivity is different for different parameters!

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In practice «NO»! sensitivity is different for different parameters!

OK, not all (just one, e.g.,  $M_{H}$ ) and not from a single observable....(See below).

### Consistency check of the SM

 Given N>N<sub>par</sub> observables one can check the consistency of the model, since the latter predicts the relations between observables.

 $O_i(E) = O_i(E, \alpha, \alpha_s, M_Z, M_W, M_H, M_t, \dots)$ 

- In practice: use as many observables as possible and do a fit
  - e.g. ZFITTER or GFITTER for ElectroWeak Precision Observables (EWPO)





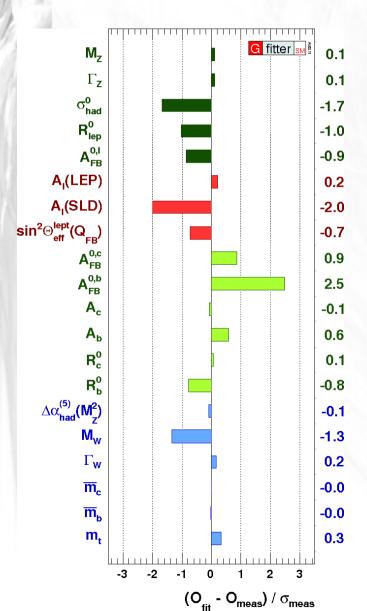
### **Precision SM observables**

- Low energy
  - Fermi  $G_{\mu}$  constant measured in muon decay
  - Anomalous magnetic moment of muon  $(g-2)_{\mu}$
- Z-pole observables
  - Z-boson mass and width  $M_Z, \Gamma_Z$
  - Effective couplings of Z to fermions  $g_V, g_A$
  - Effective Weinberg angle  $\sin^2 \theta_{lept}$
- LEP 2, Tevatron, LHC:
  - W-boson mass  $M_W$
  - Top quark mass  $M_t$

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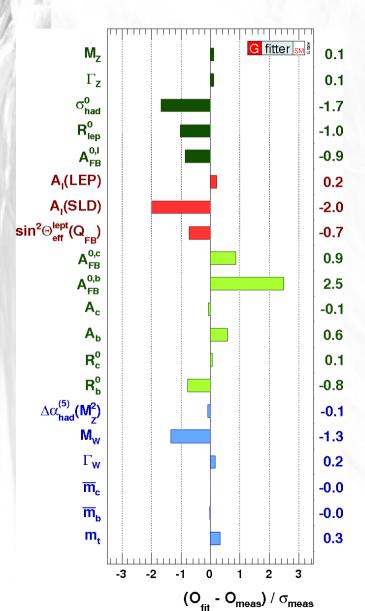
## Gfitter fit :)



Parameter	Input value	Free	Results from global EW fits:		Complete fit w/o
		in fit	Standard fit	Complete fit	exp. input in line
$M_Z$ [GeV]	$91.1875 \pm 0.0021$	yes	$91.1874 \pm 0.0021$	$91.1877 \pm 0.0021$	$91.1983 \substack{+0.0133 \\ -0.0155}$
$\Gamma_Z$ [GeV]	$2.4952 \pm 0.0023$	-	$2.4959 \pm 0.0015$	$2.4955 \pm 0.0014$	$2.4951 \substack{+0.0017 \\ -0.0016}$
$\sigma_{\rm had}^0$ [nb]	$41.540\pm0.037$	_	$41.478\pm0.014$	$41.478\pm0.014$	$41.469 \pm 0.015$
$R^0_\ell$	$20.767\pm0.025$	_	$20.743\pm0.018$	$20.741 \pm 0.018$	$20.718 \substack{+0.027 \\ -0.026}$
$A_{ m FB}^{0,\ell}$	$0.0171 \pm 0.0010$	_	$0.01641 \pm 0.0002$	$0.01620 \substack{+0.0002 \\ -0.0001}$	$0.01606 \pm 0.0001$
$A_{\ell}^{(\star)}$	$0.1499\pm0.0018$	_	$0.1479 \pm 0.0010$	$0.1472^{+0.0009}_{-0.0006}$	_
$A_c$	$0.670\pm0.027$	-	$0.6683^{+0.00044}_{-0.00043}$	$0.6680 \substack{+0.00040 \\ -0.00028}$	$0.6679^{+0.00042}_{-0.00025}$
$A_b$	$0.923 \pm 0.020$	_	$0.93470 \substack{+0.00009 \\ -0.00008}$	$0.93463 \substack{+0.00008\\-0.00005}$	0.93463+0.00007
$A_{ m FB}^{0,c}$	$0.0707 \pm 0.0035$	_	$0.0741 \pm 0.0005$	$0.0737 \substack{+0.0005 \\ -0.0004}$	$0.0738 \pm 0.0004$
$A_{\rm FB}^{0,b}$	$0.0992 \pm 0.0016$	_	$0.1037 \pm 0.0007$	$0.1035 \substack{+0.0003 \\ -0.0004}$	$0.1038^{+0.0003}_{-0.0005}$
$R_c^0$	$0.1721 \pm 0.0030$	_	$0.17226 \pm 0.00006$	$0.17226 \pm 0.00006$	$0.17226 \pm 0.0000$
$R_b^0$	$0.21629 \pm 0.00066$	_	$0.21578 \substack{+0.00005 \\ -0.00008}$	$0.21577 \substack{+0.00005\\-0.00008}$	0.21577+0.00005
$\sin^2 \theta_{eff}^{\ell}(Q_{FB})$	$0.2324 \pm 0.0012$	_	$0.23141 \pm 0.00012$	$0.23150 \substack{+0.00008 \\ -0.00011}$	$0.23152 \substack{+0.00006\\-0.00013}$
$M_H$ [GeV] <sup>(o)</sup>	Likelihood ratios	yes	$95^{+30[+74]}_{-24[-43]}$	$125^{+8[+21]}_{-10[-11]}$	$95^{+30[+74]}_{-24[-43]}$
M <sub>W</sub> [GeV]	$80.399 \pm 0.023$	_	$80.382 \substack{+0.014 \\ -0.015}$	$80.368\substack{+0.007\\-0.010}$	$80.360 \substack{+0.012 \\ -0.011}$
Γ <sub>W</sub> [GeV]	$2.085\pm0.042$	-	$2.093 \pm 0.001$	$2.092\pm0.001$	$2.091 \substack{+0.002 \\ -0.001}$
m_[GeV]	$1.27^{+0.07}_{-0.11}$	yes	$1.27^{+0.07}_{-0.11}$	$1.27 \substack{+0.07 \\ -0.11}$	_
m <sub>b</sub> [GeV]	$4.20^{+0.17}_{-0.07}$	yes	$4.20^{+0.16}_{-0.07}$	$4.20 \substack{+0.16 \\ -0.07}$	_
$m_t$ [GeV]	$173.2 \pm 0.9$	yes	$173.3 \pm 0.9$	$173.5 \pm 0.9$	$177.2^{+2.9}_{-3.1}(\bigtriangledown)$
$\Delta \alpha_{\rm had}^{(5)}(M_Z^2)^{(\dagger \bigtriangleup)}$	$2749 \pm 10$	yes	$2750 \pm 10$	$2748 \pm 10$	$2716^{+60}_{-45}$
$\alpha_s(M_Z^2)$	_	yes	$0.1192 \pm 0.0028$	$0.1193\pm0.0028$	$0.1193 \pm 0.0028$
$\delta_{ m th} M_W$ [MeV]	$[-4,4]_{theo}$	yes	4	4	_
$\delta_{\rm th} \sin^2 \theta_{\rm eff}^{\ell}$ (†)	$[-4.7, 4.7]_{theo}$	yes	4.7	4.7	_

<sup>(\*)</sup>Average of LEP ( $A_{\ell} = 0.1465 \pm 0.0033$ ) and SLD ( $A_{\ell} = 0.1513 \pm 0.0021$ ) measurements. The fit w/o the LEP (SLD) measurement but with the direct Higgs searches gives  $A_{\ell} = 0.1471 \stackrel{+0.0010}{_{-0.0000}} (A_{\ell} = 0.1467 \stackrel{+0.0007}{_{-0.0004}})$ . <sup>(o)</sup>In brackets the  $2\sigma$ . <sup>(†)</sup>In units of  $10^{-8}$ . <sup>( $\Delta$ )</sup>Rescaled due to  $\alpha_s$  dependency. <sup>( $\nabla$ )</sup>Ignoring a second less significant minimum, cf. fig. ?? and the result of eq. (??).

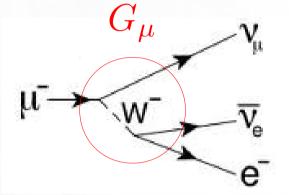
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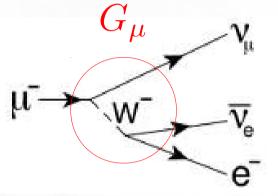
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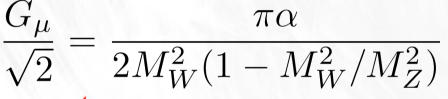
 $G_F = 1.16637(1) \times 10^{-5} \text{ GeV}^{-5}$  $\frac{G_{\mu}}{\sqrt{2}} = \frac{g}{2\sqrt{2}} \frac{1}{M_W^2} \frac{g}{2\sqrt{2}}$ 

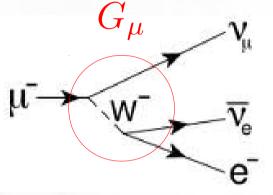


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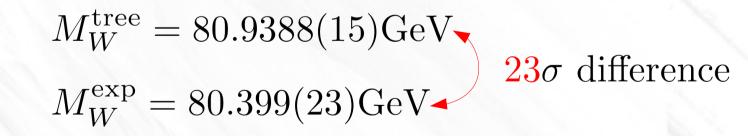


$$G_F = 1.16637(1) \times 10^{-5} \text{ GeV}^{-5}$$





Tree-level



We need to include high order effects!

 $\alpha = 1/137.035999679, M_Z = 91.1876(28) \text{ GeV}$ 

At high orders one needs to specify a renormalization scheme to define the renormalized parameters

#### We were considering

- physical masses  $M_Z, M_W$ 
  - $\rightarrow$  ON-SHELL renormalization prescription
- $\alpha(q^2=0)$  fine-structure constant in MOM-scheme

Fermi theory

 $\frac{d\mu}{\sqrt{2}} = \frac{\pi d}{2M_W^2 (1 - M_W^2 / M_Z^2)}$ 

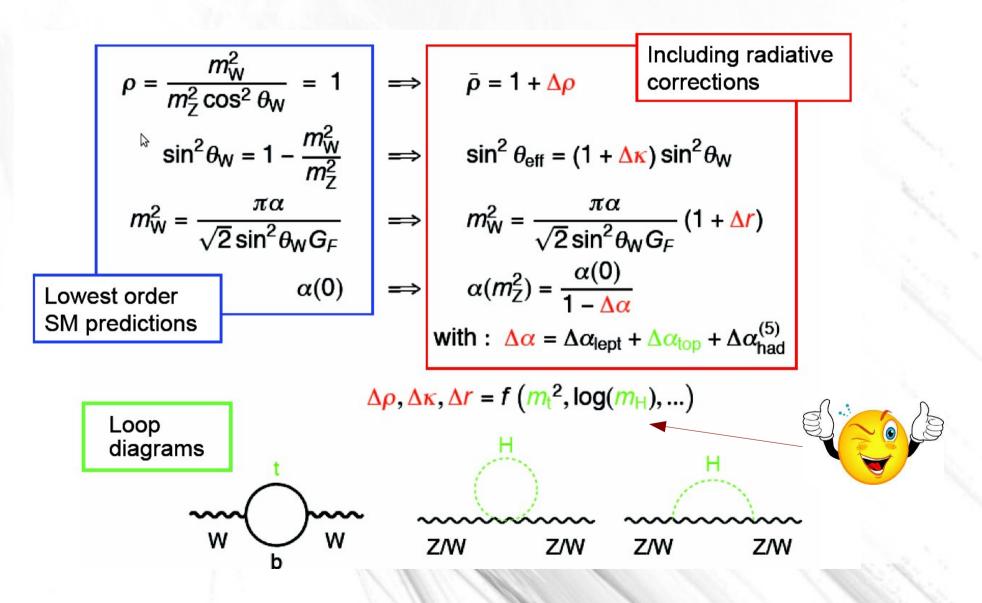
 $\pi \alpha$ 

Typical MATCHING problem

Comparison of observables In «effective» and «fundamental» theories

SM

 $G_{\mu}$ 



2

#### with loop contributions

$$\frac{G_F}{\sqrt{2}} = \frac{\pi\alpha}{M_W^2 \left(1 - M_W^2 / M_Z^2\right)} \cdot \left(1 + \Delta r\right)$$

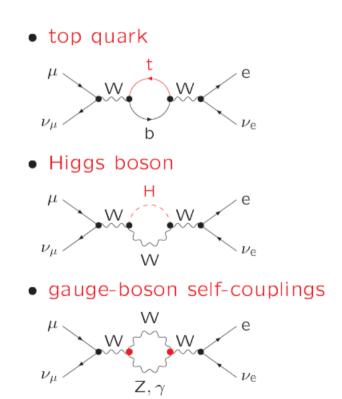
 $\Delta r$ : quantum correction  $\Delta r = \Delta r(m_t, M_H)$ 

#### determines W mass

 $M_W = M_W(\alpha, G_F, M_Z, m_t, M_H)$ 

complete at 2-loop order

1-loop examples



full structure of SM

#### with loop contributions

$$\frac{G_F}{\sqrt{2}} = \frac{\pi\alpha}{M_W^2 \left(1 - M_W^2 / M_Z^2\right)} \cdot \left(1 + \Delta r\right)$$



 $\Delta r$ : quantum correction

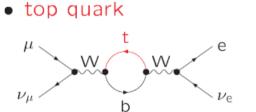
 $\Delta r = \Delta r(m_t, M_H, m_Z, m_W, \dots)$ 

#### determines W mass

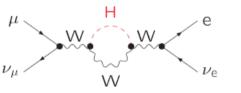
 $M_W = M_W(\alpha, G_F, M_Z, m_t, M_H)$ 

complete at 2-loop order

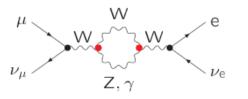
#### 1-loop examples



• Higgs boson



gauge-boson self-couplings

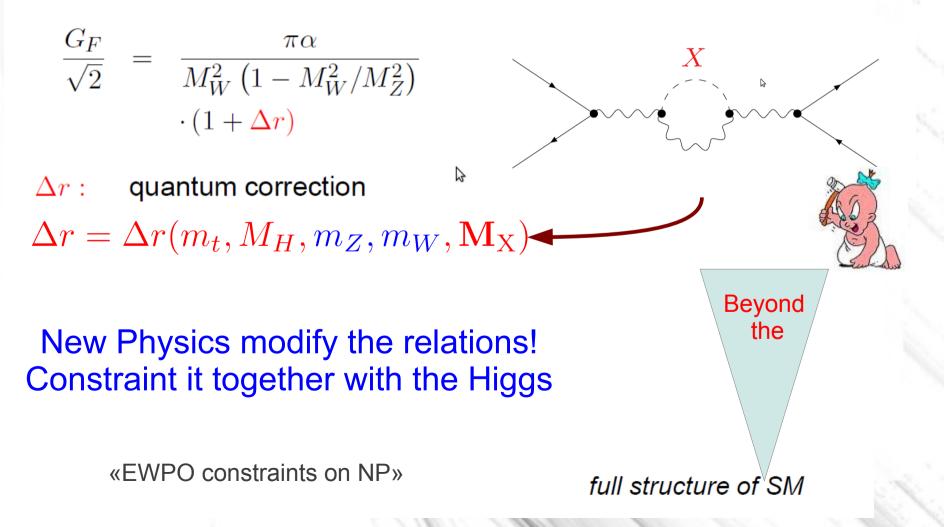


Try to FIT it!

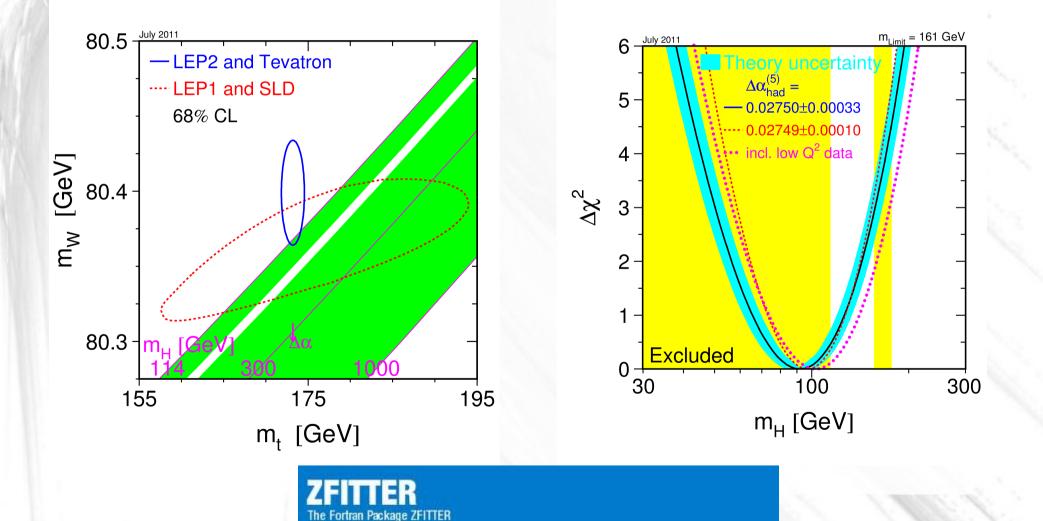
full structure of SM

#### with loop contributions

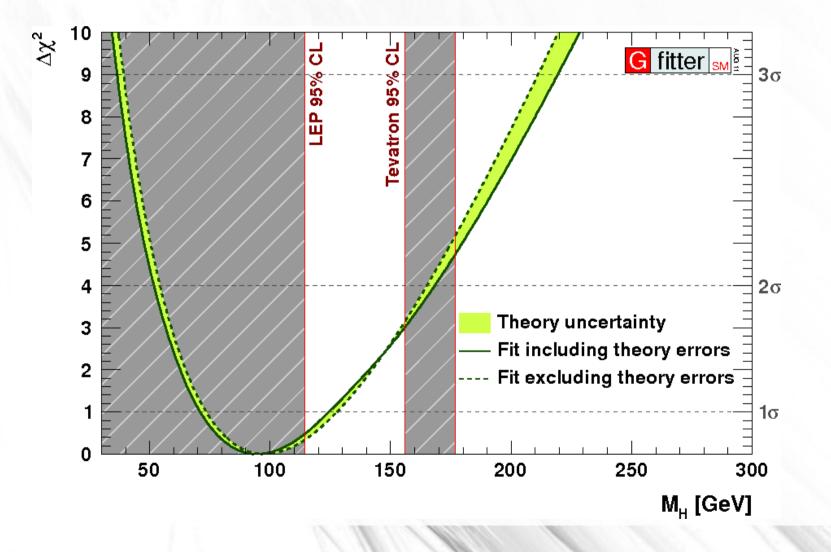
1-loop examples



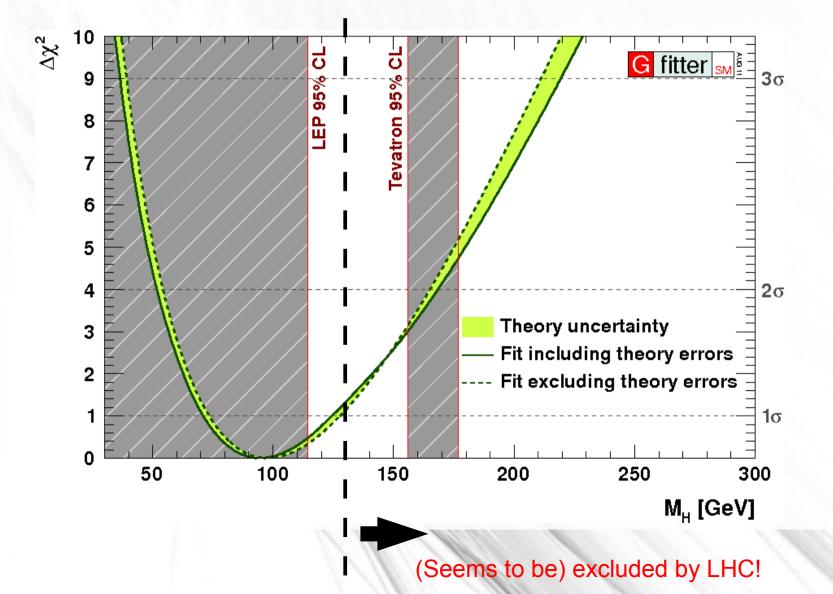
## (Indirect) bounds on Higgs Mass



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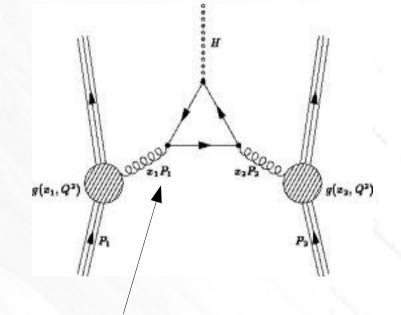


## More on importance of Loop corrections

## Some processes in the SM are ONLY due to loops!

MSTW 2008 NLO PDFs (68% C.L.)

Very important to the LHC Higgs searches



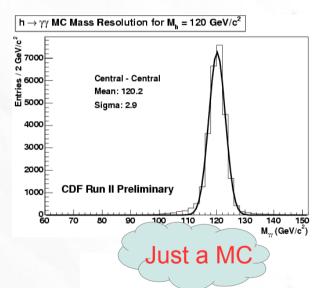
Since we expect that  $M_h <<1$  TeV, x should be small

## The missing piece: Higgs particle

Higgs mass is a free parameter of the SM!

Should be measured in a direct experiment (e.g. as a Peak in some disrtibution)

- But we have indirect bounds
  - from precision corrections
  - from theoretical consistency



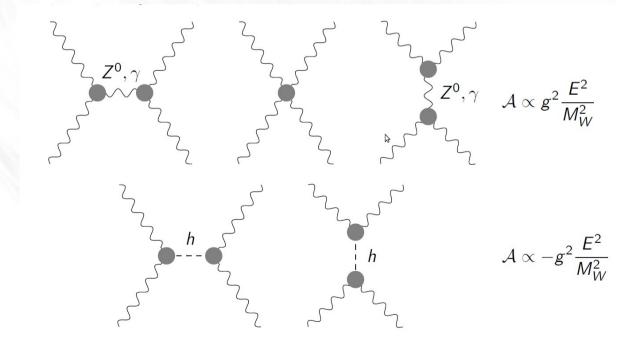


## Theoretical constraints on the Higgs boson mass

Unitarity problem in WW → WW scattering

Longitudinal

polarization of W-boson



Without Higgs WW scattering violate unitarity (cross-section grows with energy)

TeV scale favoured

## **Theoretical constraints** on the Higgs boson mass

Triviality and Stability

For large  $M_{\mu}$ :

For small  $M_{\mu}$ :

$$M_H^2 = 2\lambda v^2 \qquad \qquad \frac{d\lambda}{dt} = \frac{1}{16\pi^2} \left(12\lambda^2 - 3g_t^4 + 6\lambda g_t^2 + \cdots\right)$$

For large M<sub>H</sub>:  

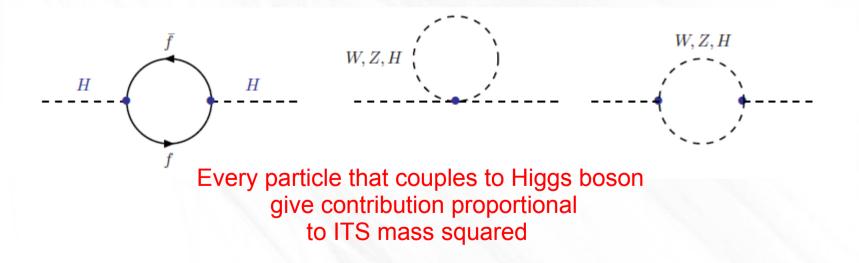
$$\lambda(Q) = \frac{M_H^2}{2v^2 - \frac{3}{2\pi^2}M_H^2 \ln \frac{Q}{v}}$$
For small M<sub>H</sub>:  

$$\lambda(Q) = \lambda(v) - \frac{\frac{3}{8\pi^2}y_t^4(v) \ln \frac{Q}{v}}{1 - \frac{9}{16\pi^2}y_t^2(v) \ln \frac{Q}{v}}$$

$$\lambda(Q) = \lambda(v) - \frac{3}{6\pi^2}y_t^4(v) \ln \frac{Q}{v}}{1 - \frac{9}{16\pi^2}y_t^2(v) \ln \frac{Q}{v}}$$

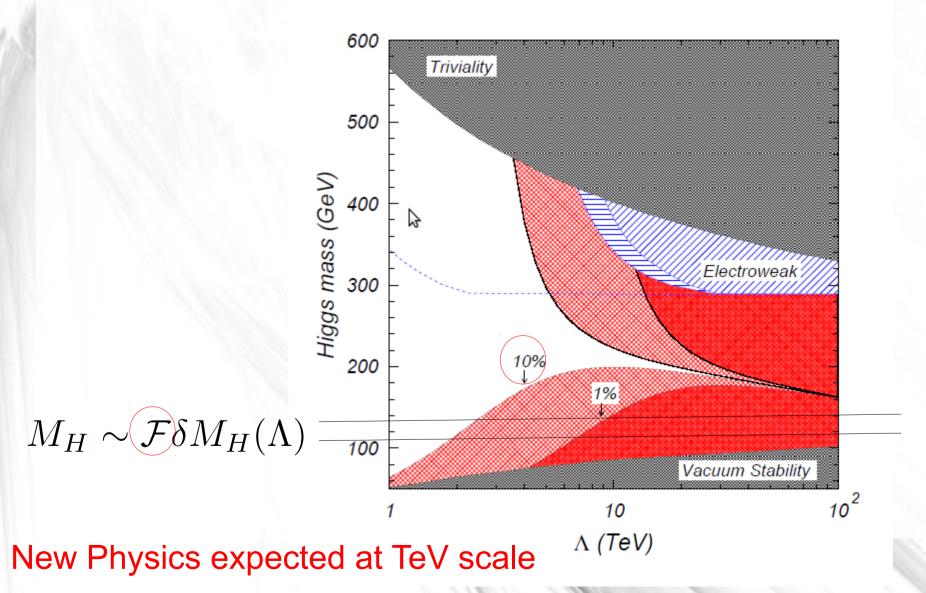
## Theoretical constraints on the Higgs boson mass

- Naturalness (hierarchy problem)
  - No symmetry to protect the Higgs mass from large radiative corrections (scalar particle)



NB: For vector bosons: gauge symmetry, for fermions — chiral symmetry

## Theoretical constraints on the Higgs boson mass

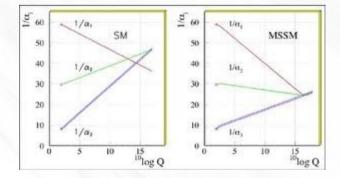


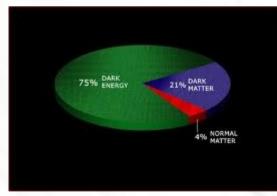
## Issues of the SM

- Higgs is missing (EWSB mehanism)
- No Dark Matter candidate
- Gauge coupling unification is problematicty
- Large number of free parameters
- Flavor problem
- Gravity?

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### Lecture 4 summary

- The SM exhibits (almost) perfect agreement with data in High Energy Physics experiments!
- Possible New Physics **HAS** to reproduce it as a low-energy effective theory!
- Still, there are some issues that prevent us from saying that the SM is the unltimate theory
- We are waiting for NEW data from LHC to find the last ingridient of the model — the Higgs boson (test the EWSB mechansm).

# Topics NOT covered in the lectures

• QCD

(A.V. Nesterenko, O.V. Teryaev)

Flavor Physics in Lepton sector

(V.A. Naumov and S.M. Bilenky)

Renormalization

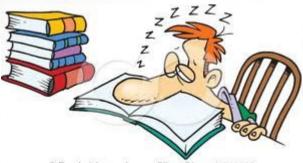
(A.A. Vladimirov)

Top Physics





## Thank you!



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