LFV SUSY searches with Rp conservation

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<u>OUTLINE</u>

Basic physics

Input / bounds from FV processes

(emphasis on rare charged lepton decays/conversions)

Effects of Quantum Corrections

Constraining non-universality from symmetries

Cosmological data

LFV collider signatures

Recent Reviews:

- Flavour Physics of Leptons & Dipole Moments, M. Raidal et al., hep-ph/0801.1826
- Collider aspects of flavour physics at high Q,
 F. del Aguila et al.,hep-ph/0801.1800
- RGEs:

See i.e. Hisano & Nomura, hep-ph/9810479.

MSSM content & parameters

- SM particles & superpartners
- 2 Higgs fields with coupling μ , ratio of v.e.v.s = tan β
- SUSY-breaking parameters:

Scalar masses m_0 Gaugino masses $m_{1/2}$

- Trilinear soft terms A_{λ} Bilinear soft terms B_{μ}
- Assume universality? constrained MSSM = CMSSM Single m_0 , single $m_{1/2}$, single A_{λ} , B_{μ}
- CMSSM different from mSUGRA

where we have additional relations, as functions of $m_{3/2}$

Strong bounds from flavour violating processes motivate Universality in soft scalar masses

$$egin{aligned} BR(\mu
ightarrow e \gamma) &\lesssim 1.2 \cdot 10^{-11} \ BR(au
ightarrow e \gamma) &\lesssim 1.1 \cdot 10^{-7} \ BR(au
ightarrow lll') &\lesssim O(10^{-8}) \ (l,l'=e,\mu) \end{aligned}$$

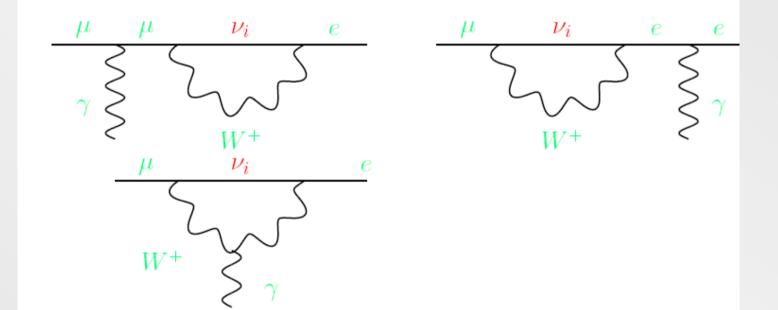
$$BR(\mu^- \to e^- e^+ e^-) \lesssim 10^{-12}$$
$$BR(\tau \to \mu\gamma) \lesssim 6.8 \cdot 10^{-8}$$

Very good expected future BR sensitivities:

$$\mu \to e\gamma$$
 10^{-14}

$$\mu^- Ti \rightarrow e^- Ti \ 10^{-18}$$



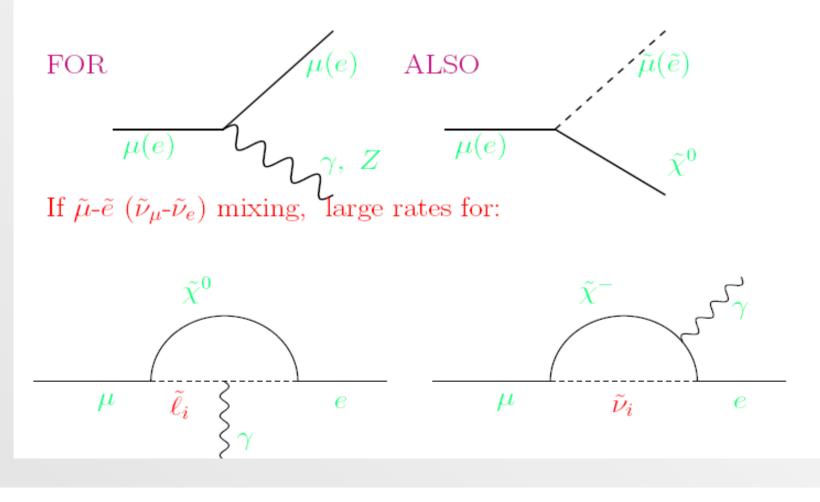


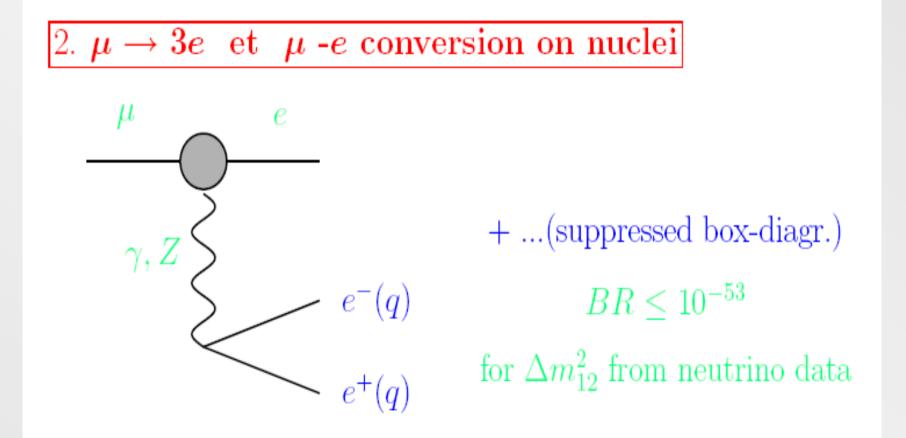
 $\nu_i = \nu_\mu \cos\theta + \nu_e \sin\theta, \ \Gamma = \frac{1}{16} \frac{G_F^2 \ m_\mu^5 \ \alpha}{128 \ \pi^4} \left(\frac{m_2^2 - m_1^2}{m_W^2}\right) \sin^2\theta \cos^2\theta$ $BR \le 10^{-50}, \text{ for } \Delta m_{12}^2 \text{ from neutrino data too small!}$

LFV in minimal SUSY

MSSM: For each SM vertex, also the one with

2 particles \rightarrow superparticles





Quite some space for deviations from universality in consistency with all bounds

Universality violations possible at a high scale? If yes, we need to predict their magnitude (possibly through flavour symmetries)

Even if we have universality at a high scale, it may be broken at low energies due to <u>quantum corrections</u>

Let us now look at all these in more detail:

Generic SLEPTON mass matrices

In the unrotated basis $\tilde{\ell}_i = (\tilde{e}_L, \tilde{\mu}_L, \tilde{\tau}_L, \tilde{e}_R^*, \tilde{\mu}_R^*, \tilde{\tau}_R^*)$, the slepton mass matrix reads as:

$${\cal L}_M=-rac{1}{2} ilde{\ell}^\dagger M_{ ilde{\ell}}^2 ilde{\ell}, \ \ M_{ ilde{\ell}}^2=\left(egin{array}{cc} M_{LL}^2&M_{LR}^2\ M_{RL}^2&M_{RR}^2\end{array}
ight),$$

where

$$\begin{split} M_{LL}^2 &= \frac{1}{2} m_{\ell}^{\dagger} m_{\ell} + M_L^2 - \frac{1}{2} (2m_W^2 - m_Z^2) \cos 2\beta \ I \\ M_{RR}^2 &= \frac{1}{2} m_{\ell}^{\dagger} m_{\ell} + M_R^2 - (m_Z^2 - m_W^2) \cos 2\beta \ I \\ M_{LR}^2 &= (A^e - \mu \tan \beta) \ m_{\ell} \\ M_{RL}^2 &= (M_{LR}^2)^{\dagger} \end{split}$$

$$A^e_{ij} = A_0 \cdot \delta_{ij}, \quad M^2_L = M^2_R = m^2_0 \ I$$

$$\delta_{XX}^{ij} = (M_{XX}^2)^{ij} / (M_{XX}^2)^{ii}$$
 (X = L, R)

Universal soft terms at GUT(mSUGRA models)

In a basis such that m_{l} is diagonal:

$$M_{\tilde{\ell}}^2 = egin{pmatrix} m_{\tilde{\ell}_L}^2 & 0 & 0 & ar{A}_{ ilde{e}} \cdot m_e & 0 & 0 \ 0 & m_{ ilde{\mu}_L}^2 & 0 & 0 & ar{A}_{ ilde{\mu}} \cdot m_\mu & 0 \ 0 & 0 & m_{ ilde{ au}_L}^2 & 0 & 0 & ar{A}_{ ilde{ au}} \cdot m_ au \ egin{pmatrix} 0 & 0 & m_{ ilde{ au}_L}^2 & 0 & 0 \ ar{A}_{ ilde{ au}} \cdot m_e & 0 & 0 & m_{ ilde{ au}_R}^2 & 0 \ 0 & ar{A}_{ ilde{ au}} \cdot m_\mu & 0 & 0 & m_{ ilde{ au}_R}^2 & 0 \ egin{pmatrix} 0 & 0 & M_{ ilde{ au}_R}^2 & 0 & 0 \ eta & 0 & ar{A}_{ ilde{ au}} \cdot m_\mu & 0 & 0 & m_{ ilde{ au}_R}^2 & 0 \ eta & 0 & 0 & ar{A}_{ ilde{ au}} \cdot m_\mu & 0 & 0 & m_{ ilde{ au}_R}^2 & 0 \ eta & 0 & 0 & ar{A}_{ ilde{ au}} \cdot m_\mu & 0 & 0 & m_{ ilde{ au}_R}^2 & 0 \ eta & 0 & 0 & ar{A}_{ ilde{ au}} \cdot m_\mu & 0 & 0 & m_{ ilde{ au}_R}^2 & 0 \ eta & 0 & eta & ar{ au}_R & eta & et$$

The 1st and 2nd generation sleptons are almost degenerate:

$$m_{ ilde{ au}_L} < m_{ ilde{ ee}_L} = m_{ ilde{\mu}_L}$$
; $m_{ ilde{ au}_R} < m_{ ilde{ ee}_R} = m_{ ilde{\mu}_R}$

$$m_{ ilde{ au}_R} < m_{ ilde{ au}_L}$$

Simplest Scheme: 2-3 Lepton Flavour Violation

(motivated by large Atmospheric Neutrino Mixing)

If large 2-3 mixing, in a basis such that m_1 is diagonal:

$$M_{\tilde{\ell}}^{2} = \begin{pmatrix} m_{\tilde{e}_{L}}^{2} & 0 & 0 & \bar{A}_{\tilde{e}} \cdot m_{e} & 0 & 0 \\ 0 & m_{\tilde{\mu}_{L}}^{2} & M_{LL}^{2} & 0 & \bar{A}_{\tilde{\mu}} \cdot m_{\mu} & 0 \\ 0 & M_{LL}^{2} & m_{\tilde{\tau}_{L}}^{2} & 0 & 0 & \bar{A}_{\tilde{\tau}} \cdot m_{\tau} \\ \hline \bar{A}_{\tilde{e}} \cdot m_{e} & 0 & 0 & m_{\tilde{e}_{R}}^{2} & 0 & 0 \\ 0 & \bar{A}_{\tilde{\mu}} \cdot m_{\mu} & 0 & 0 & m_{\tilde{\mu}_{R}}^{2} & M_{RR}^{2} \\ 0 & 0 & \bar{A}_{\tilde{\tau}} \cdot m_{\tau} & 0 & M_{RR}^{2} & m_{\tilde{\tau}_{R}}^{2} \end{pmatrix}$$

Flavor mixing entries are defined as

$$\delta_{XX}^{ij} = (M_{XX}^2)^{ij} / (M_{XX}^2)^{ii} \quad (X = L, R).$$

LFV through RGE effects

Including See-Saw neutrinos

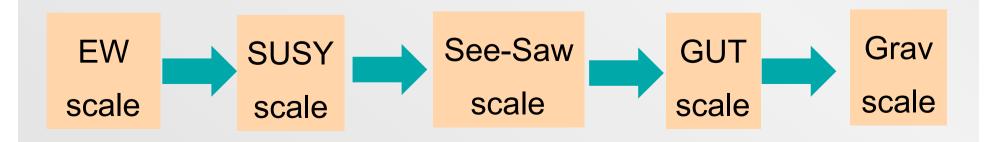
Leptonic Sector of Superpotential (up to See-Saw scale):

$$W = N_i^c (\lambda_{\nu})_{ij} L_j H_2 - E_i^c (\lambda_e)_{ij} L_j H_1 + \frac{1}{2} N^c{}_i \mathcal{M}_{ij} N_j^c + \mu H_2 H_1$$

Superpotential of Effective Low Energy Theory:

$$W_{eff} = L_i H_2 \left(\lambda_{\nu}^T \left(\mathcal{M}^D \right)^{-1} \lambda_{\nu} \right)_{ij} L_j H_2 - E_i^c (\lambda_e)_{ij} L_j H_1$$

RGE evolution: Various Steps



Even if:

$$M_{\text{GUT}}: m_{\tilde{\ell}, \tilde{\nu}} \propto \begin{pmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1 \end{pmatrix}, \quad \text{REGs} \longrightarrow \begin{pmatrix} 1 & \star & \star \\ \star & 1 & \star \\ \star & \star & 1 \end{pmatrix}$$

• RGEs for the charged-lepton mass matrix

$$t\frac{d}{dt}(m_{\tilde{\ell}}^2)_i^j = \frac{1}{16\pi^2} \left\{ (m_{\tilde{\ell}}^2 \lambda_\ell^\dagger \lambda_\ell)_i^j + (m_{\tilde{\ell}}^2 \lambda_\nu^\dagger \lambda_\nu)_i^j + \dots \right\}$$

The corrections in the basis where $(\lambda_{\ell}^{\dagger}\lambda_{\ell})_{i}^{j}$ is diagonal, are:

$$\delta m_{\tilde{\ell}} \propto \frac{1}{16\pi} \ln \frac{M_{\rm GUT}}{M_N} \lambda_{\nu}^{\dagger} \lambda_{\nu} m_{\rm SUSY}^2$$

(And similar corrections for $\delta m_{\tilde{\nu}}$)

For big μ -e lepton mixing, big rates for $\mu \to e\gamma$

Provided the neutrino Yukawas are sufficiently large (larger MN in See-Saw)

Corrections from Massive Neutrinos to Yukawa Couplings (from MGUT to MN)

$$\begin{split} 16\pi^2 \mu \frac{d}{d\mu} f_{e_{ij}} &= \left\{ -\frac{9}{5} g_1^2 - 3g_2^2 + 3 \mathrm{Tr}(f_d f_d^{\dagger}) + \mathrm{Tr}(f_e f_e^{\dagger}) \right\} f_{e_{ij}} \\ &\quad + 3(f_e f_e^{\dagger} f_e)_{ij} + (f_e f_{\nu}^{\dagger} f_{\nu})_{ij}, \\ 16\pi^2 \mu \frac{d}{d\mu} f_{\nu_{ij}} &= \left\{ -\frac{3}{5} g_1^2 - 3g_2^2 + 3 \mathrm{Tr}(f_u f_u^{\dagger}) + \mathrm{Tr}(f_{\nu} f_{\nu}^{\dagger}) \right\} f_{\nu_{ij}} \\ &\quad + 3(f_{\nu} f_{\nu}^{\dagger} f_{\nu})_{ij} + (f_{\nu} f_e^{\dagger} f_e)_{ij}. \end{split}$$

 $\frac{(\text{very simple at 1-loop, small tan}\beta)}{16\pi^2 \frac{d}{dt}\lambda_{\tau}} = (\lambda_N^2 - G_E)\lambda_{\tau}$ $16\pi^2 \frac{d}{dt}\lambda_N = (4\lambda_N^2 + 3\lambda_t^2 - G_N)\lambda_N$

<u>Slepton contributions from runs if Mgrav ~MGUT</u> <u>RGEs from Mgrav ~ $M_{GUT} \rightarrow M_{N}$ </u>

$$egin{aligned} &(m_{ ilde{L}}^2)_{ij}&\simeq&-rac{1}{8\pi^2}(3m_0^2+a_0^2)V_{Dki}^*V_{Dkj}f_{
u_k}^2\lograc{M_{ ext{grav}}}{M_{
u_k}}\ &(m_{ ilde{e}}^2)_{ij}&\simeq&0,\ &A_e^{ij}&\simeq&-rac{3}{8\pi^2}a_0f_{e_i}V_{Dki}^*V_{Dkj}f_{
u_k}^2\lograc{M_{ ext{grav}}}{M_{
u_k}}, \end{aligned}$$

Contribute only to LL slepton mixing

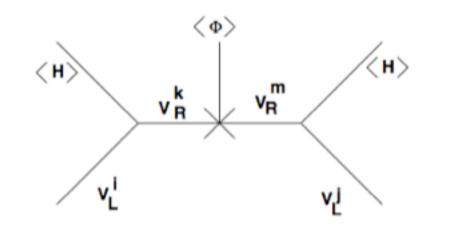
Slepton contributions from runs if Mgrav > MGUT

Renormalisation effects from $M_{GUT} \rightarrow M_G (= M_P?)$

$$\begin{array}{lll} (m_L^2)_{ij} &\simeq & -\frac{2}{(4\pi)^2} U_{ik} f_{\nu_k}^2 U_{jk}^{\star} (3m_0^2 + A_0^2) \log \frac{M_G}{M_{N_k}}, \\ (m_{\overline{E}}^2)_{ij} &\simeq & -\frac{6}{(4\pi)^2} \mathrm{e}^{-i\varphi_{d_i}} V_{ki}^{\star} f_{u_k}^2 V_{kj} \mathrm{e}^{i\varphi_{d_j}} (3m_0^2 + A_0^2) \log \frac{M_G}{M_{GUT}}, \end{array}$$

Contribute also to RR slepton mixing

Keep also in mind renormalisation of effective neutrino operator



$$16\pi^2 \frac{d}{dt} \sin^2 2\theta_{23} = -2\sin^2 2\theta_{23} (1 - \sin^2 2\theta_{23}) (Y_\tau^2) \frac{m_{eff}^{33} + m_{eff}^{22}}{m_{eff}^{33} - m_{eff}^{22}}$$

$$\frac{m_{eff}^{ij}}{m_{eff,0}^{ij}} = exp\left\{\frac{1}{8\pi^2}\int_{t_0}^t \left(-c_i g_i^2 + 3\lambda_t^2 + \frac{1}{2}(\lambda_i^2 + \lambda_j^2)\right)\right\} \quad I_{\lambda_i} = exp\left[\frac{1}{8\pi^2}\int_{t_0}^t \lambda_i^2 dt\right]$$
$$= I_g \cdot I_t \cdot \sqrt{I_{\lambda_i}} \cdot \sqrt{I_{\lambda_j}}$$

SU(5) - small tan β

 $(large \ top, \ \underline{small} \ bottom \ and \ tau \ Yuka$

Observables	Numerical Values
$m_{ u_1}/m_{ u_3}$	0.16(0.09)
$m_{ u_2}/m_{ u_3}$	0.20(0.37)
M_{1}/M_{3}	0.06(0.16)
M_{2}/M_{3}	0.12(0.42)
θ_{23}	0.91(1.13)
θ_{12}	0.56(0.10)
θ_{13}	0.21(0.07)
δ	0.43(-0.45)
J_{CP}	0.0088(0.00066)
J_{CP} ϵ_1	0.0088(0.00066) 0.00046(0.0039)
ϵ_1	0.00046(0.0039)
ϵ_1 ϵ_2	0.00046(0.0039) 0.0007(0.00014)
ϵ_1 ϵ_2 ϵ_3	0.00046(0.0039) 0.0007(0.00014) 0.0003(0.000001)

[Ellis, Gomez, SL]

LFV through non-Universality at high energies

SU(5) unification

$$5^*: \qquad \begin{pmatrix} d_1^c \\ d_2^c \\ d_3^c \\ e^- \\ -\nu_e \end{pmatrix}_L \qquad , \qquad \begin{pmatrix} s_1^c \\ s_2^c \\ s_3^c \\ \mu^- \\ -\nu_\mu \end{pmatrix}_L \qquad , \qquad \begin{pmatrix} b_1^c \\ b_2^c \\ b_3^c \\ \tau^- \\ -\nu_\mu \end{pmatrix}_L$$

SU(5) to SU(3) X SU(2) X U(1) decomposition:

<u>5*: (3*, 1, 1/3) + (1, 2*, -1/2)</u>

SU(5) unification

$$10: \quad \begin{pmatrix} 0 & u_3^c & -u_2^c & | & -u^1 & -d^1 \\ -u_3^c & 0 & u_1^c & | & -u^2 & -d^2 \\ u_2^c & -u_1^c & 0 & | & -u^3 & -d^3 \\ \hline u^1 & u^2 & u^3 & | & 0 & -e^+ \\ d^1 & d^2 & d^3 & | & e^+ & 0 \end{pmatrix}_L, \quad \begin{pmatrix} c^i & \\ \mu^+ & c_i^c \\ s^i & \end{pmatrix}_L, \quad \begin{pmatrix} t^i & \\ \tau^+ & t_i^c \\ b^i & \end{pmatrix}_L$$

SU(5) to SU(3) X SU(2) X U(1) decomposition:

 $10: (3^*, 1, -2/3) + (3, 2, 1/6) + (1, 1, 1)$



(i) Assume the family symmetry is combined with SU(5)
(ii) Use the GUT structure ONLY to constrain U(1) charges
Under this group we have the following relations:

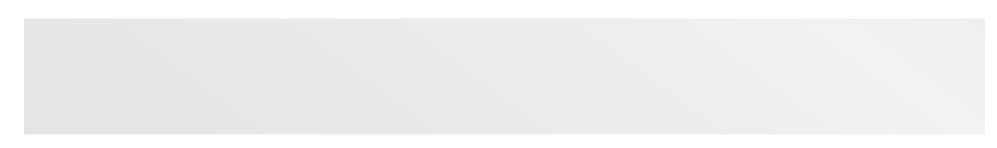
 $egin{array}{rcl} Q_{(q,u^c,e^c)_i} &=& Q_i^{10} \ Q_{(l,d^c)_i} &=& Q_i^{\overline{5}} \ Q_{(
u_R)_i} &=& Q_i^{
u_R} \end{array}$

• M_{up} symmetric • $M_{\ell^{\pm}} = M_{down}^T$

• L lepton mixing \approx R down-quark one

Can we obtain acceptable patterns of masses/mixings? i.e.

$$\begin{split} \frac{M_u}{m_t} &= \begin{pmatrix} \bar{\epsilon}^6 & \bar{\epsilon}^5 & \bar{\epsilon}^3 \\ \bar{\epsilon}^5 & \bar{\epsilon}^4 & \bar{\epsilon}^2 \\ \bar{\epsilon}^3 & \bar{\epsilon}^2 & 1 \end{pmatrix}, \frac{M_{down}}{m_b} = \begin{pmatrix} \bar{\epsilon}^4 & \bar{\epsilon}^3 & \bar{\epsilon}^3 \\ \bar{\epsilon}^3 & \bar{\epsilon}^2 & \bar{\epsilon}^2 \\ \bar{\epsilon} & 1 & 1 \end{pmatrix} \\ \\ \frac{M_\ell}{m_\tau} &= \begin{pmatrix} \bar{\epsilon}^4 & \bar{\epsilon}^3 & \bar{\epsilon} \\ \bar{\epsilon}^3 & \bar{\epsilon}^2 & 1 \\ \bar{\epsilon}^3 & \bar{\epsilon}^2 & 1 \end{pmatrix}, m_{eff} \propto \begin{pmatrix} \bar{\epsilon}^2 & \bar{\epsilon} & \bar{\epsilon} \\ \bar{\epsilon} & 1 & 1 \\ \bar{\epsilon} & 1 & 1 \end{pmatrix} \end{split}$$



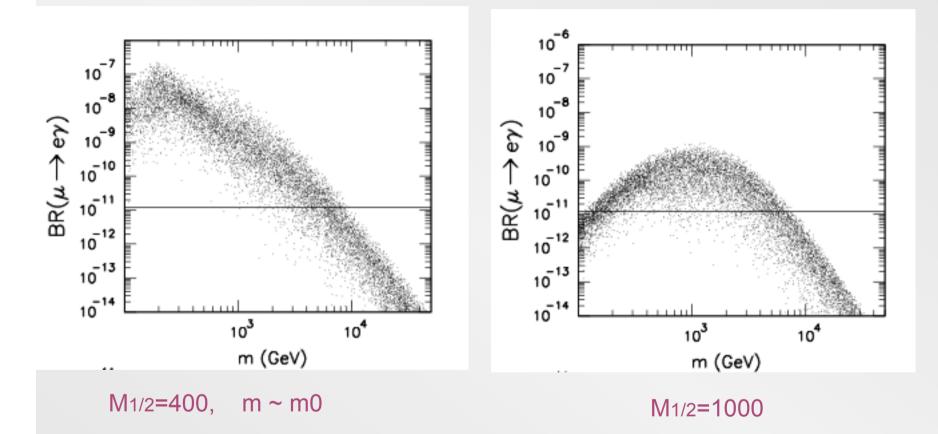
Flavour symmetries determine soft SUSY terms

$$\mathcal{L}_{m^{2}} = m_{0}^{2}(\phi_{1}^{*}\phi_{1} + \phi_{2}^{*}\phi_{2} + \phi_{3}^{*}\phi_{3} + \left(\frac{\langle\theta\rangle}{M_{\mathrm{fl}}}\right)^{q_{2}-q_{1}}\phi_{1}^{*}\phi_{2} + \left(\frac{\langle\theta\rangle}{M_{\mathrm{fl}}}\right)^{q_{3}-q_{1}}\phi_{1}^{*}\phi_{3} + \left(\frac{\langle\theta\rangle}{M_{\mathrm{fl}}}\right)^{q_{3}-q_{2}}\phi_{2}^{*}\phi_{3} + \mathrm{h.c.}).$$

L-R symmetric	(1	$\tilde{\epsilon}^{ a+2b }$	$\tilde{\epsilon}^{ a+b }$	
		$\widetilde{\epsilon}^{ a+2b }$	1	$\widetilde{\epsilon}^{ b }$	
	l	$\widetilde{\epsilon}^{ a+b }$	$\widetilde{\epsilon}^{ b }$	1)

$$\frac{\mathrm{SU}(5)}{\mathrm{E}_{L} \sim \begin{pmatrix} 1 & \lambda^{2} & \lambda^{2} \\ \lambda^{2} & 1 & 1 \\ \lambda^{2} & 1 & 1 \end{pmatrix}} \qquad \mathrm{E}_{R} \sim \begin{pmatrix} 1 & \lambda & \lambda^{3} \\ \lambda & 1 & \lambda^{2} \\ \lambda^{3} & \lambda^{2} & 1 \end{pmatrix}$$
$$\frac{Q_{(q,u^{c},e^{c})_{i}} = Q_{i}^{10}}{Q_{(l,d^{c})_{i}} = Q_{i}^{5}} \qquad \qquad \frac{M_{\ell}}{m_{\tau}} = \begin{pmatrix} \overline{\epsilon}^{4} & \overline{\epsilon}^{3} & \overline{\epsilon} \\ \overline{\epsilon}^{3} & \overline{\epsilon}^{2} & 1 \\ \overline{\epsilon}^{3} & \overline{\epsilon}^{2} & 1 \end{pmatrix}}{\mathrm{R}: (3,2,0)}$$
$$\mathrm{E}_{R}: (3,2,0)$$

[Chankowski,Kowalska,Lavignac,Pokorski]





Viable models with non-Abelian flavour symmetries

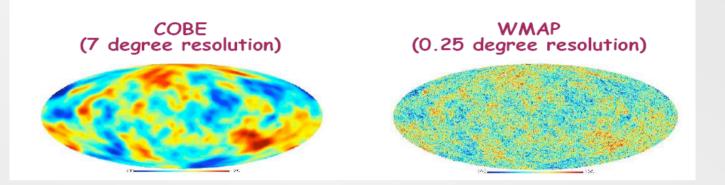
i.e. SU(3) family symmetry [Antusch,King,Malinsky]

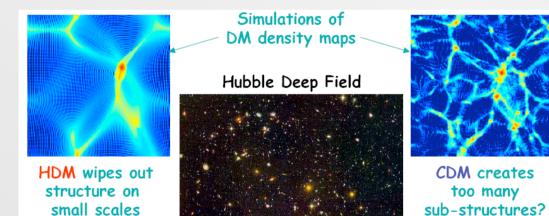
$$\begin{split} \hat{m}_{\nu^{c}}^{2} &\approx m_{0}^{2} \left[b_{0}^{\nu^{c}} \mathbbm{1} + \begin{pmatrix} \varepsilon^{2} \overline{\varepsilon}^{2} \, b_{1}^{\nu^{c}} \, \varepsilon^{2} \overline{\varepsilon}^{2} \, b_{1}^{\nu^{c}} e^{i\phi_{1}} \, \varepsilon^{2} \overline{\varepsilon}^{2} \, b_{1}^{\nu^{c}} e^{i\phi_{2}} \\ \cdot \, \varepsilon^{2} b_{2}^{\nu^{c}} \, \varepsilon^{2} b_{2}^{\nu^{c}} e^{i\phi_{3}} \\ \cdot \, \cdot \, \varepsilon^{2} b_{2}^{\nu^{c}} \, \varepsilon^{2} b_{3}^{\nu^{c}} \end{pmatrix} + \dots \right] \\ \hat{m}_{e^{c}}^{2} &\approx m_{0}^{2} \left[b_{0}^{e^{c}} \mathbbm{1} + \begin{pmatrix} \overline{\varepsilon}^{4} b_{1}^{e^{c}} \, \overline{\varepsilon}^{4} b_{1}^{e^{c}} e^{i\phi_{1}} \, \overline{\varepsilon}^{4} b_{1}^{e^{c}} e^{i\phi_{2}} \\ \cdot \, \varepsilon^{2} b_{2}^{e^{c}} \, \varepsilon^{2} b_{2}^{e^{c}} e^{i\phi_{3}} \\ \cdot \, \varepsilon^{2} b_{2}^{e^{c}} \, \varepsilon^{2} b_{2}^{e^{c}} e^{i\phi_{3}} \\ \cdot \, \varepsilon^{2} b_{2}^{e^{c}} \, \varepsilon^{2} b_{3}^{e^{c}} \end{pmatrix} + \dots \right] . \end{split}$$

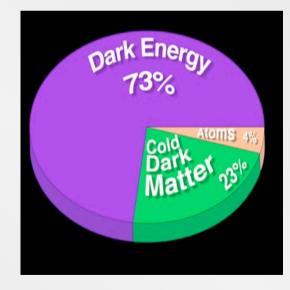
Viable predictions for smaller SUSY masses

Including Dark Matter Considerations in SUSY with massive neutrinos

COSMOLOGICAL OBSERVATIONS







Dark Matter- What can it be?

Has to be Weakly Interacting (*if strong or EM would couple to matter and be detectable*)

- Baryonic? (i.e. Neutron Stars, Black Holes)
- -<u>Neutrinos?</u> (simpest schemes excluded by WMAP)
- Axions?

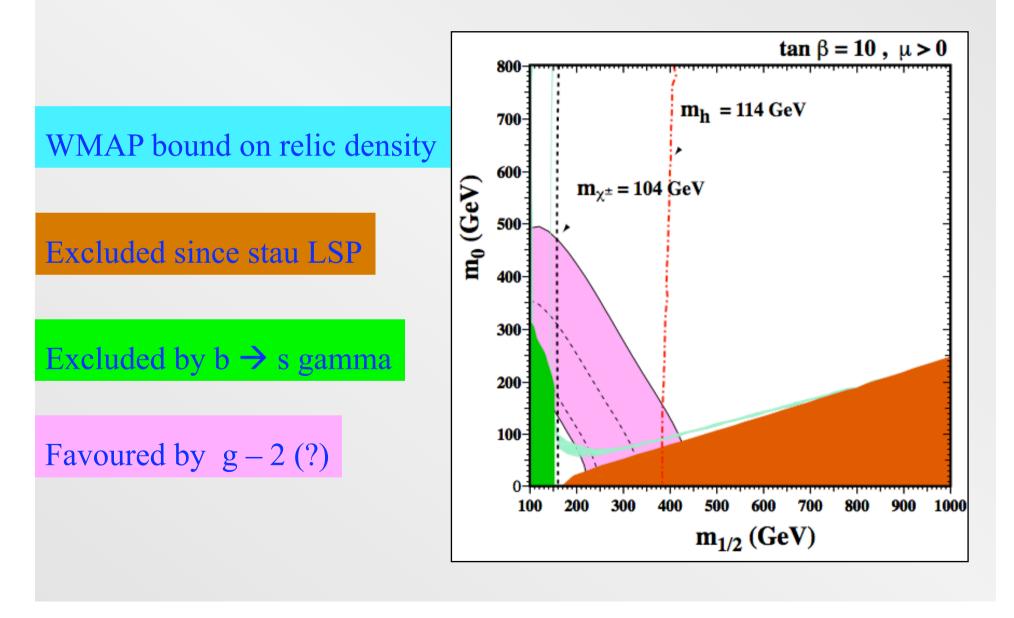
-Lightest SUSY particle (LSP)?

- o <u>Neutralino?</u>
- <u>S-neutrino</u> (excluded by LEP direct searches)
- Gravitino? (would be really hard to detect)

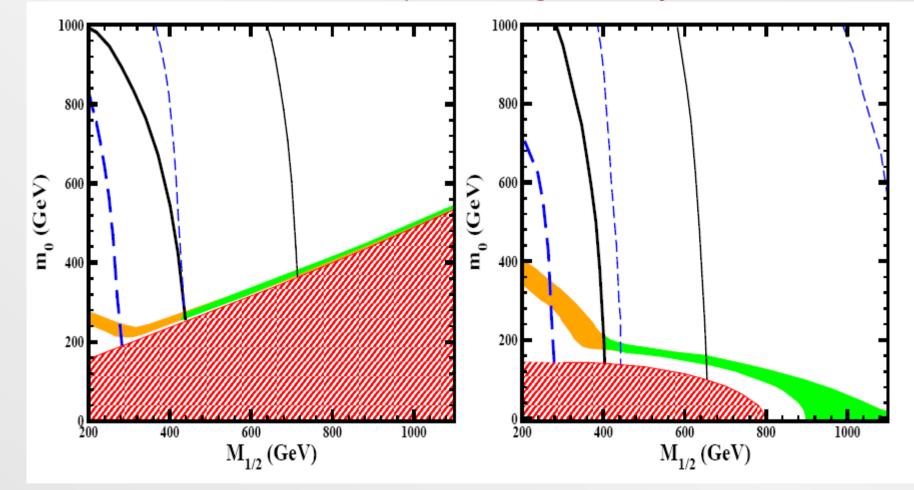
The parametric space favoured by Cosmology also constrains the allowed channels in Colliders (for instance SUSY cascade chains)

WMAP on CMSSM (stable neutralino LSP)

[Ellis, Olive, Santoso, Spanos]

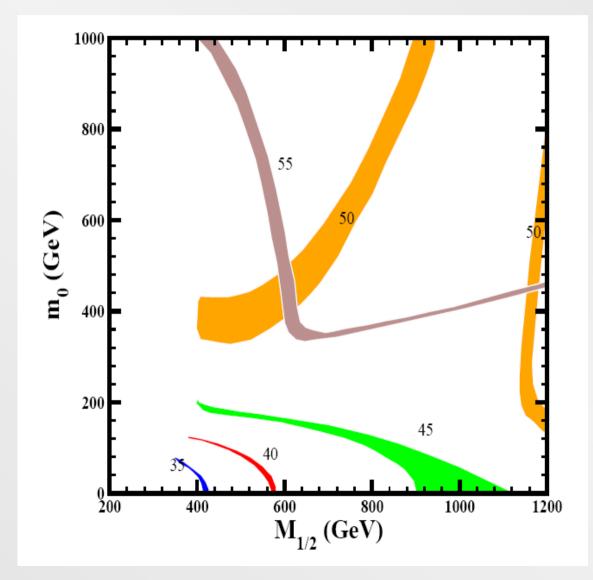


If massive neutrinos with large Yukawa couplings and/or additional GUT corrections, picture significantly modified!



Effect of quantum corrections to Yukawas & (s-)particle masses [Gomez, SL, Naranjo,Rodriguez-Quintero] Allowed region (linked to sparticle spectra)

very sensitive to $tan\beta$



LFV in Colliders

Bounds from Rare processes very strong

Can the LHC or the ILC see LFV?

In which channels?

For which area of the SUSY parameter space?

IT TURNS OUT THAT:

♦In general, the LHC and the ILC good for probes for a heavy sparticle spectrum

Also for areas where cancellations take place in the loop diagrams for rare decays

[Gomez, Leontaris, SL, Vergados]

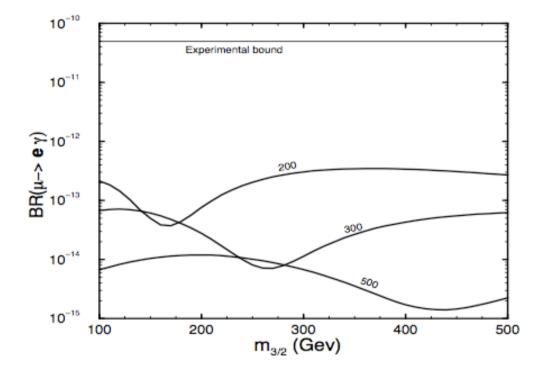
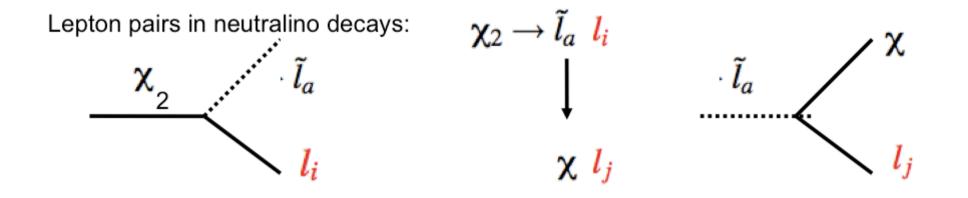


Figure 3: $BR(\mu \rightarrow e\gamma)$ for a range of values of $m_{1/2}$ (labeled above). Universal soft masses at the GUT scale are considered ($\Delta = 0$). The curves are obtained using $\tan \beta = 7$ and $A_0 = -1.5m_{3/2}$ as input parameters.

Example of FC versus LFV

[Hisano, Kitano, Nojiri]

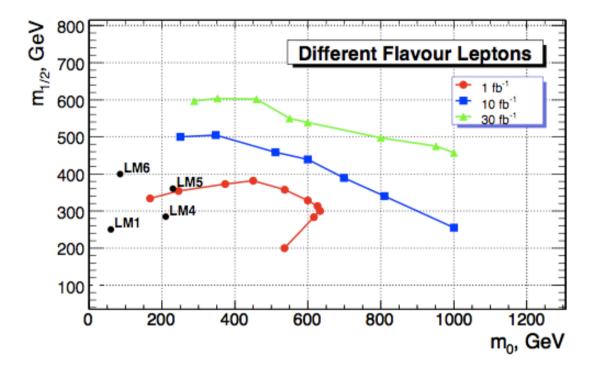


LFV: Decays involving leptons of different flavour

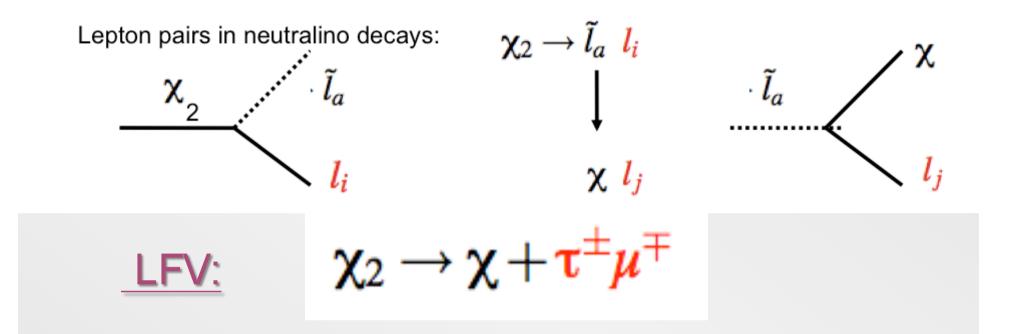
How large LFV?

Depends on sfermion mixing (back to MODEL BUILDING)

LFV in µ-e channel [Andreev, Bityukov, Krasnikov, Toropin]

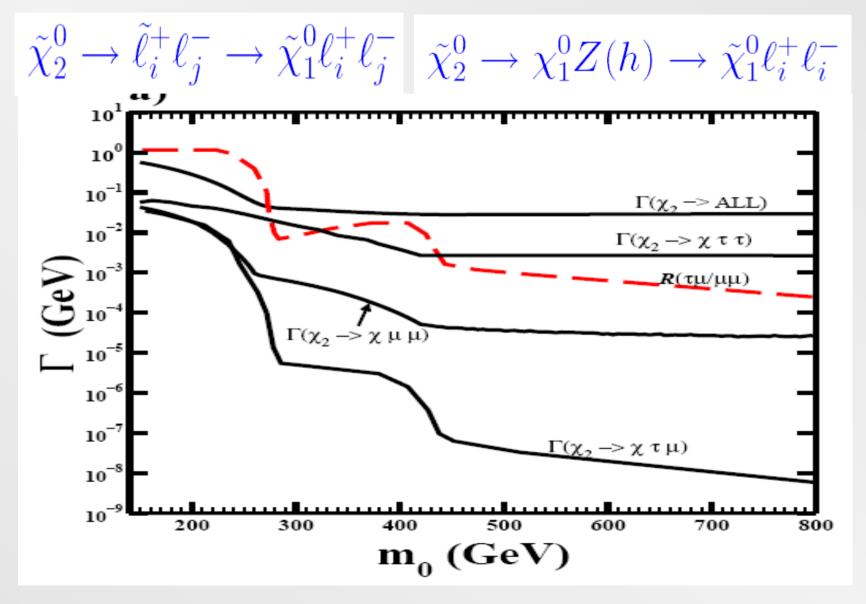


LFV in the τ –μ channel

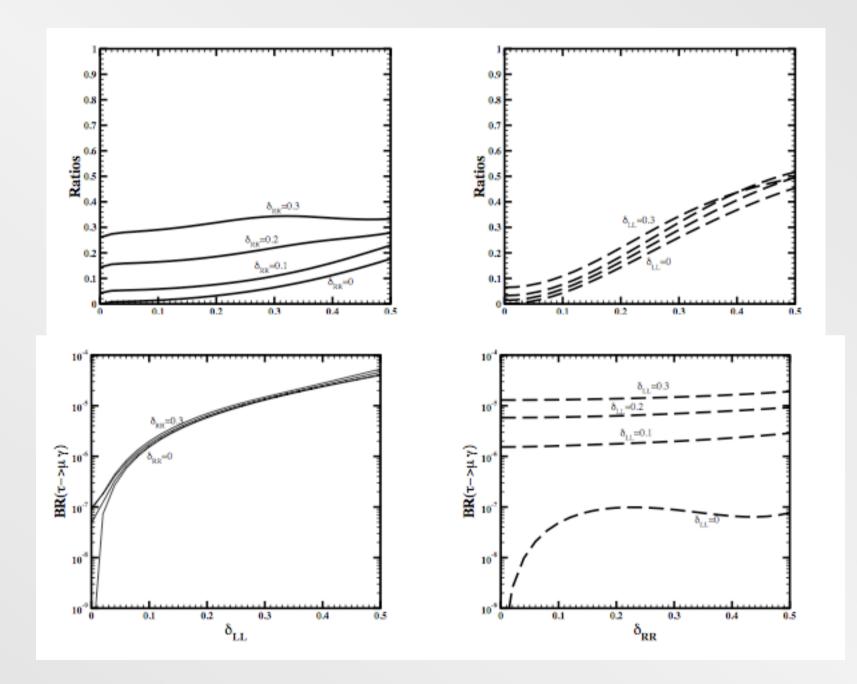


[Hinchliffe,Paige]

LFV at the LHC

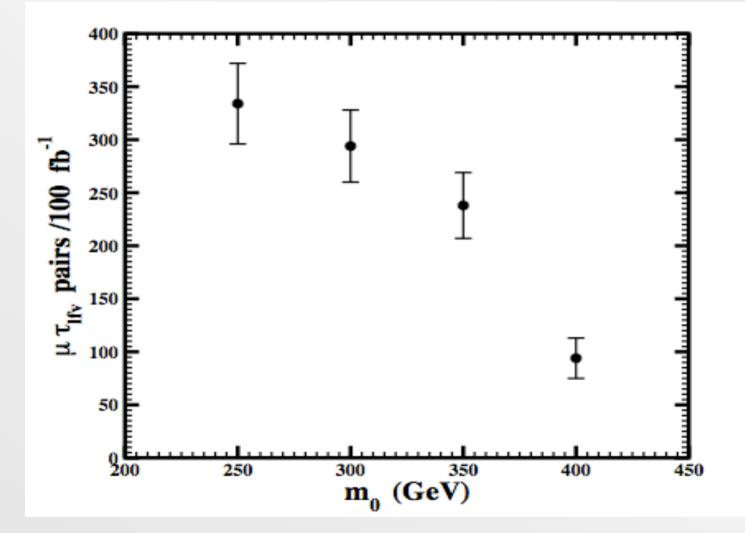


[Ellis, Carvahlo, Gomez, SL, Romao]



[Carquin, Ellis, Gomez, SL, Rodriguez-Quintero]

Results for Varying m_0 at Fixed $M_{1/2}$



LFV at a LC

$$\begin{array}{rcl} e^+e^- & \rightarrow & \tilde{\ell}_i^-\tilde{\ell}_j^+ \rightarrow \tau^\pm \mu^\mp \tilde{\chi}_1^0 \tilde{\chi}_1^0 \\ e^+e^- & \rightarrow & \tilde{\nu}_i \tilde{\nu}_j^c \rightarrow \tau^\pm \mu^\mp \tilde{\chi}_1^+ \tilde{\chi}_1^- \\ e^+e^- & \rightarrow & \tilde{\chi}_2^\pm \tilde{\chi}_1^\mp \rightarrow \tau^\pm \mu^\mp \tilde{\chi}_1^+ \tilde{\chi}_1^- \\ e^+e^- & \rightarrow & \tilde{\chi}_2^0 \tilde{\chi}_1^0 \rightarrow \tau^\pm \mu^\mp \tilde{\chi}_1^0 \tilde{\chi}_1^0 \end{array}$$

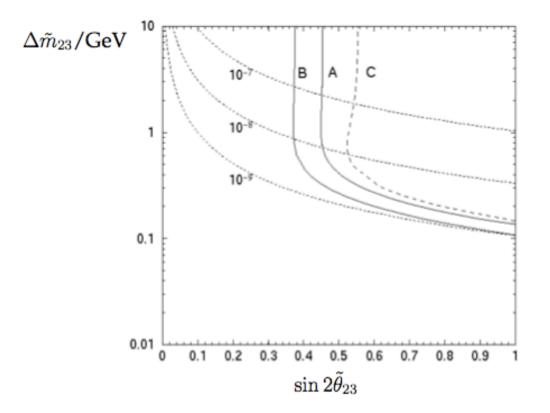


Figure 1: Various 3σ significance contours in the $\Delta \tilde{m}_{23} - \sin 2\tilde{\theta}_{23}$ plane, for the SUSY point mentioned in the text. The contours A and B show the integrated signals (8–9) at $\sqrt{s} = 500 \text{ GeV}$ and for 500 fb⁻¹ and 1000 fb⁻¹, respectively. The contour C shows the $\tilde{\nu}\tilde{\nu}^{c}$ contribution separately for 500 fb⁻¹ [6]. The dotted lines indicate contours for $Br(\tau \rightarrow \mu\gamma)=10^{-7}$, 10^{-8} and 10^{-9} [11].

[Deppish,Kalinowski,Pas,Redelbach,Ruckl]

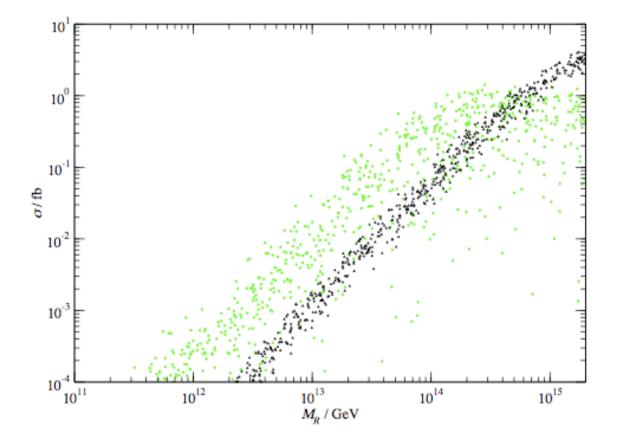


Figure 2: Cross-sections at $\sqrt{s} = 500$ GeV for $e^+e^- \rightarrow \mu^+e^- + 2\tilde{\chi}_1^0$ (circles) and $e^+e^- \rightarrow \tau^+\mu^- + 2\tilde{\chi}_1^0$ (triangles) in scenario B.

Scenario	$m_{1/2}/\text{GeV}$	m_0/GeV	aneta	\tilde{m}_6/GeV	$\tilde{\Gamma}_6/\text{GeV}$	$m_{ ilde{\chi}_1^0}/{ m GeV}$
В	250	100	10	208	0.32	98

CONCLUSIONS

✓ Neutrino data point towards SM extensions with LFV

LFV strongly bounded by various experimental processes, but there is significant space for LFV physics searche

 Predictions differ sufficiently enough to give input for Model Building Aspects and sources of LFV
 Non-universalities versus RGE effects
 Minimal versus non-Minimal GUT schemes
 R-conserving versus R-violating SUSY

<u>A very exciting ERA to come!</u>