

Supercurrents in $\mathcal{N}=1$ Minimal Supergravity in the Superconformal Formalism

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Outline

- *Review of Callan-Coleman-Jackiw (CCJ) improved currents*
- *Non-linear form of Ferrara-Zumino equations from 1975*
- *Generalized expressions for Ward identities of Einstein tensor and Supercurrent multiplets using superconformal approach*
 - *Old Minimal supergravity*
 - *Old Minimal with FI term*
 - *New Minimal supergravity*
- *Conclusions*

- In general for every rigid symmetry there is a conserved current upon using equations of motion

- Currents from gauge couplings

For example $S = \int d^4x \left[-\frac{1}{4g^2} F_{\mu\nu} F^{\mu\nu} + A^\mu J_\mu + \dots \right] \rightarrow$

$$\partial^\mu F_{\mu\nu} = J_\nu \quad \partial^\mu J_\mu \approx 0$$

Pure gravity $G_{\mu\nu} \approx 0$

- Symmetric energy momentum tensor is due to Lorentz rotation invariance:
 $T_{\mu\nu} = T_{\nu\mu}$

- When there is a conformal symmetry *C. G. Callan, Jr., S. R. Coleman and R. Jackiw, Annals Phys. 59 (1970) 42-73* proved that there is a traceless improved current $\Theta_{\mu\nu} = T_{\mu\nu} + \text{improvement term}$

Example:

$$S = \int d^4x \sqrt{-g} \left(\frac{1}{2} \partial_\mu \varphi \partial^\mu \varphi - V(\varphi) \right)$$

has a symmetric and conserved, but not traceless canonical energy-momentum tensor

$$T_{\mu\nu} = \partial_\mu \varphi \partial_\nu \varphi + g_{\mu\nu} \mathcal{L}_M, \quad \partial^\mu T_{\mu\nu} = 0, \quad T_\mu{}^\mu = -\partial^\mu \varphi \partial_\mu \varphi + 4V$$

The improved energy momentum tensor

$$\Theta_{\mu\nu} = T_{\mu\nu} - \frac{1}{6} (\partial_\mu \partial_\nu - g_{\mu\nu} \square) \varphi^2$$

$$\partial^\mu \Theta_{\mu\nu} = 0, \quad \Theta_\mu{}^\mu \approx 0 \quad \text{quartic } V$$

\approx means equation of motions $\square \varphi \approx -V_\varphi$ is satisfied.

The improved energy-momentum tensor can also be obtained from the conformal invariant gravity-coupled action.

A conventional gravity theory

$$S = \int d^4x \sqrt{-g} \left[\frac{1}{2} \kappa^{-2} R + \mathcal{L}_M \right] \quad G_{\mu\nu} \approx \kappa^2 T_{\mu\nu}$$

To have to conformal and Weyl symmetry, replace $\kappa^{-2} \rightarrow \kappa^{-2} - \frac{1}{6} \varphi^2$. For the action

$$\mathcal{S} = \int d^4x \sqrt{-g} \left[\frac{1}{2} (\kappa^{-2} - \frac{1}{6} \varphi^2) R + \frac{1}{2} g^{\mu\nu} \partial_\mu \varphi \partial_\nu \varphi - V(\varphi) \right]$$

$$G_{\mu\nu} \approx \kappa^2 \Theta_{\mu\nu}^c, \quad \Theta_{\mu\nu}^c = T_{\mu\nu} - \frac{1}{6} (\nabla_\mu \partial_\nu - g_{\mu\nu} \nabla^\rho \partial_\rho) \varphi^2 + \frac{1}{6} \varphi^2 G_{\mu\nu}$$

$$\nabla^\mu \Theta_{\mu\nu}^c \approx 0, \quad \Theta_{\mu}^{\mu} \approx 0 \quad \text{for quartic } V(\varphi).$$

Rigid Poincaré

$$T_{\mu\nu} = \partial_\mu\varphi\partial_\nu\varphi + g_{\mu\nu}L$$

Rigid conformal

$$\Theta_{\mu\nu} = T_{\mu\nu} - \frac{1}{6}(\partial_\mu\partial_\nu - g_{\mu\nu}\square)\varphi^2$$

Local conformal

$$\Theta_{\mu\nu}^c = T_{\mu\nu} - \frac{1}{6}(\nabla_\mu\partial_\nu - g_{\mu\nu}\nabla^\rho\partial_\rho)\varphi^2 + \frac{1}{6}\varphi^2 G_{\mu\nu}$$

From conformal action

These formulations can be obtained from a conformal action, containing apart from the physical field φ also a compensating scalar φ_0 . Both have then Weyl weight 1

$$\begin{aligned}\mathcal{S} &= \int d^4 \sqrt{g} \left[-\frac{1}{2} \varphi_0 \square^C \varphi_0 + \frac{1}{2} \varphi \square^C \varphi + \lambda \varphi^4 \right] \\ &= \int d^4 \sqrt{g} \left[\frac{1}{2} \partial_\mu \varphi_0 \partial^\mu \varphi_0 - \frac{1}{2} \partial_\mu \varphi \partial^\mu \varphi + \frac{1}{12} (\varphi_0^2 - \varphi^2) R + \lambda \varphi^4 \right]\end{aligned}$$

The Einstein frame: $\varphi_0^2 = \varphi^2 + 6\kappa^{-2} \rightarrow \mathcal{N}_0$ conformal part in the action, not traceless energy-momentum tensor.

Conformal frame: $\varphi_0^2 = 6\kappa^{-2} \rightarrow$ Conformal action, traceless energy-momentum tensor.

$$\text{Old Minimal } \bar{D}^{\dot{\alpha}} E_{\alpha\dot{\alpha}} = D_{\alpha} \mathcal{R}, \quad \text{New Minimal } \bar{D}^{\dot{\alpha}} E_{\alpha\dot{\alpha}} = W_{\alpha} \\ \nabla^{\mu} G_{\mu\nu} = 0$$

At the linearized level

$$E_{\alpha\dot{\alpha}} + \kappa^2 J_{\alpha\dot{\alpha}} \approx 0 \quad \rightarrow \quad G_{\mu\nu} + \kappa^2 T_{\mu\nu} \approx 0$$

S. Ferrara and B. Zumino, Nucl. Phys. B134 (1978) 301-326, Z. Komargodski and N. Seiberg, JHEP 1007 (2010) 017

$$\text{Old Minimal } \bar{D}^{\dot{\alpha}} J_{\alpha\dot{\alpha}} \approx D_{\alpha} Y, \quad \text{New Minimal } \bar{D}^{\dot{\alpha}} J_{\alpha\dot{\alpha}} \approx \omega_{\alpha} \\ \bar{D}^{\dot{\alpha}} J_{\alpha\dot{\alpha}} \approx D_{\alpha} Y + \omega_{\alpha}.$$

S. Ferrara and B. Zumino, Nucl. Phys. B134 (1978) 301-326, S. M. Kuzenko, Eur. Phys. J. C71 (2011) 1513.

The trace equations:

$$\mathcal{R} \approx -\kappa^2 Y, \quad W_{\alpha} \approx -\kappa^2 \omega_{\alpha} \quad \rightarrow \quad R \approx -\kappa^2 T_{\mu}^{\mu}$$

Supercurrent and Einstein tensor from superconformal approach.

Action and field equations

The actions of chiral multiplets in the superconformal setup are symbolically obtained from

$$\mathcal{S} = \left[N(X^I, \bar{X}^{\bar{I}}) \right]_D + \left[\mathcal{W}(X^I) \right]_F$$

X^I are chiral multiplets with $(1, 1)$ Weyl and chiral weights. $I = 0, \dots, n$. n is the number of physical multiplets, 0 is for compensating multiplet.

$N = (2, 0)$ is real. $\mathcal{W} = (3, 3)$ is holomorphic.

$$S^0 = X^0, \quad S^i = \frac{X^i}{X^0}, \quad i = 1, \dots, n,$$

$$N(X, \bar{X}) = S^0 \bar{S}^{\bar{0}} \Phi(S, \bar{S}), \quad \mathcal{W}(X) = (S^0)^3 W(S)$$

Pure supergravity is obtained for

$$N = -3S^0\bar{S}^{\bar{0}}, \quad \Phi = -3.$$

This allows us to define the matter coupling function Φ_M as $\Phi(S, \bar{S}) = -3 + 3\Phi_M(S, \bar{S})$ such that

$$N(X, \bar{X}) = S^0\bar{S}^{\bar{0}}(-3 + 3\Phi_M(S, \bar{S})), \quad \mathcal{W}(X) = (S^0)^3W(S)$$

The field equation for the compensating multiplet, $I = 0$ can be written as

$$\mathcal{R} + \frac{Y}{(S^0)^2} \approx 0 \quad \text{with} \quad Y \equiv -2(S^0)^3\Delta W + S^0T(\bar{S}^{\bar{0}}\Delta K)$$

$\mathcal{R} \equiv \frac{1}{S^0}T(\bar{S}^{\bar{0}})$ *T. Kugo and S. Uehara, Prog.Theor.Phys. 73 (1985) 235.* The operation T is the superconformal version of the superspace operation \bar{D}^2 . In the conformal frame $S^0 = \kappa^{-1}$, we have $\mathcal{R} + \kappa^2Y \approx 0$.

Defined

$$\Delta K \equiv -\frac{1}{3\bar{S}^0} \left(N_0 + 3\bar{S}^0 \right) = S^i \Phi_{M i} - \Phi_M$$
$$\Delta W \equiv \frac{1}{3(S^0)^2} \mathcal{W}_0 = W - \frac{1}{3} S^i W_i .$$

Conformal case: $\Delta K = \Delta W = 0$. Thus W is homogeneous of rank 3 and Φ_M is homogeneous of rank 1 both in S^i and \bar{S}^i , $\Phi_{M i \bar{j}}$ has degree zero. One field $\Phi_M = S \bar{S}$ corresponds to the conformally coupled scalar of CCJ.

Non-linear form of Ferrara-Zumino equations 1975

We want to generalize the FZ equation $\bar{D}^{\dot{\alpha}} E_{\alpha\dot{\alpha}} = D_{\alpha} \mathcal{R}$ to the non linear level. In the superconformal formulation nonlinearities come from two sources: compensator dependence, coupling to matter. The non linear version of $E_{\alpha\dot{\alpha}}$ we denote as $\mathcal{E}_{\alpha\dot{\alpha}}$.

$$\bar{D}^{\dot{\alpha}} \mathcal{E}_{\alpha\dot{\alpha}} = (S^0)^{k_1} (\bar{S}^{\bar{0}})^{k_2} D_{\alpha} \left(\frac{\mathcal{R}}{S^0} \right)$$

Matching the Weyl and chiral weights on both sides we find

$$k_1 + k_2 = w, \quad k_1 - k_2 = 3$$

We found that the non-linear version is the tensor $\mathcal{E}_{\alpha\dot{\alpha}}$ and \mathcal{R} satisfy

$$\bar{D}^{\dot{\alpha}} \mathcal{E}_{\alpha\dot{\alpha}} = (S^0)^3 \mathcal{D}_{\alpha} \left(\frac{\mathcal{R}}{S^0} \right)$$

Generalized Bianchi identity. The scalar curvature multiplet \mathcal{R} , with chiral and Weyl weights (1, 1)

$$\mathcal{E}_{\alpha\dot{\alpha}} = -4i \bar{S}^{\bar{0}} \overleftrightarrow{\partial}_{\alpha\dot{\alpha}} S^0 - 2(D_{\alpha} S^0)(\bar{D}_{\dot{\alpha}} \bar{S}^{\bar{0}})$$

$$\mathcal{R} \equiv (S^0)^{-1} T(\bar{S}^{\bar{0}}).$$

In the components $\{X^0, \Omega^0, F^0\}$, using $\mathcal{E}_{\alpha\dot{\alpha}} = \frac{1}{4}i(\gamma^\mu)_{\alpha\dot{\alpha}}\mathcal{E}_\mu$

$$\mathcal{E}_\mu = 4iX^0\mathcal{D}_\mu\bar{X}^{\bar{0}} - 4i\bar{X}^{\bar{0}}\mathcal{D}_\mu X^0 + 2i\bar{\Omega}^0 P_L \gamma_\mu \Omega^{\bar{0}}$$

$$\mathcal{D}_\mu X^I = (\partial_\mu - b_\mu - iA_\mu) X^I - \frac{1}{\sqrt{2}}\bar{\psi}_\mu \Omega^I.$$

$$\begin{aligned} \mathcal{E}_\mu &= -8A_\mu\bar{X}^{\bar{0}}X^0 - 4i\bar{X}^{\bar{0}}\overset{\leftrightarrow}{\partial}_\mu X^0 \\ &+ 2i\bar{\Omega}^0 P_L \gamma_\mu \Omega^{\bar{0}} + 2i\sqrt{2}\bar{\psi}_\mu \left(\bar{X}^{\bar{0}}\Omega^0 - X^0\Omega^{\bar{0}} \right) \end{aligned}$$

Since we have used an explicit splitting in the action we can write

$$[N(X, \bar{X})]_D = S^0 \bar{S}^0 (1 - \Phi_M(S, \bar{S}))$$

using the relation between \mathcal{R} and Y and $\mathcal{E}_{\alpha\dot{\alpha}} + J_{\alpha\dot{\alpha}} \approx 0$ one finds the conservation law for the supercurrent

$$\begin{aligned} \bar{\mathcal{D}}^{\dot{\alpha}} J_{\alpha\dot{\alpha}} &\approx (S^0)^3 \mathcal{D}_\alpha \left(\frac{Y}{(X^0)^3} \right) \\ &\approx -(S^0)^3 \mathcal{D}_\alpha \left(2\Delta W - (S^0)^{-2} T \left(\bar{S}^0 \Delta K \right) \right) \end{aligned}$$

$\bar{\mathcal{D}}^{\dot{\alpha}} J_{\alpha\dot{\alpha}} \approx 0$ in the conformal frame.

For the compensator $S^0 = \{X^0, \chi^0, F^0\}$ and for the physical multiplets $S^i = \{S^i, \chi^i, F^i\}$

$$\begin{aligned}
 J_\mu &= -\Phi_M \mathcal{E}_\mu \\
 &+ 2i\bar{X}^{\bar{0}}\Phi_{M\bar{i}}\bar{X}^{\bar{0}}\gamma_\mu\chi^{\bar{i}} + 2iX^0\Phi_{Mi}\bar{X}^i\gamma_\mu\chi^0 \\
 &+ 2iX^0\bar{X}^{\bar{0}}\left[2(\Phi_{Mi}\mathcal{D}_\mu S^i - \Phi_{M\bar{i}}\mathcal{D}_\mu \bar{S}^{\bar{i}}) - \Phi_{M\bar{i}j}\bar{X}^i\gamma_\mu\chi^{\bar{j}}\right] \\
 \mathcal{E}_\mu &= 4iX^0\mathcal{D}_\mu\bar{X}^{\bar{0}} - 4i\bar{X}^{\bar{0}}\mathcal{D}_\mu X^0 + 2i\bar{X}^{\bar{0}}\gamma_\mu\chi^{\bar{0}}
 \end{aligned}$$

Choosing a frame:

In the Einstein frame : $-3\kappa^{-2} = N = 3X^0\bar{X}^{\bar{0}}(-1 + \Phi_M)$,

$$0 = N_I\Omega^I = 3\bar{X}^{\bar{0}}\left[(-1 + \Phi_M)\chi^0 + \Phi_{M\bar{i}}\chi^{\bar{i}}\right]$$

Supergravity and matter fields get mixed.

In the conformal frame : $X^0 = \kappa^{-1}$, $\Omega^0 = 0$

In this frame \mathcal{E}_μ does not depend on matter fields, and we have

$$\mathcal{E}_\mu = -8\kappa^{-2} A_\mu ,$$

$$J_\mu = 8\kappa^{-2} \Phi_M A_\mu + 2i\kappa^{-2} \left[2(\Phi_{M i} \mathcal{D}_\mu S^i - \Phi_{M \bar{i}} \mathcal{D}_\mu \bar{S}^{\bar{i}}) - \Phi_{M i \bar{j}} \bar{\chi}^i \gamma_\mu \chi^{\bar{j}} \right]$$

This is the generalization of the CCJ.

The bosonic part: improved currents and a modified Einstein equation with a matter energy-momentum tensor that contains the gravity part $G_{\mu\nu} \Phi_M$ and a $U(1)$ part, such that it is conserved and traceless due to the equations of motions.

A_μ field equation

We observe that

$$-\frac{3}{4}(\mathcal{E}_\mu + J_\mu) = iN_{\bar{I}}\mathcal{D}_\mu\bar{X}^{\bar{I}} - iN_I\mathcal{D}_\mu X^I + \frac{1}{2}iN_{I\bar{J}}\bar{\Omega}^I\gamma_\mu\Omega^{\bar{J}}$$

Since $[\mathcal{W}]_F$ does not involve A_μ (the gauge field of the R -symmetry in the conformal approach, and is the auxiliary field in the super-Poincaré action) the A_μ field equation

$$e^{-1}\frac{\delta}{\delta A^\mu}[N(X, \bar{X})]_D = iN_{\bar{I}}\mathcal{D}_\mu\bar{X}^{\bar{I}} - iN_I\mathcal{D}_\mu X^I + \frac{1}{2}iN_{I\bar{J}}\bar{\Omega}^I\gamma_\mu\Omega^{\bar{J}}$$
$$\mathcal{E}_\mu + J_\mu \approx 0 \quad \bar{\mathcal{D}}^{\dot{\alpha}}(\mathcal{E}_{\alpha\dot{\alpha}} + J_{\alpha\dot{\alpha}}) \approx 0$$

This expression is a superconformal primary, and can be used as first component of a superconformal multiplet.

Old minimum supergravity with a FI term

$$\mathcal{L} = \left[-3S^0 e^{\xi V} \bar{S}^0 \right]_D$$

ξ is the (dimensionless) FI constant. S^0 transforms under an abelian gauge group gauged by a real multiplet V .

$$S^0 \rightarrow S^0 e^{-\xi V}, \quad V \rightarrow V + \Lambda + \bar{\Lambda}$$

$$\bar{\mathcal{D}}^{\dot{\alpha}} \mathcal{E}_{\alpha\dot{\alpha}} = (e^{\xi V} S^0)^3 \mathcal{D}_{\alpha} \left[\frac{e^{-3\xi V} T(e^{\xi V} \bar{S}^0)}{(S^0)^2} \right] + 3\xi S^0 e^{\xi V} \bar{S}^0 W_{\alpha}.$$

Chirally extended supergravity in [B. de Wit and P. van Nieuwenhuizen, Nucl. Phys. B139 \(1978\) 216-220](#).

$\xi \rightarrow 0$ gives the pure old minimal.

$$\mathcal{L} = \left[N(X^I, \bar{X}^I) e^{\xi V} \right]_D$$

$$N(X^I, \bar{X}^I) = S^0 \bar{S}^0 \Phi(S^i, \bar{S}^i) = S^0 \bar{S}^0 (-3 + \Phi_M)$$

The super-Einstein equations are obtained by variation of the auxiliary field A_μ

$$\frac{1}{4} i \gamma_{\alpha\dot{\alpha}}^\mu e^{-1} \frac{\delta \mathcal{L}}{\delta A^\mu} = \mathcal{E}_{\alpha\dot{\alpha}}(S^0, V) + J_{\alpha\dot{\alpha}}(S^0, V, S^i) \approx 0$$

from which

$$\bar{D}^{\dot{\alpha}} \mathcal{E}_{\alpha\dot{\alpha}} + \bar{D}^{\dot{\alpha}} J_{\alpha\dot{\alpha}} \approx 0$$

$$\bar{D}^{\dot{\alpha}} J_{\alpha\dot{\alpha}} \approx \frac{1}{3} (e^{\xi V} S^0)^3 \mathcal{D}_\alpha \left[\frac{e^{-3\xi V} T \left(e^{\xi V} N_0^M \right)}{(S^0)^2} \right] - \xi W_\alpha N^M e^{\xi V}$$

which is the generalization of $R \approx -\kappa^2 T_\mu^\mu$ of general relativity.

$$\mathcal{L}^{nm} = \left[3L \ln \frac{L}{S^0 \bar{S}^0} \right]_D$$

$$\bar{D}^{\dot{\alpha}} \mathcal{E}_{\alpha \dot{\alpha}}^L = L W_{\alpha}^L$$

Adding Matter

$$\mathcal{L}^{nm} = \left[3L \ln \frac{L}{S^0 \bar{S}^0} \right]_D + [L K]_D$$

K is the matter Kähler potential. The field equation of χ_{α} , the fermionic component of the linear multiplet, which defines $W_{\alpha}^L, W_{\alpha}^K$

$$W_{\alpha}^L + W_{\alpha}^K \approx 0$$

The A_{μ} field equation is split as

$$\mathcal{E}_a^L + J_a \approx 0$$

$$\bar{D}^{\dot{\alpha}} J_{\alpha \dot{\alpha}} \approx L W_{\alpha}^K$$

- *We have generalized FZ multiplet for any $\mathcal{N} = 1$ $d = 4$ curved background using superconformal approach*
- *Generalized Ward identities for Supercurrent and Einstein tensor multiplet from superconformal approach*
 - *Pure Old Minimal*
 - *+ matter*
 - *Pure Old Minimal + FI*
 - *Pure New Minimal*
 - *+ matter*
- *Important for*
 - *Nonlinear realization of supersymmetry, nilpotent fields,...*
 - *Cosmological applications*
 - *Supergravity backgrounds for localization techniques*
 - *...*

Thank you

The auxiliary fields are

old minimal: $(A_a, u = \kappa \bar{F}^0)$,

new minimal: $(A_\mu, a_{\mu\nu}), \quad H^\mu \equiv e^{-1} \varepsilon^{\mu\nu\rho\sigma} \left(\kappa^2 \partial_\nu a_{\rho\sigma} - \frac{1}{4} \bar{\psi}_\nu \gamma_\rho \psi_\sigma \right)$

The actions for pure Poincaré supergravity are

$$\kappa^2 e^{-1} \mathcal{L}^{\text{om}} = \frac{1}{2} R - \frac{1}{2} \bar{\psi}^\mu \gamma^{\mu\nu\rho} D_\nu \psi_\rho - \frac{1}{3} u \bar{u} + 3 A^a A_a$$

$$\begin{aligned} \kappa^2 e^{-1} \mathcal{L}^{\text{nm}} = & \frac{1}{2} R - \frac{1}{2} \bar{\psi}^\mu \gamma^{\mu\nu\rho} (D_\nu - \frac{3}{2} i \gamma_* A_\nu) \psi_\rho + \frac{3}{4} H_a H^a \\ & + \frac{3}{2} \kappa^2 \varepsilon^{\mu\nu\rho\sigma} (2A_\mu - H_\mu) \partial_\nu a_{\rho\sigma} \end{aligned}$$

the Weyl multiplet: $\{e_{\mu}^a, \psi_{\mu}, b_{\mu}, A_{\mu}\}$.

chiral multiplets: $\{X^I, \Omega^I, F^I\}$.

a real vector multiplet: $\{V, \zeta, \mathcal{H}, W_{\alpha}, \lambda, D\}$, which in Wess–Zumino (WZ) gauge is reduced to $\{W_{\mu}, \lambda, D\}$.

a real linear multiplet: $\{L, a_{\mu\nu}, \chi\}$.

$$\begin{aligned} \overline{\mathcal{D}}^{\dot{\alpha}} \mathcal{E}_{\alpha\dot{\alpha}} = & \frac{1}{6} \sqrt{2} \left[X^I \Theta(\Omega)_{I\alpha} + 2\Omega_{\alpha}^I \Theta(F)_I \right] - \\ & i\zeta_{\alpha} \Theta(\mathcal{H}) - 2W_{\alpha} \Theta(D) + W_{\alpha}^{\chi} L \end{aligned}$$