# **D** = 4 Black Holes From Geodesics

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Introduction	D=3 Description as Geodesics	The Seed Geodesic	The Issue of Nilpotent Orbits and an Example	Conclusions

# Outline

# Introduction

- Black Holes in Extended D = 4 Supergravity
- D=3 Description as Geodesics
  - The global symmetry in *D* = 3
- 3 The Seed Geodesic
  - Seed Geodesic in Universal Submanifold
- 4 The Issue of Nilpotent Orbits and an Example

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Seed solution: Simplest solution with all duality invariant properties of the most general one

Black Holes in Extended D = 4 Supergravity

## Static, Asymtotically Flat Black Holes in D=4 SUGRAS

#### **Bosonic field content**

- $n_S$  scalar fields  $\phi^r$   $(r = 1, ..., n_S)$
- $n_V$  vector fields  $A^{\Lambda}_{\mu}$   $(\Lambda = 0, \dots, n_V 1)$
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#### The ansatz

$$ds^2 = -e^{2U} dt^2 + e^{-2U} \left[ rac{c^4}{\sinh^4(c\, au)} d au^2 + rac{c^2}{\sinh^2(c\, au)} \left( d heta^2 + \sin( heta) darphi^2 
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• 
$$\phi^r = \phi^r(\tau), \ U = U(\tau), \quad \frac{d\tau}{dr} = \frac{\sinh^2(c\,\tau)}{c^2} = \frac{1}{(r-r_0)^2-c^2};$$

- *c* extremality parameter, two horizons:  $r_{\pm} = r_0 \pm c$
- electric and magnetic charges  $e_{\Lambda}$ ,  $m^{\Lambda}$ :  $\Gamma^{M} \equiv (m^{\Lambda}, e_{\Lambda})$
- Extreme solutions c = 0:  $\lim_{\tau \to -\infty} e^{-2U} = \frac{A_H}{4\pi} \tau^2$

Black Holes in Extended D = 4 Supergravity

#### Seed solution in maximal SUGRA

- 70 scalar fields  $\phi^r \in \mathcal{M}_{scal} = \frac{G_4}{H_4} = \frac{E_{7(7)}}{SU(8)}$
- 28 vector fields A<sup>A</sup><sub>µ</sub>
- Duality group is  $E_{7(7)}$ ;  $\Gamma^{M} = (m^{\Lambda}, e_{\Lambda}) \in 56$  symplectic representation
- Parameters of a black-hole encoded in central charges computed at infinity:
   Z<sub>AB</sub>(φ<sub>∞</sub>, Γ) ∈ 28 of SU(8)

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$$Z_{AB} \stackrel{\text{SU(8)}}{\longrightarrow} \begin{pmatrix} Z_1 \ \epsilon & 0 & 0 & 0 \\ 0 & Z_2 \ \epsilon & 0 & 0 \\ 0 & 0 & Z_3 \ \epsilon & 0 \\ 0 & 0 & 0 & Z_4 \ \epsilon \end{pmatrix}$$

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- (Z<sub>k</sub>) can be identified with the charges (Z, Z̄<sub>s</sub>, Z̄<sub>t</sub>, Z̄<sub>u</sub>) of a
   N = 2 STU truncation
- Five SU(8) invariants:  $\rho_k = |Z_k|, \ \theta = \operatorname{Arg}(Z_1 Z_2 Z_3 Z_4)$
- Seed solution, also solution to the STU truncation, has 5 parameters

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- Invariant measure along the geodesic coincides with the extremality parameter:  $G_{IJ}(\phi) \dot{\phi}^{I} \dot{\phi}^{J} = 2 c^{2}$

## Mathematical description of the geodesic

# Definitions...

• Let  $\mathfrak{g}$ ,  $\mathfrak{H}$  be the Lie algebras of G and H. Involution  $\sigma(\mathfrak{H}) = -\eta \mathfrak{H}^T \eta = \mathfrak{H}$  induces

the (pseudo-) Cartan decomposition:

$$\mathfrak{g}=\mathfrak{H}\oplus\mathfrak{K}$$

with  $\sigma(\mathfrak{K}) = -\mathfrak{K}$ 

• Given coset representative  $\mathcal{V}(\phi^{l}) \in e^{Solv}$  and geodesic  $\phi^{l}(\tau)$ , define  $\mathcal{V}(\tau) \equiv \mathcal{V}(\phi^{l}(\tau))$ :

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- *Q* is the Noether charge matrix:  $Q = 2 \mathcal{V}^{-T} V^T \mathcal{V}^T$



• Geodesic uniquely defined by initial point  $\phi_0^I = \phi^I(\tau = 0)$ and initial velocity  $V_0 = V(\tau = 0) \in \Re$ 

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- Action of G on a geodesic (φ<sub>0</sub>, V<sub>0</sub>)
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### Some algebra....

• G is larger than  $G_4$ :  $G_4 \times SL(2, \mathbb{R})_E \subset G$ . Its algebra decomposes as follows:

$$\mathfrak{g}=\mathfrak{sl}(2,\mathbb{R})_{\textit{E}}\oplus\mathfrak{g}_{4}\oplus(2,\textbf{R})$$

where  ${\bf R}$  is the (symplectic) representation of the electric and magnetic charges under  ${\it G}_4$ 

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- $\mathcal{N} = 8$  example:

 $G_4 = E_{7(7)}$ ,  $H_4 = SU(8)$ ,  $G = E_{8(8)}$ ,  $H = SO^*(16)$ 

 $\bm{R}=\bm{56} \text{ of } \mathrm{E}_{7(7)} \text{ and } \hat{\bm{R}}=\bm{28}_{-1}+\overline{\bm{28}}_{+1} \text{ of } \mathrm{U}(1)_{\textit{E}}\times\textit{H}_{4}=\mathrm{U}(8)$ 

Seed Geodesic in Universal Submanifold

Any element of 
 <sup>(R)</sup>
 <sup>(R)</sup>
 or 
 <sup>(R)</sup>
 can be rotated by U(1)<sub>E</sub> × H<sub>4</sub> into minimal *abelian* subalgebras 
 <sup>(N)</sup>
 = Span(k<sub>k</sub>) and 
 <sup>(N)</sup>
 = Span(J<sub>k</sub>), where k = 0,... p - 1 and

$$p = \operatorname{rank}\left(\frac{H}{\mathrm{U}(1)_E \times H_4}\right)$$

Together with  $H_k = [J_k, k_k]$  they generate  $SL(2, \mathbb{R})^p \subset G$
Seed Geodesic in Universal Submanifold

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True for all V<sub>0</sub>-diagonalizable cases. V<sub>0</sub> non-diagonalizable (e.g. extremal solutions) only geodesics originating from regular solutions (A<sub>H</sub> > 0).

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 Regular extremal solutions, c<sup>2</sup> ∝ tr(V<sub>0</sub><sup>2</sup>) = 0, are characterized by a nilpotent Lax matrix V<sub>0</sub><sup>k</sup> = 0, k ≤ k<sub>0</sub>. Unfolds in the (dS<sub>2</sub>)<sup>p</sup> factor

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- Construction of the seed geodesic within a universal submanifold common to a broad class of models. E.g. *p* = 4: *N* = 8, *N* = 2 with rank-3 symmetric SK manifolds (*STU*) etc.

## The Issue of Nilpotent Orbits

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- D = 4, N = 2 SUGRA coupled to 1 vector multiplet
- Complex scalar t in  $\frac{SL(2,\mathbb{R})}{SO(2)} [\mathcal{F}(t) = t^3]$  coupled to 2 vectors; 4 charges  $\Gamma = (m^0, m^1, e_0, e_1)$  [(D6, D4, D0, D2) brane-charges]

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• Explicit matrix representation:  $J = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}$ ,  $k = \begin{pmatrix} 0 & 1 \\ -1 & 0 \end{pmatrix}$ ,  $\mathcal{H} = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}$ 

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• Solution in terms of  $U, t, Z^M$ :

$$e^{-2U} = \sqrt{H_0(H_1)^3}; \ t = -i\sqrt{\frac{H_0}{H_1}} \ , \ \ \mathcal{Z}^0 = \frac{\varepsilon_0 a_0 \, \tau}{H_0}; \ \mathcal{Z}_1 = \sqrt{3} \, \frac{\varepsilon_1 a_1 \, \tau}{H_1}$$

 $\mathbf{H}_k = 1 - \sqrt{2} a_k \tau$ . Charges are:  $e_0 = \varepsilon_0 a_0, m^1 = -\varepsilon_1 a_1$ 

• Regular solution:  $a_k > 0 \Rightarrow \beta$ -label =  $\gamma$ -label.

• At the horizon  $\tau \to -\infty$ :

$$arphi_i o arphi_i^{ ext{fix}}$$
 (stable attractor) ,  $e^{-2U} o rac{A_H}{4\pi} au^2$ 

where

$$\frac{A_H}{4\pi} = \sqrt{4 a_0 (a_1)^3} = \sqrt{4 \varepsilon e_0 (m^1)^3} = \sqrt{\varepsilon I_4(e,m)}$$

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•  $|a_k| = 1$  modulo action of SO(1, 1)<sup>2</sup>  $\subset$  H

$$\begin{array}{c|c} G_{\mathbb{C}^-} \\ \text{orbit} \end{array} & \beta \text{-labels} \\ \hline \\ g_{\mathbb{C}} \\ g$$

 Introduction
 D=3 Description as Geodesics
 The Seed Geodesic
 The Issue of Nilpotent Orbits and an Example
 Conclusions

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### **Other Orbits and Tensor Classifiers**

- Other two orbits with  $I_4 = 0$ , X step-2 nilpotent:
  - **a**  $\mathcal{O}_1$ :  $m^1 \rightarrow 0$  (doubly-critical) small-bh
  - **b**  $\mathcal{O}_2$ :  $e_0 \rightarrow 0$  (lightlike) small-bh
- Last orbit: X step-7 nilpotent. No regular single-center, regular multicenter solution [Bossard, Ruef, 1106.5806]
- Describe orbits through H-invariants from Lax matrix. Define symmetric H-covariant tensors whose signature is H-invariant [Fre', Sorin, M.T., 1103.0848]

#### **Tensor Classifiers**

$$V_0 = \Delta^{lpha A} K_{lpha A}, \qquad \Delta^{lpha A} \in (\mathbf{2}, \mathbf{4}) ext{ of } H = \mathrm{SL}(2, \mathbb{R})_1 imes \mathrm{SL}(2, \mathbb{R})_2$$

$$\begin{split} \mathcal{T}^{xy} &= \epsilon_{\alpha\beta} \Delta^{\alpha A} \Delta^{\beta B} \Pi^{xy}_{AB} \in (\mathbf{1},\mathbf{3}) \times_{s} (\mathbf{1},\mathbf{3}) \\ \mathfrak{T}^{xy} &= (s^{a})_{\alpha\beta} (s_{a})_{\gamma\delta} (t^{x})_{AB} (t^{y})_{CD} \Delta^{\alpha A} \Delta^{\beta B} \Delta^{\gamma C} \Delta^{\delta D} \in (\mathbf{1},\mathbf{3}) \times_{s} (\mathbf{1},\mathbf{3}) \\ \mathbb{T}^{ab} &= (s^{a})_{\alpha\beta} (s^{b})_{\gamma\delta} (t^{x})_{AB} (t_{x})_{CD} \Delta^{\alpha A} \Delta^{\beta B} \Delta^{\gamma C} \Delta^{\delta D} \in (\mathbf{3},\mathbf{1}) \times_{s} (\mathbf{3},\mathbf{1}) \end{split}$$

where  $\mathfrak{sl}(2,\mathbb{R})_1 = \operatorname{Span}\{s_a\}, \, \mathfrak{sl}(2,\mathbb{R})_2 = \operatorname{Span}\{t_x\}. \text{ BPS solution} \Leftrightarrow \mathcal{T}^{xy} \equiv 0$ 

## Conclusions

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- Apply analysis to multicenter and rotating solutions, characterizing their seed solutions in *D* = 3 within universal truncations

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- $\mathscr{M}_{scal}$  is globally isometric to a solvable group:  $\mathscr{M}_{scal} \sim e^{Solv_4[\phi']}$
- $\mathscr{M}_{scal}^{(3)}$ , being pseudo-Riemannian, is only *locally* isometric to a solvable group:  $\mathscr{M}_{scal}^{(3)} \sim e^{Solv[\phi']}$
- Physical fields φ<sup>l</sup> are *local* coordinates (physical patch) for *M*<sup>(3)</sup><sub>scal</sub>, while φ<sup>r</sup> are global coordinates on *M*<sub>scal</sub>

# Example $dS_2 \equiv \frac{SL(2,\mathbb{R})}{SO(1,1)}$ • $-(X^0)^2 + (X^1)^2 + (X^2)^2 = 2$ • Solvable coords. $e^{-\phi} = X^0 + X^1 > 0, e^{-\phi}\chi = \sqrt{2}X^2$ • Metric: $ds^2 = -2d\phi^2 + \frac{1}{2}e^{-2\phi}d\chi^2$

#### Parametrization of the scalar manifold

- $\mathscr{M}_{scal}$  is globally isometric to a solvable group:  $\mathscr{M}_{scal} \sim e^{Solv_4[\phi']}$
- $\mathscr{M}_{scal}^{(3)}$ , being pseudo-Riemannian, is only *locally* isometric to a solvable group:  $\mathscr{M}_{scal}^{(3)} \sim e^{Solv[\phi']}$
- Physical fields φ<sup>l</sup> are *local* coordinates (physical patch) for *M*<sup>(3)</sup><sub>scal</sub>, while φ<sup>l</sup> are global coordinates on *M*<sub>scal</sub>

