

# Matrix models, noncommutative gauge theory, and gravity

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Dubna, July 2011

# Motivation

- “**fundamental**” d.o.f. for quantum gravity might be different from **macroscopic** ones
- “**emergence**”, quantum structure of space-time
- cosm. const. problem: need new mechanism ?

candidate:

Matrix Models

- related to string theory (“nonpert. def.”)
- describes dynamical NC space-time (branes)
- accessible, novel tools

NC gauge theory  $\leftrightarrow$  gravity

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# Matrix Models

IKKT (IIB) model

Ishibashi, Kawai, Kitazawa and Tsuchiya 1996

$$S[X] = -\text{Tr} \left( [X^a, X^b][X^{a'}, X^{b'}] g_{aa'} g_{bb'} + \bar{\Psi} \gamma_a [X^a, \Psi] \right)$$

$$X^a = X^{a\dagger} \in \text{Mat}(N, \mathbb{C}), \quad a = 0, \dots, 9$$

$$N \rightarrow \infty$$

gauge symmetry  $X^a \rightarrow UX^aU^{-1}$ ,  $SO(9, 1)$ , SUSY, etc.

as  $\left\{ \begin{array}{l} 1) \text{ nonpert. def. of IIB string theory (on } \mathbb{R}^{10}) \\ 2) \mathcal{N} = 4 \text{ SUSY Yang-Mills gauge thy. on } \mathbb{R}_\theta^4 \end{array} \right. \quad (\text{IKKT})$

dynamical NC branes  $\mathcal{M} \subset \mathbb{R}^{10}$

carry (NC) gauge theory **coupled to gravity** (H.S. 2007 ff)

**(brane-world scenarios)**

# Space-time & gravity from matrix models:

e.o.m.:  $\delta S = 0 \Rightarrow [X^a, [X^{a'}, X^{b'}]]\eta_{aa'} = 0$

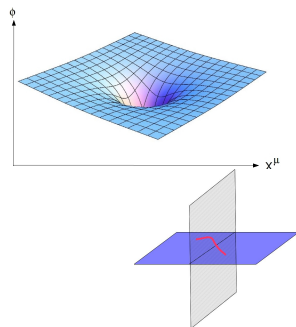
solutions:

- $[X^a, X^b] = i\theta^{ab} \mathbf{1}$ ,
  - $[X^a, X^b] \sim i\theta^{ab}(x)$ ,
- **space-time** as  
3+1-dim. **brane solution**

$$X^a \sim x^a : \mathcal{M}^4 \hookrightarrow \mathbb{R}^{10}$$

- intersecting branes, stacks  
(as in string theory)
- compact extra dim  $\mathcal{M}^4 \times T^2$ , etc.

“quantum plane”  $\mathbb{R}_\theta^4$   
generic quantum space



(H.S., arXiv:1106.6153)

## Moyal-Weyl quantum plane $\mathbb{R}_\theta^4$

$$[\bar{X}^\mu, \bar{X}^\nu] = i\bar{\theta}^{\mu\nu} \mathbf{1}, \quad \mu, \nu = 1, \dots, 4$$

... Heisenberg algebra  $\mathcal{A} = \text{Mat}(\infty, \mathbb{C}) =$  **space of functions on  $\mathbb{R}_\theta^4$**

$$\Delta \bar{X}^\mu \Delta \bar{X}^\nu \geq |\bar{\theta}^{\mu\nu}|$$

relation with classical  $\mathbb{R}^4$ :

$$\phi(x) = \int d^4 k \tilde{\phi}(k) e^{ik_\mu x^\mu} \leftrightarrow \int d^4 k \tilde{\phi}(k) e^{ik_\mu \bar{X}^\mu} =: \Phi(\bar{X}) \in \text{Mat}(\infty, \mathbb{C})$$

$\Phi(\bar{X}^\mu) \in \text{Mat}(\infty, \mathbb{C})$  ... general function on  $\mathbb{R}_\theta^4$   
 $\bar{X}^\mu \in \text{Mat}(\infty, \mathbb{C})$  ... quantized coordinate functions on  $\mathbb{R}_\theta^4$

$[\bar{X}^\mu, \Phi(\bar{X})] =: i\bar{\theta}^{\mu\nu} \partial_\nu \Phi(\bar{X}) \sim i\bar{\theta}^{\mu\nu} \partial_\nu \phi(x) \rightarrow$  NC field theory

## Quantized symplectic manifolds

$(\mathcal{M}, \theta^{\mu\nu}(x))$  ...  $2n$ -dim. manifold with Poisson structure

Its **quantization** is NC algebra  $\mathcal{A}$  such that

$$\begin{aligned} \mathcal{I}: \mathcal{C}(\mathcal{M}) &\rightarrow \mathcal{A} \subset \mathcal{L}(\mathcal{H}) \\ f(x) &\mapsto \hat{f}(X) \end{aligned}$$

such that

$$\begin{aligned} \hat{f}\hat{g} &= \mathcal{I}(fg) + \mathcal{O}(\theta) \\ [\hat{f}, \hat{g}] &= \mathcal{I}(i\{f, g\}) + \mathcal{O}(\theta^2) \end{aligned}$$

(“nice“)  $\Phi \in \text{Mat}(\infty, \mathbb{C}) \leftrightarrow$  quantized function on  $\mathcal{M}$

examples:

fuzzy sphere  
fuzzy torus  
etc.

$$\left. \begin{array}{l} \mathbb{S}_N^2 \\ T_N^{2n} \end{array} \right\}$$

... built-in UV cutoff,  $\dim \mathcal{H} \sim \text{Vol}(\mathcal{M})$

small deformation of  $\mathbb{R}^4$ :

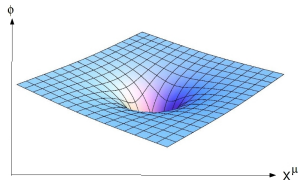
$$X^a = \bar{X}^a + A^a = \begin{pmatrix} \bar{X}^\mu + A^\mu(\bar{X}^\mu) \\ \phi^i(\bar{X}^\mu) \end{pmatrix}$$

- “old” interpretation:  $U(1)$  NC gauge theory on  $\mathbb{R}^4$
- more appropriate interpretation:

$X^a \sim x^a$ :  $\mathcal{M} \hookrightarrow \mathbb{R}^{10}$       quantized embedding map

$[X^\mu, X^\nu] \sim i\{x^\mu, x^\nu\}$       ...Poisson bracket

... defines embedded NC space  $\mathcal{M}$





consider scalar field  $\varphi$  (“test particle”) on generic NC brane

$$X^a \sim x^a: \quad \mathcal{M} \hookrightarrow \mathbb{R}^{10}$$

$\varphi \in \mathcal{A}$ , action

$$S[\varphi] = \text{Tr} [X^a, \varphi][X^b, \varphi] \eta_{ab} \quad (\text{gauge inv.!!})$$

$$\sim \int d^4x \sqrt{|\theta_{\mu\nu}^{-1}|} \theta^{\mu'\mu} \partial_{\mu'} x^a \partial_\mu \varphi \theta^{\nu'\nu} \partial_{\nu'} x^b \partial_\nu \varphi \eta_{ab}$$

$$\sim \int d^4x \sqrt{|G_{\mu\nu}|} G^{\mu\nu}(x) \partial_\mu \varphi \partial_\nu \varphi$$

(use  $[f, \varphi] \sim i\theta^{\mu\nu}(x) \partial_\mu f \partial_\nu \varphi$ )

where

$$G^{\mu\nu}(x) = e^{-\sigma} \theta^{\mu\mu'}(x) \theta^{\nu\nu'}(x) g_{\mu'\nu'}(x) \quad \text{effective metric (cf. open string m.)}$$

$$g_{\mu\nu}(x) = \partial_\mu x^a \partial_\nu x^b \eta_{ab} \quad \text{induced metric on } \mathcal{M}_\theta^4 \text{ (cf. closed string m.)}$$

$$e^{-2\sigma} = \frac{|\theta_{\mu\nu}^{-1}|}{|g_{\mu\nu}|}$$

fluctuations of matrices  $X^a, \Psi$  around background  
 → gauge fields, scalar fields, fermions on  $\mathcal{M}^4$

(NOT 10 dim!)

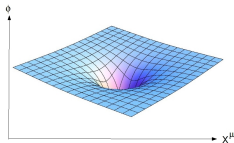
all fields couple to metric  $G^{\mu\nu}(x)$   
 determined by  $\theta^{\mu\nu}(x)$ , embedding (scalar fields)  
 dynamical ⇒ (“emergent”) gravity

can show

$$\square := [X^a, [X^b, \cdot]] g_{ab} \sim e^\sigma \square_G$$

matrix e.o.m  $[X^a, [X^{a'}, X^b]] \eta_{aa'} = 0 \iff$

$$\begin{aligned} \square_G X^a &= 0, & \text{“minimal surface”} \\ \nabla^\mu (e^\sigma \theta_{\mu\nu}^{-1}) &= e^{-\sigma} G_{\rho\nu} \theta^{\rho\mu} \partial_\mu \eta \\ & \eta \sim G^{\mu\nu} g_{\mu\nu} \end{aligned}$$



covariant formulation in semi-classical limit

(H.S. 2007/08)

# Gauge fields

basic observation:

M.M. is **gauge invariant**:

$$X^a \rightarrow U^{-1} X^a U$$

→ **fluctuations**  $X^\mu = \bar{X}^\mu + \theta^{\mu\nu} A_\nu$  transform as  
 $A_\mu \rightarrow U^{-1} A_\mu U + U^{-1} \partial_\mu U$  (nonabelian) gauge field!

note

$$\begin{aligned} [X^\mu, X^\nu] &= i\theta^{\mu\nu} + i\theta^{\mu\mu'} \theta^{\nu\nu'} (\partial_{\mu'} A_{\nu'} - \partial_{\nu'} A_{\mu'} + [A_{\mu'}, A_{\nu'}]) \\ &= i\theta^{\mu\nu} + i\theta^{\mu\mu'} \theta^{\nu\nu'} F_{\mu'\nu'} \end{aligned}$$

(use  $[\bar{X}^\mu, \phi] \sim i\theta^{\mu\nu} \partial_\nu \phi$ )

however:

carefully distinguish  $U(1)$  sector (geometry) from  $SU(n)$  sector (YM)!

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$su(n)$  gauge fields: same model, new vacuum

$$Y^a = \begin{pmatrix} Y^\mu \\ Y^i \end{pmatrix} = \begin{pmatrix} X^\mu \otimes \mathbf{1}_n \\ \phi^i \otimes \mathbf{1}_n \end{pmatrix}$$

( $n$  coinciding branes)

include fluctuations:

$$Y^a = (1 + A^\rho \partial_\rho) \begin{pmatrix} X^\mu \otimes \mathbf{1}_n \\ \phi^i \otimes \mathbf{1}_n + \Phi^i \end{pmatrix}$$

where

$$\begin{aligned} \mathcal{A}^\mu &= -\theta^{\mu\nu} A_{\nu,\alpha} \otimes \lambda^\alpha, & \lambda^\alpha &\in su(n) \\ \Phi^i &= \phi_\alpha^i \otimes \lambda^\alpha \end{aligned}$$

$\Rightarrow$  effective action:

$$S_{YM} = \int d^4x \sqrt{G} e^\sigma G^{\mu\mu'} G^{\nu\nu'} \text{tr} F_{\mu\nu} F_{\mu'\nu'} + 2 \int \eta(x) \text{tr} F \wedge F$$

(H.S., JHEP 0712:049 (2007), JHEP 0902:044,(2009) )

...  $su(n)$  Yang-Mills coupled to metric  $G^{\mu\nu}(x)$

note:

- $SU(n)$  fluctuations ... nonabelian gauge fields
- $U(1)$  fluctuations  $\rightarrow$  **geometry**, metric  $G \sim \theta\theta g$

$$\left. \begin{array}{l} F_{\mu\nu}(x) \\ \phi^i(x) \end{array} \right\} \rightarrow \left\{ \begin{array}{l} \theta^{\mu\nu}(x), \\ g_{\mu\nu}(x) = \eta_{\mu\nu} + \partial_\mu \phi^i \partial_\nu \phi_i \end{array} \right.$$

- generic 4D metric on  $\mathcal{M}^4 \subset \mathbb{R}^{10}$  (almost-Kähler)
- explains & takes advantage of **UV/IR mixing**  
(= new divergences in all NC models **except IKKT**)

# Quantization

Quantization of matrix model:

$$Z = \int dX^a d\Psi e^{-S[X]-S[\Psi]} = e^{-S_{\text{eff}}} \quad \text{“path integral”}$$

2 interpretations:

- ① as NC gauge theory on  $\mathbb{R}_\theta^4$ :  $X^\mu = \bar{X}^\mu + \bar{\theta}^{\mu\nu} A_\nu$   
 → new divergences (UV/IR mixing) in  $U(1)$  sector  
 except  $\mathcal{N} = 4$  SUSY ( $\equiv$  IKKT)
- ② on  $\mathcal{M}^4 \subset \mathbb{R}^{10}$ :  $U(1)$  absorbed in  $\theta^{\mu\nu}(x)$ ,  $g_{\mu\nu}$   
 → induced **gravity** action (Sakharov 1967)

$$S_{\text{eff}} \sim \int d^4x \sqrt{|G|} (\Lambda^4 + c\Lambda_4^2 R[G] + \dots)$$

- 
- explanation UV/IR mixing,  $U(1)$  entanglement (=gravity)
  - IKKT model  $\Rightarrow$  good quantum theory! ( $\mathcal{N} = 4$  SUSY)

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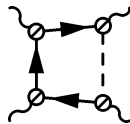
● **explanation** UV/IR mixing,  $U(1)$  entanglement (=gravity)

● IKKT model  $\Rightarrow$  **good quantum theory** ! ( $\mathcal{N} = 4$  SUSY)



example: 1-loop eff. action due to fermions:  
 (all terms of dim  $\leq 6$ ):

$$\begin{aligned}
 \Gamma_{\text{eff}} = & \frac{\Lambda^4}{\Lambda_{\text{NC}}^4} \int \frac{d^4x}{(2\pi)^2} \left( g^{\alpha\beta} D_\alpha \varphi^i D_\beta \varphi_i \right. \\
 & - \frac{1}{2} \Lambda_{\text{NC}}^4 (\bar{\theta}^{\mu\nu} F_{\nu\mu} \bar{\theta}^{\rho\sigma} F_{\sigma\rho} + (\bar{\theta}^{\sigma\sigma'} F_{\sigma\sigma'})(F\bar{\theta}F\bar{\theta})) \\
 & \left. - 2\bar{\theta}^{\nu\mu} F_{\mu\alpha} g^{\alpha\beta} \partial_\nu \varphi^i \partial_\beta \varphi_i + \frac{1}{2} (\bar{\theta}^{\mu\nu} F_{\mu\nu}) g^{\alpha\beta} \partial_\beta \varphi^i \partial_\alpha \varphi_i + \text{h.o.} \right) \\
 & + \frac{\Lambda^2}{\Lambda_{\text{NC}}^4} \int \frac{d^4x}{(2\pi)^2} \left( - \frac{11}{2} F_{\rho\eta} \square_g F_{\sigma\tau} \bar{G}^{\rho\sigma} \bar{G}^{\eta\tau} - 12 \square_g \varphi^i \square \varphi_i \right. \\
 & \left. + \frac{1}{2} \Lambda_{\text{NC}}^4 (\bar{\theta}^{\mu\nu} F_{\mu\nu}) \bar{\square}_G (\bar{\theta}^{\rho\sigma} F_{\rho\sigma}) + \dots \right) \\
 & + \frac{\Lambda^6}{\Lambda_{\text{NC}}^8} \int \frac{d^4x}{(2\pi)^2} (\dots) + \dots
 \end{aligned}$$



(all of this is due to UV/IR mixing !)

(D. Blaschke, H.S., M. Wohlgenannt JHEP 1103 (2011) )

# effective generalized matrix model:

re-assemble as **matrix model**:  $X^a = \begin{pmatrix} \bar{X}^\mu \\ 0 \end{pmatrix} + \begin{pmatrix} -\bar{\theta}^{\mu\nu} A_\nu \\ \phi^i \end{pmatrix}$

$$\Gamma_L[X] = \text{Tr} \frac{L^4}{\sqrt{\frac{1}{2} H^2 - H^{ab} H_{ab} + \frac{1}{2} \mathcal{L}_{10, \text{curv}}[X] + \dots}} \sim \int d^4x \Lambda^4(x) \sqrt{g(x)}$$

$$\mathcal{L}_{10, \text{curv}}[X] = c_1 [X^c, H^{ab}] [X_c, H_{ab}] + c_2 H^{cd} [X_c, [X^a, X^b]] [X_d, [X_a, X_b]] + \dots$$

$$H^{ab} = [X^a, X^c] [X^b, X_c] + (a \leftrightarrow b), \quad H = H^{ab} \eta_{ab}$$

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$SO(D)$  manifest, broken by background (e.g.  $\mathbb{R}_\theta^4$ )

$\Rightarrow$  highly non-trivial predictions for (NC) gauge theory

expect generalization to nonabelian  $\mathcal{N} = 4$  SYM: full  $SO(9, 1)$  !

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**= powerful tool for (NC) gauge theory and (emergent) gravity**

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# (1-loop) finiteness of IKKT model

background field method  $X^a \rightarrow X^a + Y^a$ :

$$\begin{aligned} \Gamma_{1\text{-loop}} &= \frac{1}{2} \text{Tr} \left( \log(\mathbf{1} + \Sigma_{ab}^{(10)} \square^{-1} [\Theta^{ab}, \cdot]) - \frac{1}{2} \left( \log(\mathbf{1} + \Sigma_{ab}^{(16)} \square^{-1} [\Theta^{ab}, \cdot]) \right) \right) \\ &= O(\text{Tr}(\Sigma_{ab} \square^{-1})^4), \quad \text{due to } \mathcal{N} = 4 \\ \square &= [X^a, [X^a, \cdot]] \\ \Theta^{rs} &= [X^r, X^s], \quad \Sigma_{rs} \dots SO(9, 1) \text{ generator} \end{aligned}$$

fully  $SO(9, 1)$  covariant

( IKKT, Chepelev & Tseytlin )

background  $\mathbb{R}_\theta^4$ :  $\equiv \mathcal{N} = 4$  SYM on  $\mathbb{R}_\theta^4$ , no UV div.

$SO(9, 1)$  invariant formalism, broken **spontaneously** through  $\mathbb{R}_\theta^4$   
NC essential.

background  $\mathcal{M}^4 \times K$ :  $\mathcal{N} = 4$  broken

induced gravity on  $\mathcal{M}$  below  $\Lambda_K$

(work in progress, [D. Blaschke, H.S.](#))

# higher-order terms $\mathcal{L}_{\text{curv}}[X]$ : curvature

$$H^{ab} := [X^a, X^c][X^b, X_c]$$

$$T^{ab} := H^{ab} - \frac{1}{4}\eta^{ab}H, \quad H := H^{ab}\eta_{ab},$$

$$\square X := [X^b, [X_b, X]]$$

result:

for 4-dim.  $\mathcal{M} \subset \mathbb{R}^D$  with  $g_{\mu\nu} = G_{\mu\nu}$ :

$$\text{Tr}(2T^{ab}\square X_a\square X_b - T^{ab}\square H_{ab}) \sim \frac{2}{(2\pi)^2} \int d^4x \sqrt{g} e^{2\sigma} R$$

$$\text{Tr}([X^a, X^c], [X_c, X^b])[X_a, X_b] - 2\square X^a\square X_a$$

$$\sim \frac{1}{(2\pi)^2} \int d^4x \sqrt{g} e^\sigma \left( \frac{1}{2} e^{-\sigma} \theta^{\mu\eta} \theta^{\rho\alpha} R_{\mu\eta\rho\alpha} - 2R + \partial^\mu \sigma \partial_\mu \sigma \right)$$

Blaschke, H.S. CQG 27 (2010)

(cf. Arnlind, Hoppe, Huisken 2010, 2011)

cf. Einstein-Hilbert, **pre-geometric**

## issues for (emergent) gravity

- complicated dynamics, not well understood
- bare M.M. action:

$$\text{NC } U(1) \text{ gauge fields } \partial^\mu F_{\mu\nu} = 0 \quad \Rightarrow \quad R_{\mu\nu}[\bar{G} + h] = 0$$

(Rivelles 2002)

compactif. **extra dim**  $\Rightarrow$  additional Ricci-flat d.o.f.

- 1-loop

$\Rightarrow$  induced E-H action,

also modification of bulk metric:

e.g.  $AdS^5 \times S^5$  near stack of branes, cf. DBI,  $\mathcal{N} = 4$  SYM

(Tseytlin, Maldacena ...)

- Minkowski signature:

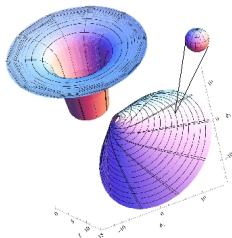
either complexified  $\theta^{\mu\nu}$

or  $G, g$  have different causality structures

## illustrative examples:

- 1) “harmonic branch”  
 near-realistic cosmolog. solutions (big bounce)  
 (NC) minimal surfaces  $\mathcal{M} \subset \mathbb{R}^D$       D. Klammer, H.S. PRL 102 (2009)  
 insensitive to vacuum energy, SN Ia without fine-tuning !

- 2) “Einstein branch”  
 Example: Schwarzschild geometry  
 Blaschke, H.S. CQG 27 (2010)  
 embedding  $\mathcal{M}^4 \subset \mathbb{R}^{10}$ ,  $\theta^{\mu\nu}(x)$

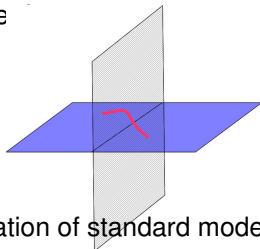


# Relation with particle physics

intersecting brane solutions

**chiral fermions** at intersection = 4D space  
(as in string theory)

$$\begin{pmatrix} X_{(11)}^a & \psi_{(12)} \\ \psi_{(21)} & X_{(22)}^a \end{pmatrix}$$



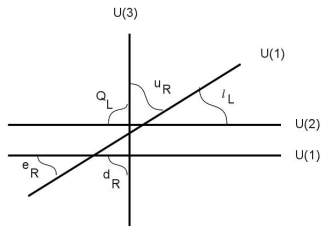
stacks of intersecting branes → realization of standard model

A. Chatzistavrakidis, H.S., G. Zoupanos arXiv:1107.0265

cf. string theory

**clear-cut, predictive framework**

1-loop → intersecting branes can form  
bound system!





# realization of standard model

4 intersecting D7 branes  $\rightarrow U(3)_C \times U(2)_L \times U(1) \times U(1)$

intersections  $M^4 \times K_{ab}^2$ , flux on  $K_{ab}^2$ , chiral bifund. fermions

Intersection	Representation	Particle	flux
$D_a \cap D_b$	$(\bar{3}, 2)(-1, 1, 0, 0)$	$Q_L$	$N'_\beta - N_\beta$
$D_a \cap D_c$	$(\bar{3}, 1)(-1, 0, 1, 0)$	$d_R$	$N''_\beta - N_\beta$
$D_a \cap D_d$	$(\bar{3}, 1)(-1, 0, 0, 1)$	$u_R$	$N'_\alpha - N_\alpha$
$D_d \cap D_b$	$(1, 2)(0, 1, 0, -1)$	$l_L$	$N_\gamma - N''_\gamma$
$D_d \cap D_c$	$(1, 1)(0, 0, 1, -1)$	$e_R$	$N'_\gamma - N''_\gamma$

correct chiral particle spectrum, families from fluxes

1-loop effective action

$\rightarrow$  intersecting branes may form bound state (dep. on flux)

# summary and outlook

- matrix models  $Tr\left([X^a, X^b][X_a, X_b] + \bar{\Psi}\not{D}\Psi\right)$   
 dynam. NC spaces (gravity), gauge theory & matter
- NC gauge theory ( $\mathcal{N} = 4$  SYM)  $\leftrightarrow$  “stringy” (IKKT)  
 $\rightarrow$  candidate for quantum theory of fund. interactions
- QFT methods for gravity / “string theory”!  
 (integration over geometries)
- **intersecting branes**  $\rightarrow$  particle physics:  
**brane-world** scenarios
- ... lots of interesting issues!

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$U(1)$  gauge fields as gravitons

(fix embedding)

$$G^{\mu\nu}(x) = \bar{\eta}^{\mu\nu} - h^{\mu\nu} \quad (+O(F^2))$$

$F_{\mu\nu}(x)$  ...  $u(1)$  field strength

find

$$h_{\mu\nu} = \bar{\eta}_{\nu\nu'} \bar{\theta}^{\nu'\rho} F_{\rho\mu} + \bar{\eta}_{\mu\mu'} \bar{\theta}^{\mu'\eta} F_{\eta\nu} - \frac{1}{2} \bar{\eta}_{\mu\nu} (\bar{\theta}^{\rho\eta} F_{\rho\eta})$$

... linearized metric fluctuation

e.o.m:

$$\begin{aligned} [X^\mu, [X^\nu, X^{\mu'}]] \eta_{\mu\mu'} &= 0 \\ \Rightarrow \partial^\mu F_{\mu\nu} &= 0 \\ \Rightarrow R_{\mu\nu}[G] &= 0 \quad (\partial^\mu h_{\mu\nu} = 0 \dots \text{harm. gauge}) \end{aligned}$$

cf. [Rivelles \[hep-th/0212262\]](#)

while  $R_{\mu\nu\rho\eta} \neq 0$

$\Rightarrow$  on-shell d.o.f. of gravitational waves on Minkowski space

i.e.: NC  $U(1)$  on  $\mathbb{R}_\theta^4$  as gravitons

# Geometrical e.o.m.

assume effective action

$$S = \int d^4x \sqrt{|g|} (-2\Lambda^4 + \Lambda_4^2 R) + S_{\text{matter}}$$

e.o.m.

$$\begin{aligned} \delta S &= \int d^4x \sqrt{|g|} \delta g_{\mu\nu} (-\Lambda^4 g^{\mu\nu} + 8\pi T^{\mu\nu} - \Lambda_4^2 \mathcal{G}^{\mu\nu}) \\ &= -2 \int \delta\phi^i \partial_\mu (\sqrt{|g|} (-\Lambda^4 g^{\mu\nu} + 8\pi T^{\mu\nu} - \Lambda_4^2 \mathcal{G}^{\mu\nu})) \partial_\nu \phi^i \end{aligned}$$

since  $g_{\mu\nu} = \eta_{\mu\nu} + \partial_\mu \phi^i \partial_\nu \phi^i$

① “Einstein branch”

$$\Lambda^4 g^{\mu\nu} + \Lambda_4^2 \mathcal{G}^{\mu\nu} = 8\pi T^{\mu\nu}$$

② “harmonic branch”

$$\Lambda^4 \square_g \phi = (8\pi T^{\mu\nu} - \Lambda_4^2 \mathcal{G}^{\mu\nu}) \nabla_\mu \partial_\nu \phi$$

prototype: **flat space**  $\mathbb{R}_0^4 \subset \mathbb{R}^{10}$ , **even for  $\Lambda \gg 0!$**