

CORRELATIONS BETWEEN $\langle p_T \rangle$
AND JET MULTIPLICITIES FROM THE BFKL POMERON

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HIGH-ENERGY HADRONIC COLLISIONS :

→ STRONG CORRELATIONS ARE OBSERVED EXPERIMENTALLY BETWEEN THE $\langle p_T \rangle$ AND THE MULTIPLICITIES OF THE PRODUCED PARTICLES.

→ $\langle p_T \rangle$ GROWS WITH MULTIPLICITY.

SEVERAL EXPLANATIONS → IT IS ASSUMED THAT WITH ONLY ONE HARD COLLISION (ONE PAIR OF NON-INTERACTING STRINGS) THERE ARE NO CORRELATIONS BETWEEN $\langle p_T \rangle$ AND MULTIPLICITY.

THIS ASSUMPTION CAN BE TESTED WITHIN THE BFKL DYNAMICS, WHICH PRESENTS A DETAILED DESCRIPTION OF PARTICLE (JET) PRODUCTION AT HIGH ENERGIES.

● OUR AIM: TO SEE WHETHER THERE EXIST CORRELATIONS BETWEEN $\langle p_T \rangle$ AND THE NUMBER OF PRODUCED JETS IN THE HARD POMERON DESCRIBED BY THE BFKL CHAIN OF INTERACTING REGGEIZED GLUONS.

THIS STUDY IS RELATED TO THAT BY KWIECINSKI, LEWIS & MARTIN (PRD 54(1996)6664) WHO CALCULATED THE EXCLUSIVE PROBABILITIES TO OBSERVE A GIVEN NUMBER OF PRODUCED JETS FROM THE EXCHANGED HARD POMERON.

FORMALISM USED BY KLM

LET $f(y, k)$ BE THE AMPUTATED BFKL AMPLITUDE,
WHERE : $y \longrightarrow$ RAPIDITY.
 $k \longrightarrow$ TWO-DIMENSIONAL TRANSVERSE MOMENTUM
OF THE VIRTUAL (REGGEIZED) GLUON.

$\frac{f(y, k)}{k^2} \longrightarrow$ UNINTEGRATED GLUON DISTRIBUTION.

THE BFKL EQUATION FOR f :

$$f(y, k) = f^{(0)}(y, k) + \bar{\alpha}_s \int_0^y dy_1 \int \frac{d^2 k_1}{\pi q^2} \cdot \left(\frac{k^2}{k_1^2} f(y_1, k_1) - f(y, k) \cdot \theta(k^2 - q^2) \right)$$

HERE : $\bar{\alpha}_s = \frac{3 \cdot \alpha_s}{\pi}$

$q = k - k_1 \longrightarrow$ THE TRANSVERSE MOMENTUM OF
THE EMITTED (REAL) GLUON.

$f^{(0)}(y, k) \longrightarrow$ THE IMPACT FACTOR OF THE
TARGET.

TO SUPPRESS THE PHYSICALLY UNKNOWN INFRARED DOMAIN
AND MAKE THE EQUATION NUMERICALLY TRACTABLE, THE
INTEGRATION OVER k_1 IS CONSTRAINED TO THE INTERVAL:

$$\underline{Q_0 (1 \text{ GeV}/c) < k_1 < Q_2 (100 \text{ GeV}/c)}$$

→ DEFINITION: AN OBSERVABLE SET IS A REAL GLUON WITH $q^2 \geq \mu^2$.

THIS DEFINITION SPLITS THE INTEGRATION OVER MOMENTA INTO TWO PARTS.

→ TWO PARTS INTO THE INTEGRATION KERNEL:

- K_R (RESOLVED ONE) → EMITTED GLUONS WITH $q^2 > \mu^2$.
- K_{UV} (UNRESOLVED ONE) → IT COMBINES EMISSION OF GLUONS WITH $q^2 < \mu^2$ AND THE SUBTRACTION TERM.

EXPLICITLY:

$$(K_R f)(k) = \bar{\alpha}_s \cdot k^2 \int \frac{d^2 k_1}{\pi q^2 k_1^2} \theta(q^2 - \mu^2) \cdot f(k_2)$$

$$(K_{UV} f)(k) = \bar{\alpha}_s \cdot k^2 \int \frac{d^2 k_1}{\pi q^2 k_1^2} \left(\theta(\mu^2 - q^2) \cdot f(k_2) - \frac{k_1^2}{k^2} \theta(k^2 - q^2) \cdot f(k) \right)$$

→ EXCLUSIVE PROBABILITIES TO PRODUCE n JETS ARE OBTAINED BY INTRODUCING n OPERATORS K_R BETWEEN THE GREEN'S EQUATIONS WITH KERNEL K_{UV} .

EXCLUSIVE PROBABILITIES TO PRODUCE m JETS

IF ONE WRITES

$$f(y) = \sum_{n=0} f_n(y)$$

CONTRIBUTION TO THE
FULL GLUON DISTRIBUTION
FROM THE PRODUCTION OF
 m JETS.

THEN ONE GETS A RECURSIVE RELATION

$$\rightarrow f_n(y) = \int_0^y dy_1 k(y-y_1) \cdot f_{n-1}(y_1) \quad (*)$$

WHERE

$$k(y) = e^{y \cdot k_{UV}} \cdot k_R$$

THIS RELATION ONE TO SUCCESSIVELY CALCULATE
THE RELATIVE PROBABILITIES TO PRODUCE $m=0, 1, 2, \dots$ JETS
STARTING FROM THE NON-JET CONTRIBUTION

$$f_0(y) = e^{y \cdot k_{UV}} \cdot f^{(0)}(0) + \int_0^y dy_1 \cdot e^{(y-y_1) \cdot k_{UV}} \cdot \frac{d f^{(0)}(y_1)}{dy_1}$$

$f^{(0)}(y, k)$ IS CHOSEN TO VANISH AT $y=0$.

THE DISTRIBUTIONS $f_m(y, k)$ ARE NOT OBSERVABLES.

- PHYSICAL PROBABILITIES ARE OBTAINED BY CONVOLUTING f_m WITH THE GLUON DISTRIBUTION IN THE PROJECTILE (THE PROJECTILE IMPACT FACTOR).

- PERTURBATIVE QCD \longrightarrow VIRTUAL PHOTON AS A PROJECTILE.

- HADRONIC COLLISIONS \longrightarrow UNPERTURBATIVE IMPACT FACTOR

IN BOTH CASES THE EXCLUSIVE PROBABILITIES TO OBSERVE n JETS ARE GIVEN BY

$$\longrightarrow P_n(y) = \frac{\int (d^2k/k^4) h(k) f_m(y, k)}{\int (d^2k/k^4) h(k) f(y, k)}$$

WHERE $h(k)$ IS THE IMPACT FACTOR OF THE PROJECTILE.

BOTH IMPACT FACTORS, $f^{(0)}(R)$ OF THE TARGET AND $h(k)$ OF THE PROJECTILE, SHOULD VANISH AS $k \rightarrow 0$.

- VIRTUAL PHOTON IMPACT FACTOR BY NIKOLAIEV & ZAKHAROV
Z. PHYS. C 49, 607(1991).

- HADRONIC IMPACT FACTOR:

$$f^{(0)}(k) = k^2 \cdot e^{-k^2/Q_0^2}$$

→ WE ARE INTERESTED IN THE $\langle q \rangle_m$ OF THE OBSERVED SETS, PROVIDED THEIR NUMBER m IS FIXED.

SINCE THE MOMENTUM q OF THE EMITTED GLUON IS HIDDEN INSIDE THE KERNEL K_R , TO FIND AN AVERAGE OF ANY QUANTITY $\phi(q)$ ONE HAS TO CHANGE K_R TO THE KERNEL K_{av} DEFINED BY

$$(K_{av} f)(k) = \bar{\alpha}_s k^2 \int \frac{d^2 k_1}{\pi q^2 k_1^2} \mathcal{D}(q^2 - \mu^2) \cdot \phi(q) \cdot f(k_1)$$

→ WITH m SETS ONE HAS TO:

- SUBSTITUTE ONE OF THE m OPERATORS K_R WHICH GENERATE THE SETS BY K_{av} .
- TAKE A SUM OF ALL SUCH SUBSTITUTIONS, AND DIVIDE BY m .
- INTEGRATE OVER ALL MOMENTA OF THE VIRTUAL GLUON k MULTIPLIED BY THE PROJECTIVE IMPACT FACTOR.
- NORMALIZE THE RESULT TO THE TOTAL PROBABILITY TO HAVE m SETS.

• HOW TO FORMALIZE THIS RECIPE?

→ INTRODUCE A GENERALIZED OPERATOR IN THE VIRTUAL GLUON MOMENTUM SPACE

$$K_{\perp}(y) = e^{y \cdot k_{uv}} [K_R + K_{uv}]$$

LET THE FUNCTION $F(y, k)$ OBEY THE EQUATION

$$F(y) = f_0(y) + \int_0^y dy_1 K_{\perp}(y - y_1) \cdot F(y_1)$$

→ ONE CAN SPLIT THE FUNCTION F INTO A SUM OF CONTRIBUTIONS F_{nm} CORRESPONDING TO THE ACTION OF n OPERATORS K_{\perp} , OUT OF WHICH $m = 0, 1, \dots, n$ ARE OPERATORS K_{uv} :

$$F(y) = \sum_{n=0}^{\infty} \sum_{m=0}^n F_{nm}(y)$$

$$F_{n0} = f_n$$

WE ARE INTERESTED IN THE CONTRIBUTION $F_{n1} \equiv f_n$ WHICH CONTAINS A SINGLE OPERATOR K_{uv} .

→ THE AVERAGE VALUE OF INTEREST IS DETERMINED BY

$$\langle \phi(q) \rangle_n = \frac{1}{n} \frac{\int (dk^2/k^4) h(k) g_n(y, k)}{\int (dk^2/k^4) h(k) f_n(y, k)}$$

AGAIN ONE EASILY SETS UP A RECURSION RELATION

FOR g_m :

$$g_m = \int_0^y dy_1 k(y-y_1) g_{m-1}(y_1) + \int_0^y dy_1 \cdot e^{(y-y_1)k_{ov}} \cdot k_{av} f_{m-1}(y_1) \quad (**)$$

WITH THE INITIAL CONDITION $g_0(y) = 0$.

→ TOGETHER WITH (*), THIS RELATION ALLOWS ONE TO CALCULATE THE FUNCTION g_m FOR $m = 1, 2, \dots$, AND THEN TO FIND THE DESIRED AVERAGES.

THE CONCRETE CHOICE OF $\phi(q)$ IS RESTRICTED BY THE CONDITION OF CONVERGENCE AT LARGE q :

$$\phi(q) < q^2, \text{ AS } q \rightarrow \infty$$

WE MAKE A NATURAL CHOICE $\phi(q) = q$:

$$(K_{av} f)(k) = \bar{\alpha}_3 \cdot k^2 \int \frac{d^2 k_1}{\pi \cdot k_1^2} \theta(q^2 - k_1^2) f(k_1)$$

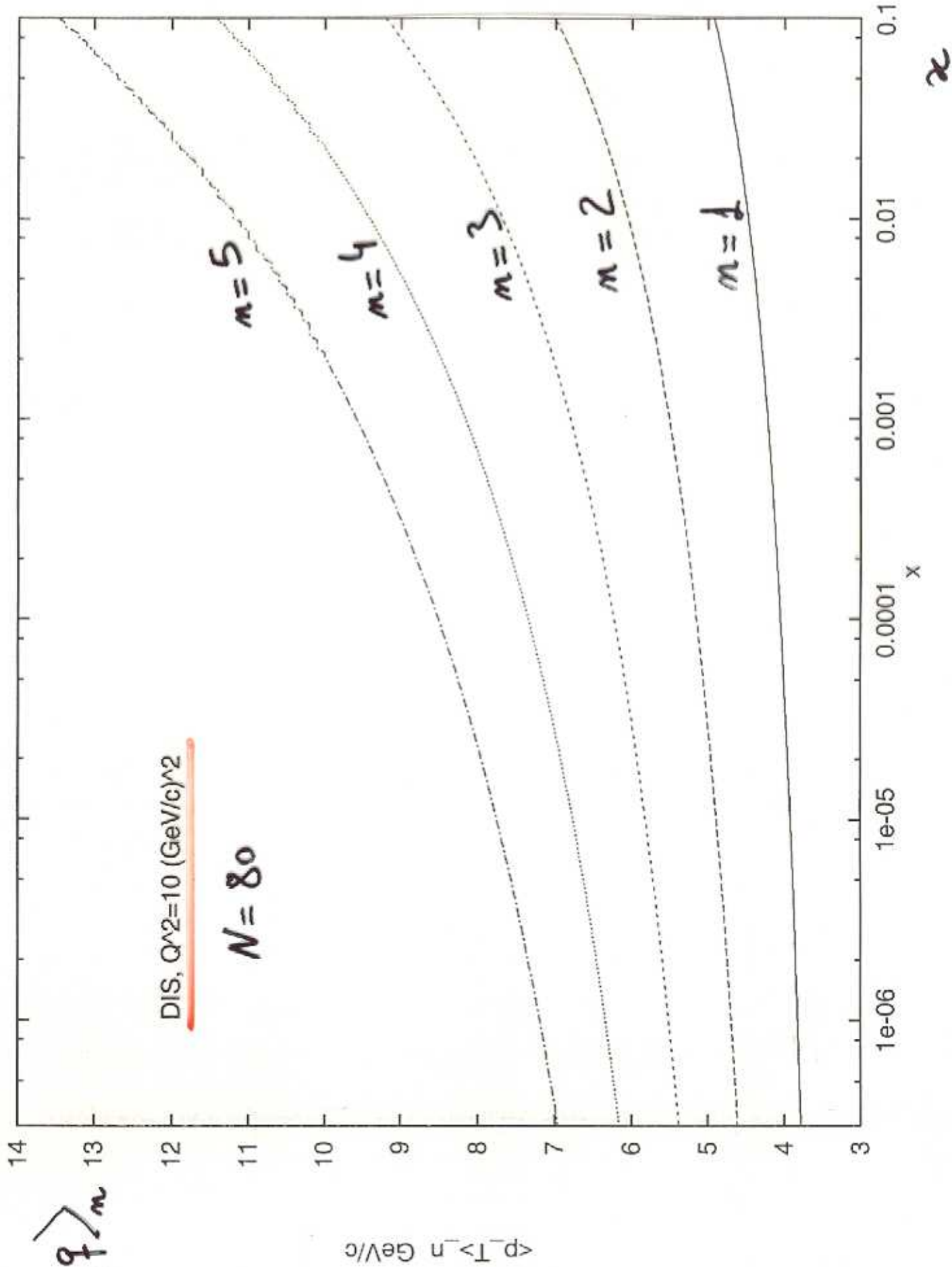
WITH THIS CHOICE THE ANGULAR INTEGRATION CAN BE DONE ANALITICALLY.

WE DEFINED OUR SETS BY TAKING $\mu = 2 \text{ GeV}/c$.

→ WE HAVE CALCULATED f_m AND g_m FROM (*) AND (**) UP TO $m = 5$ AND $n = 25$. TO DISCRETIZE THE KERNELS WE HAVE USED THE EXPANSION IN N CHEBYSHEV POLYNOMIALS

RESULTS

$\langle q \rangle_m$



$\langle f \rangle_{m=14}$

$\langle p, T \rangle_n \text{ GeV/c}$

DIS, $Q^2=100 \text{ (GeV/c)}^2$

$N = 80$

$m=5$

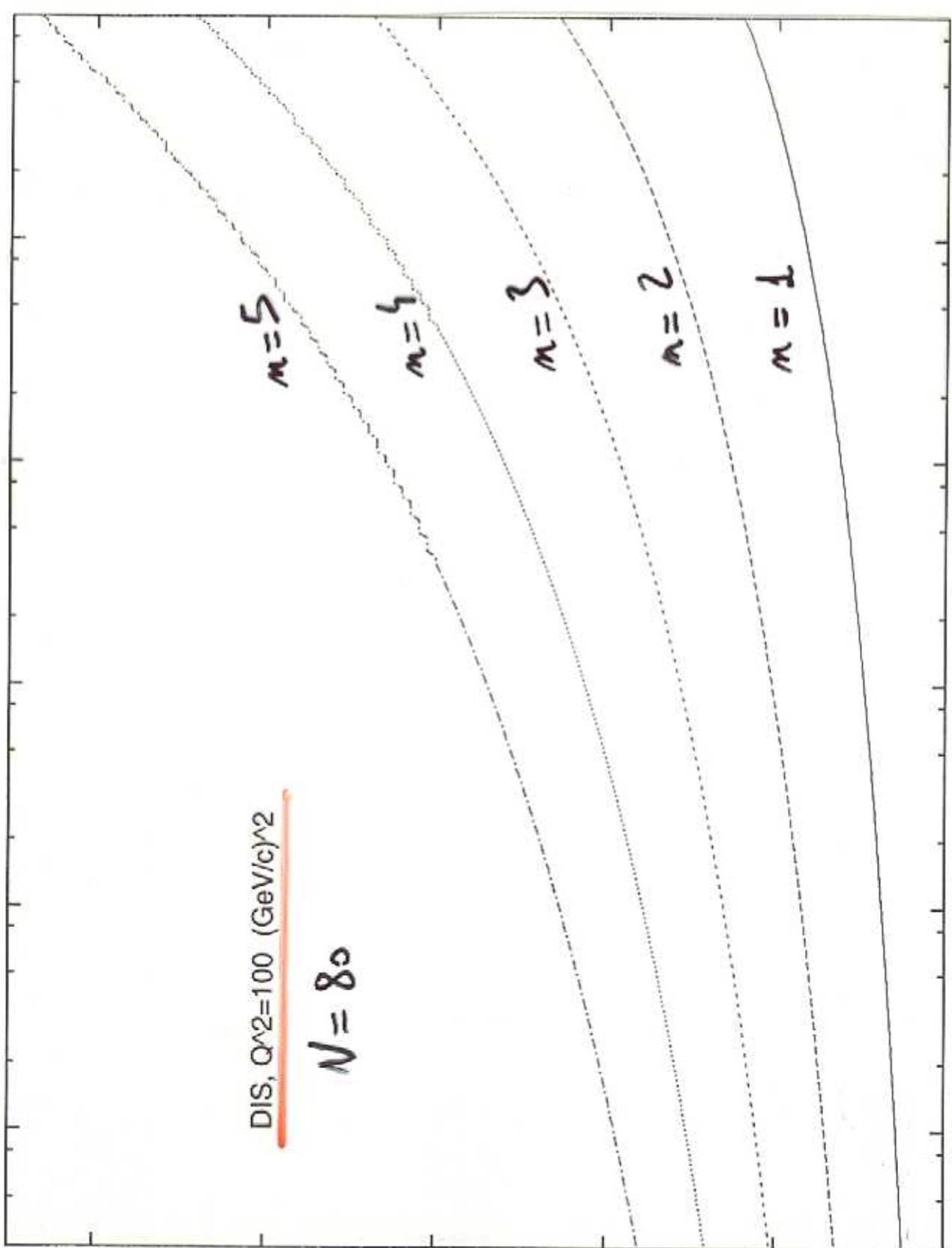
$m=4$

$m=3$

$m=2$

$m=1$

x
0.1
0.01
0.001
0.0001
1e-05
1e-06



$\langle \rho \rangle_m^{16}$

$\langle p_T \rangle_n$ GeV/c

DIS, $Q^2=1000$ (GeV/c)²

$N = 80$

$n=5$

$n=4$

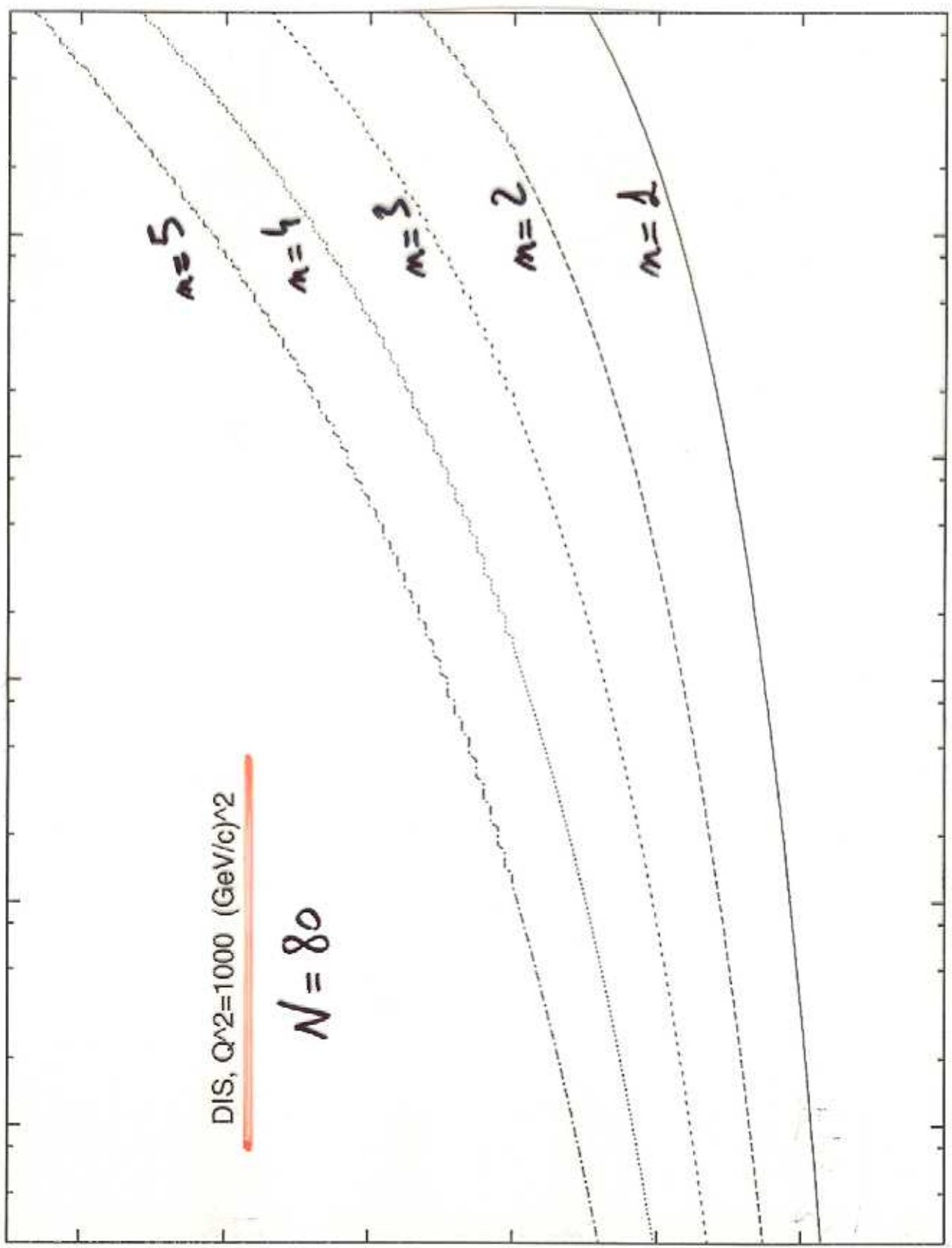
$n=3$

$n=2$

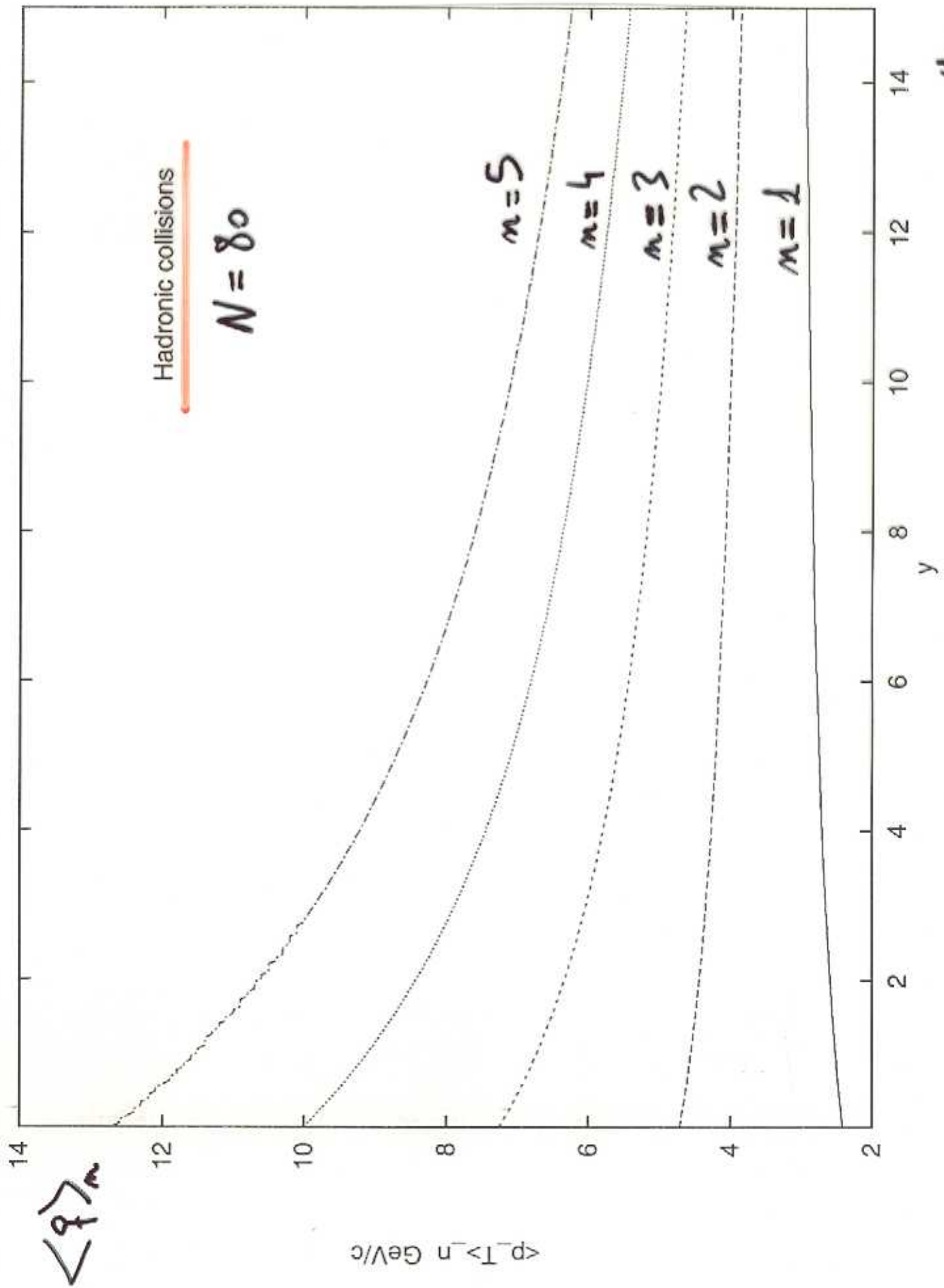
$n=1$

x

1e-06 1e-05 0.0001 0.001 0.01 0.1



\mathcal{P}



→ AS ONE OBSERVES, IN ALL CASES $\langle n \rangle_m$
STRONGLY GROWS WITH m AT ALL RAPIDITIES.
THE GROWTH IS APPROXIMATELY LINEAR

$$\langle p_T \rangle_m \approx a(y, Q^2) + b(y) \cdot m$$

WHERE BOTH a AND b DEPEND ON RAPIDITY y , BUT,
AT FIXED y , b IS UNIVERSAL.

• THE SLOPE $b(y)$ FALLS WITH y : IT IS EQUAL
TO 1.25 GeV/c AT $y = 7.5$ AND 0.8 GeV/c AT
 $y = 15$, SO THAT AT ULTRAHIGH ENERGIES ONE
MAY EXPECT THAT $\langle n \rangle_m$ WILL BECOME INDEPENDENT
OF m .

• THE DEPENDENCE ON Q^2 RESULTS IS RATHER
WEAK. IN THE BFKL MODEL AT LOW x THE BULK
OF THE Q^2 DEPENDENCE IS SEPARATED IN AN OVERALL
FACTOR WHICH CANCELS BETWEEN THE NUMERATOR AND
DENOMINATOR WHEN COMPUTING THE AVERAGES.

WE FIND THAT THE AVERAGES $\langle n \rangle_m$ GO DOWN
WITH RAPIDITY FOR ALL $m \geq 2$. THIS IS QUITE
UNEXPECTED, SINCE IN BFKL AN OVERALL AVERAGE
 $\langle n \rangle$ RAPIDLY GROWS WITH y . IT SEEMS THAT
THIS GROWTH IS TOTALLY EXPLAINED BY THE GROWTH
OF THE AVERAGE NUMBER OF SETS $\langle m \rangle$.

DISCUSSION

● EMISSIONS OF HIGH- p_T JETS IN γ^* -HADRON COLLISIONS SEEM TO BE A SUITABLE PLACE TO SEE THE BFKL SIGNATURE.

● WE OBTAIN IN SUCH EMISSIONS STRONG POSITIVE CORRELATIONS BETWEEN $\langle p_T \rangle$ AND THE NUMBER OF JETS ALREADY FOR A SINGLE POMERON EXCHANGE.

THIS INDICATES THAT THE PREDICTED CORRELATIONS DO NOT REQUIRE MULTIPLE RESCATTERINGS NOR POMERONIC INTERACTIONS, BUT THAT THEY ARE ALREADY PRESENT IN THE BASIC MECHANISM OF JET PRODUCTION.

● IF WE EXTRAPOLATE OUR RESULT TO ALL m AND ENERGIES WE WOULD GET A RELATION BETWEEN THE OVERALL AVERAGES

$$\langle p_T \rangle \propto \langle m \rangle$$

HOWEVER, THIS RELATION DOES NOT HOLD IN THE BFKL MODEL AT ASYMPTOTIC ENERGIES, WHEN $\langle p_T \rangle$ GROWS MUCH FASTER THAN $\langle m \rangle$.

THUS OUR RESULTS CANNOT BE VALID FOR ALL m AND ENERGIES AND REFER PRECISELY TO THE VALUES OF m AND ENERGIES FOR WHICH THE CALCULATIONS WERE DONE.

⑥ THE UNEXPECTED RESULT THAT $\langle P_T \rangle_m$ AT
FIXED $m \geq 2$ FALL WITH ENERGY IS NOT
DRAMATIC BUT QUITE APPRECIABLE AT ENERGIES
AT WHICH WE CAN EXPECT THE BFKL POMERON
TO BE RELEVANT ($y > 10$).

AT PRESENT WE DO NOT HAVE ANY PLAUSIBLE
EXPLANATION OF THIS PHENOMENON, WHICH DESERVES
FURTHER INVESTIGATION INCLUDING HIGHER y AND/OR m .

→ WE HOPE THAT THIS EFFECT CAN BE TESTED
EXPERIMENTALLY AS A POSSIBLE SIGNATURE OF THE
BFKL POMERON.