Geometry of supersymmetric solutions in ten dimensions

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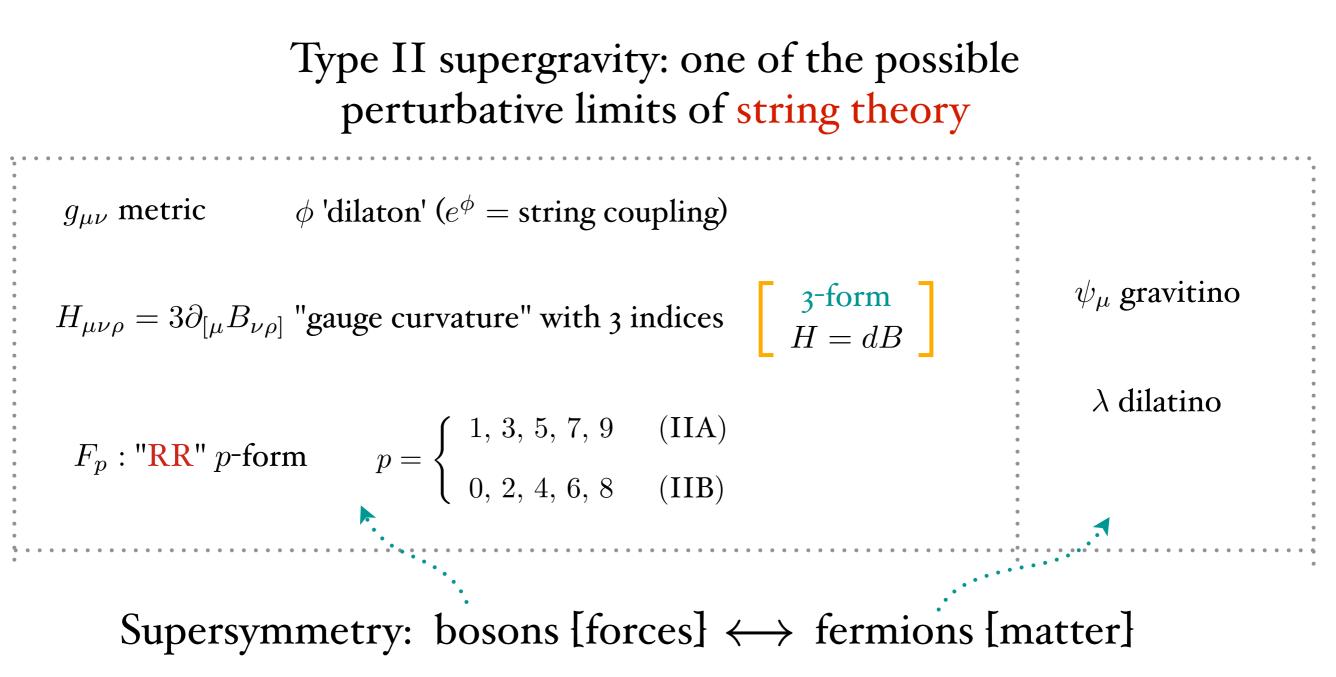








Introduction



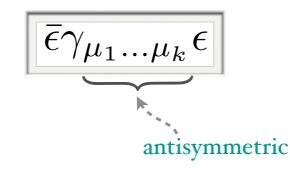
Aim of this talk: characterize supersymmetric solutions

$$\delta g_{\mu\nu} = \bar{\epsilon} \gamma_{(\mu} \psi_{\nu)} \qquad \bullet \text{ Set } \psi = 0; \Rightarrow \delta g_{\mu\nu} = 0$$

$$\delta \psi_{\mu} = \nabla_{\mu} \epsilon + \text{RR-faxes} \qquad \bullet \text{ Set } \text{RR} = 0 \Rightarrow \delta \psi_{\mu} = 0$$

$$\bigotimes$$

Finding solutions is easiest in terms of forms



"restricted holonomy"

 $\nabla_{\mu}\epsilon = 0$

Example:

$$M_{10} = \mathrm{Mink}_4 \times M_6$$

J symplectic form Ω holomorphic volume form

$$\nabla_m \epsilon = 0 \iff \begin{cases} dJ = 0\\ d\Omega = 0\\ J \wedge \Omega = 0\\ J^3 = i\Omega \wedge \bar{\Omega} \end{cases}$$

millions of Calabi-Yau's found this way [Calabi'57, Yau'77]

$$\delta g_{\mu\nu} = \bar{\epsilon} \gamma_{(\mu} \psi_{\nu)} \qquad \bullet \text{ Set } \psi = 0; \Rightarrow \delta g_{\mu\nu} = 0$$

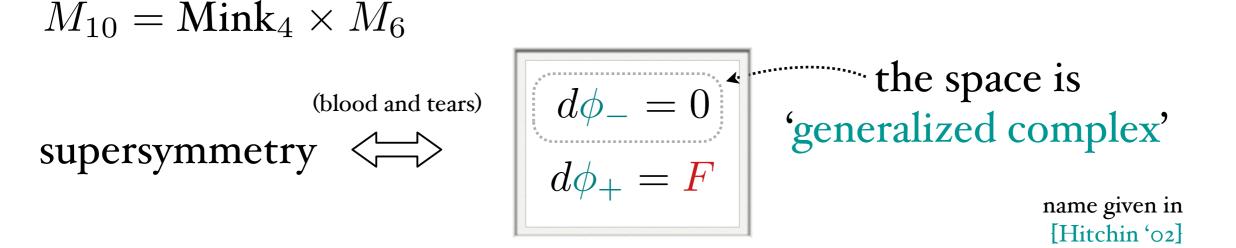
$$\delta \psi_{\mu} = \nabla_{\mu} \epsilon + \text{RR fluxes} \qquad \bullet \text{ If we don't set } \text{RR} = 0:$$

messy system! no holonomy interpretation

 \bigvee

But another interpretation exists!

[Graña, Minasian, Petrini, AT '05]



similar equations also exist for $M_{10} = \text{AdS}_4 \times M_6$ [holographically dual to CFT3]

• what about other spacetimes?

But:

(black hole)₄ × M_6 (Lifschitz)₄ × M_6

[holographically dual to nonrelativistic CFT3]

• what about different dimensions?

...

 $Mink_d \times M_{10-d}$ $AdS_d \times M_{10-d}$

In this talk, I will describe a solution to this problem:

A system in terms of forms equivalent to supersymmetry

I. Review of the system for vacuum solutions

2. The new system: algebraic equations

3. The new system: differential equations

I. Vacuum solutions

Type II theories: two supersymmetry generators $\epsilon_{1,2}$

forms: $\eta_+^{1,\dagger}\gamma_{mn}\eta_+^2$ etc.

[in IIB, from now on]

a more efficient method: $(\eta^{1}_{+} \otimes \eta^{2\dagger}_{-})_{\alpha\beta} = \eta^{1}_{+\alpha} \eta^{2\ast}_{-\beta}$ is a bispinor

'Clifford map':

bispinors forms

$$\gamma^{m_1...m_k} \longleftrightarrow dx^{m_1} \land \ldots \land dx^{m_k}$$

So
$$\phi_{\pm} \equiv \eta^1_+ \otimes \eta^2_{\pm}^{\dagger}$$
 are forms

Algebraic constraints.

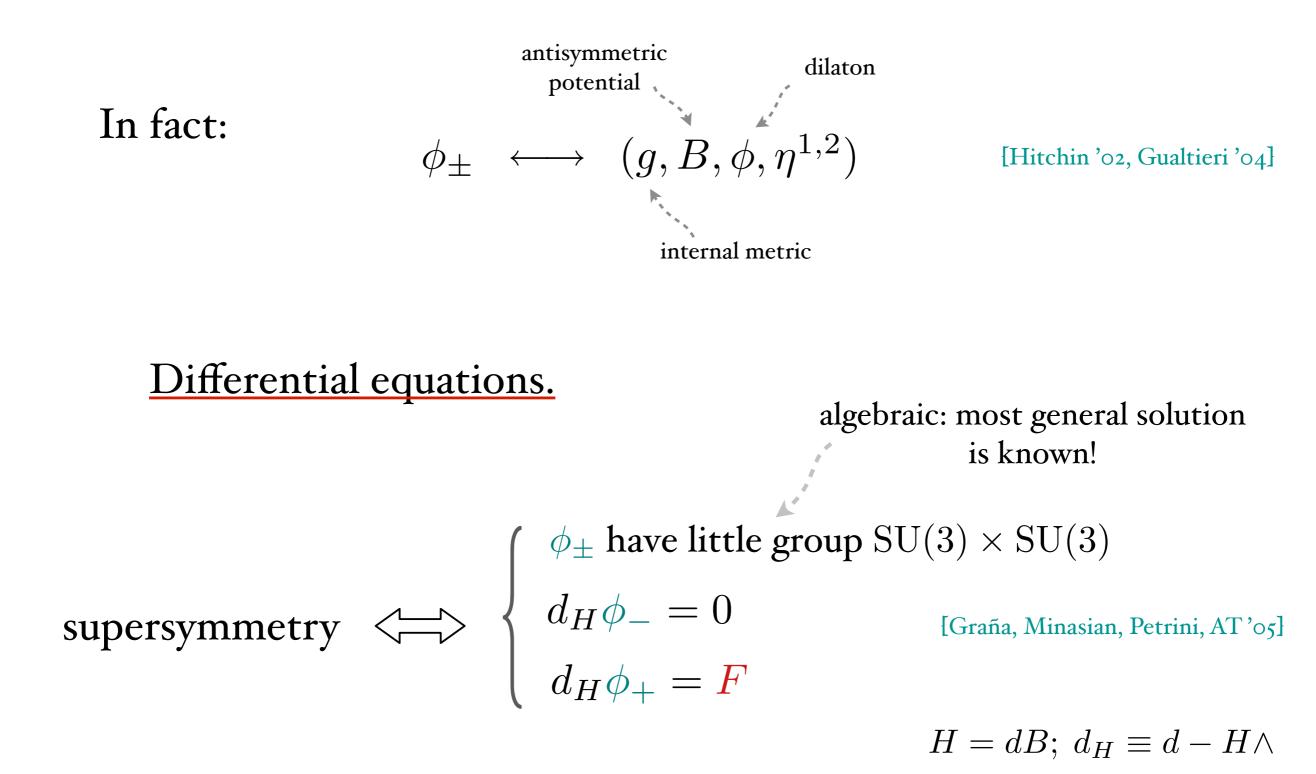
When can a pair of forms ϕ_{\pm} be written as $\eta_{\pm}^1 \otimes \eta_{\pm}^2$?

Answer:

- introduce 'generalized Lorentz group' SO(6, 6)
 - \bullet ordinary rotations $\omega^m{}_n$
 - generators: wedge by two-forms B_{mn}
 - \bullet contraction with bi-vector β^{mn}

• then: if and only if the little group $\operatorname{Stab}(\phi_{\pm})$ is $\operatorname{SU}(3) \times \operatorname{SU}(3)$ [in particular, ϕ_{\pm} are 'pure']

> Why? each η has a little group $Stab(\eta) = SU(3)$



II. Ten dimensions: algebraic equations

We now want to generalize the system we just saw to any supersymmetric solution in 10d

 $\begin{bmatrix} \text{not just} & \text{Mink}_4 \\ \text{AdS}_4 \end{bmatrix} \times M_6$

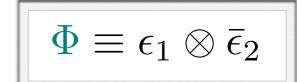
First question: what is the little group of a 10d spinor ϵ ?

 $K^M = \bar{\epsilon} \gamma^M \epsilon$ light-like vector: $K^M K_M = 0$

 ϵ also defines a 5-form: $\epsilon \otimes \overline{\epsilon} = K + \Omega_5 + *K$ $\iota_K \Omega$ breaks SO(8) to Spin(7)

$$\operatorname{Stab}(K) = \operatorname{SO}(8) \ltimes \mathbb{R}^{8}$$
$$\operatorname{Stab}(\epsilon) = \operatorname{Spin}(7) \ltimes \mathbb{R}^{8}$$

With two spinors $\epsilon_{1,2}$, we can define bilinear



Algebraic constraints.

When can a form Φ be written as $\epsilon_1 \otimes \overline{\epsilon}_2$?

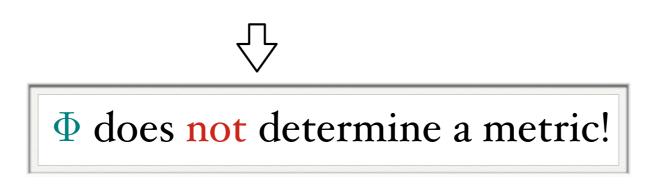
• Again look at 'generalized Lorentz' SO(10, 10)

From $\epsilon_{1,2}$, one would expect $(\text{ISpin}(7))^2$

Unfortunately the right answer is:

• if and only if $\operatorname{Stab}(\Phi)$ is $(\operatorname{Spin}(7)^2 \times \operatorname{Sl}(2,\mathbb{R})) \ltimes \operatorname{Heis}_{33}$ [Φ is not pure!] • I listed all possible solutions to this constraint.

• $\operatorname{Stab}(\Phi = \epsilon_1 \otimes \overline{\epsilon}_2)$ bigger than $\operatorname{Stab}(\epsilon_1) \times \operatorname{Stab}(\epsilon_2)$



 $K_1 = e_{-}$

 e_{+1}

 e_1,\ldots,e_8

each ϵ defines an "incomplete vielbein";

Possibility: complete them with vectors e_{+1} , e_{+2}

$$(\Phi, e_{\pm_1}, e_{\pm_2}) \mapsto \operatorname{metric}$$

II. Ten dimensions: differential equations

two lightlike vectors

 $K_i^M = \bar{\epsilon}_i \gamma^M \epsilon_i$

$$K = \frac{1}{2}(K_1 + K_2)$$

$$\tilde{K} = \frac{1}{2}(K_1 - K_2)$$

$$K_1$$

timeliko - 1

two old equations:

$$L_{\mathbf{K}}g = 0$$
$$\iota_{\mathbf{K}}H = d\tilde{K}$$

[*K* is Killing vector]

$$\Box > L_{\mathbf{K}} H = \{d, \iota_{\mathbf{K}}\} H = \iota_{\mathbf{K}} dH = 0$$

[K preserves H]

 K_2

[Hackett-Jones, Smith '04; Koerber, Martucci '07; Figueroa-O'Farrill, Hackett-Jones, Moutsopoulos '07]

What about Φ ?

$$L_{\mathbf{K}}g = 0 \qquad \iota_{\mathbf{K}}H = d\tilde{K}$$
$$d_{H}\Phi = (\tilde{K} \wedge + \iota_{K})\mathbf{F}$$

[AT '11]

This equation reproduces the ones for vacua:

supersymmetry
$$\langle - \rangle$$

 $\begin{pmatrix} L_K g = 0 & \iota_K H = d\tilde{K} \\ \hline d_H \Phi = (\tilde{K} \wedge + \iota_K) F & \text{[AT'II]} \\ (\psi \otimes \overline{\epsilon_2} e_{+_2}, d_H (\Phi e_{+_2}) - F) = 0 \\ (e_{+_1} \epsilon_1 \otimes \overline{\psi}, d_H (\Phi e_{+_2}) - F) = 0 \\ \forall \psi \end{pmatrix}$

 $\{(,) \text{ natural inner product}\}$



- no Ansatz necessary
- same system in IIA and IIB
- no covariant derivatives



- forms $e_{+_1}\epsilon_1\otimes\overline{\psi}$ should be characterized without spinors
- appearance of e_+ (inevitable?)

It is possible to apply it now to 4d susy solutions

[work in progress with my student Dario Rosa]

• one can apply similar methods to $\mathcal{N} = 2$, d = 4 supergravity

- ullet a 4d spinor describes an \mathbb{R}^2 structure: again an "incomplete vielbein"
- we need again to introduce e_+ 's
- but when K is timelike, \exists nicer system

• we are now putting the 10d system in the same form

• when K is timelike, we traded the ugly 'pairing equations' for $dK = \ldots$

Conclusions

• We saw how supersymmetry in type II string theory can be expressed in terms of forms.

• There is room for improvement, but potentially the system we found is more fundamental than the one for 4d compactifications.

• It still uses the geometry of $T \oplus T^*$, but it does not involve generalized complex structures or pure spinors.